

# GRACE Follow-On orbit and gravity field determination using GPS carrier phase and SLR observations

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## Introduction

LEO (Low-Earth Orbiter) satellites equipped both with GNSS (Global Navigation Satellite System) antennas and laser retroreflectors enable the combination of these two space geodetic techniques, and thus, can serve as so-called **space ties**.

Taking the example of the two **GRACE Follow-On** LEO satellites (Landerer et al., 2020), which are dedicated to the determination of the Earth's gravity field, **GPS code and carrier phase** observations and **SLR** observations can be **combined** to derive **reduced-dynamic orbits** together **with Earth's gravity field parameters**. In this case, the two satellites act as a space tie connecting the SLR and GPS observations techniques, and the Earth's gravity field serves as a global tie.

## Methodology

Satellite **orbit parameters** of both GRACE Follow-On satellites (GF1 and GF2) and **Earth's gravity field** parameters are **simultaneously** estimated using the Celestial Mechanics Approach (Beutler et al., 2010), which is implemented in the Bernese GNSS Software (Dach et al., 2015).

**Daily reduced-dynamic** orbits are estimated using **pseudo-stochastic** orbit parameters to absorb unknown deficiencies in the modelling. The jointly estimated **Earth's gravity field** parameters are set up to **degree and order 70** and solved for **monthly snapshots**. The time-variable gravity field recovery **relies on a set of a priori force models**, which account for forces perturbing the satellites trajectories.

The parameter estimation is based on 10 s sampled GPS carrier phase observations to the two GRACE Follow-On satellites and SLR two-way range observations (for the **year 2019**).

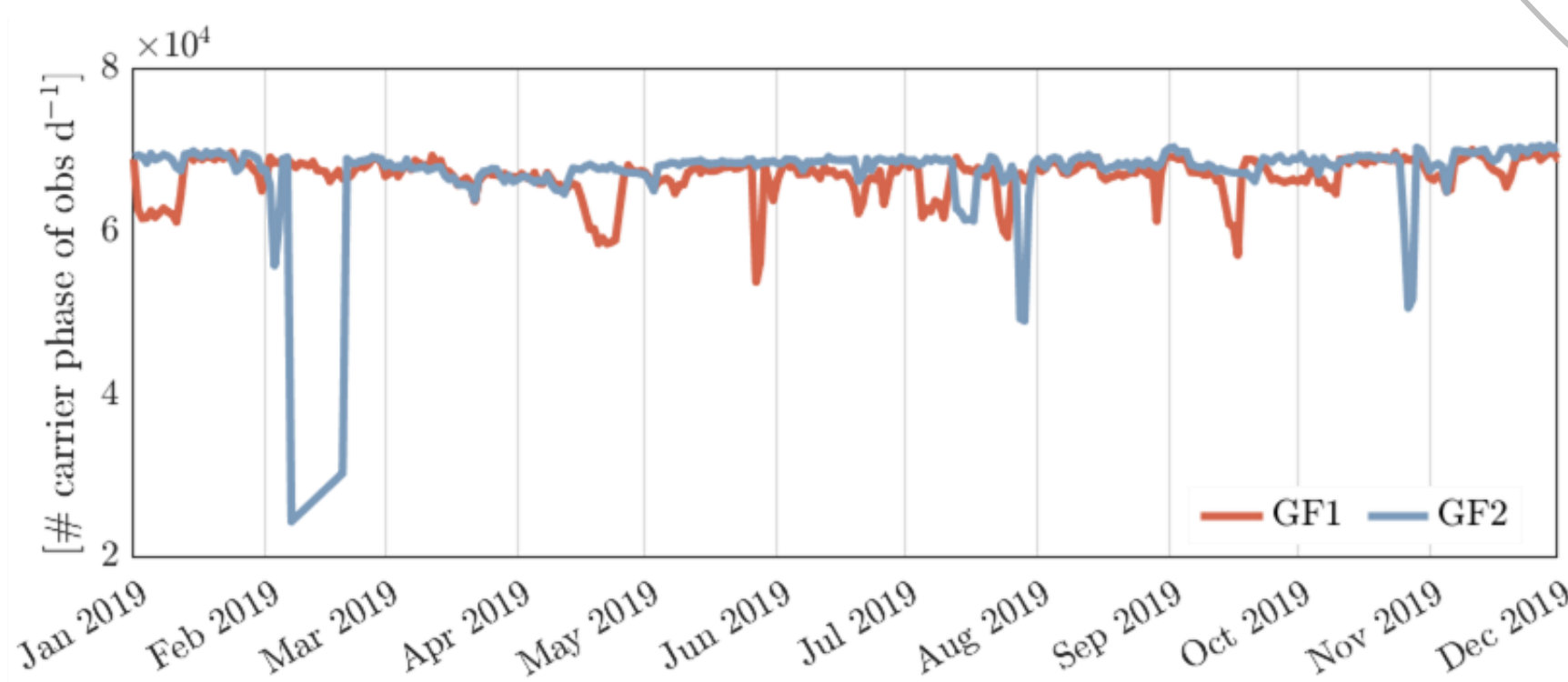
## Parametrization

<b>Orbit parameters for GRACE Follow-On:</b>	# parameters per arc
Initial conditions:	2x [6]
Constant accelerations:	2x [3]
15 min pulses in	
radial direction	2x[96]
along-track direction	2x[96]
cross-track direction	2x[96]
satellite clock correction	per epoch
carrier phase ambiguities	
<b>Gravity field parameters:</b>	
degree/order 2 .. 70	

## GPS Observations

Total number of GPS observations in use to:  
GF1: ~2 M per month  
GF2: ~2 M per month

Elevation mask: 10°



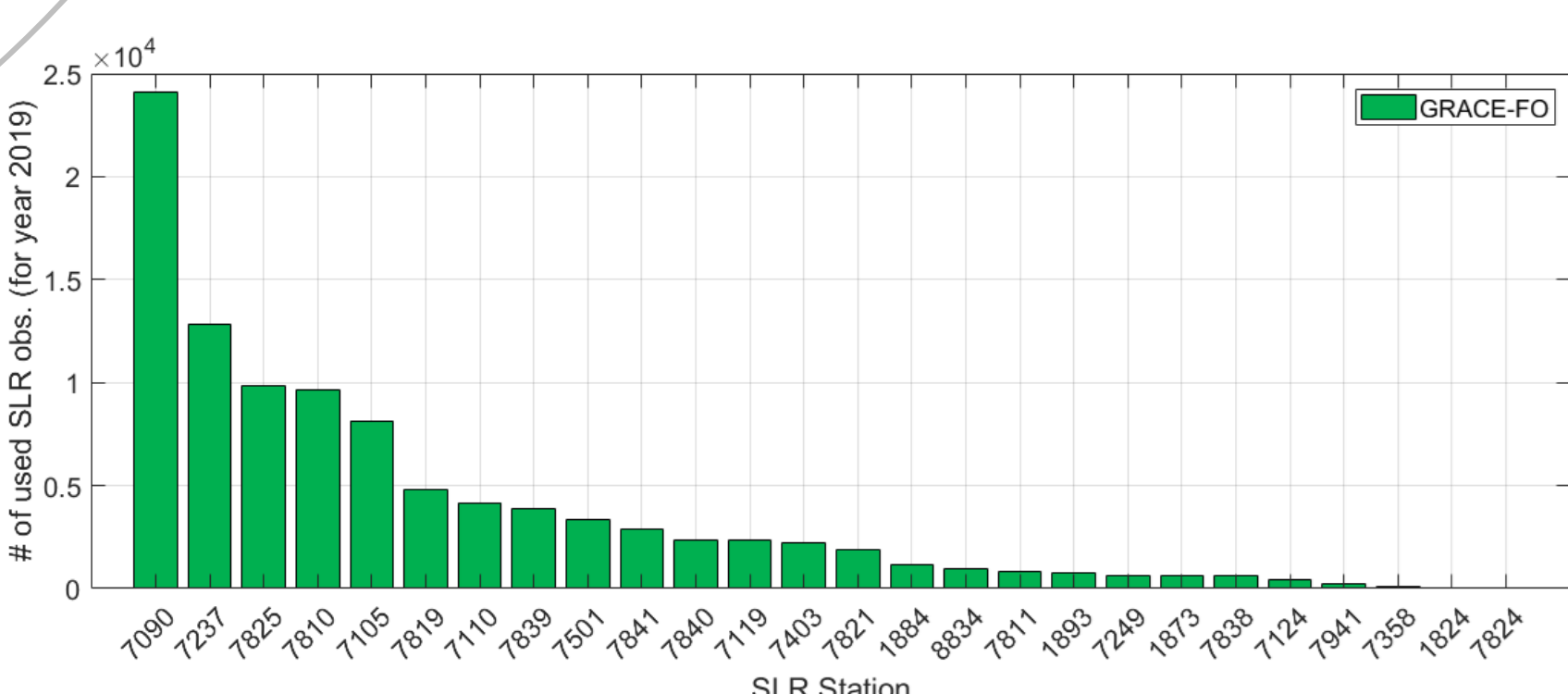
## Background Models

<b>Gravity field</b>	EIGEN-GRGS RL05
<b>Astronomic bodies</b>	JPL DE405
<b>Mean pole</b>	Linear
<b>Solid Earth tides</b>	IERS2010
<b>Solid Earth pole tides</b>	IERS2010
<b>Ocean tides</b>	FES2014b
<b>Ocean pole tides</b>	IERS2010
<b>Atmospheric tides</b>	AOD RL06
<b>De-aliasing</b>	AOD RL06
<b>Relativistic effects</b>	IERS2010
<b>Reference frame</b>	ITRF/SLRF 14
<b>ERP</b>	IERS-14-C04

## SLR Observations

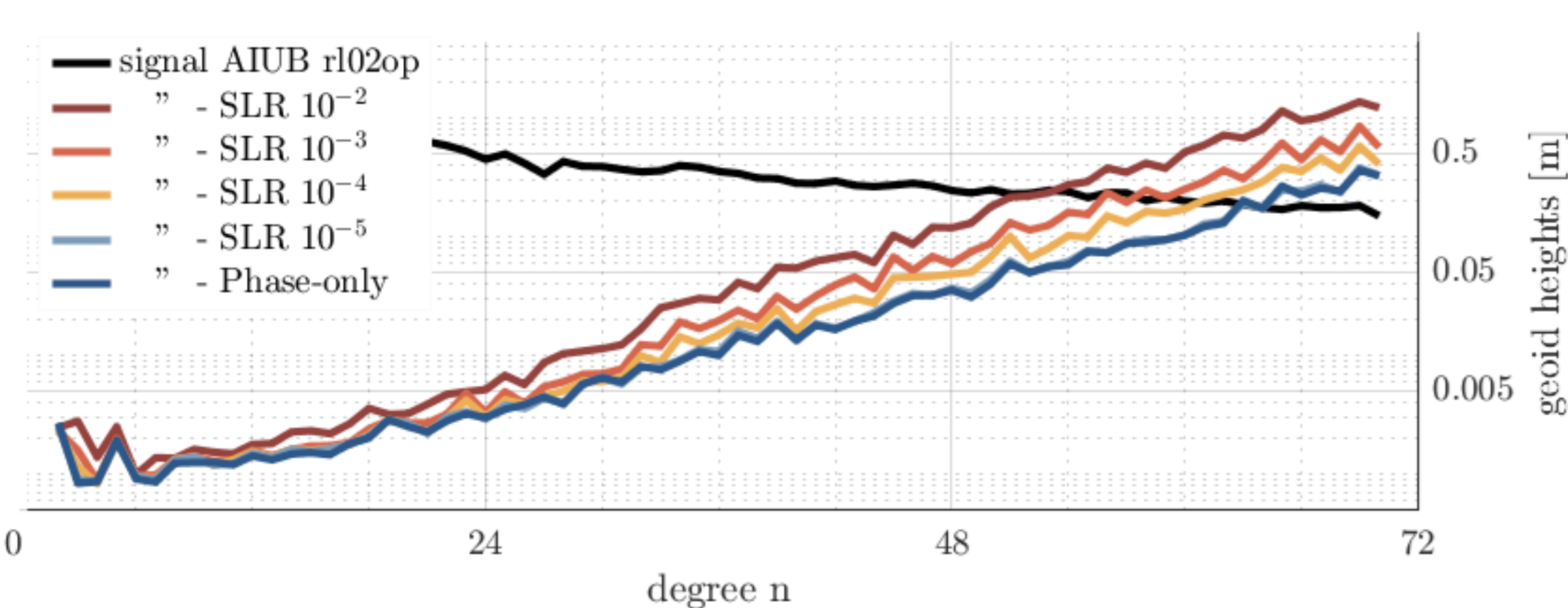
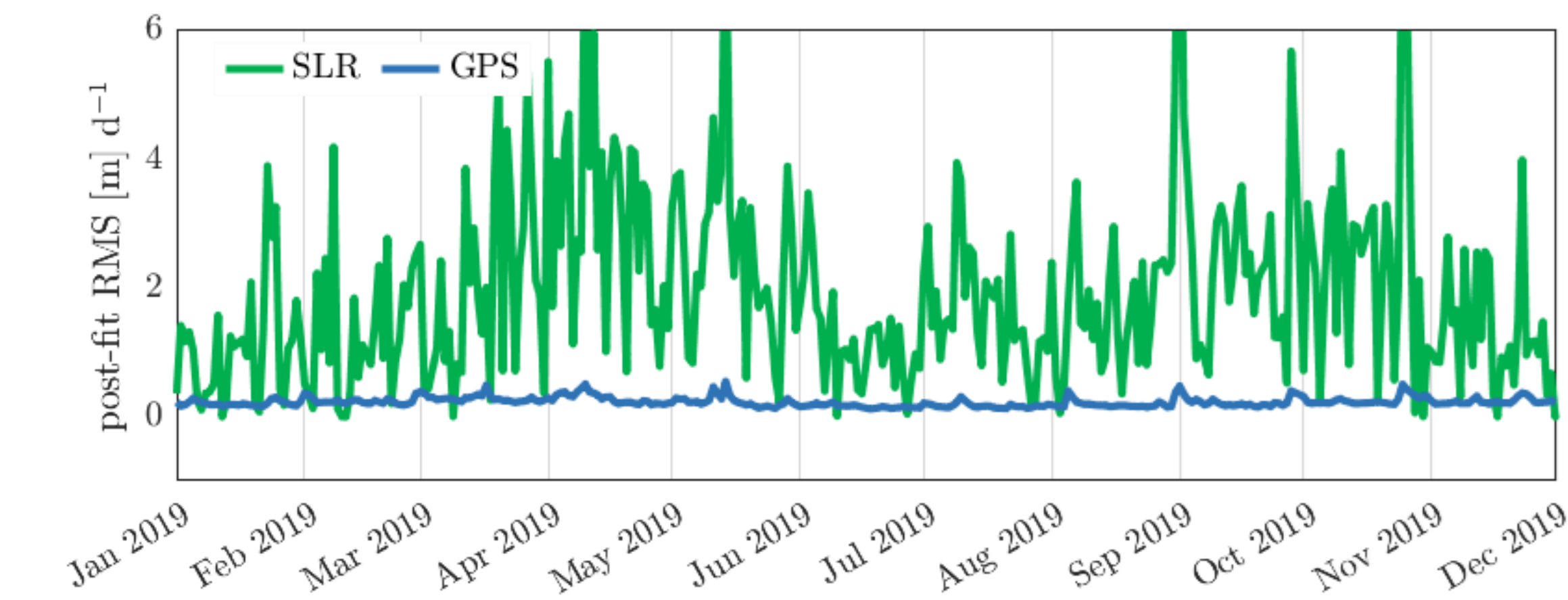
Total number of SLR normal point observations to:  
GF1: ~4'000 per month  
GF2: ~4'000 per month  
Tracked by 26 SLR stations

Elevation mask: 3°



## Results

The daily dynamic orbit estimation, i.e., 6 initial conditions for each GRACE-FO satellite, results in **post-fit RMS** values of on average **20 cm** using only GPS observations and **2 m** with only SLR observations. In the **combination of GPS and SLR** observations to GRACE-FO, the **two different space geodetic techniques have to be weighted** relative to each other. In this study a weight according to **the ratio of these post-fit RMS values squared** is introduced (case SLR 10<sup>-2</sup>) and besides, additional weighting schemes are investigated, where SLR is further down-weighted. For the gravity field estimation, the **SLR contribution** to the estimation of the **pseudo-stochastic parameters is neglected** due to the low number of observations. The relative weighting scheme plays a crucial role in determining the gravity field solution.



## References

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