

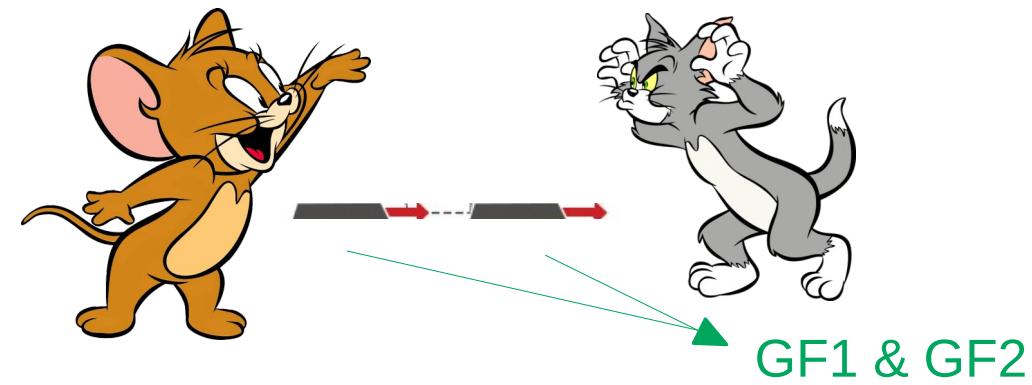
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GRACE Follow-On Orbit and Gravity Field Determination from GPS Carrier Phase Observations

Martin Lasser, Ulrich Meyer, Daniel Arnold and Adrian Jäggi

22nd Swiss Geoscience Meeting 2024, 8 – 9 November 2024, Basel, Switzerland

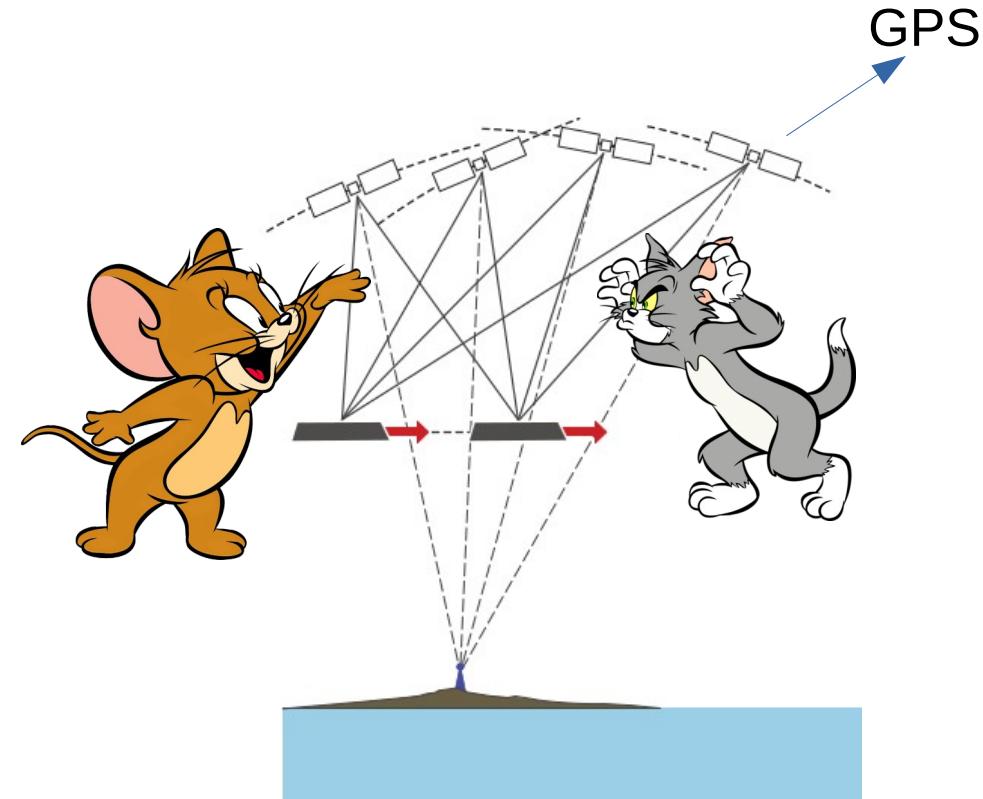
u^b GRACE Follow-On
Observation concept



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GRACE Follow-On

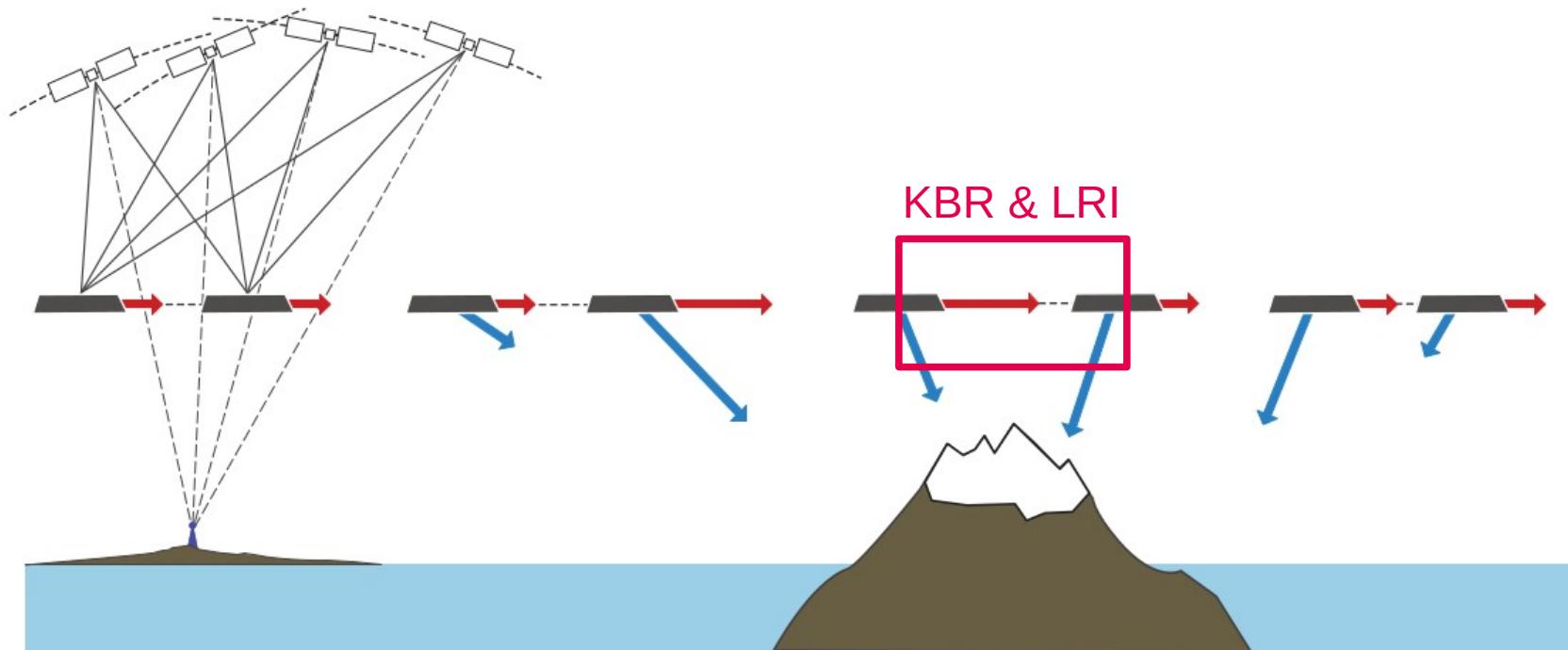
Observation concept



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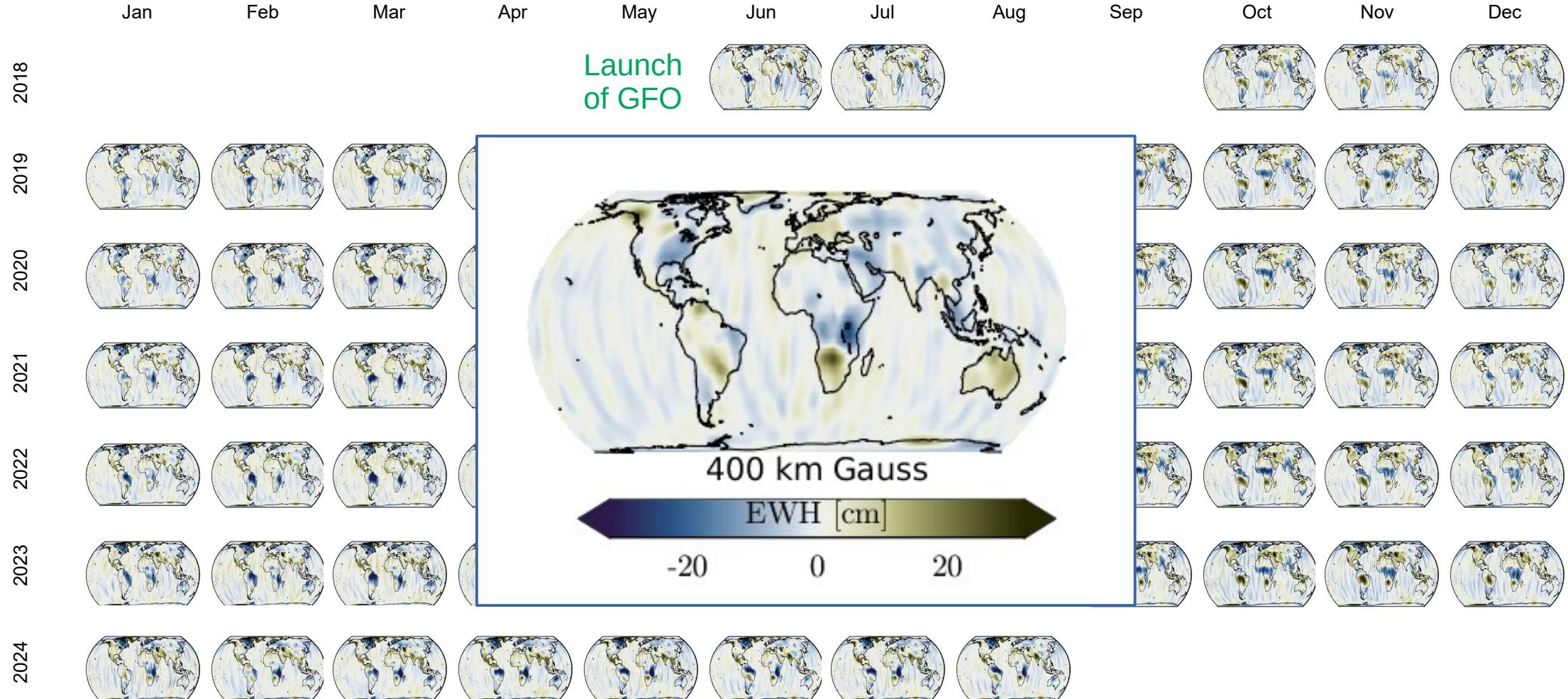
GRACE Follow-On

Observation concept



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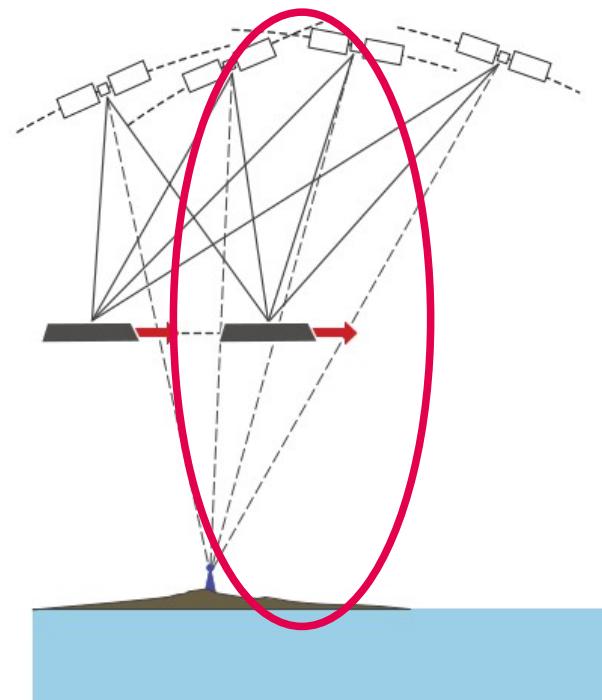
GRACE Follow-On Results



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GRACE Follow-On

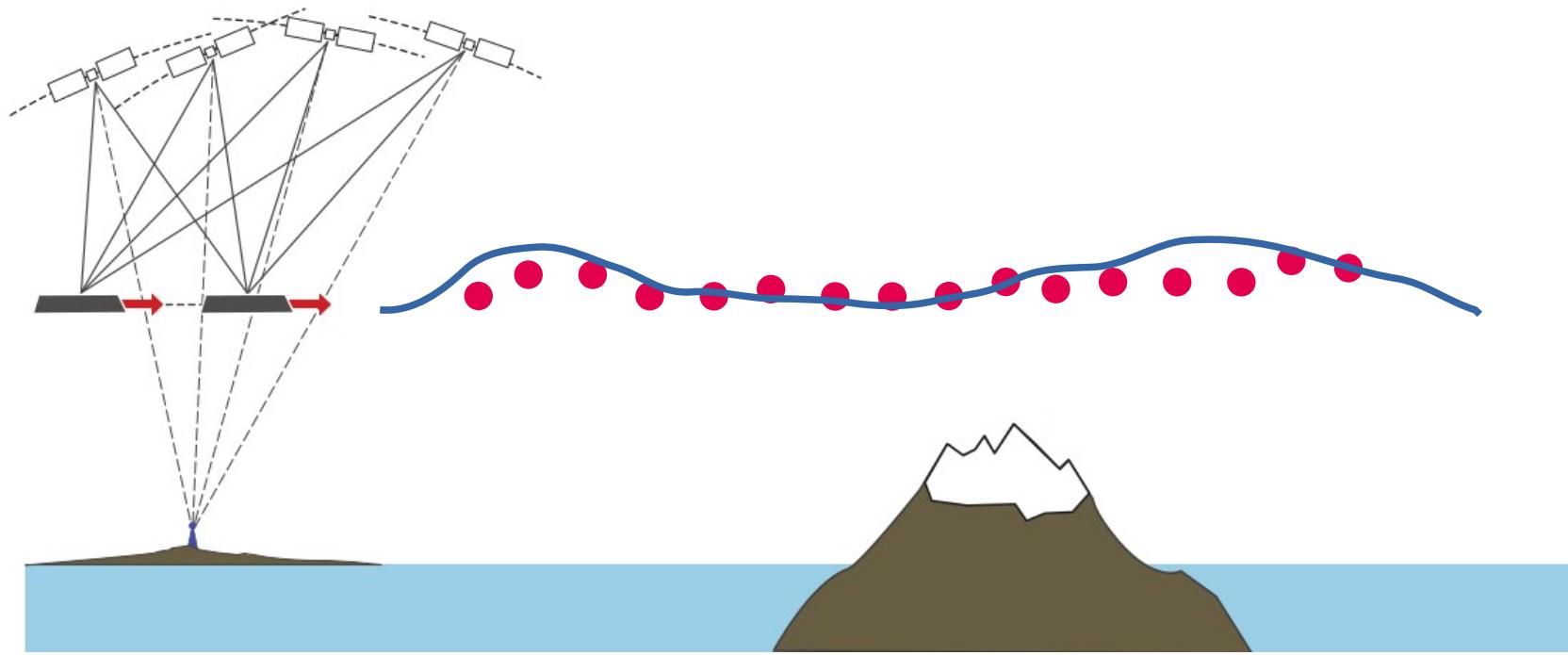
Orbit representation



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GRACE Follow-On

Orbit representation

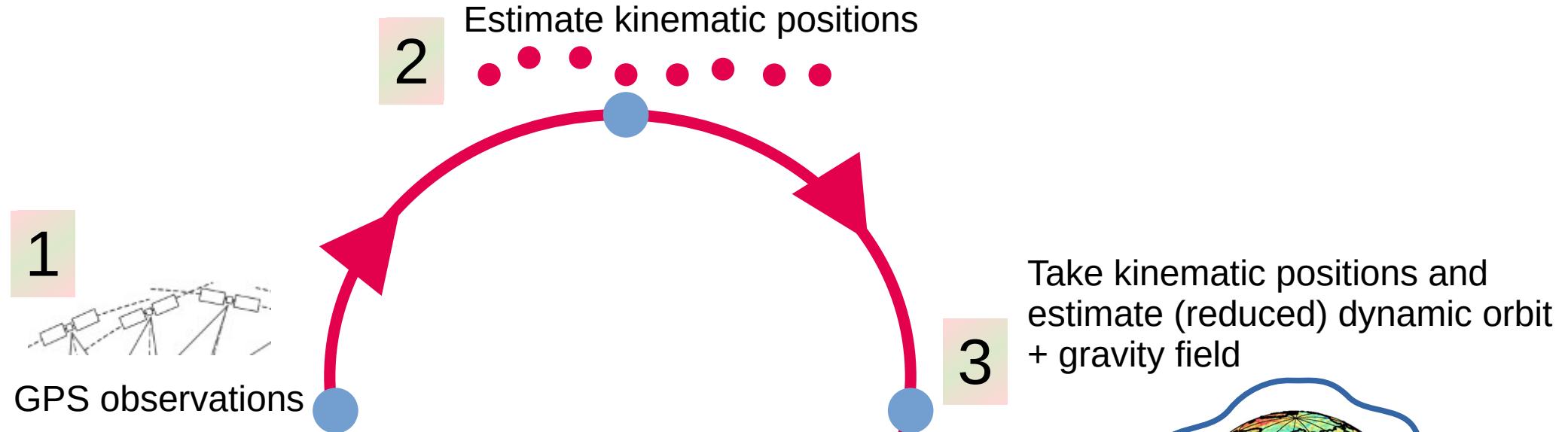


Kinematic **orbit**
(Reduced)
dynamic orbit

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Gravity Field Recovery

Detour



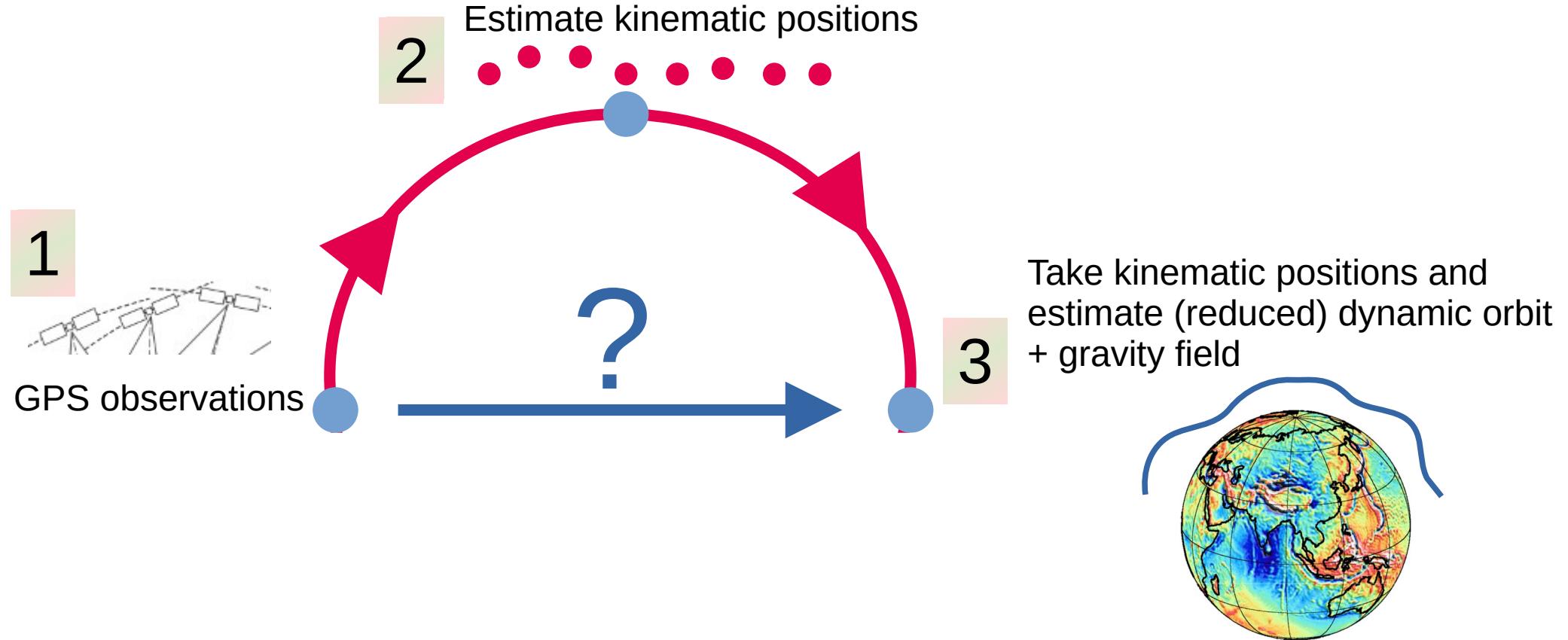
Background information

- Celestial Mechanics Approach (CMA, Beutler et al., 2010) applied
- CODE GNSS products
- PCV maps used
- Ambiguities float or integer-fixed

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Gravity Field Recovery

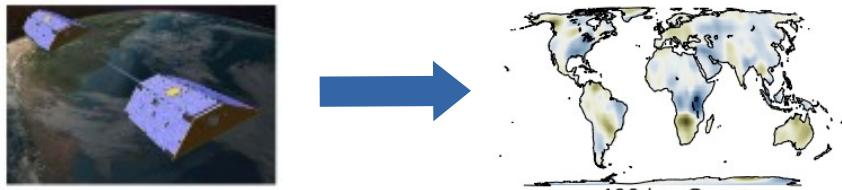
Shortcut



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Parametrisation

For estimating gravity fields



Basic parametrisation

- Initial conditions
- Bias in radial, along-track, cross-track
- Clock per epoch
- Ambiguities

Additional parameters

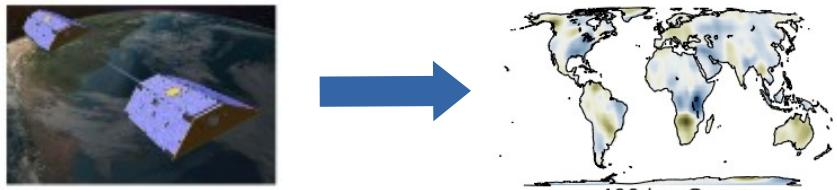
- 15 min PCA per satellite in
 - radial
 - along-track
 - cross-track

in daily arcs

+ gravity field d/o=2..70

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Parametrisation For estimating gravity fields



Basic parametrisation

- Initial conditions
- Bias in radial, along-track, cross-track
- Clock per epoch
- Ambiguities

Additional parameters

- 15 min PCA per satellite in
 - radial
 - along-track
 - cross-track

+ gravity field

Force models

Gravity field	GOCO06s
Astronomic bodies	JPL DE421 (all planets)
Mean pole	Linear
Solid Earth tides	IERS2010
Solid Earth pole tides	IERS2010
Ocean tides	FES2014b (+ admittances from TUG)
Ocean pole tides	Desai
Atmospheric tides	AOD RL06
Atmospheric & oceanic dealiasing	AOD RL06
Relativistic effects	IERS2010

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The Detour in a Nutshell

A bit of math ...

Least-squares to estimate kinematic positions

$$\begin{aligned}\hat{\mathbf{x}}_{\text{KIN}} &= (\mathbf{A}^T \mathbf{P}_{\text{ph}} \mathbf{A})^{-1} \mathbf{A}^T \mathbf{P}_{\text{ph}} \boldsymbol{\ell}_{\text{ph}} \quad \text{with} \quad \mathbf{A} := \frac{\partial \boldsymbol{\ell}_{\text{ph}}}{\partial \mathbf{x}_{\text{KIN}}} \\ &= \mathbf{N}_{\text{ph}}^{-1} \mathbf{A}^T \mathbf{P}_{\text{ph}} \boldsymbol{\ell}_{\text{ph}} .\end{aligned}$$

Least-squares to estimate orbit and gravity field parameters
based on kinematic positions

$$\begin{aligned}\hat{\mathbf{p}} &= (\mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \mathbf{B})^{-1} \mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \boldsymbol{\ell}_{\text{KIN}} \quad \text{with} \quad \mathbf{B} := \frac{\partial \boldsymbol{\ell}_{\text{KIN}}}{\partial \mathbf{p}} \quad \mathbf{Q}_{\text{KIN}} := \mathbf{N}_{\text{ph}}^{-1} \\ &= \mathbf{N}_{\text{p}}^{-1} \mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \hat{\mathbf{x}}_{\text{KIN}} .\end{aligned}$$

u^b The Detour in a Nutshell

A bit of math ...

Or summarise...

$$\mathbf{M} := \mathbf{AB}$$

$$\hat{\mathbf{p}} = (\mathbf{M}^T \mathbf{P}_{ph} \mathbf{M})^{-1} \mathbf{M}^T \mathbf{P}_{ph} \boldsymbol{\ell}_{ph}$$

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Stochastic Properties of Kinematic Positions

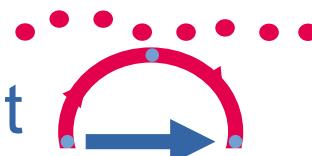
And how to consider them

Recall kinematic positions as observations

$$\hat{\mathbf{p}} = (\mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \mathbf{B})^{-1} \mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \boldsymbol{\epsilon}_{\text{KIN}} \quad \text{with} \quad \mathbf{B} := \frac{\partial \ell_{\text{KIN}}}{\partial \mathbf{p}}$$

$$= \mathbf{N}_p^{-1} \mathbf{B}^T \mathbf{Q}_{\text{KIN}}^{-1} \hat{\mathbf{x}}_{\text{KIN}} .$$

$$\mathbf{Q}_{\text{KIN}} := \mathbf{N}_{\text{ph}}^{-1}$$



Remark for the shortcut

When directly starting with phase observations all correlations are implicitly considered.

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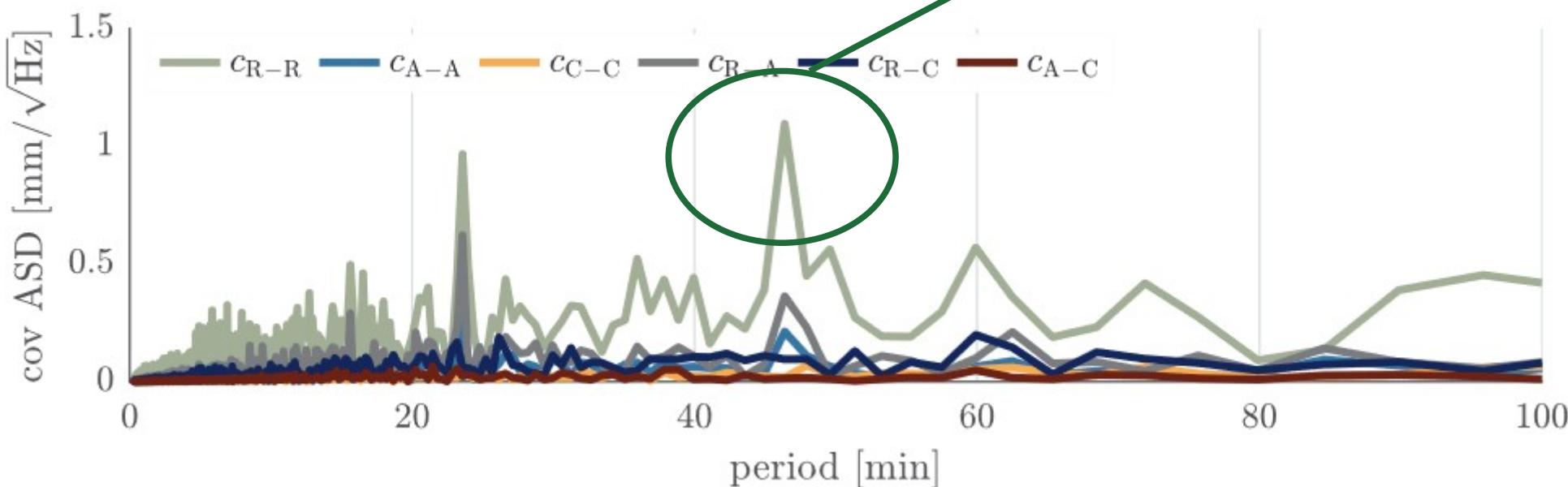
Stochastic Properties Kinematic Positions

Power-spectrum – epoch-wise

$$\mathbf{Q}_{\text{KIN}} := \mathbf{N}_{\text{ph}}^{-1}$$

$$\mathbf{C}_{\text{epo}}^{\text{KIN}} = \begin{bmatrix} c_{xx} & c_{xy} & c_{xz} \\ c_{xy} & c_{yy} & c_{yz} \\ c_{xz} & c_{yz} & c_{zz} \end{bmatrix} \rightarrow \begin{array}{l} \text{radial} \\ \text{along-track} \\ \text{cross-track} \end{array}$$

2/rev peak due to polar gap
of GPS constellation



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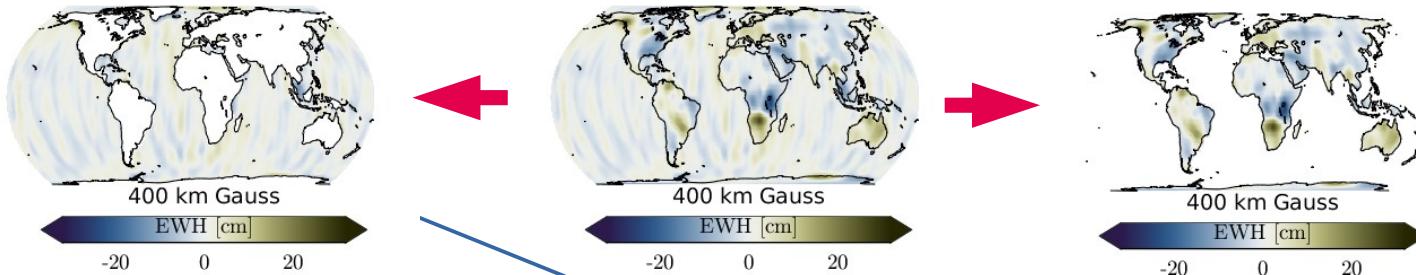
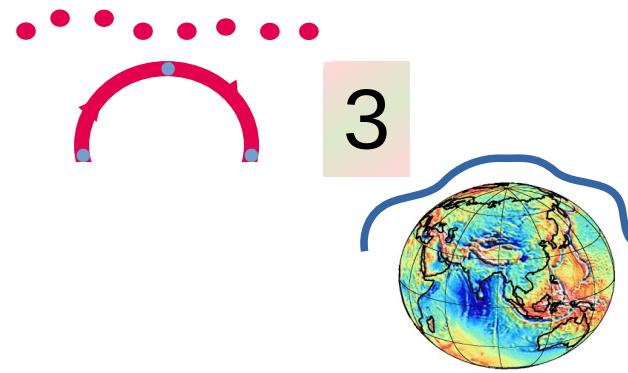
Results Reduced Dynamic Orbit

A priori orbit (pre-fit)

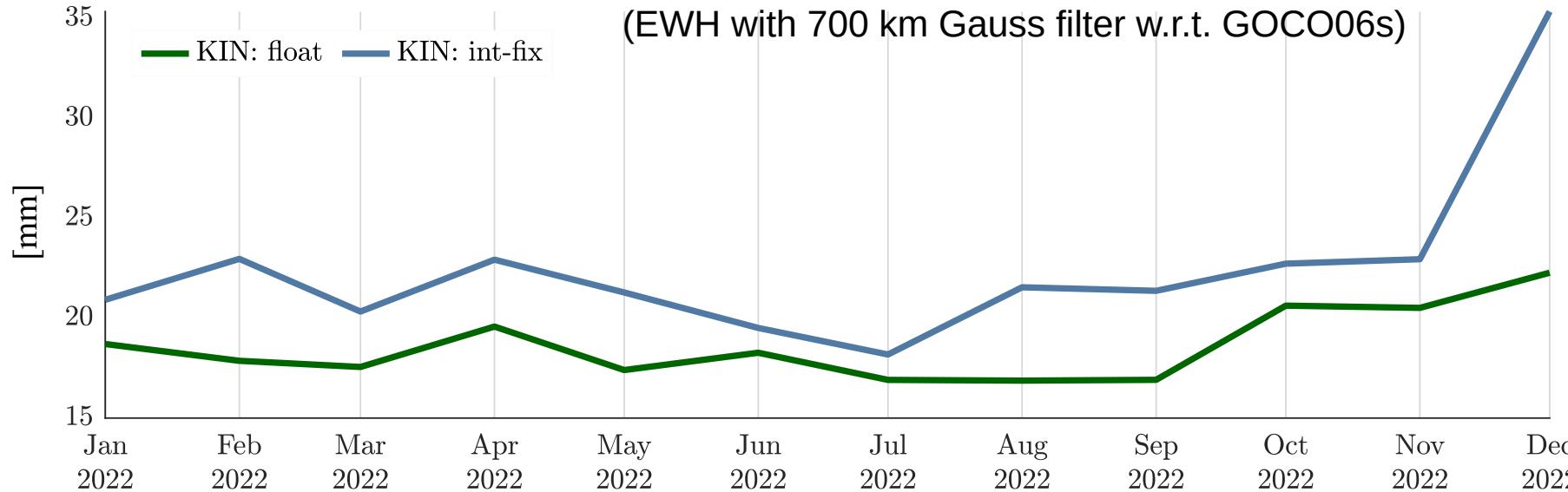
	float	fixed
MAD carrier phase residuals [cm]	0.21	0.29
MAD w.r.t. kinematic positions [cm]	radial along cross 1.05 – 0.84 – 0.69	radial along cross 0.63 – 0.26 – 0.19
MAD of K-band validation [cm]	0.17	0.14

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Results Gravity Field From kinematic positions



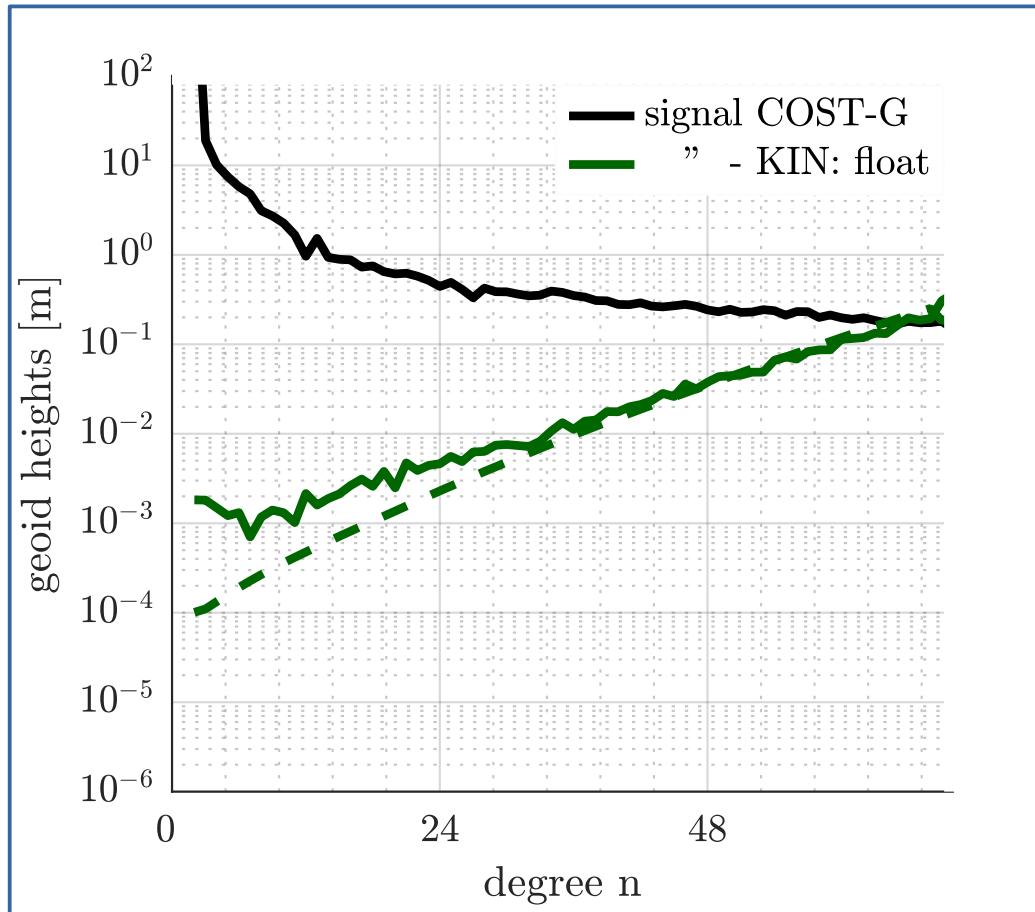
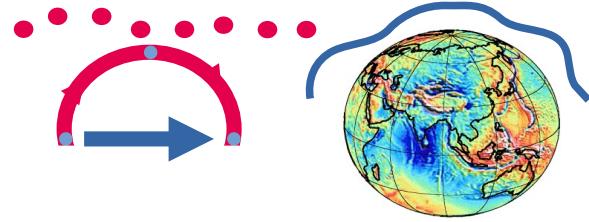
RMS over the oceans
(EWH with 700 km Gauss filter w.r.t. GOCO06s)



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Results Gravity Field

Directly from phase observations



Gravity field representation in spherical harmonics

unknown coefficients

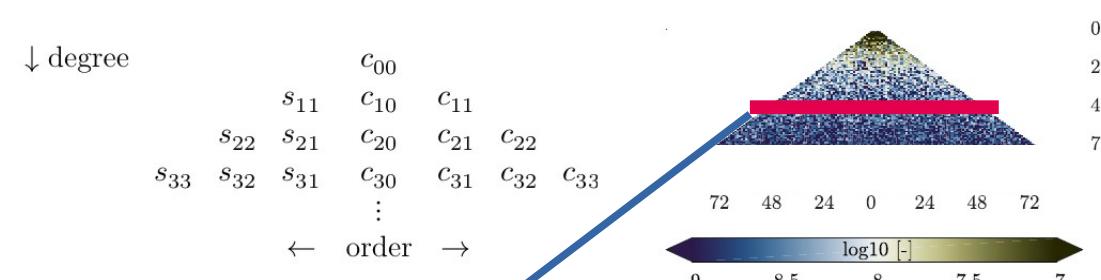
$$V = \frac{GM}{R} \sum_{n=0}^N \left(\frac{a_e}{r} \right)^n \sum_{m=0}^n P_{nm}(\sin(\varphi)) [c_{nm} \cos(m\lambda) + s_{nm} \sin(m\lambda)]$$

degree N point of evaluation (r, φ, λ)

↓ degree

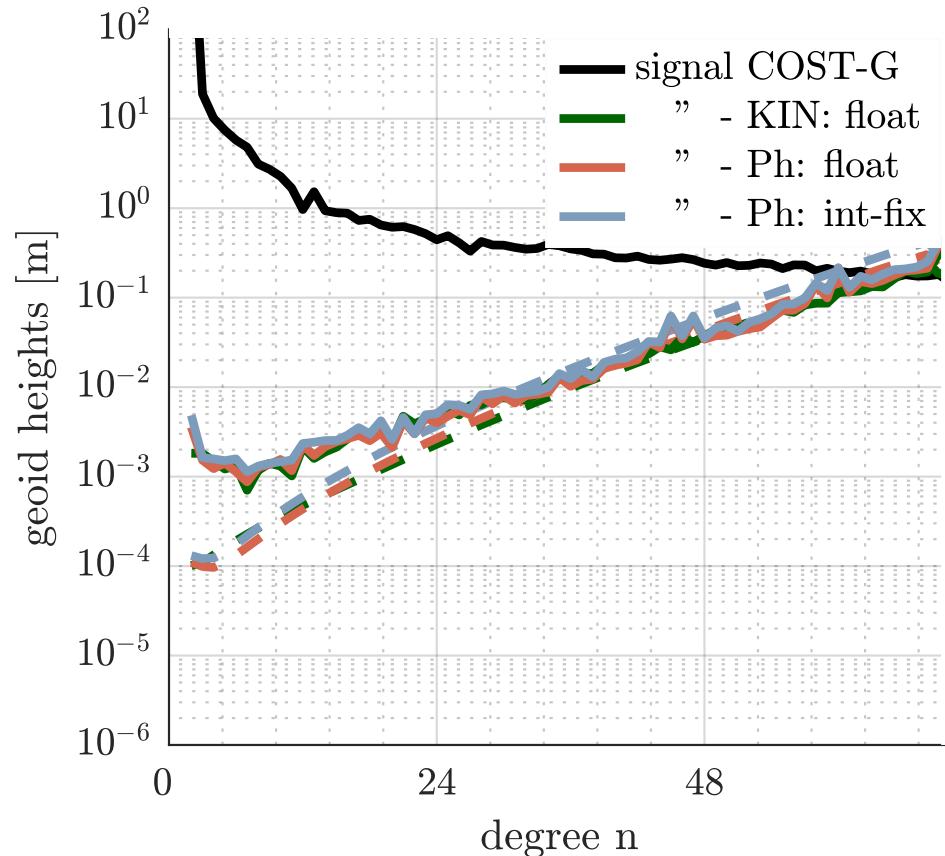
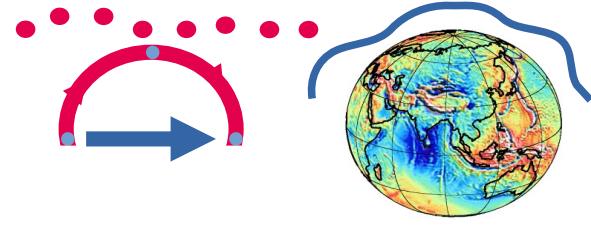
$$\begin{matrix} & c_{00} \\ & s_{11} & c_{10} & c_{11} \\ & s_{22} & s_{21} & c_{20} & c_{21} & c_{22} \\ & s_{33} & s_{32} & s_{31} & c_{30} & c_{31} & c_{32} & c_{33} \\ & & & & \vdots & & & \end{matrix}$$

← order →



$$\sigma_n = \sqrt{\sum_{m=0}^n (c_{nm}^2 + s_{nm}^2)}$$

Results Gravity Field Directly from phase observations

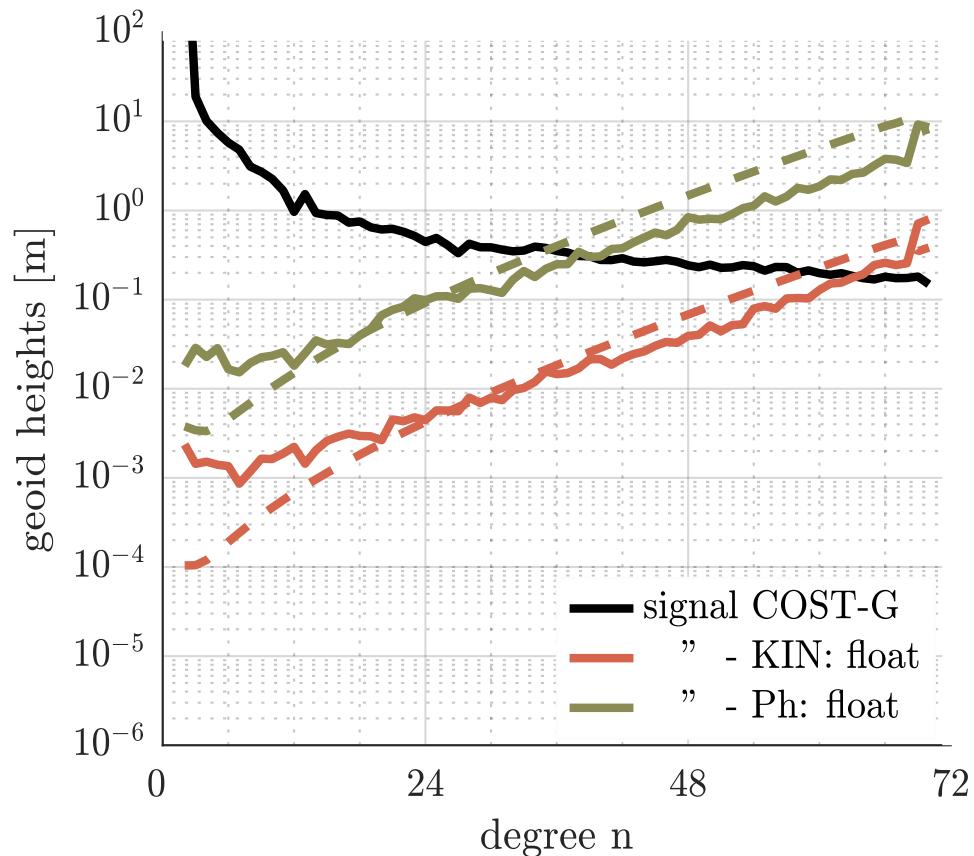
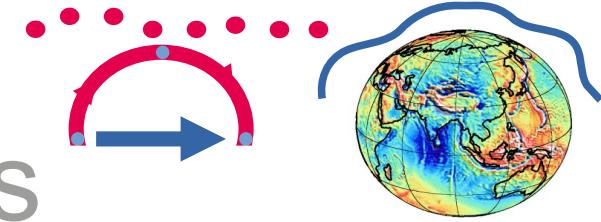


- Ambiguity-float solution almost the same for KIN or phase observations
- Ambiguity-integer-fixed solution slightly deteriorated
- Formal errors over-estimate low-degree coefficient quality

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Results Gravity Field

Robustness of kinematic positions



- Imperfect PCV model used (aka 10° elevation cut-off ignored)
- Solution directly from phase observations heavily affected
- Kinematic positions do not really suffer

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Results Reduced Dynamic Orbit

Jointly estimated with the gravity field (pre-fit & post-fit)

	float	fixed
MAD carrier phase residuals [cm]	0.21	0.29
MAD w.r.t. kinematic positions [cm]	radial along cross 1.05 – 0.84 – 0.69 1.08 – 0.86 – 0.70	radial along cross 0.63 – 0.26 – 0.19 0.64 – 0.26 – 0.20
MAD of K-band validation [cm]	0.17 0.17	0.14 0.13

Summary & Conclusions

- Gravity field determination for carrier phase observations introduced
- Challenges for ambiguity resolution strategy remain in all cases
- Computational efficiency needs to be improved
- For GRACE Follow-On:
 - Extend with inter-satellite link
 - Combine carrier phase data and K-band/LRI data

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Thank you for your attention

Contact

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