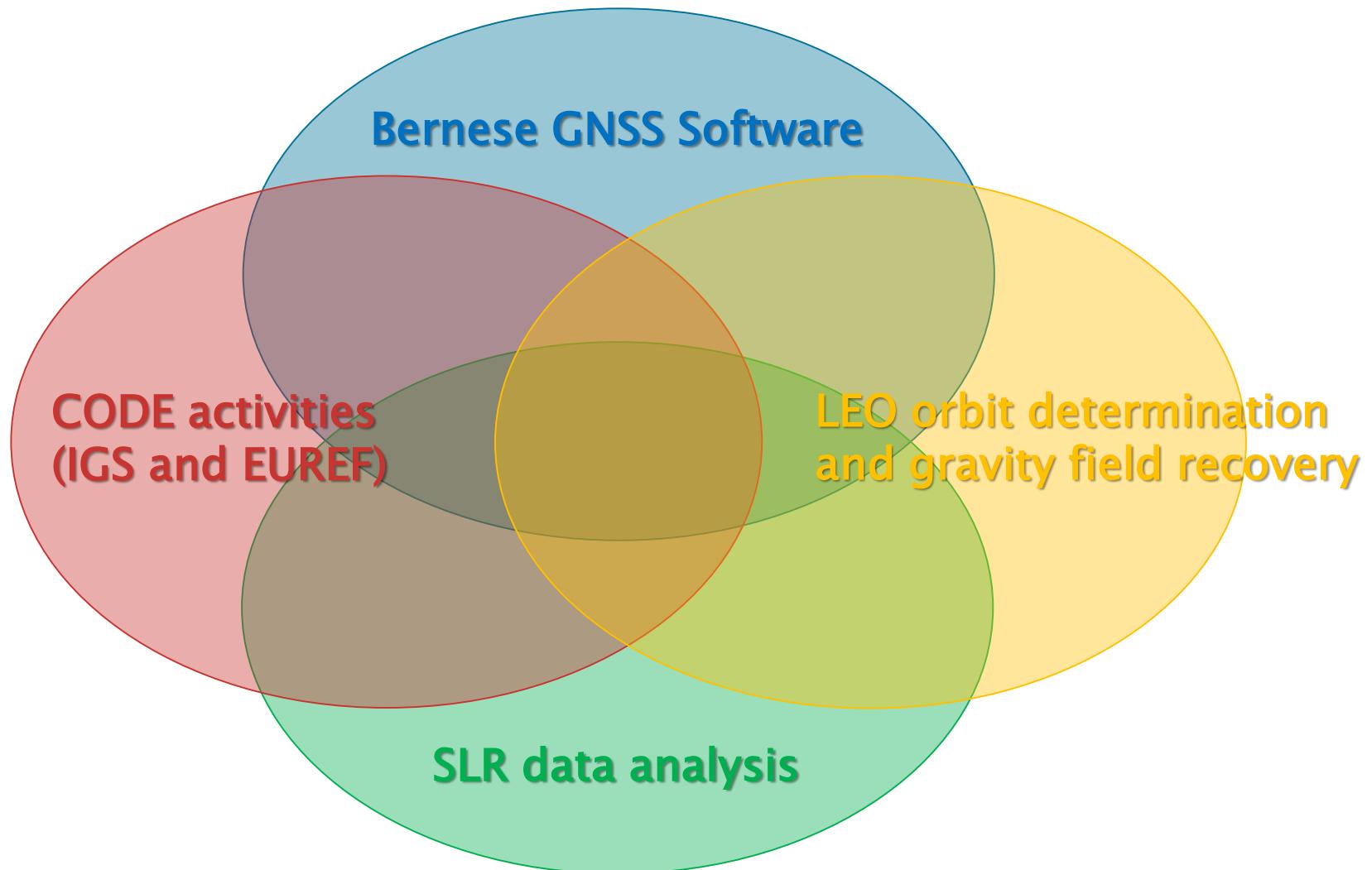


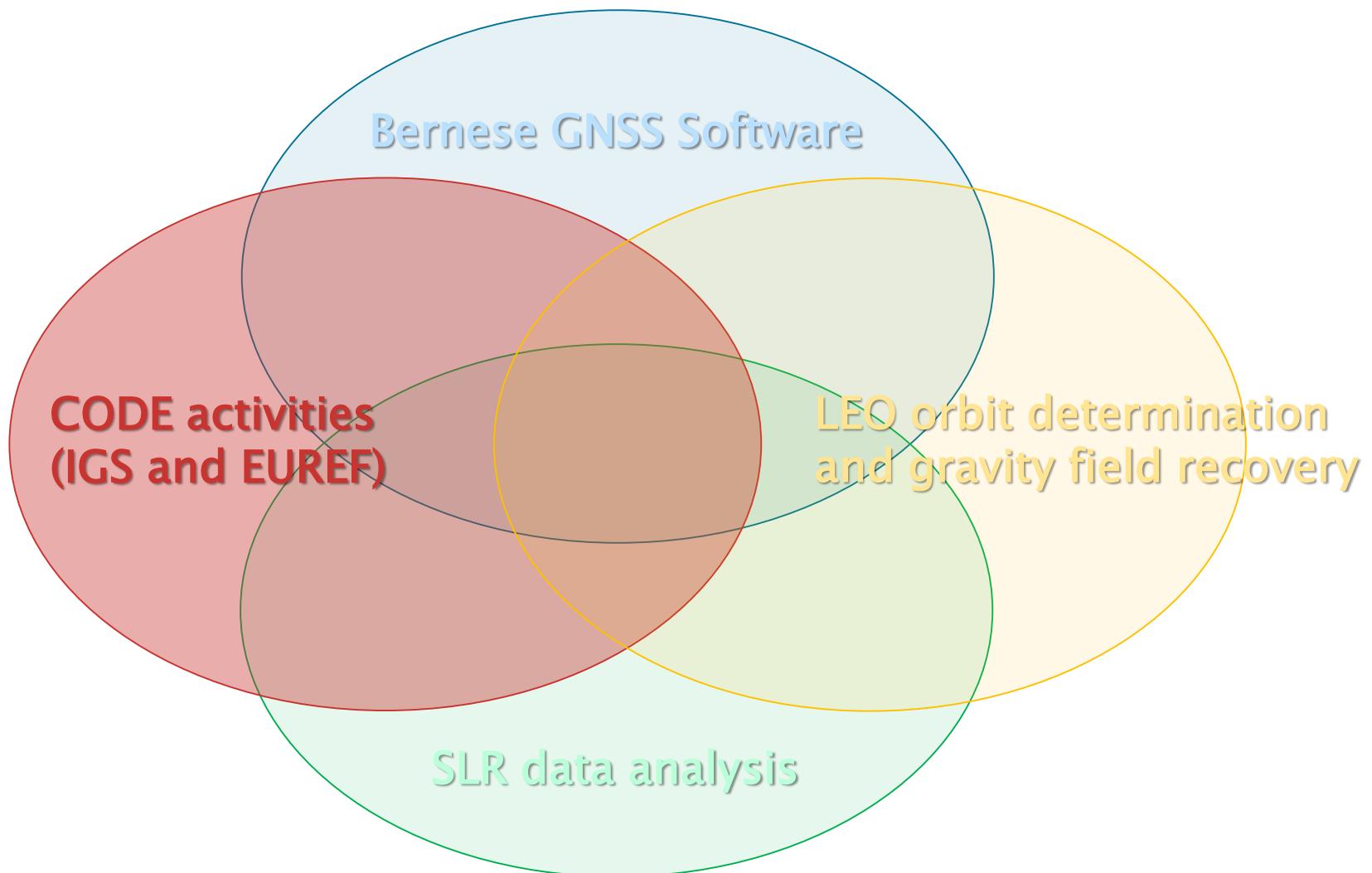
Activities in the Research Group on Satellite Geodesy at AIUB

Astronomisches Institut

Satellite Geodesy Research Group



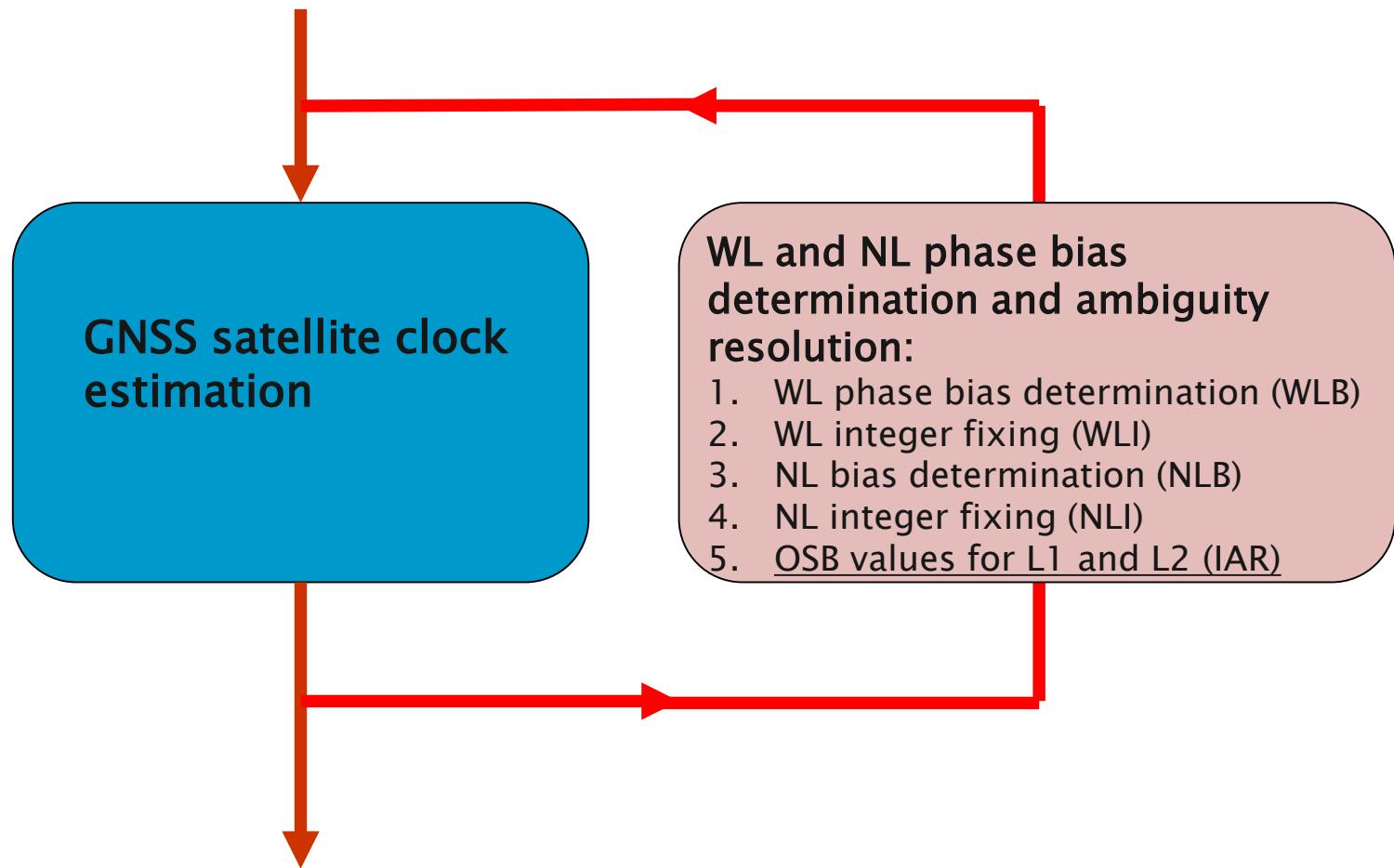
Satellite Geodesy Research Group



Clock products with resolved ambiguities

S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: **New ambiguity-fixed IGS clock analysis products at CODE**. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Procedure for clock estimation with ambiguity resolution

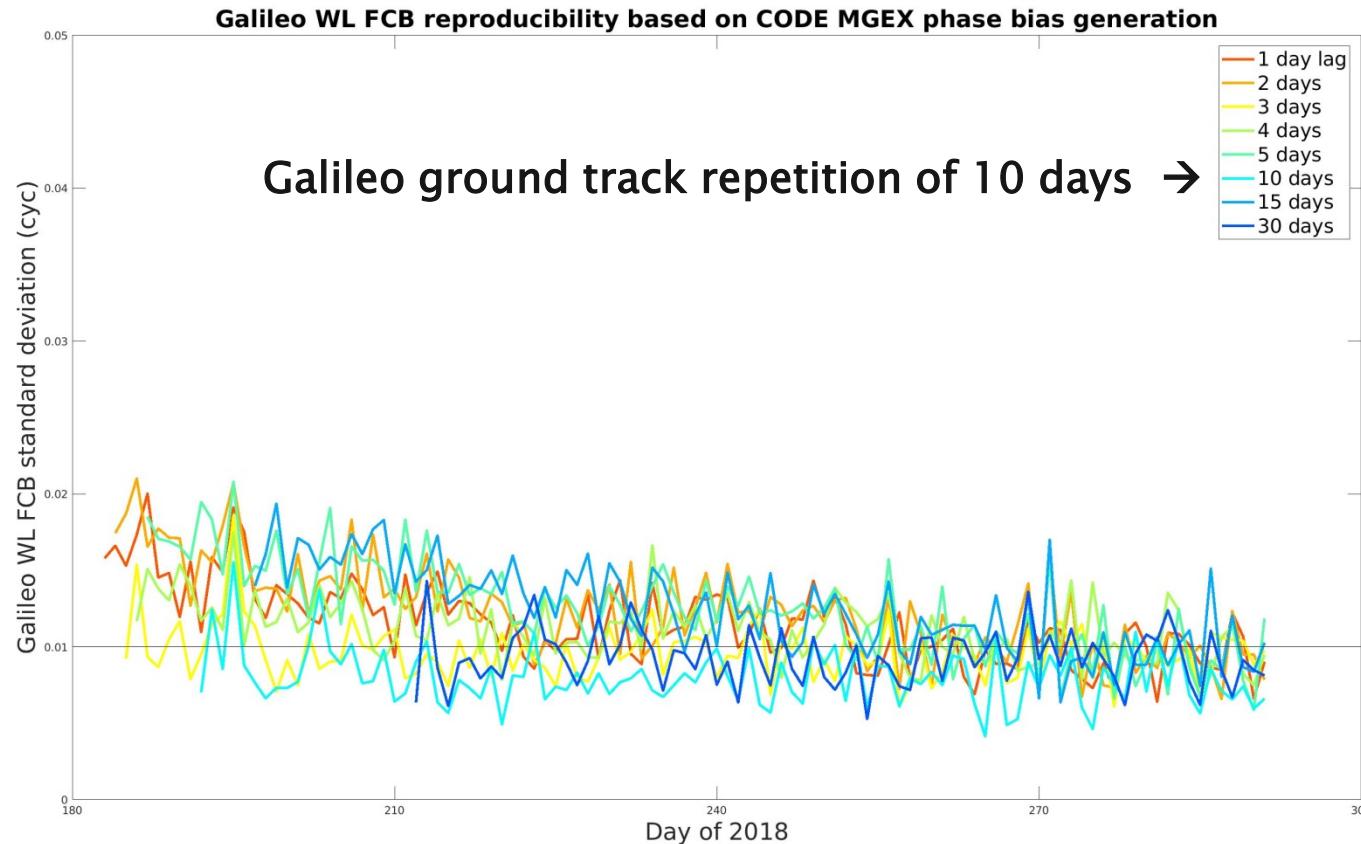


Activated for rapid, final, and MGEX clocks in July 2018

S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: New ambiguity-fixed IGS clock analysis products at CODE. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

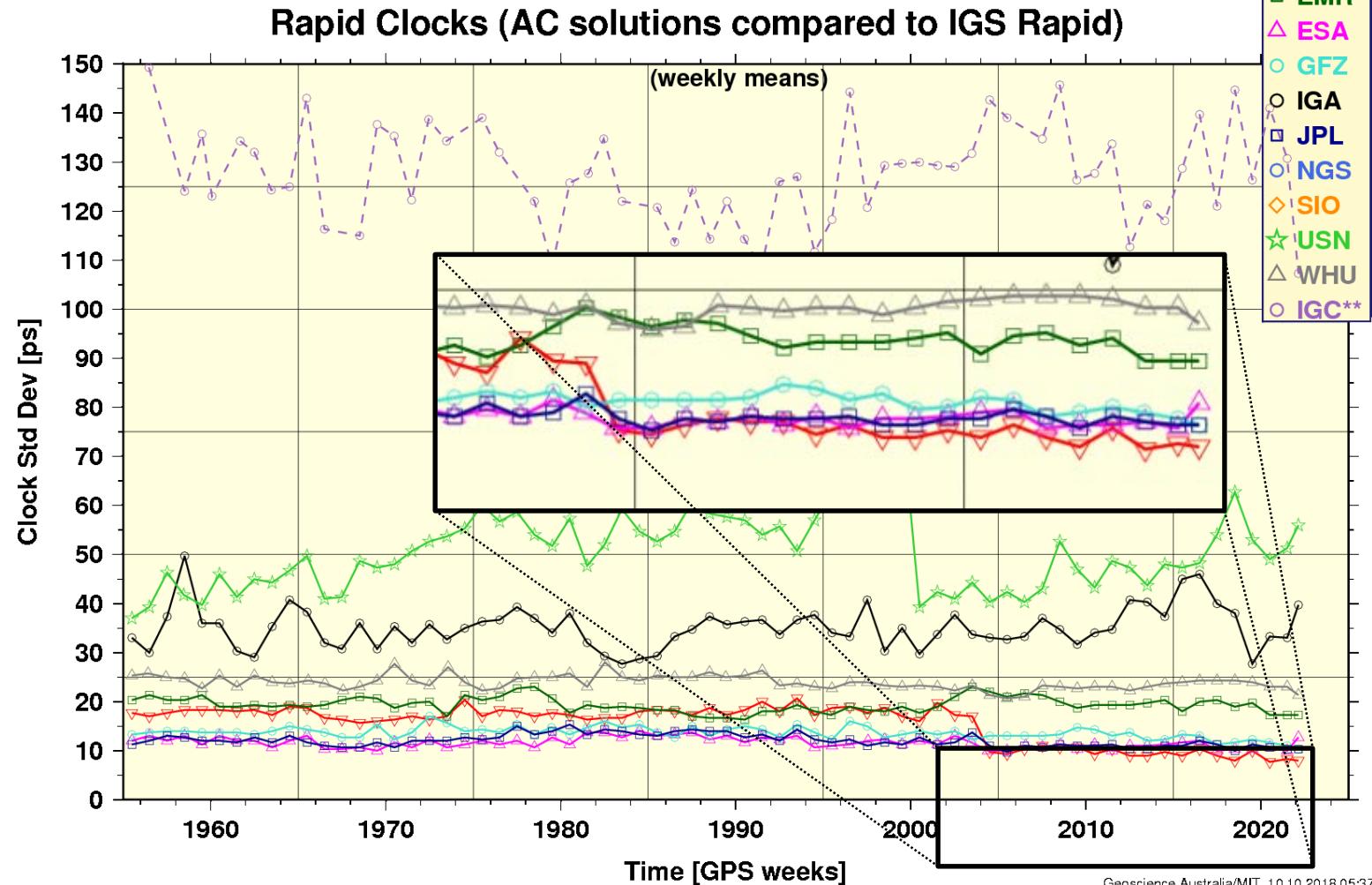
Galileo satellite widelane fractional bias results

Differences in the bias between day n and n+1, n+2, n+3, ...



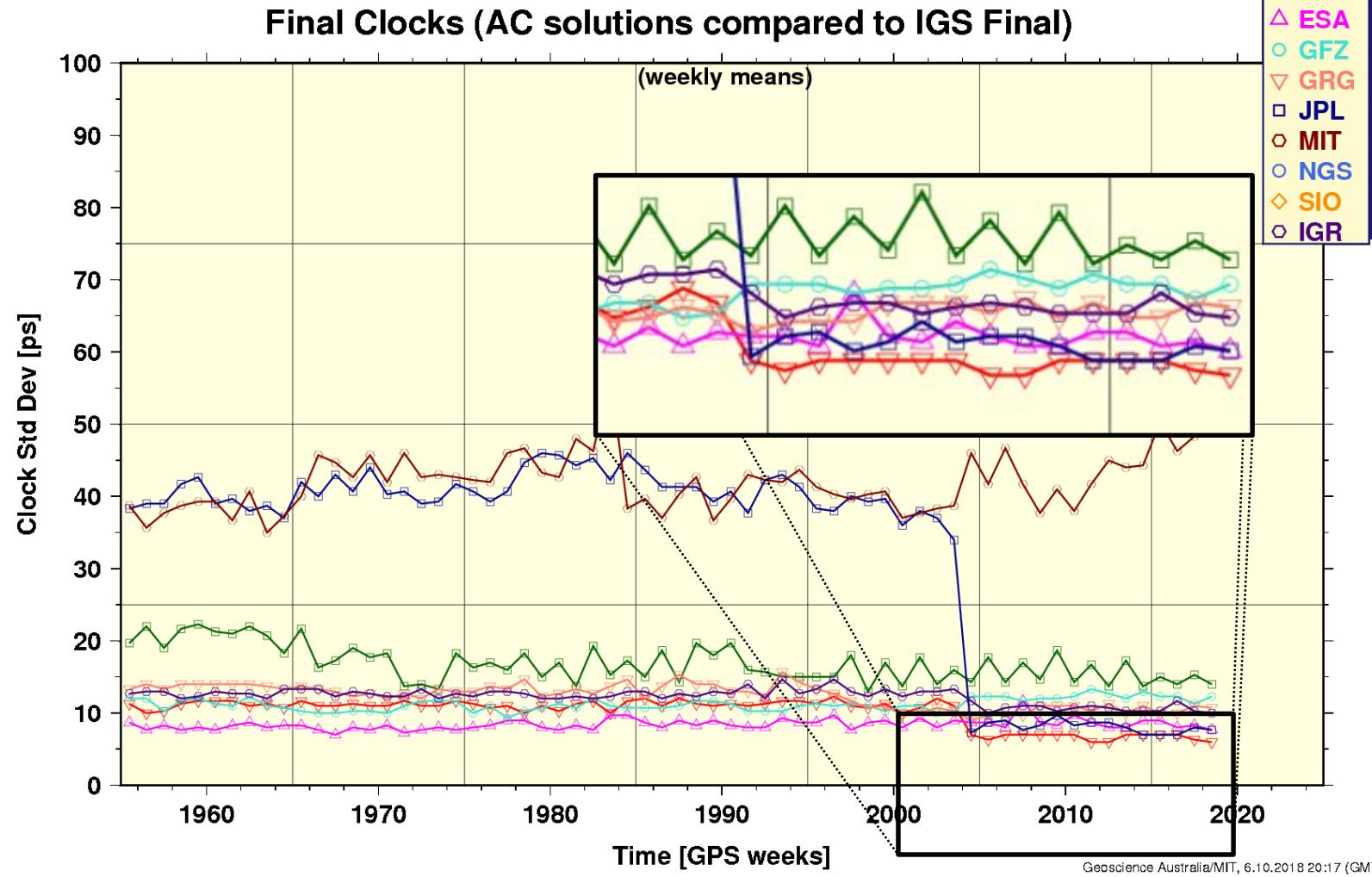
S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: **New ambiguity-fixed IGS clock analysis products at CODE**. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

IGS rapid and final clock combination



S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: New ambiguity-fixed IGS clock analysis products at CODE. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

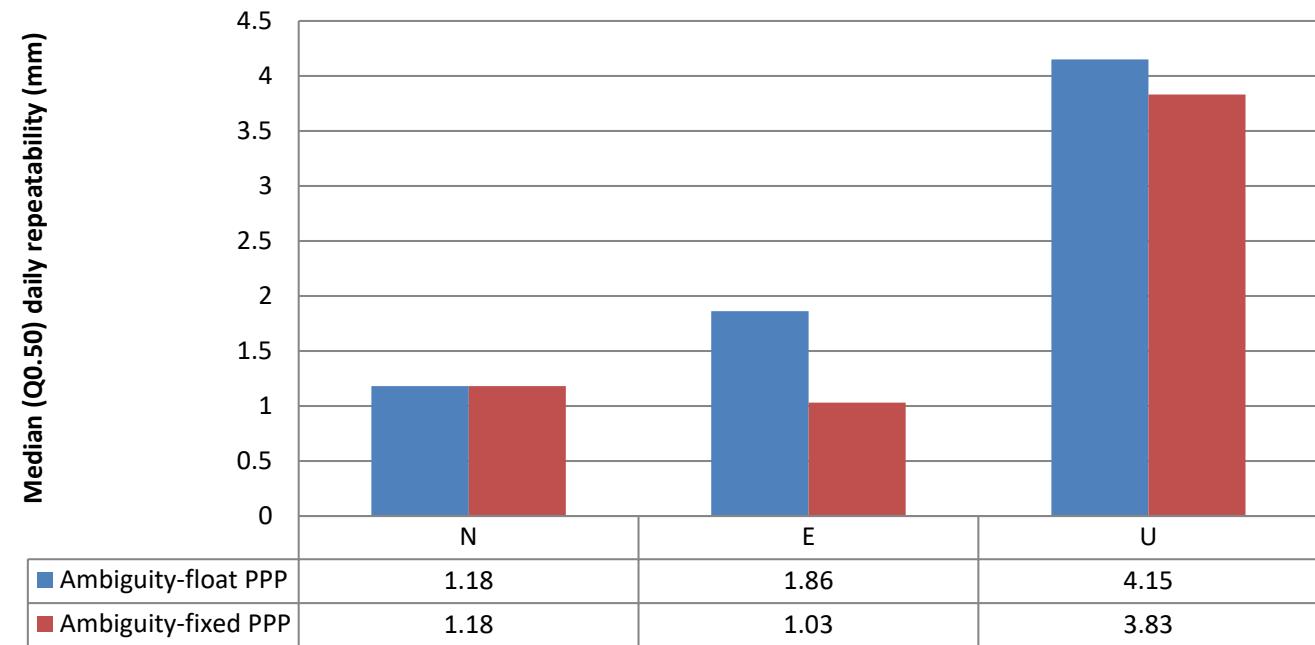
IGS final clock combination



S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: New ambiguity-fixed IGS clock analysis products at CODE. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

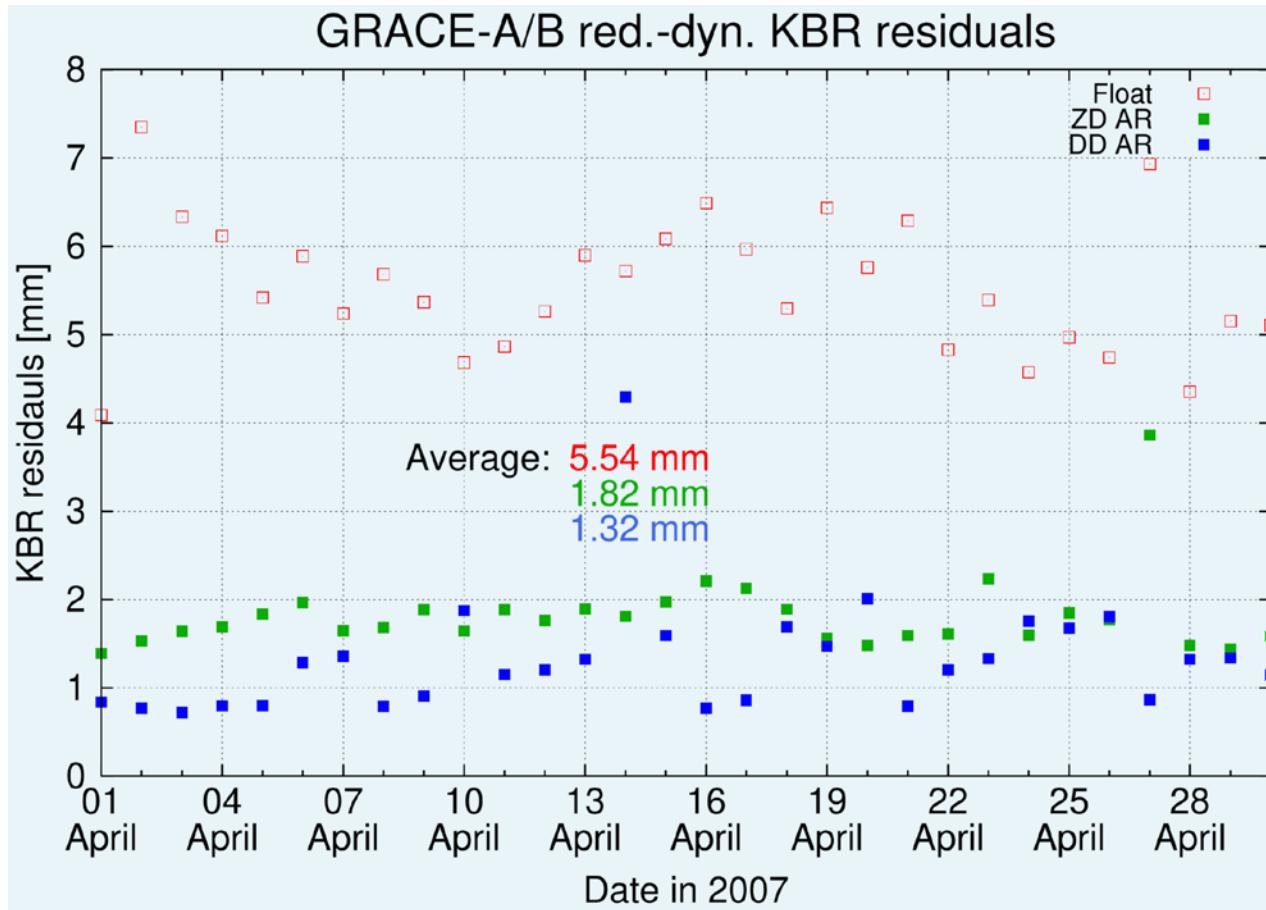
Daily PPP vs. daily IPPP

Daily PPP; CODE final product; September 2018;
295 (of 337) stations



S. Schaer, A. Villiger, D. Arnold, R. Dach, A. Jäggi, L. Prange: New ambiguity-fixed IGS clock analysis products at CODE. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

IPPP for LEO POD



Daily RMS values of K-band residuals.

D. Arnold, S. Schaer, A. Villiger, R. Dach, A. Jäggi: Undifferenced ambiguity resolution for GPS-based POD of LEO using the new CODE bias and clock product. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

IPPP for LEO POD

| Orbits | Float | | ZD AR | |
|-------------|-----------|-----------|-----------|-----------|
| | red.-dyn. | kin. | red.-dyn. | kin. |
| GRACE-A | +0.5/15.5 | +1.5/16.6 | +2.5/12.4 | +2.6/12.0 |
| GRACE-B | +0.9/12.1 | -0.5/16.9 | +3.8/8.5 | +3.7/9.6 |
| Sentinel-3A | -6.0/11.5 | -6.5/14.7 | -5.7/10.7 | -5.4/11.9 |
| Sentinel-3B | -2.9/12.4 | -4.3/15.2 | -3.5/10.4 | -3.3/11.1 |

Table 1: Mean values and standard deviations in mm of SLR residuals over April 2007 (GRACE) and September 2018 (Sentinel-3), respectively.

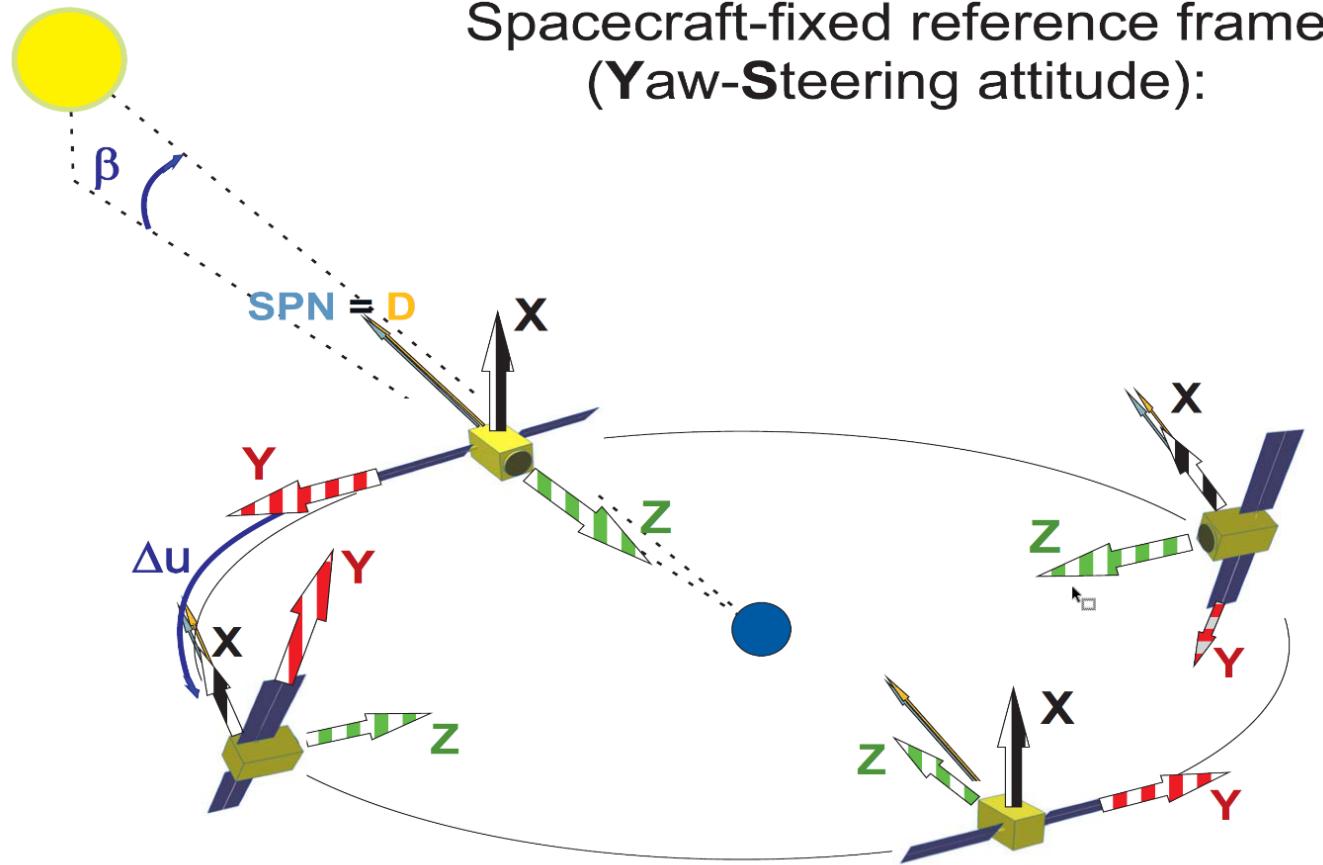
D. Arnold, S. Schaer, A. Villiger, R. Dach, A. Jäggi: Undifferenced ambiguity resolution for GPS-based POD of LEO using the new CODE bias and clock product. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

CODE contribution to IGS MGEX

Recent Improvements in the CODE MGEX

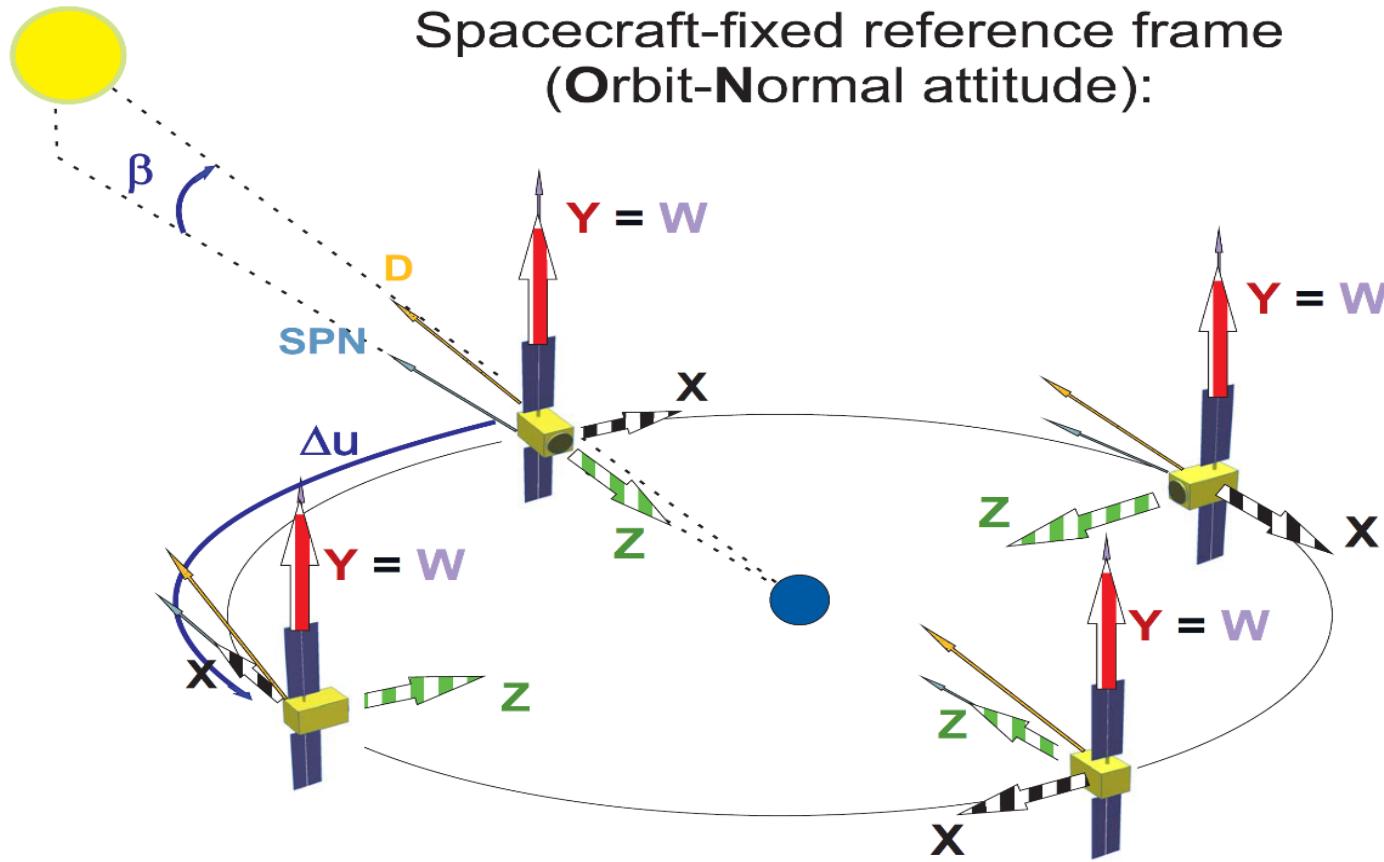
- Improved ambiguity resolution for orbit product, based on CODE's OSB product
- Activation of eclipse attitude law for Galileo
- Albedo and antenna thrust models for Galileo and QZSS
- Higher sampling of orbit (5 min) and clock products (30 s)
- Zero-diff. ambiguity resolution for clock product
- Activation of the orbit normal mode modelling

Orientation of the spacecraft



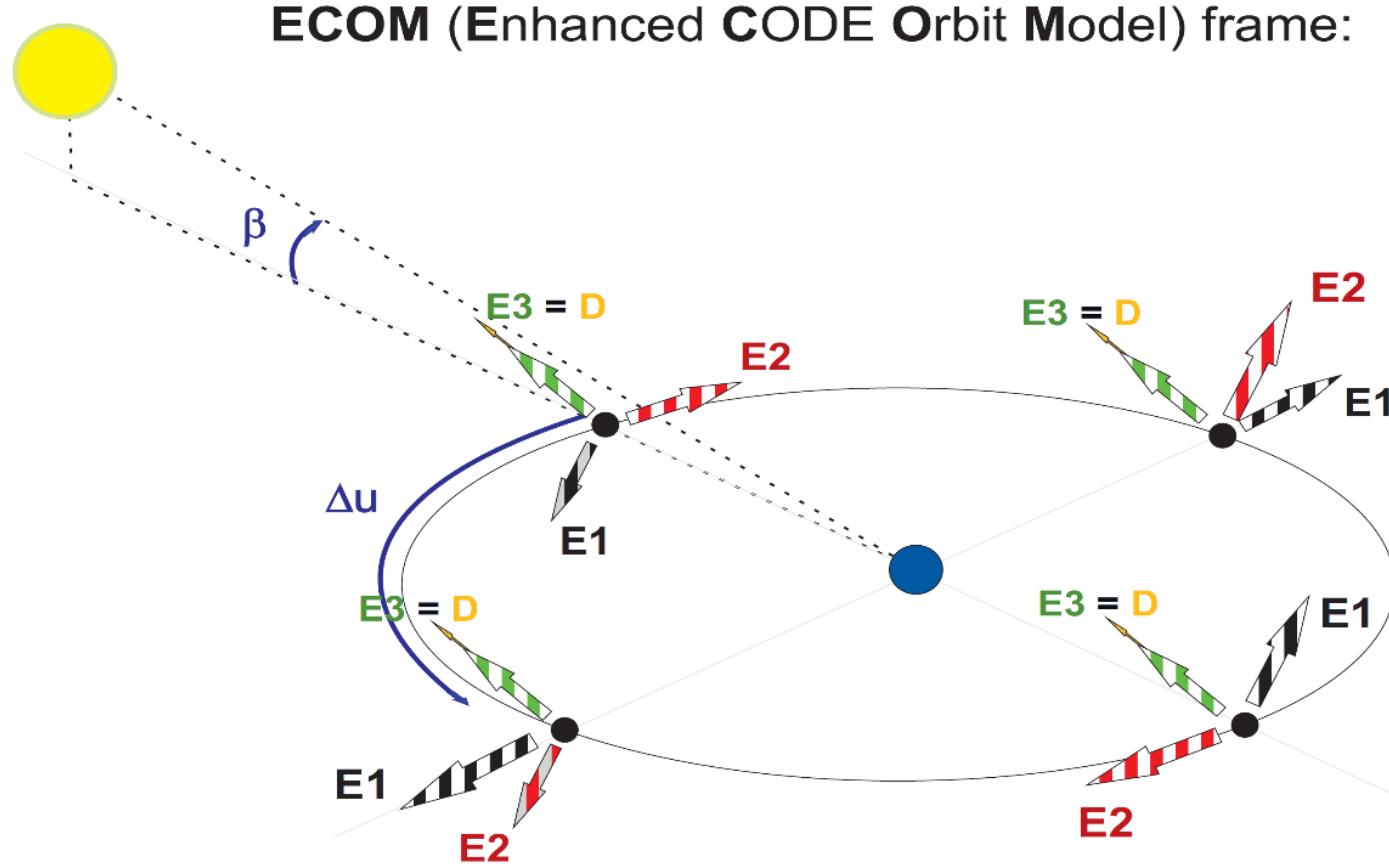
L. Prange, R. Dach, D. Arnold, G. Beutler, S. Schaer, A. Villiger, A. Jäggi: An Empirical SRP Model for the Orbit Normal Mode. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Orientation of the spacecraft



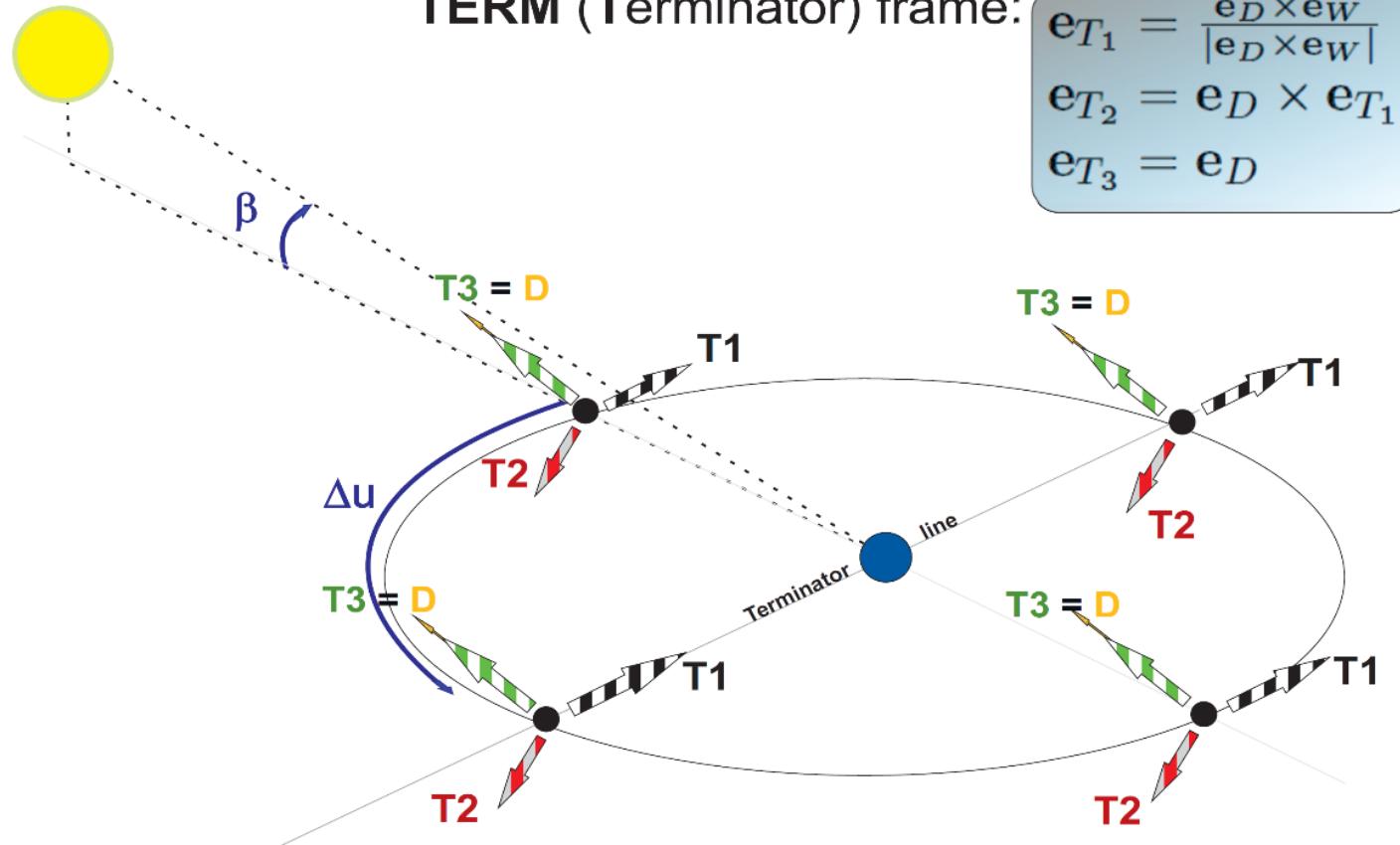
L. Prange, R. Dach, D. Arnold, G. Beutler, S. Schaer, A. Villiger, A. Jäggi: An Empirical SRP Model for the Orbit Normal Mode. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Orientation of the coordinate system



L. Prange, R. Dach, D. Arnold, G. Beutler, S. Schaer, A. Villiger, A. Jäggi: An Empirical SRP Model for the Orbit Normal Mode. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Orientation of the coordinate system



L. Prange, R. Dach, D. Arnold, G. Beutler, S. Schaer, A. Villiger, A. Jäggi: An Empirical SRP Model for the Orbit Normal Mode. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

ECOM updated for orbit normal mode

RMS from SLR residuals (IQR):

| | BDS2-MEO | BDS2-IGSO | QZSS-1 |
|-------------|----------|-----------|---------|
| Old model | 20.5 cm | 21.0 cm | 62.0 cm |
| New model | 12.2 cm | 12.2 cm | 15.2 cm |
| Improvement | 40.5 % | 41.9% | 75.5% |

Median of a linear fit of the satellite clock corrections:

| | BDS2-MEO | BDS2-IGSO | QZSS-1 |
|-------------|----------|-----------|---------|
| Old model | 1.72 ns | 1.61 ns | 1.43 ns |
| New model | 0.72 ns | 0.69 ns | 0.35 ns |
| Improvement | 58.1% | 57.1% | 75.5% |

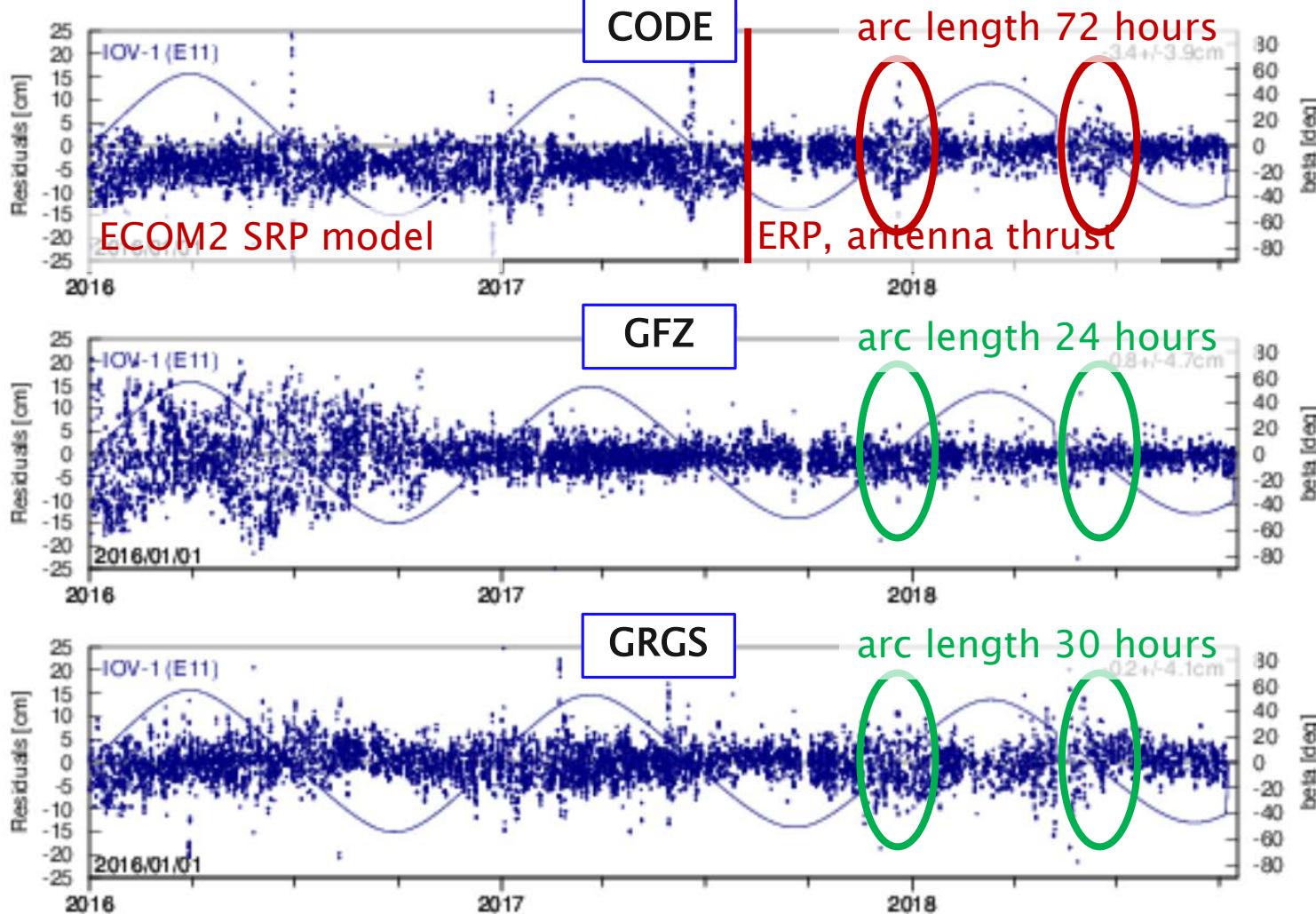
L. Prange, R. Dach, D. Arnold, G. Beutler, S. Schaer, A. Villiger, A. Jäggi: An Empirical SRP Model for the Orbit Normal Mode. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Orbit modelling during eclipse

D. Sidorov, R. Dach, L. Prange, A. Jäggi: Improved orbit modelling of Galileo satellites during eclipse seasons.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

SLR residuals for SVN 101

Comparison of MGEX solutions from <http://mgex.igs.org/analysis>

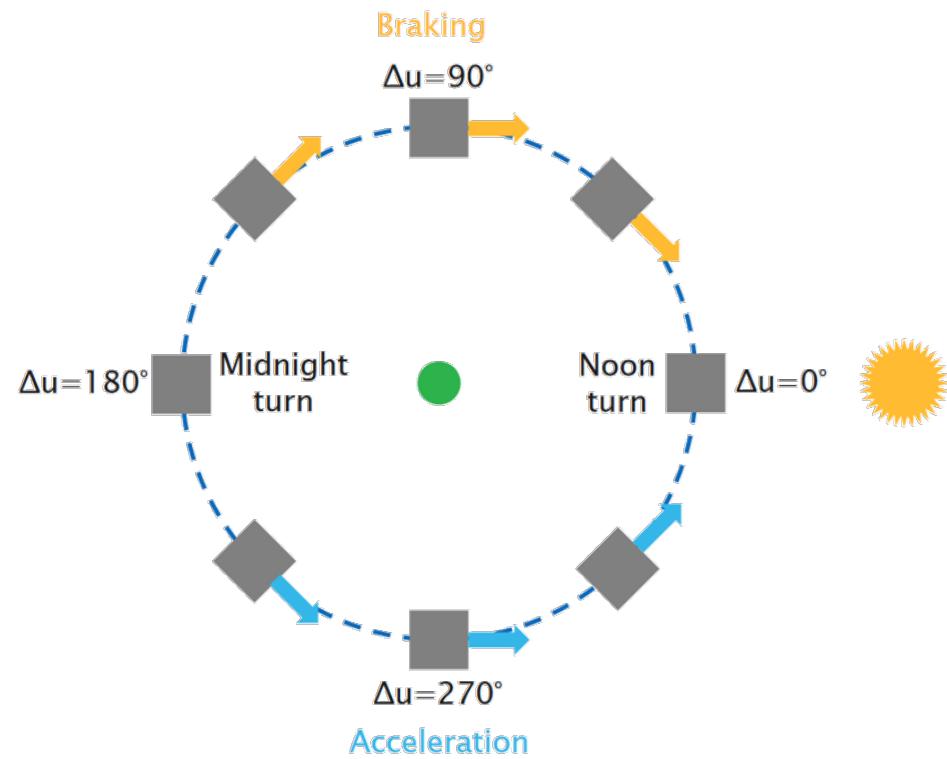
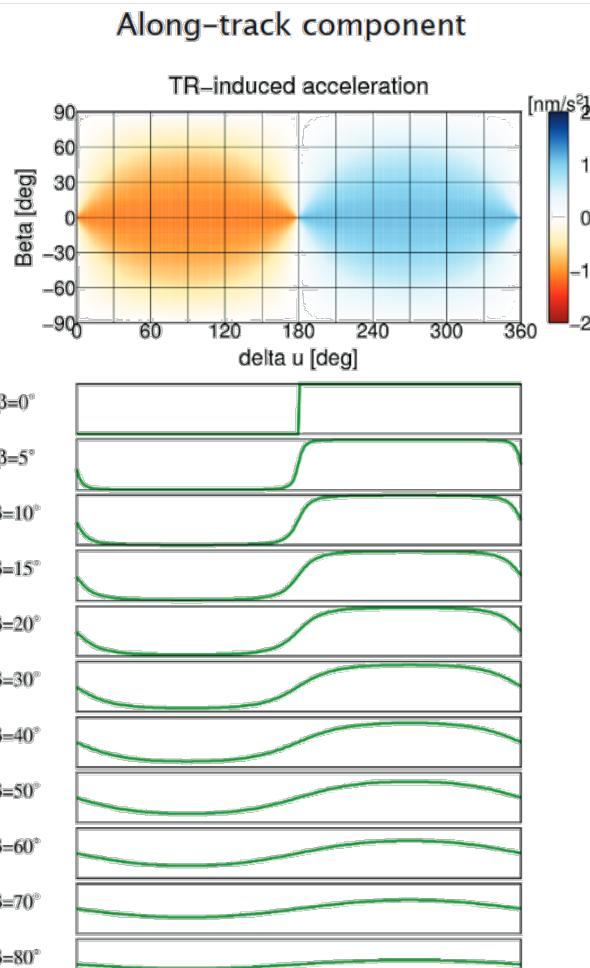


Consequences of the satellite design

- The thermal radiation from the radiators on the +Y/-Y plates cause forces compensating each other (or introducing a Y-bias).
- For Galileo we have more radiators causing additional forces (+X and -Z for FOC).
- Because of the low weight of Galileo satellites they are more sensitive to non-gravitational forces.
- Thermal radiation from the radiators are also active during eclipse periods where empirical SRP parameters are switched off.

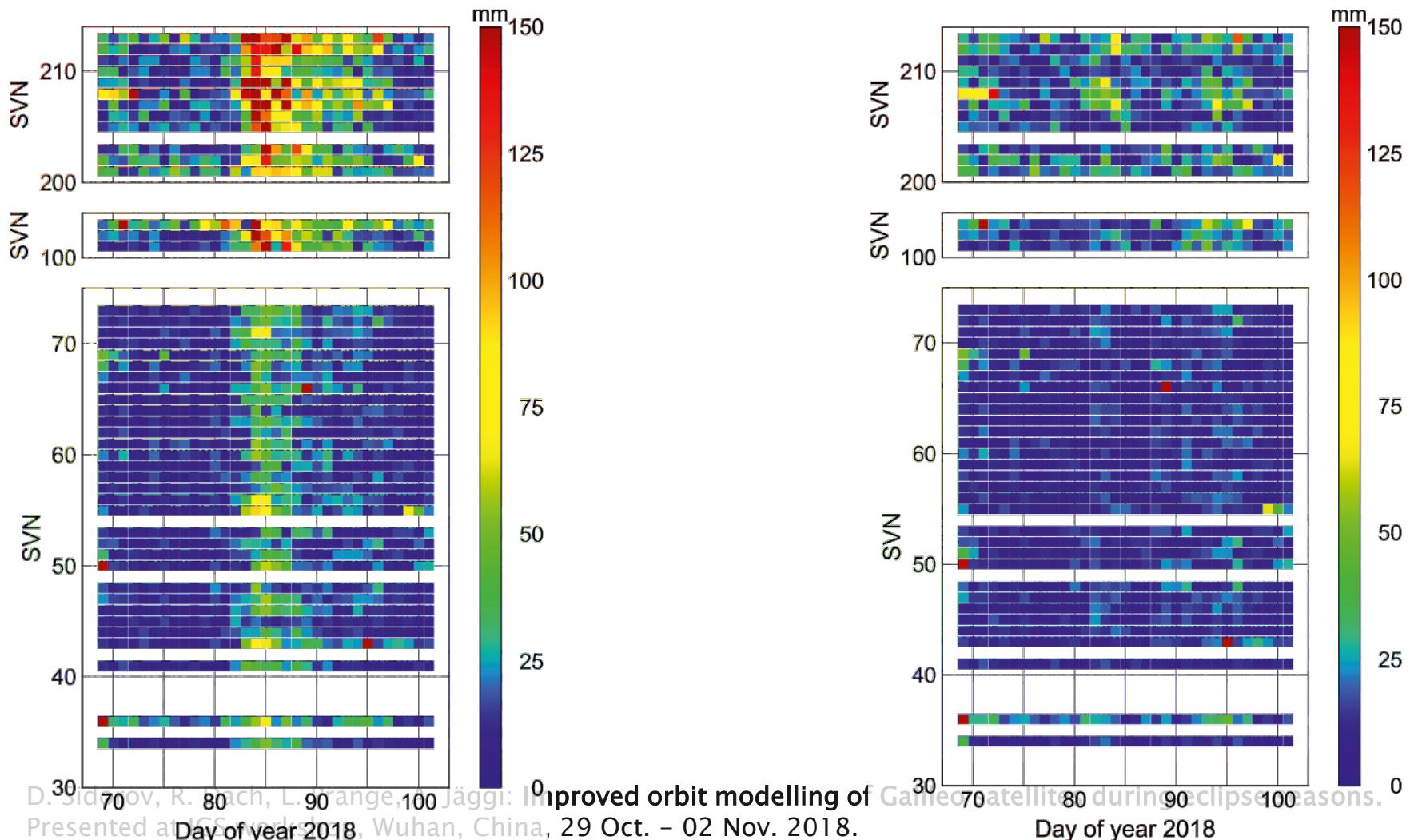
D. Sidorov, R. Dach, L. Prange, A. Jäggi: Improved orbit modelling of Galileo satellites during eclipse seasons.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Expected effect of the +X radiator

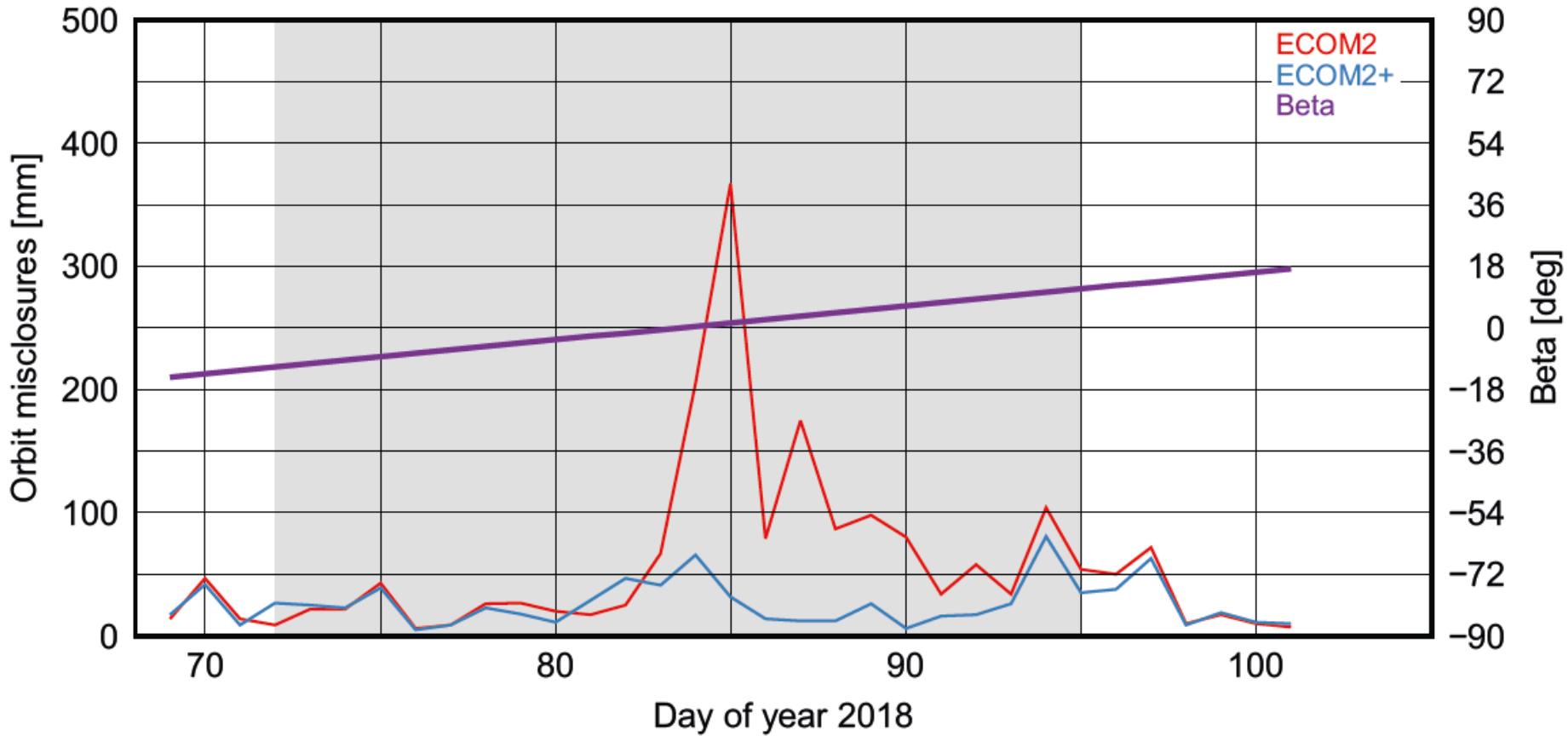


D. Sidorov, R. Dach, L. Prange, A. Jäggi: Improved orbit modelling of Galileo satellites during eclipse seasons.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Orbit misclosures at midnight



Example: IOV satellite E11



D. Sidorov, R. Dach, L. Prange, A. Jäggi: Improved orbit modelling of Galileo satellites during eclipse seasons.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Scaling factors for box-wing models

L. McNair, A. Villiger, R. Dach, A. Jäggi: Validation of boxwing models for GNSS satellites.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

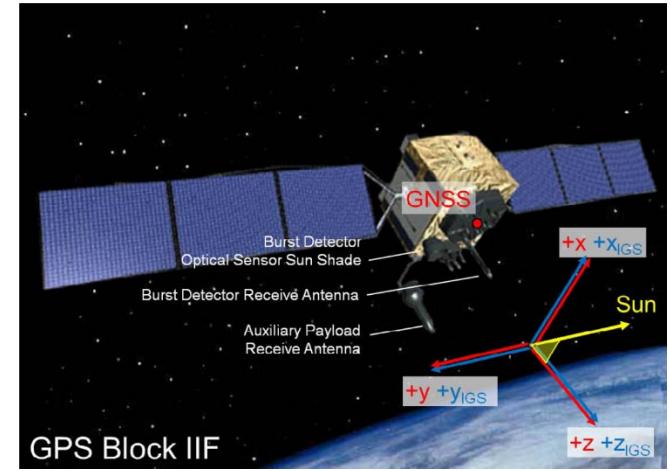
Validate boxwing model

Macromodel defines:

- Plates of the satellite with its areas and surface properties

Used to compute forces acting on the satellite because of **solar radiation pressure**.

Whether these models are correct can be assessed by **estimating scale factors** for the resulting force:

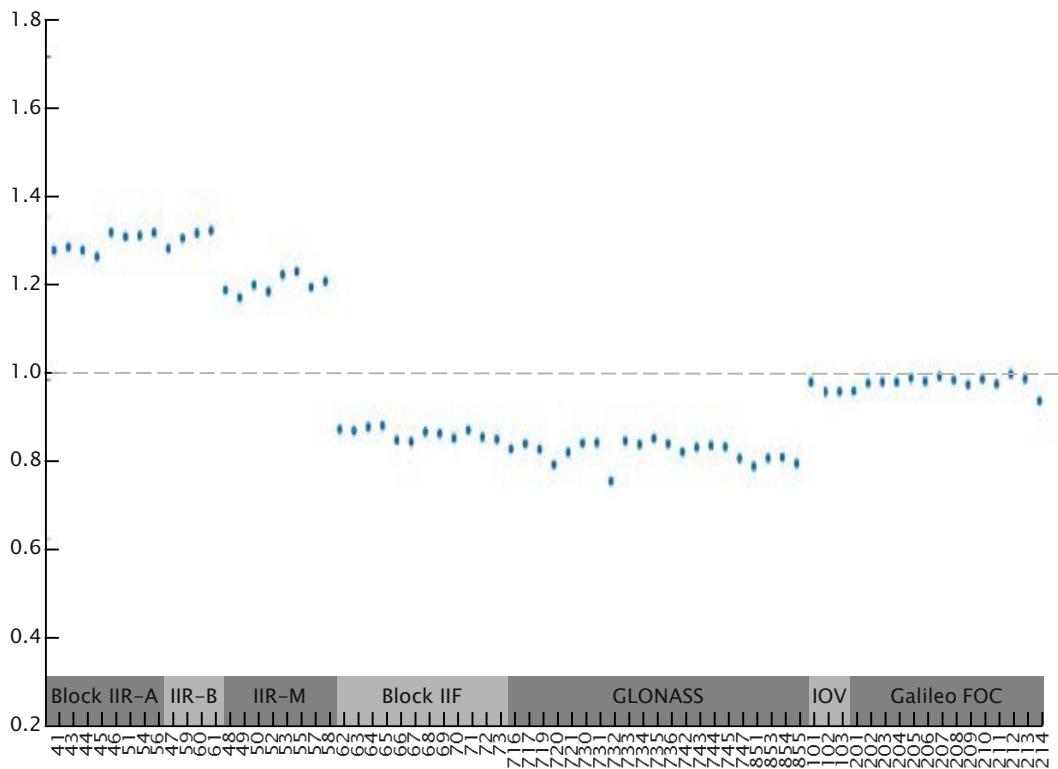


[Montenbruck et al, 2015. Adv. In Space Research]

| Plate | Mod | Area (A) [m^2] | Normal (\vec{e}_n) | Specularity (ρ) | Diffusivity (δ) | Rotation Sys. | Description |
|-------|-----|-------------------------------|------------------------|------------------------|--------------------------|-----------------|--------------------|
| 1 | 1 | 5.720 | [+1, 0, 0] | 0.112 | 0.448 | | +X |
| 2 | 1 | 5.720 | [-1, 0, 0] | 0.112 | 0.448 | | -X |
| 3 | 1 | 7.010 | [0, +1, 0] | 0.112 | 0.448 | | +Y |
| 4 | 1 | 7.010 | [0, -1, 0] | 0.112 | 0.448 | | -Y |
| 5 | 1 | 5.400 | [0, 0, +1] | 0.112 | 0.448 | | +Z |
| 6 | 1 | 5.400 | [0, 0, -1] | 0.000 | 0.000 | | -Z |
| 7 | 0 | 22.250 | [+1, 0, 0] | 0.195 | 0.035 | +SUN: [0,+1, 0] | Solar panels front |
| 8 | 0 | 22.250 | [-1, 0, 0] | 0.196 | 0.034 | -SUN: [0,+1, 0] | Solar panels back |

L. McNair, A. Villiger, R. Dach, A. Jäggi: Validation of boxwing models for GNSS satellites.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Yearly Scale Factors: Monoscale

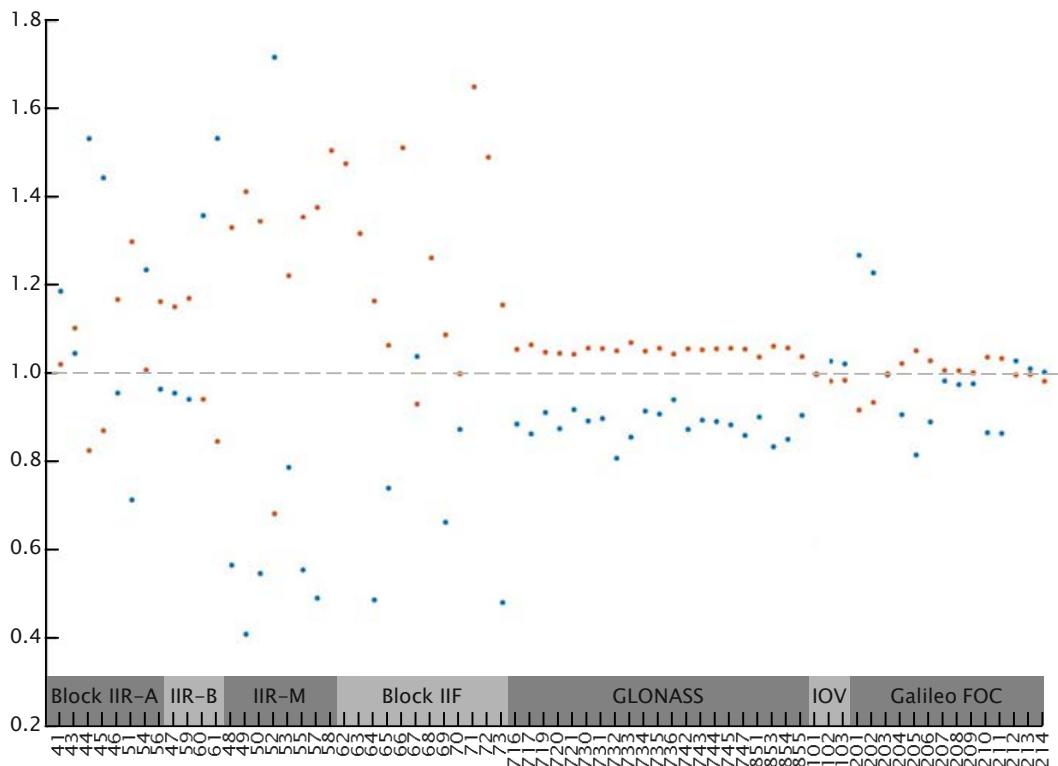


Monoscale:
(one factor per satellite)

The scale factors
show clearly the
different types of
satellites.

L. McNair, A. Villiger, R. Dach, A. Jäggi: Validation of boxwing models for GNSS satellites.
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Yearly Scale Factors: Smartscale-2



Smartscale-2:
(two factor per satellite:
solar panel and **body**)

GLONASS & Galileo:
stable scale factors
for all satellites in
same block
-> close to 1

GPS:
more variation
between satellites in
same block
-> farther away
from 1.

L. McNair, A. Villiger, R. Dach, A. Jäggi: Validation of boxwing models for GNSS satellites.
Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Deficiencies in the Receiver Antenna Calibration in an multi-GNSS environment

A. Villiger, L. Prange, R. Dach, F. Zimmermann, H. Kuhlmann, A. Jäggi: **Consistency of antenna products in the MGEX environment.** Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

IGS antenna pattern

Before Galileo and QZSS disclosed the satellite antenna corrections

| GNSS | Frq | Sat. | Rob. |
|------|-----|------|------|
| GPS | L1 | | |
| | L2 | | |
| | L5 | | |
| GLO | G1 | | |
| | G2 | | |
| | G3 | | |
| GAL | E1 | | |
| | E5a | | |
| | E5b | | |
| | E5 | | |
| | E6 | | |

| GNSS | Frq | Sat. | Rob. |
|------|-----|------|------|
| BDS | B1 | | |
| | B2 | | |
| | B3 | | |
| QZSS | L1 | | |
| | L2 | | |
| | L5 | | |

unknown estimated calibrated guess

Rob. : roboter calibrations

A. Villiger, L. Prange, R. Dach, F. Zimmermann, H. Kuhlmann, A. Jäggi: **Consistency of antenna products in the MGEX environment**. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

IGS antenna pattern

After Galileo and QZSS disclosed the satellite antenna corrections

| GNSS | Frq | Sat. | Rob. |
|------|-----|------|------|
| GPS | L1 | | |
| | L2 | | |
| | L5 | | |
| GLO | G1 | | |
| | G2 | | |
| | G3 | | |
| GAL | E1 | | |
| | E5a | | |
| | E5b | | |
| | E5 | | |
| | E6 | | |

| GNSS | Frq | Sat. | Rob. |
|------|-----|------|------|
| BDS | B1 | | |
| | B2 | | |
| | B3 | | |
| QZSS | L1 | | |
| | L2 | | |
| | L5 | | |

unknown estimated calibrated guess

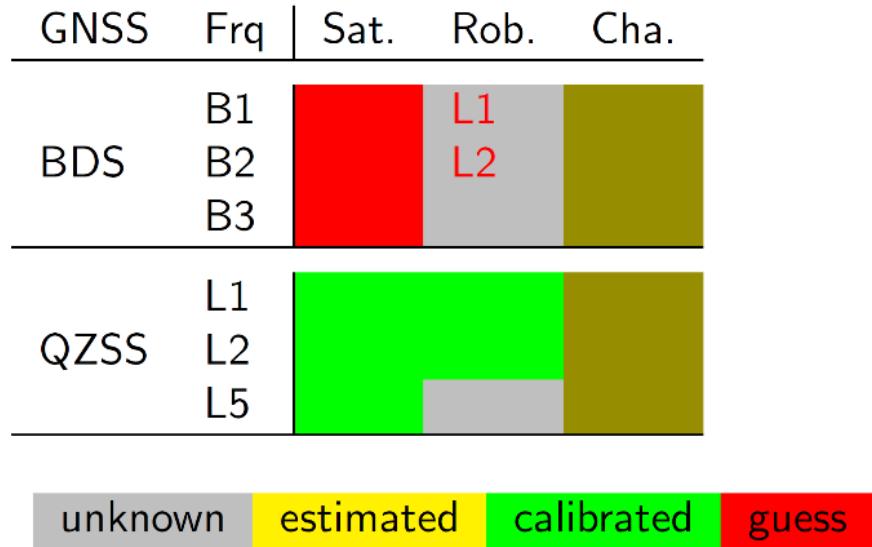
Rob. : roboter calibrations

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IGS antenna pattern

After Galileo and QZSS disclosed the satellite antenna corrections

| GNSS | Frq | Sat. | Rob. | Cha. |
|------|-----|--------|--------|-------|
| GPS | L1 | yellow | green | olive |
| | L2 | yellow | green | olive |
| | L5 | grey | | olive |
| GLO | G1 | yellow | green | olive |
| | G2 | yellow | green | olive |
| | G3 | grey | | olive |
| GAL | E1 | green | red L1 | olive |
| | E5a | green | red L2 | olive |
| | E5b | | | olive |
| | E5 | | | olive |
| | E6 | | | olive |



unknown estimated calibrated guess

Rob. : robooter calibrations

Cha. : chamber calibrations

A. Villiger, L. Prange, R. Dach, F. Zimmermann, H. Kuhlmann, A. Jäggi: **Consistency of antenna products in the MGEX environment**. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

Collection of chamber calibrations

- Antenna working group (A. Villiger, AIUB): call for chamber calibrations
- Great response from various institutions:
 - Vermessungsamt Mecklenburg–Vorpommern, Germany
 - Vermessung und Geoinformation Schleswig–Holstein, Germany
 - BKG
 - ESA
 - EUREF (publicly available)
 - GFZ

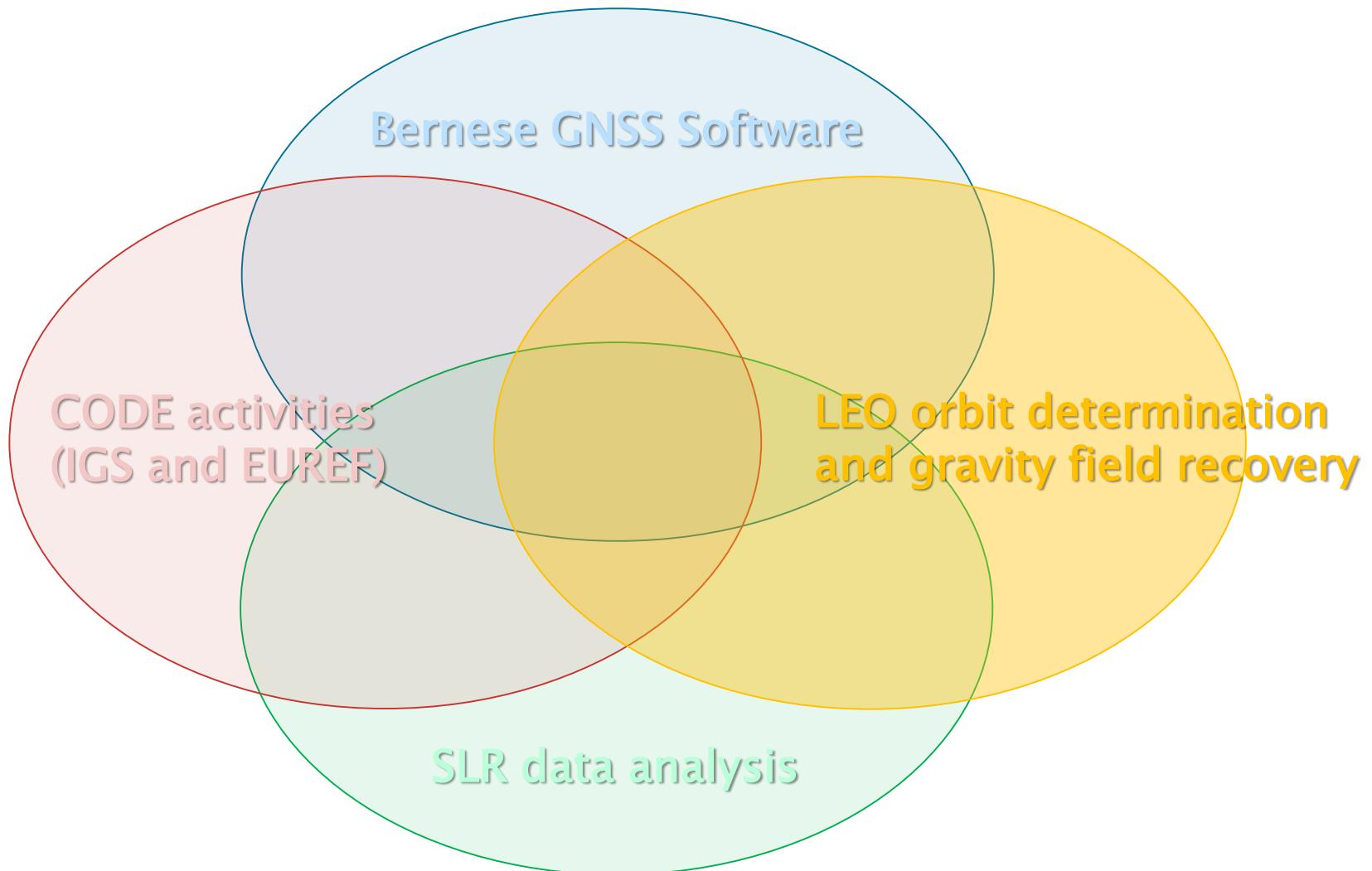
A. Villiger, A. Göttsche, M. Hugentobler, M. Hölzlmann, A. Jäggi: Consistency of antenna products in the MGEX environment. Presented at IGS workshop, Wuhan, China, 29 Oct. – 02 Nov. 2018.

ESA project related to GNSS activities

Other projects:

- TGVF/OVF: «Ground truth» for Galileo GMS
(continued with the label GRSP)
GSA-project with ESOC, BKG, GFZ, IGN
- ORBIT/SRP Modelling for Long Term Prediction
ESA-project with Airbus (defense and space)
- Improved GNSS-Based Precise Orbit Determination
by using highly accurate clocks
ESA-project with ETH Zurich and TU Munich
(finished in 2018)

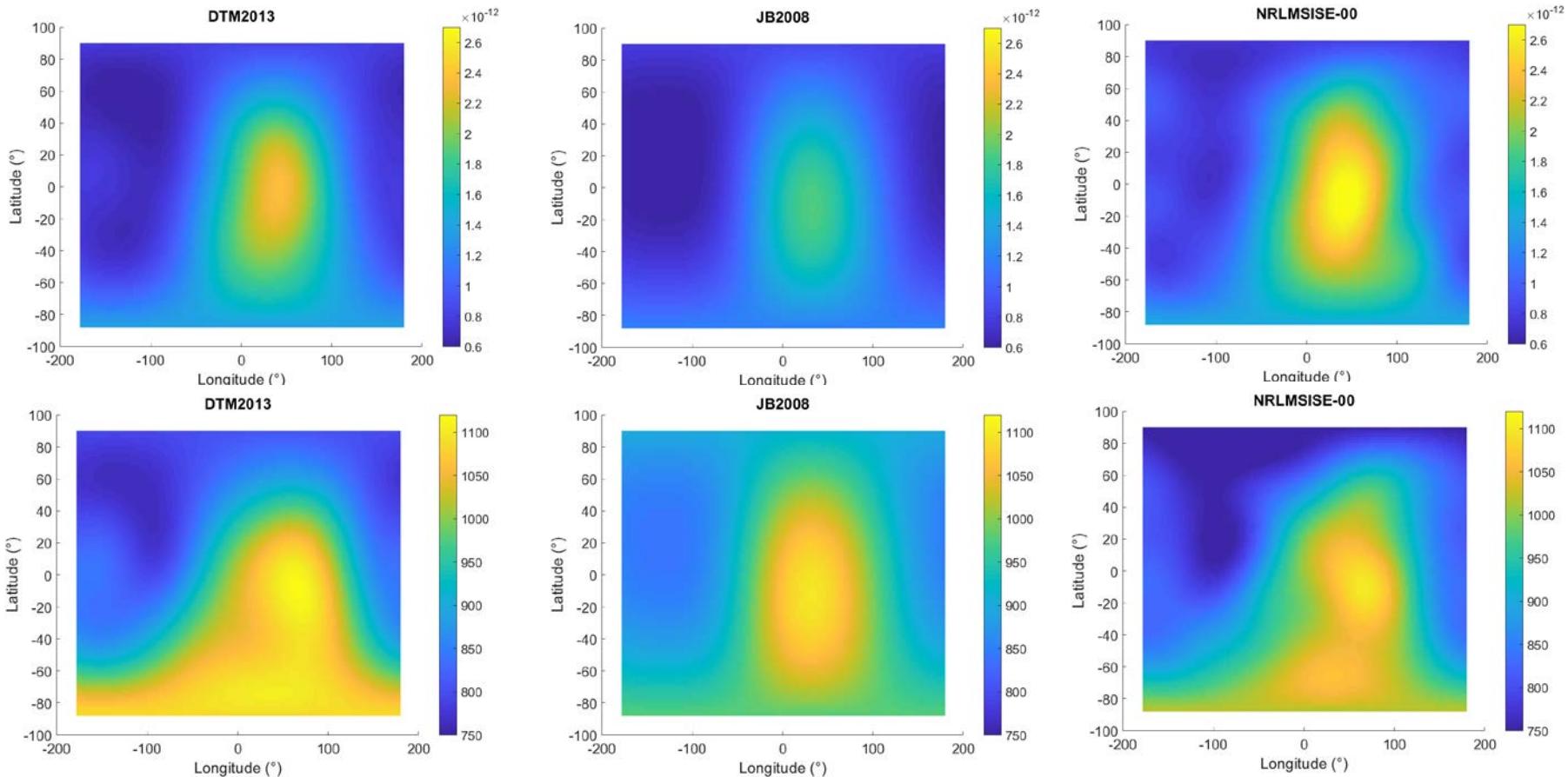
Satellite Geodesy Research Group



Atmospheric density models in LEO non-gravitational force modeling

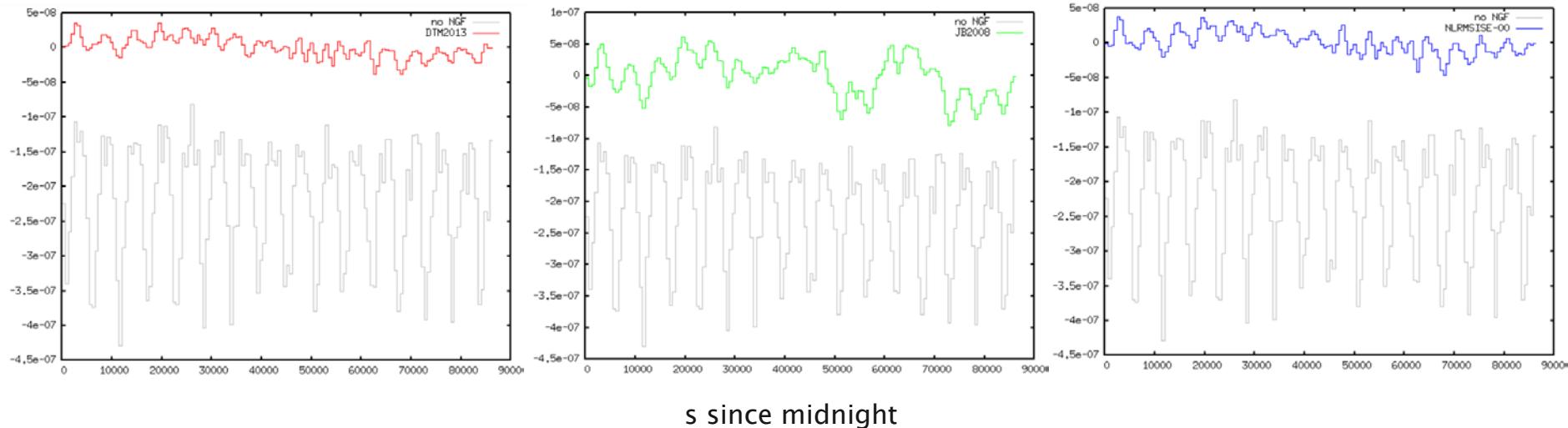
Thermospheric models

Atmospheric densities (top, kg/m³) and temperatures (bottom, K) provided by three different models at an altitude of 425 km (GRACE):



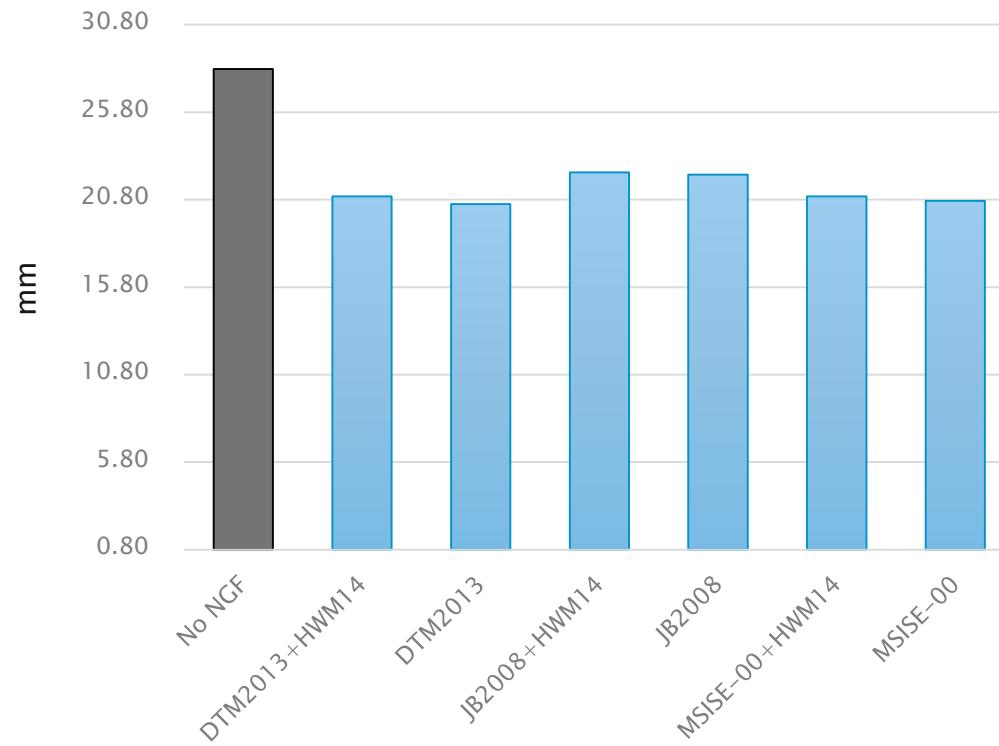
Piecewise-constant accelerations

Estimated along-track 10-min piecewise-constant accelerations (m/s^2) for GRACE -A (doy 222 of 2014) when using no non-gravitational force modeling, and **DTM2013**, **JB2008**, or **NRLMSISE-00** for aerodynamic acceleration modeling:



SLR residuals

Standard deviations of GRACE -A SLR residuals over 3 months when using different atmospheric density models, as well as with and without Horizontal Wind Model HWM14:



Copernicus POD Service



Copernicus satellite fleet

At AIUB precise orbits of all Sentinel satellites are computed



Sentinel-1A
Sentinel-1B



Sentinel-2A
Sentinel-2B

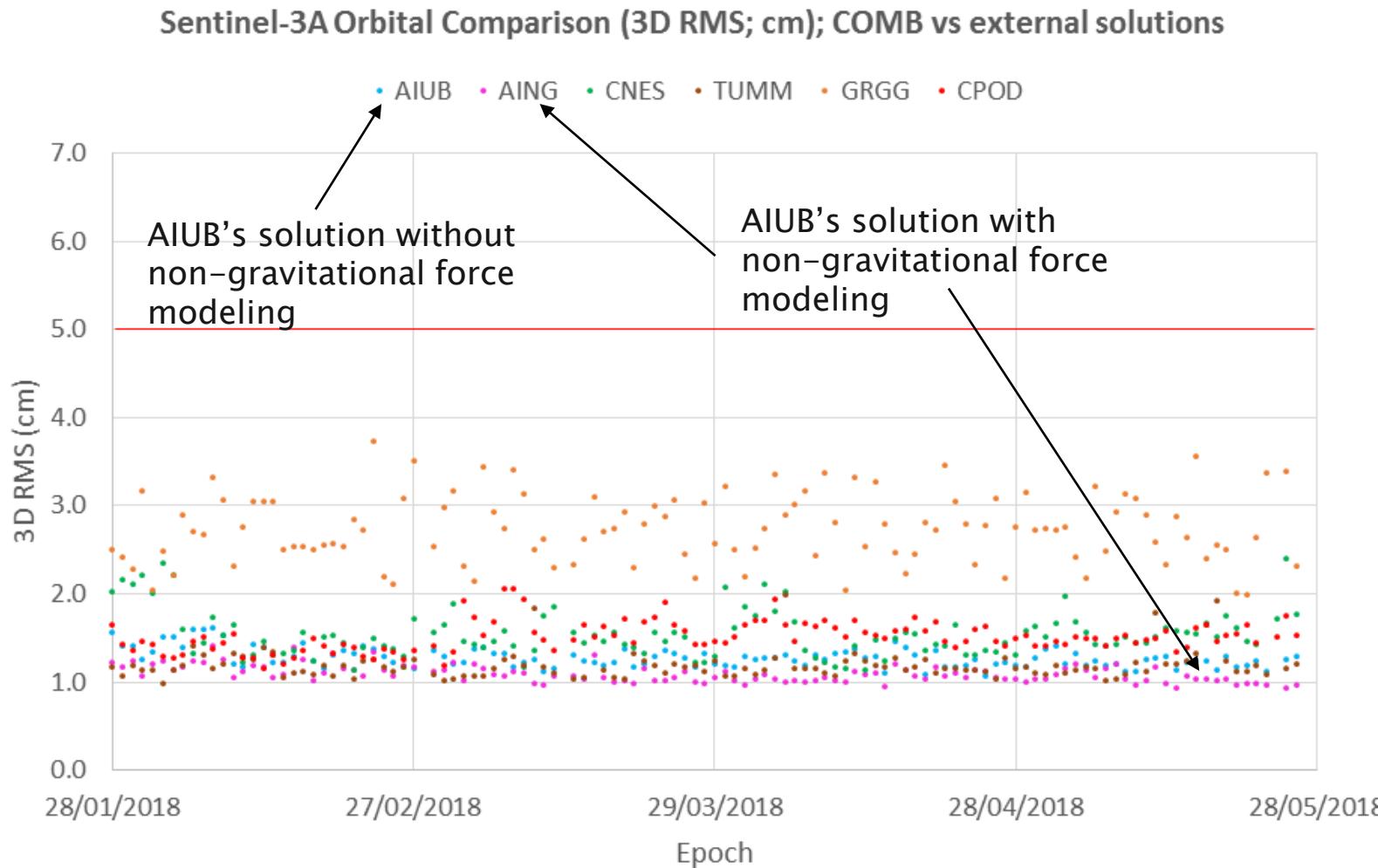


Sentinel-3A
Sentinel-3B



Courtesy: ESA

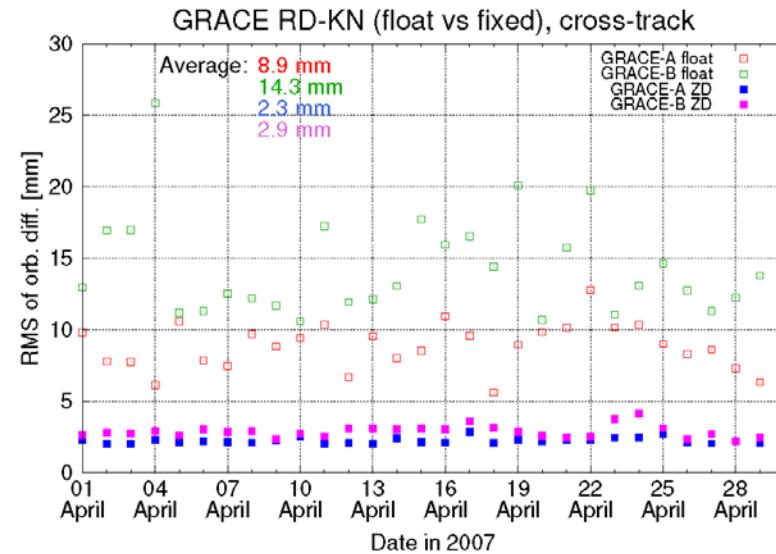
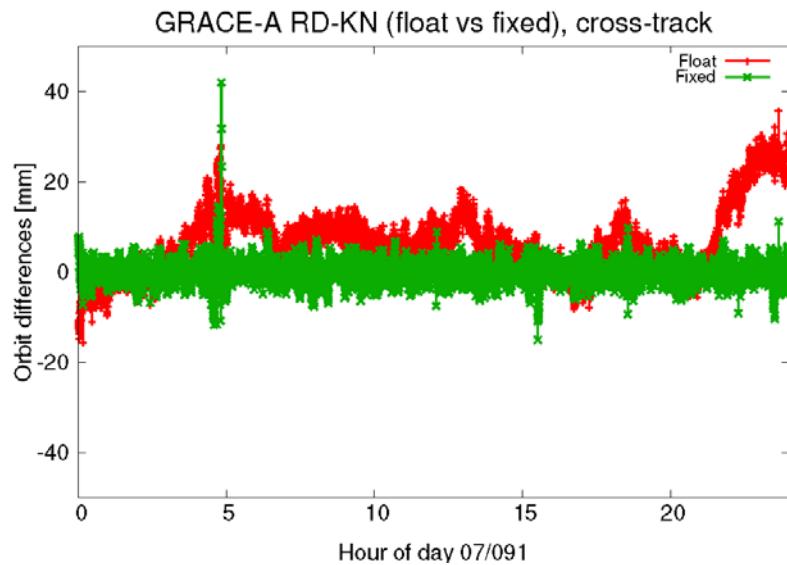
Sentinel-3A orbit comparisons



Zero-difference ambiguity resolution in LEO POD

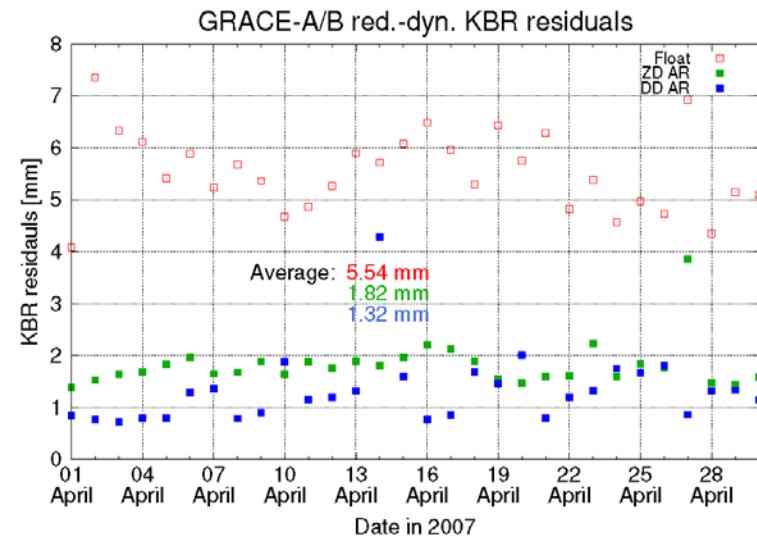
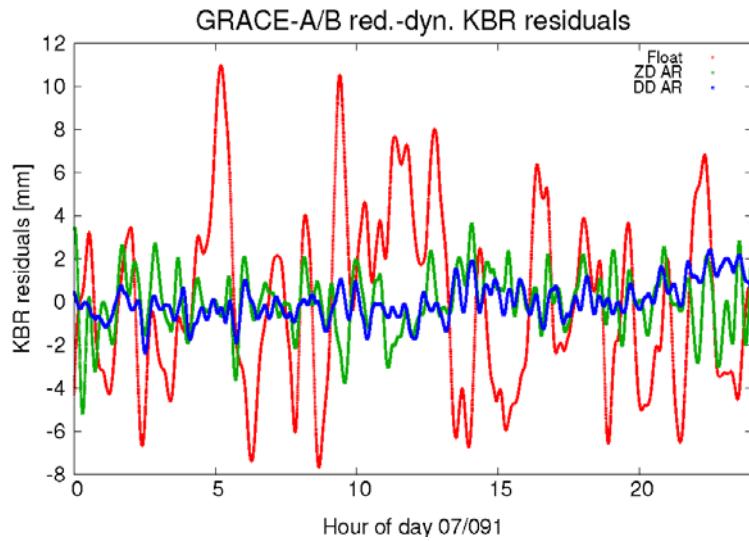
Internal orbit consistency

Based on CODE's new phase bias and clock product, zero-difference ambiguities can be resolved in LEO POD. This significantly improves the consistency between reduced-dynamic and kinematic orbit...



K-band validation

... and the K-band residuals (GRACE)

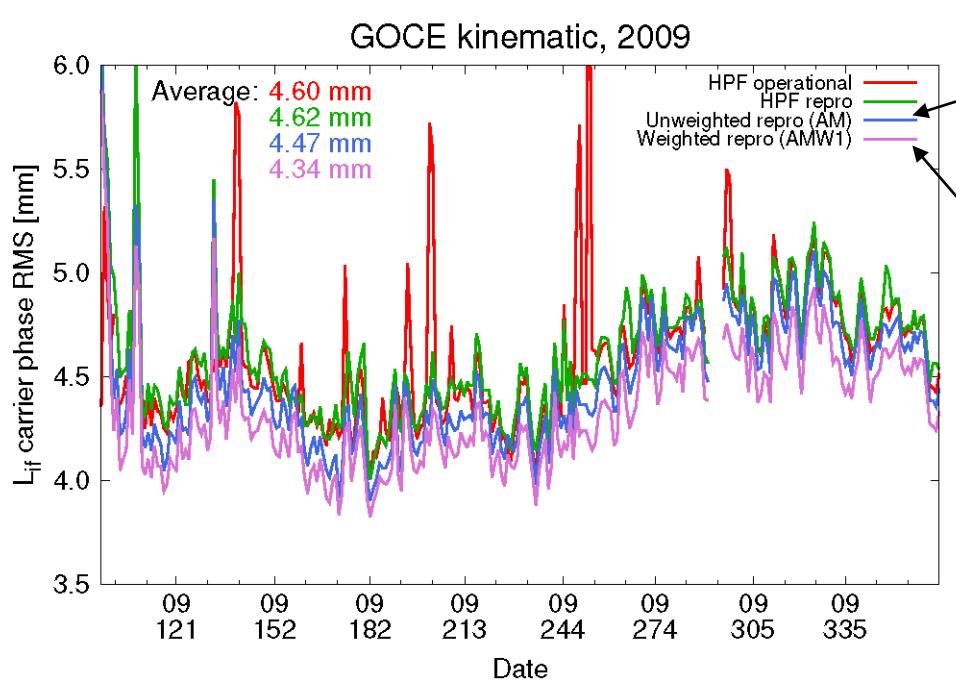


K-band residuals are comparable to what is obtained in a double-difference baseline processing with ambiguity resolution.

GOCE PSO Reprocessing

GOCE PSO reprocessing – Orbits

In the frame of an ESA–funded reprocessing of GOCE data, AIUB is responsible for the re-generation of the GOCE Precise Science Orbits (PSOs)

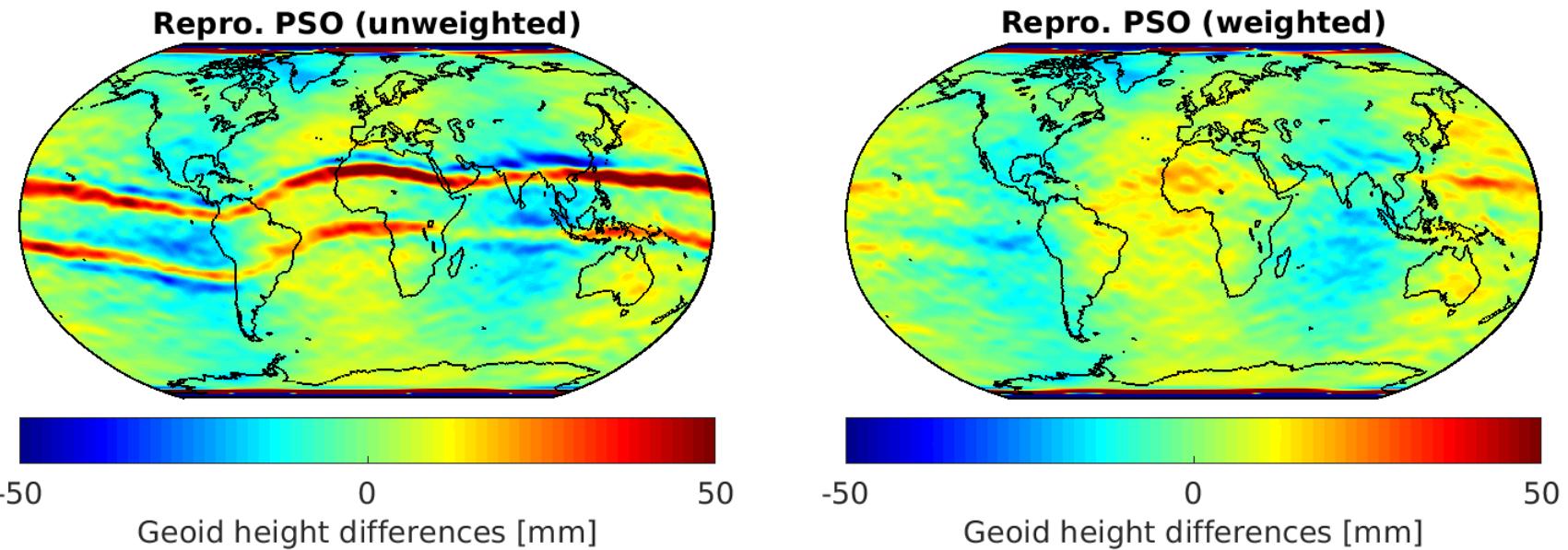


New PSOs, based on latest version of Bernese GNSS Software and homogeneous GNSS products from EGSIEM repro

In addition: Downweighting of GPS data affected by large ionospheric dynamics

GOCE PSO reprocessing – Gravity field

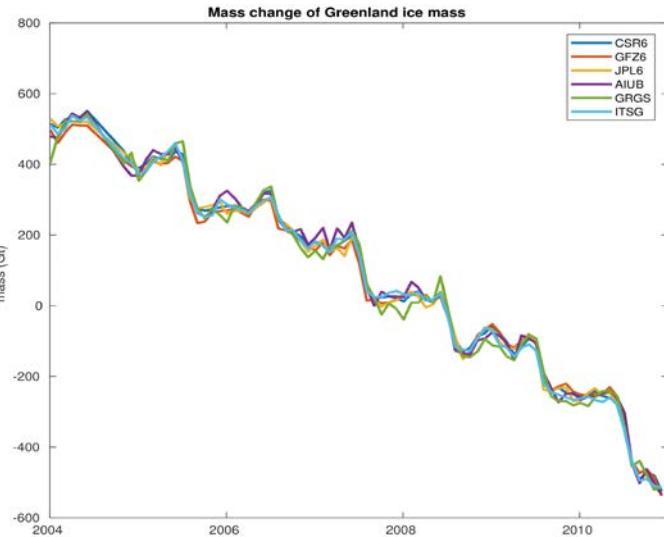
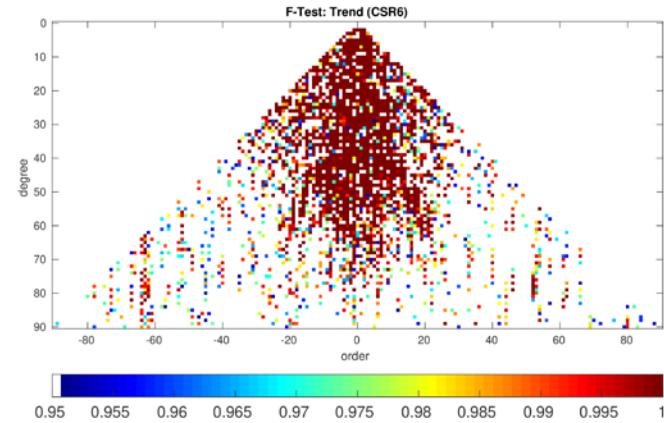
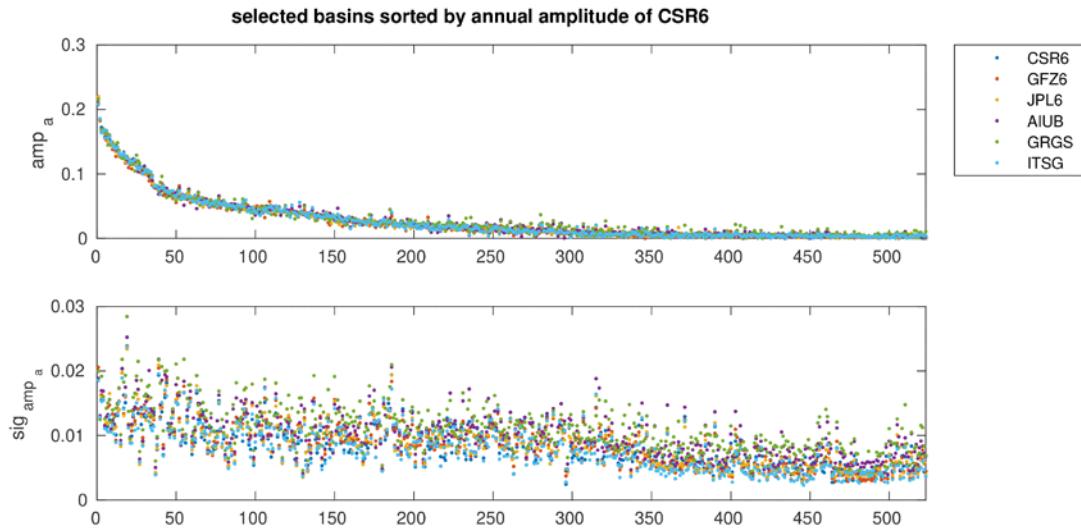
Data downweighting significantly reduces artifacts in GPS-only gravity field solutions along geomagnetic equator:



Geoid height differences of 2011 yearly GPS-only solution w.r.t. ITG-GRACE2010,
300km Gauss filter applied.

Follow-up activities from the EGSIEM project

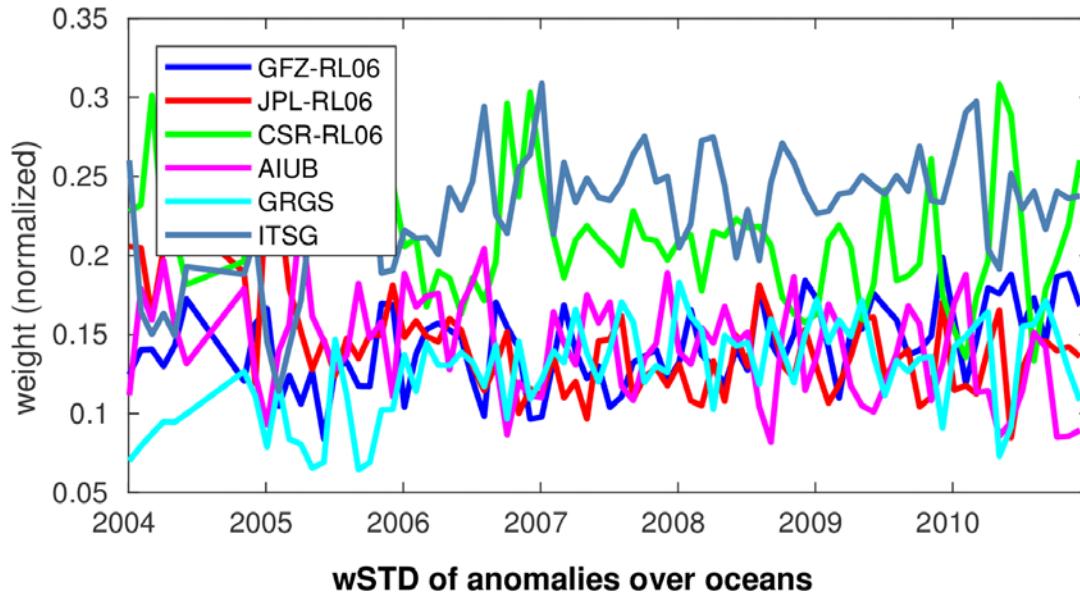
Product Center of the IAG



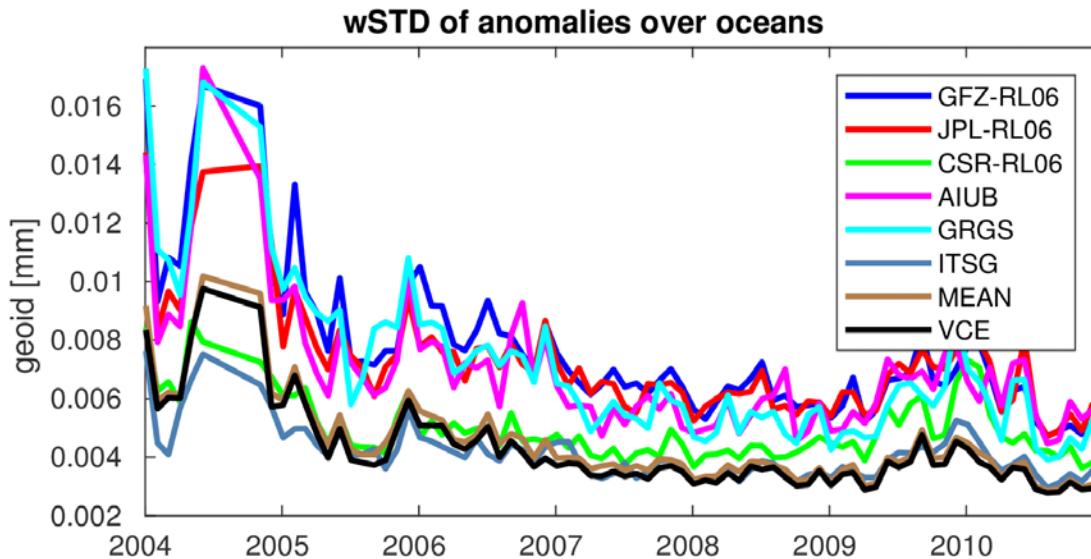
COST-G: Quality control

- Amplitudes of seasonal mass variations in river basins,
- Ice mass trends in polar regions,
- Significance tests for time variations in the spherical harmonic domain.

Product Center of the IAG

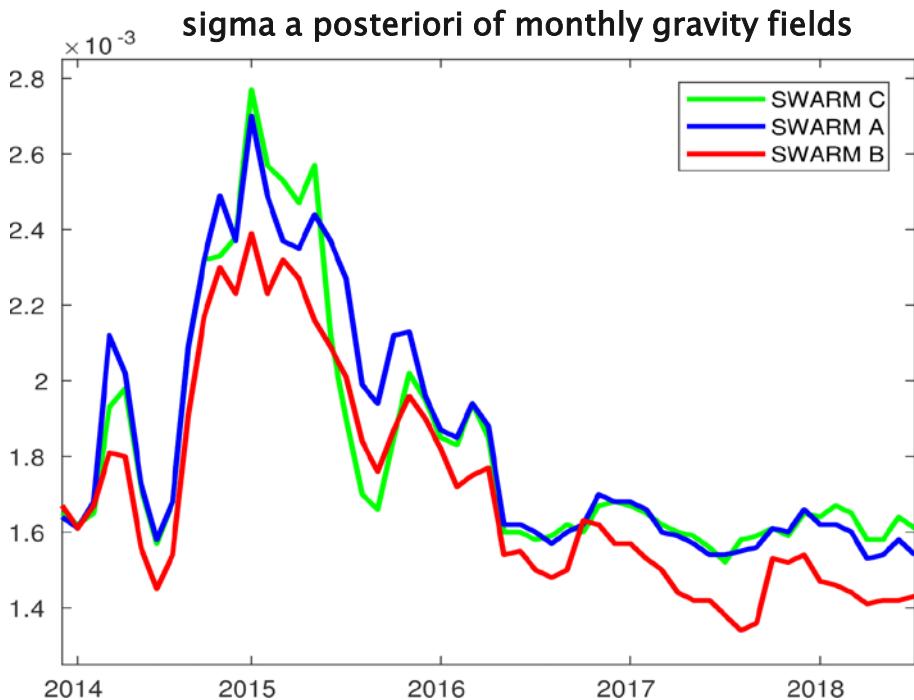


**Weighted combination
of monthly gravity
fields based on
variance component
estimation.**

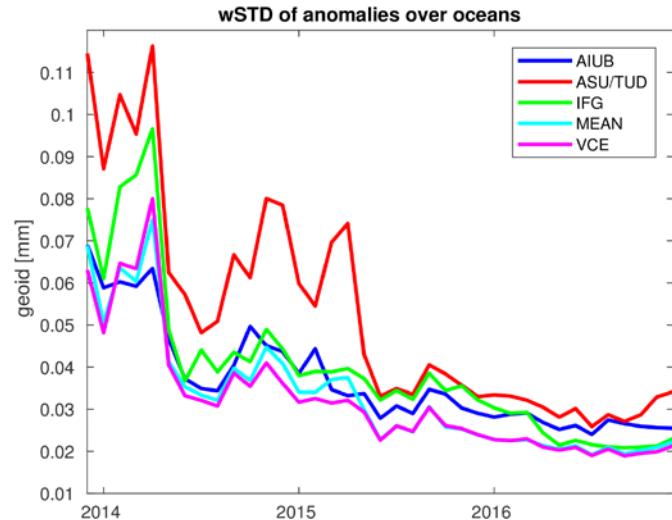
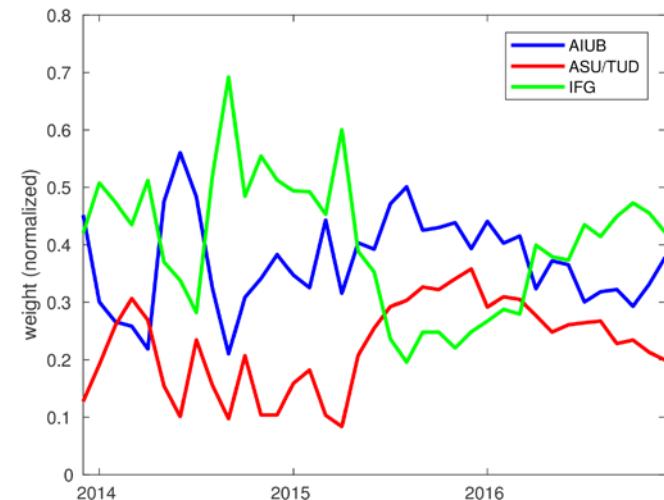


**Quality control of
individual time-series
and their combination
in terms of variability
over the oceans.**

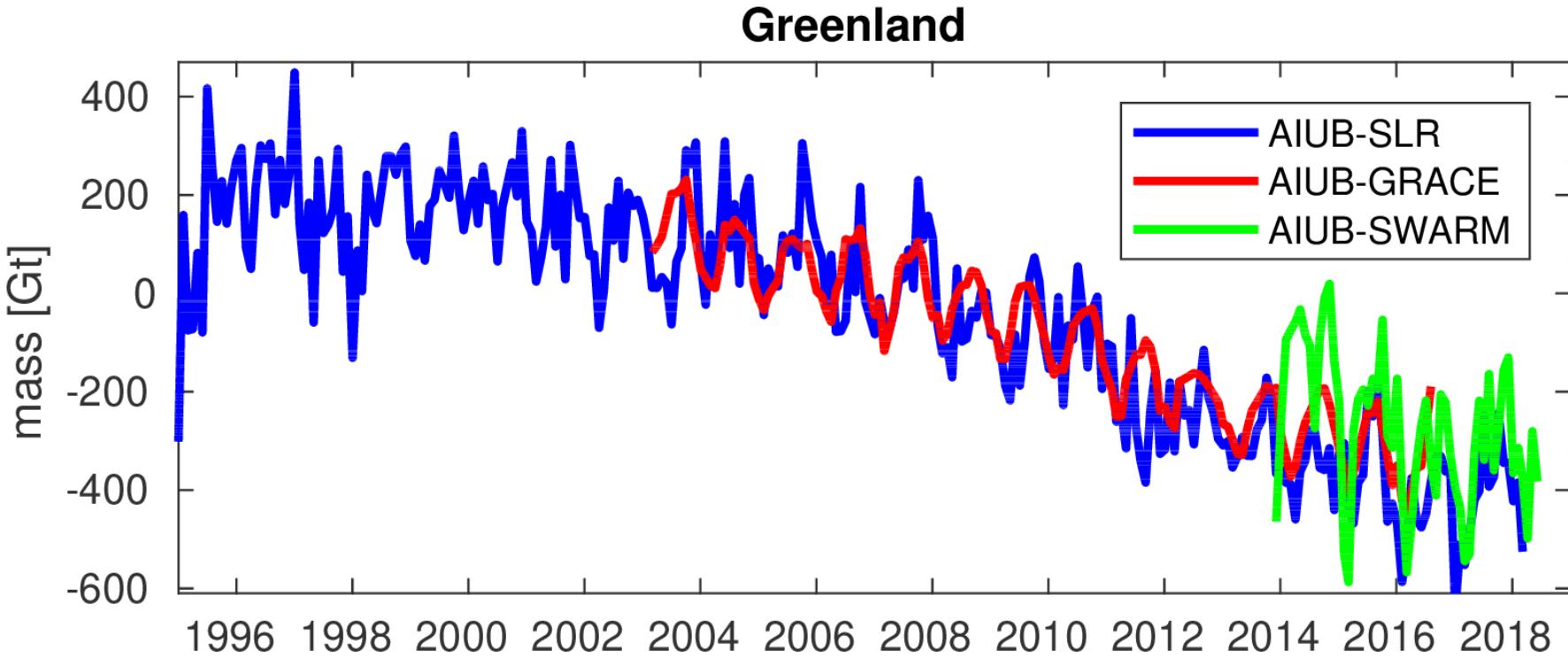
SWARM Combination



SWARM monthly gravity field
 • processing,
 • combination,
 • quality control
 in the frame of SWARM DISC.



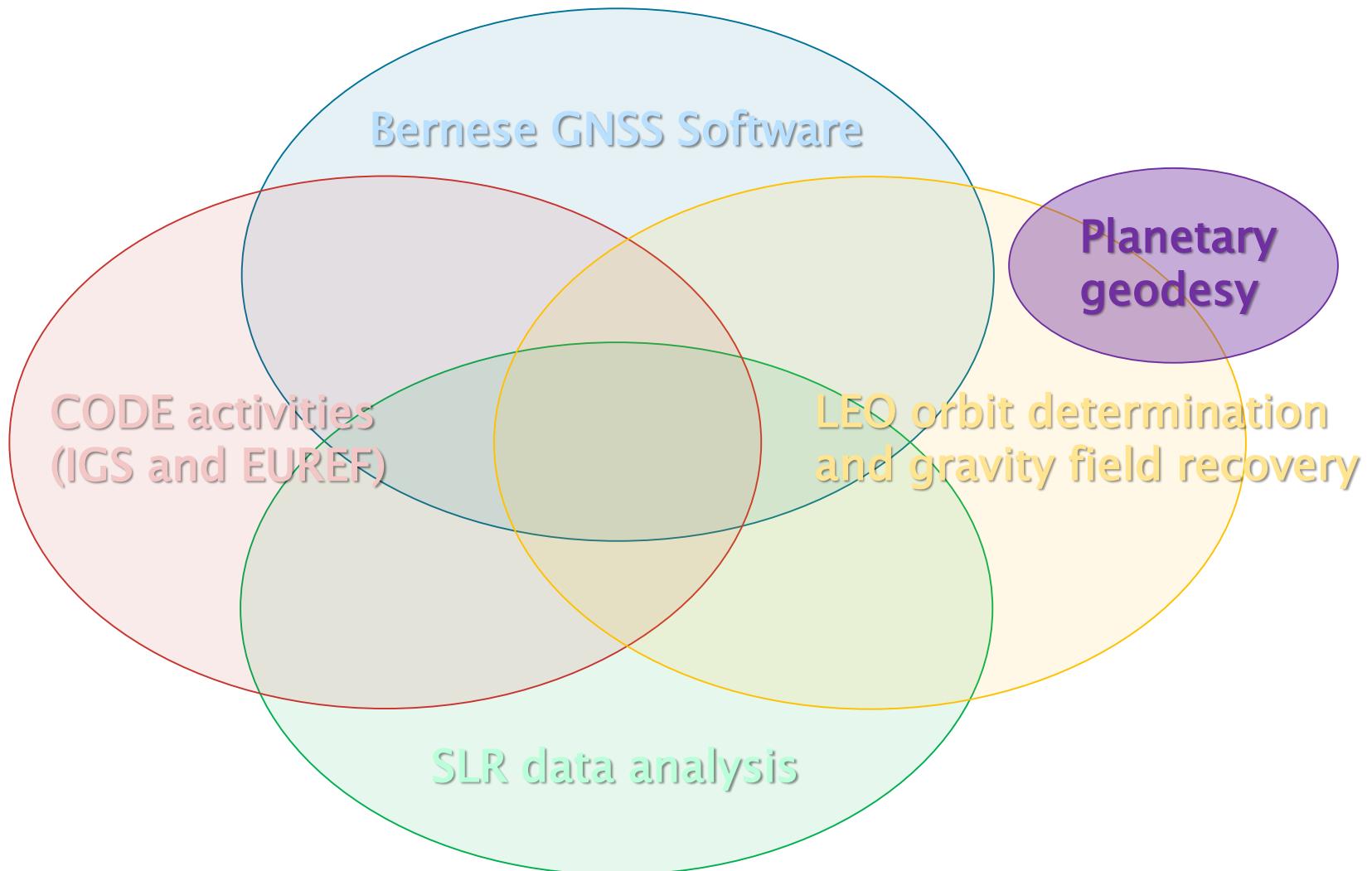
Satellite Laser Ranging



Long time-series of SLR processing (LAGEOS, SLR-LEOs)

- provide information on geocenter, Earth rotation, and the scale of the geodetic reference frame,
- extend the GRACE time-series of mass variations back to the 90s,
- and help to bridge the gap to GRACE-FO.

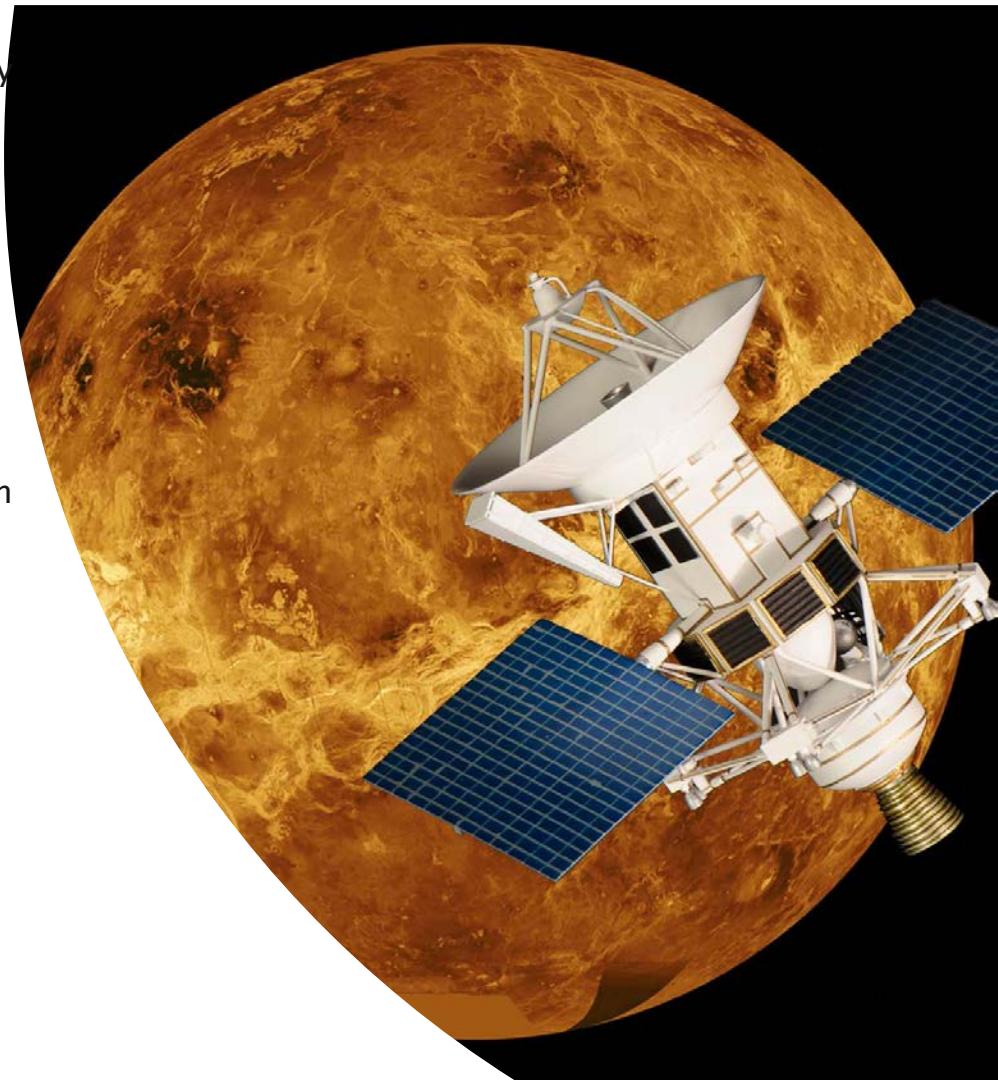
Satellite Geodesy Research Group



Non-gravitational forces for Magellan orbit determination

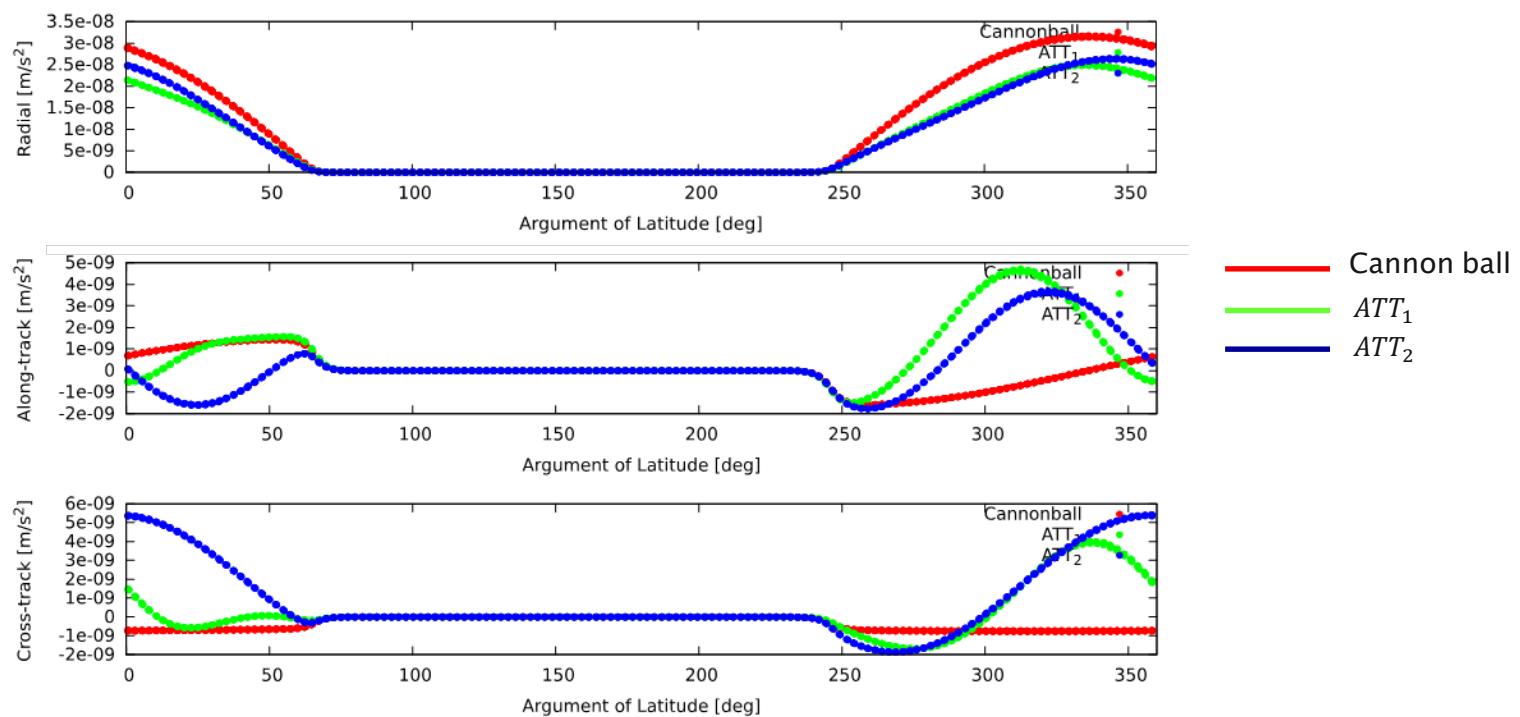
Magellan mission

- NASA mission to Venus, 1990–1994
- Best source of knowledge about Venus gravity field
- State-of-the art Venus gravity field model (MGNP180U):
 - is 20 years old
 - was derived in a non-optimal multi-step approach
 - is based on simplified modeling of non-gravitational forces
- Goal: Establish detailed spacecraft macro model and attitude laws to study impact of non-gravitational force modeling on Magellan POD



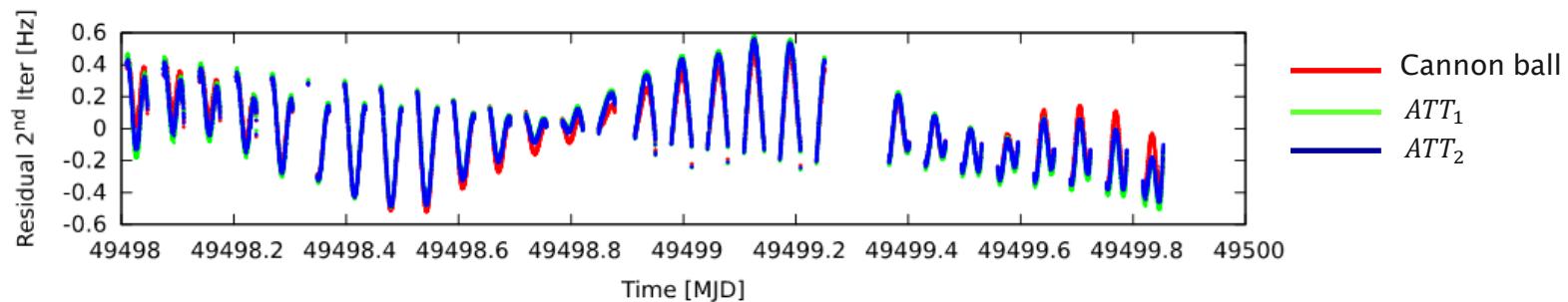
Modeled non-gravitational accelerations

Modeled reflected planetary pressure accelerations acting on Magellan for April 16, 1994 when using a **cannon ball satellite model** or a 16-plate macro model with two different attitude laws (**solar panel axis along velocity** or **solar panel perpendicular to satellite-Sun direction**, while high gain antenna is pointing towards Earth):



Doppler residuals

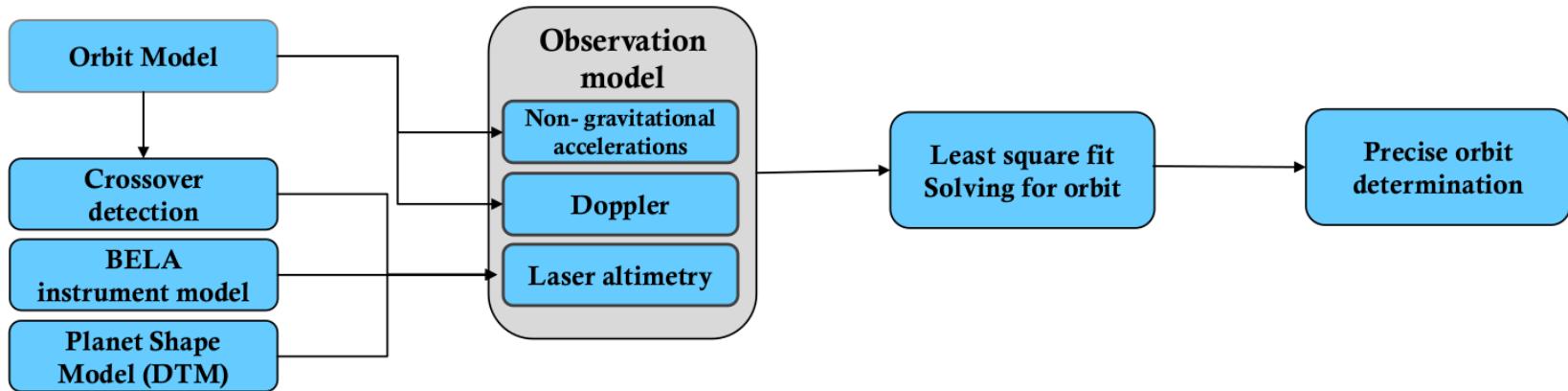
Doppler residuals after 2 iteration of Doppler-based Magellan POD on doys 146 and 147 of 1994:



| Statistics | No Surface Forces | Cannonball | ATT ₁ | ATT ₂ |
|-------------------------|-------------------|------------|------------------|------------------|
| Mean (Hz) | -0.00114 | 0.00027 | 0.00024 | 0.00019 |
| Standard Deviation (Hz) | 2.06433 | 0.22392 | 0.23624 | 0.21425 |
| RMS (Hz) | 2.06420 | 0.22391 | 0.23623 | 0.21425 |

BepiColombo MPO orbit reconstruction simulation

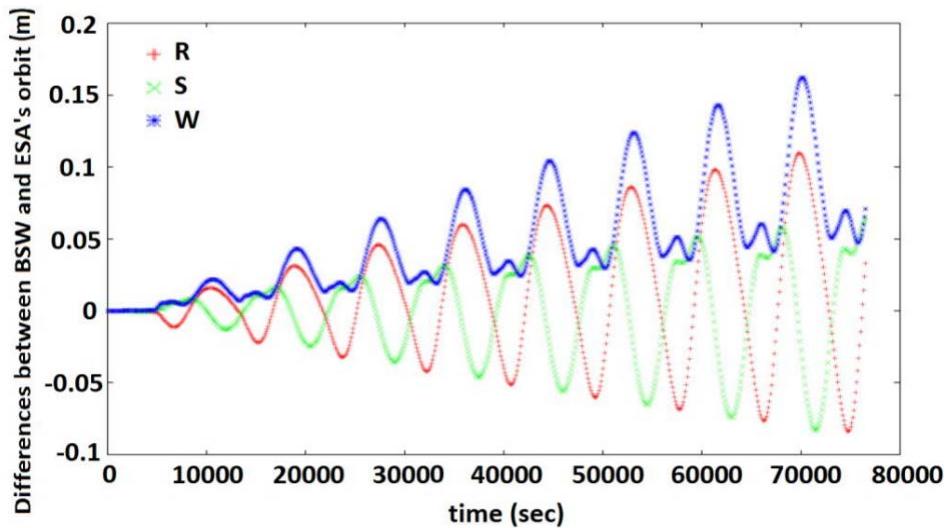
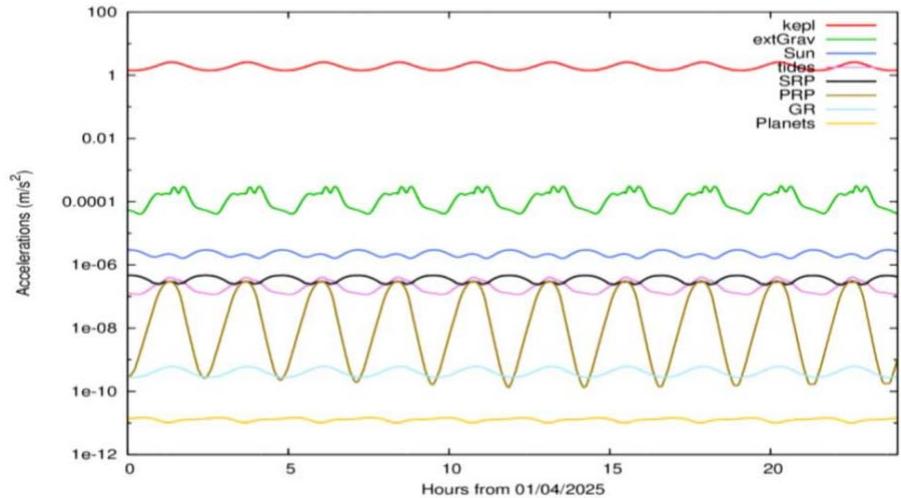
MPO orbit determination



- Simulation study for precise orbit determination of BepiColombo Mercury Planetary Orbiter (MPO) using Doppler, accelerometer and altimetry data and development version of Bernese GNSS Software.
- PhD project at Space Research and Planetary Sciences Division, University of Bern, co-supervised by AIUB

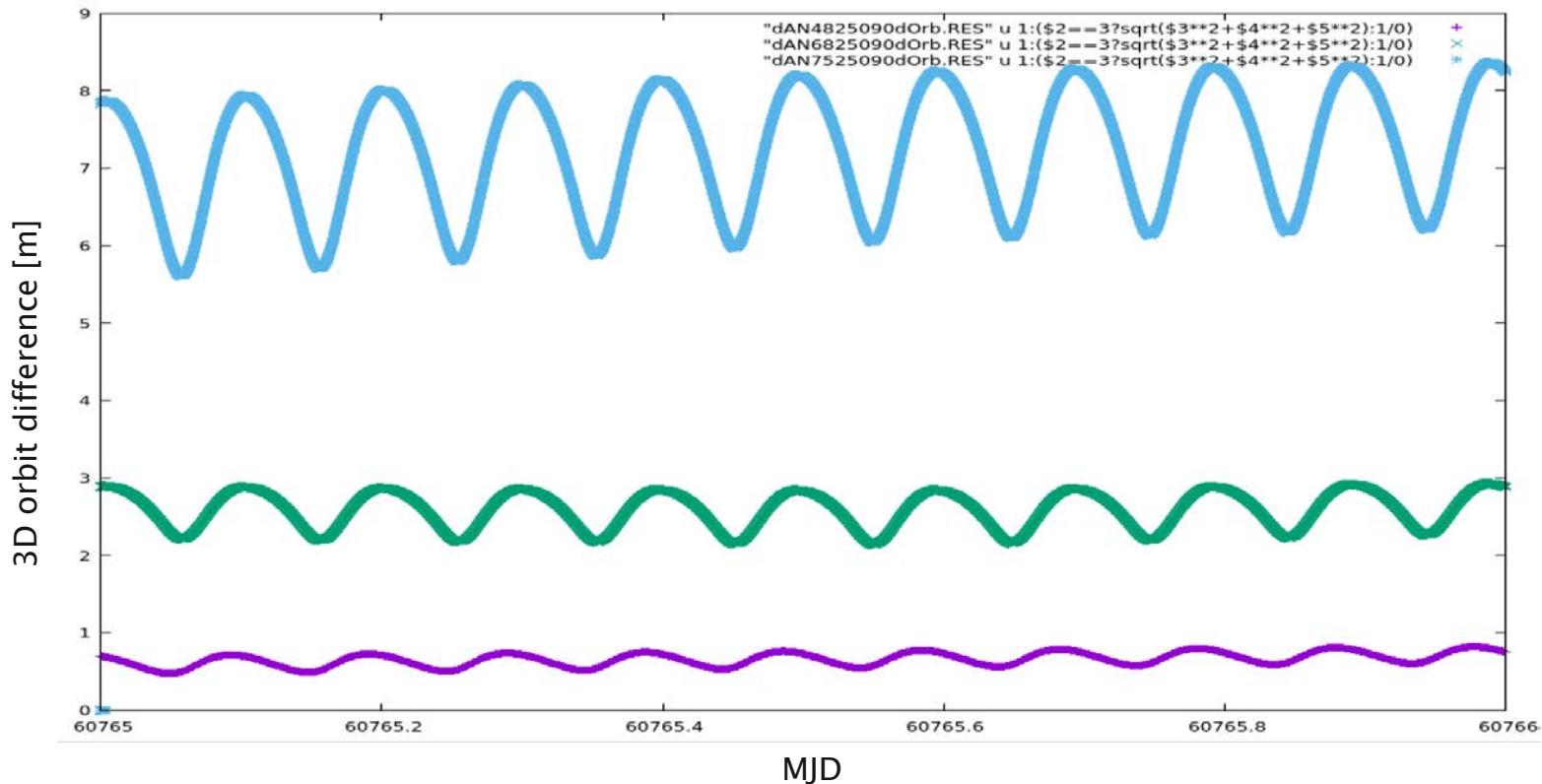
Modeling

- Force model: gravitational forces (Mercury gravity field to d/o 50, Sun, and planets), non-gravitational forces (solar and planetary radiation pressure), relativistic corrections
- Visibility conditions (occultation, elevation over Earth horizon, ground station availability)
- 15–16h tracking period and dark period
- Desaturation maneuvers (every 12h)
- Accelerometer error model



Doppler-based orbit determination

Differences to ref. orbit in a Doppler-based orbit determination when using error-free accelerometer, accelerometer noise model, and accelerometer noise model together with desaturation maneuvers introduced:



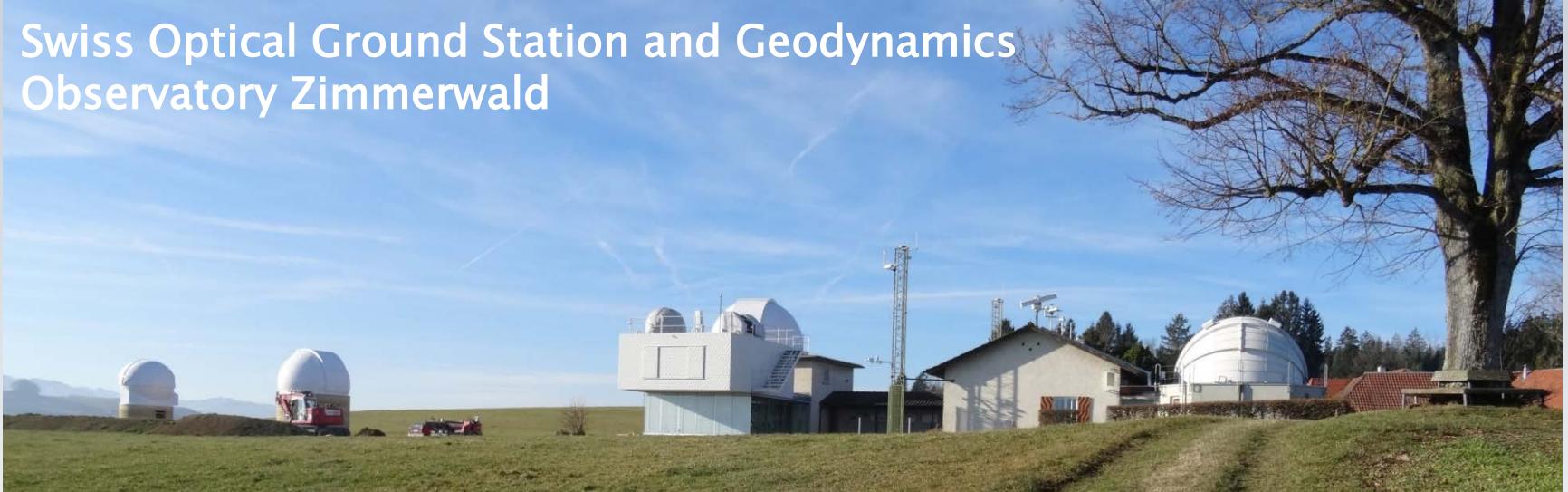
Swiss Optical Ground Station and Geodynamics Observatory Zimmerwald





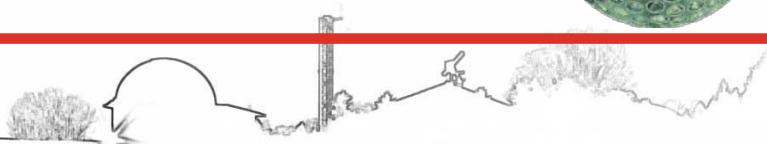
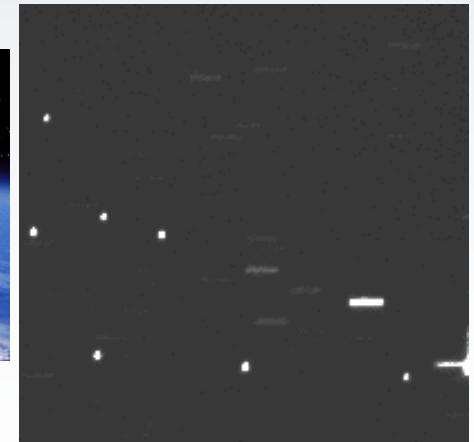
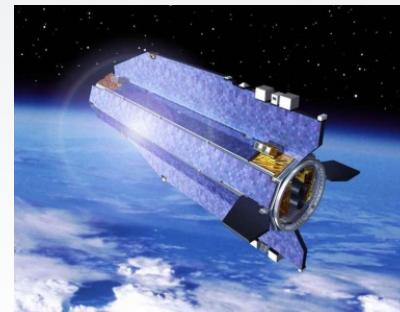
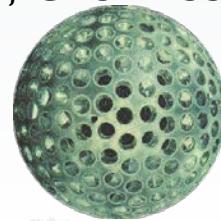
24/7 Betrieb

Swiss Optical Ground Station and Geodynamics Observatory Zimmerwald



24/7 Operations

- **Optical Observations**
space debris (SSA), asteroids, comets
- **Satellite Laser Ranging**
- **Satellite Receivers (GPS-, GLONASS- and Galileo)**



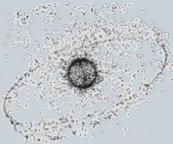


Telecopes

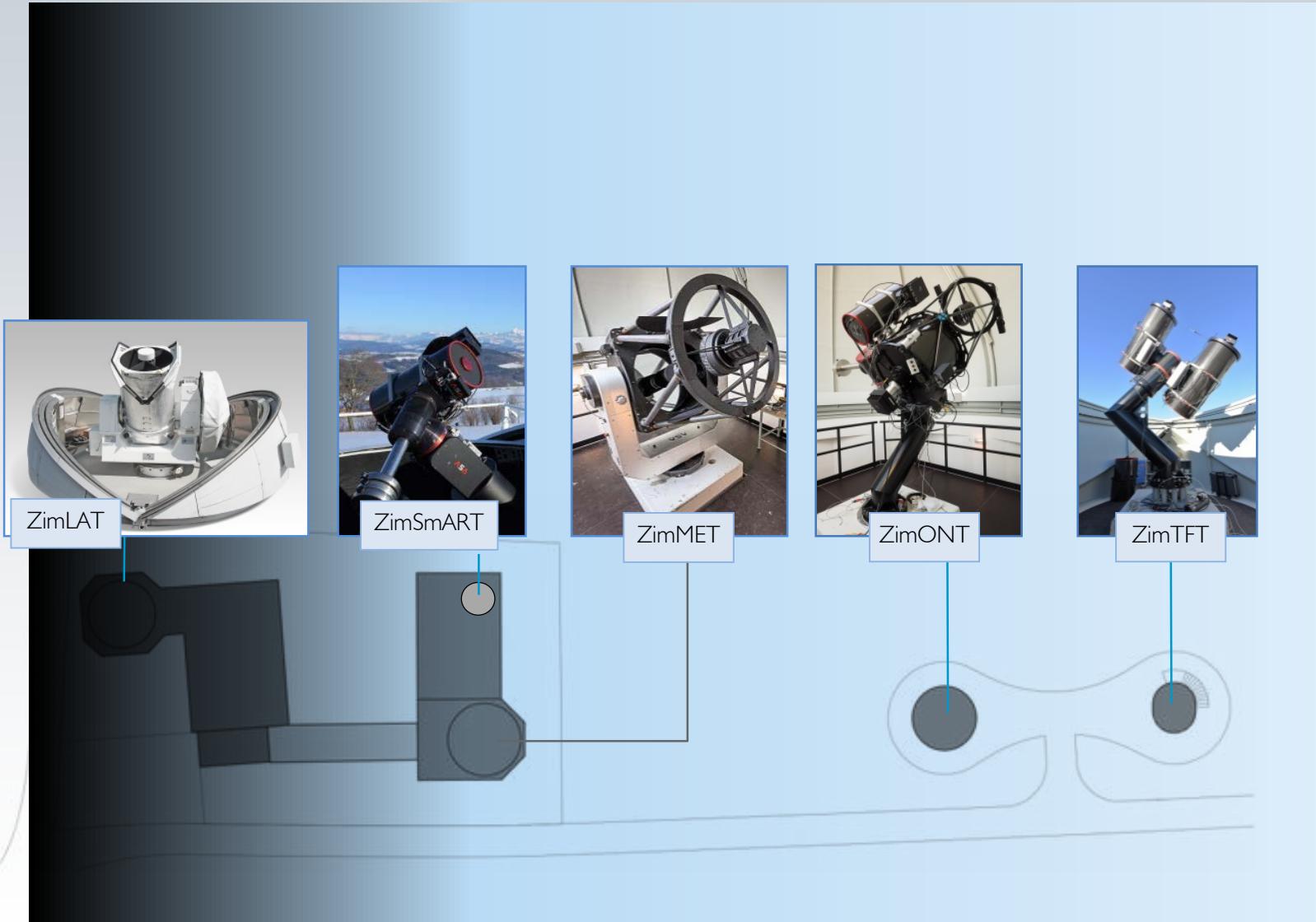
NEW DOMES/TELESCOPES

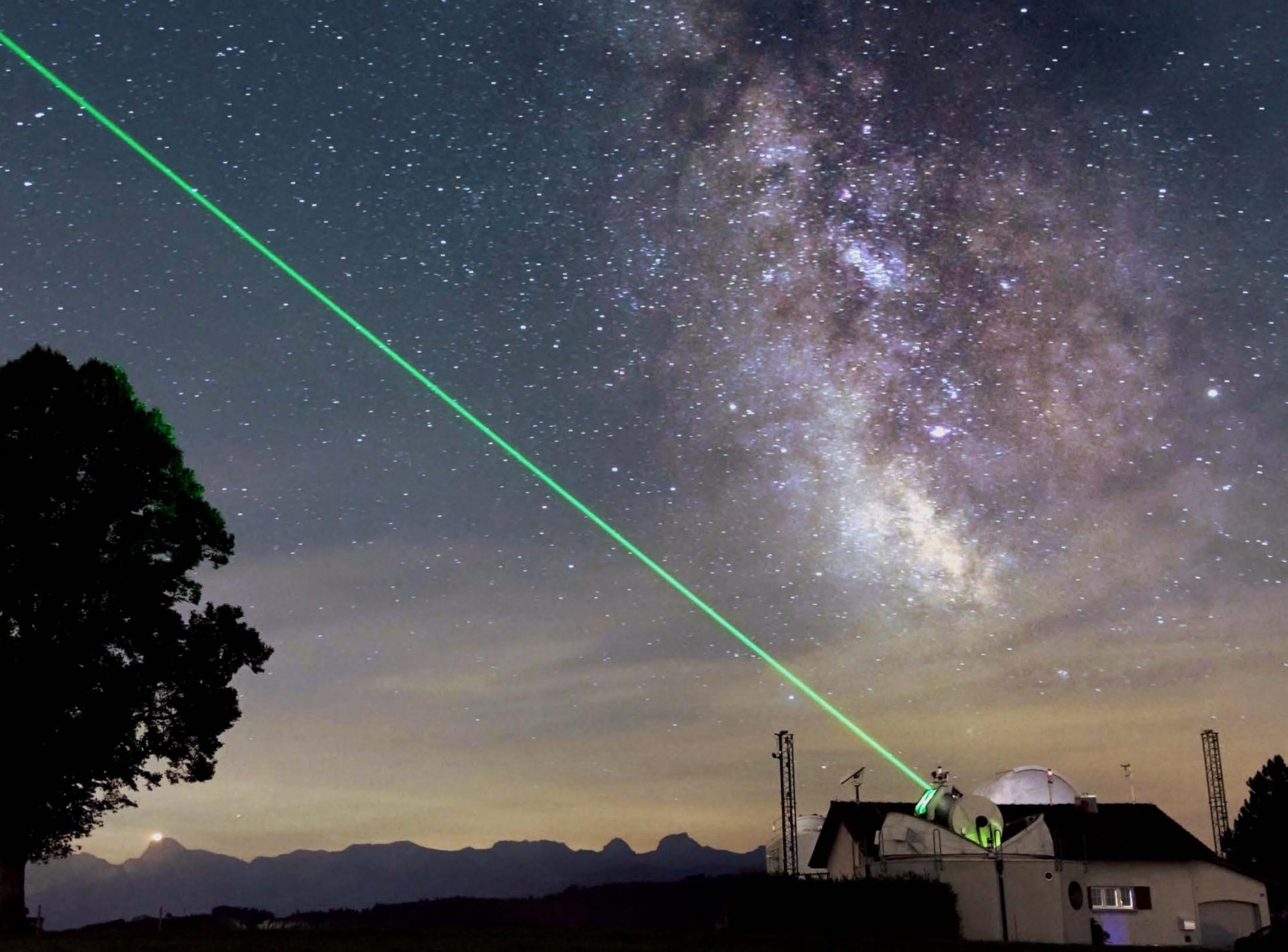
Official inauguration event on May 29
Public Day on June 2





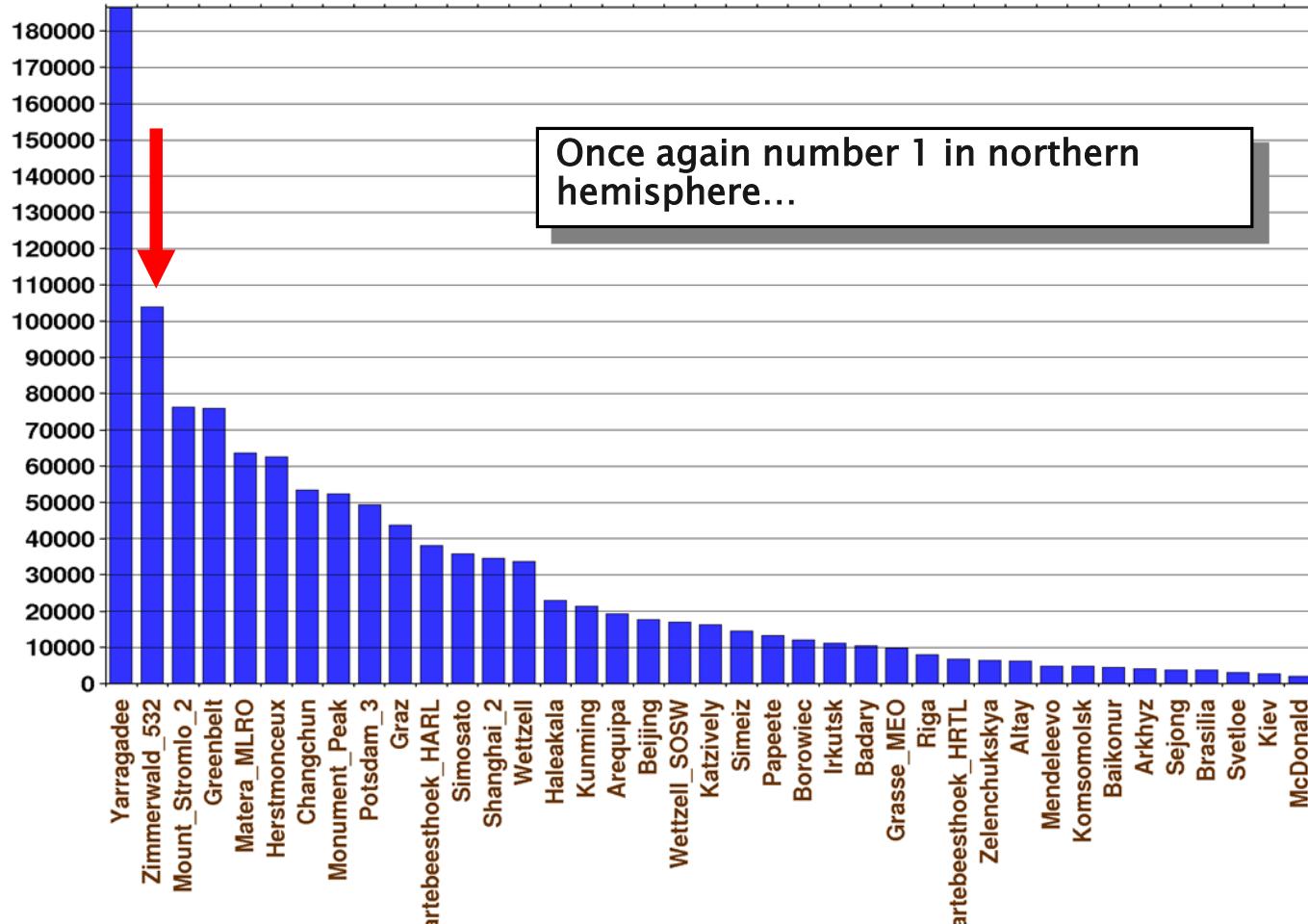
6 Operational Telescopes!





ILRS Station Performance

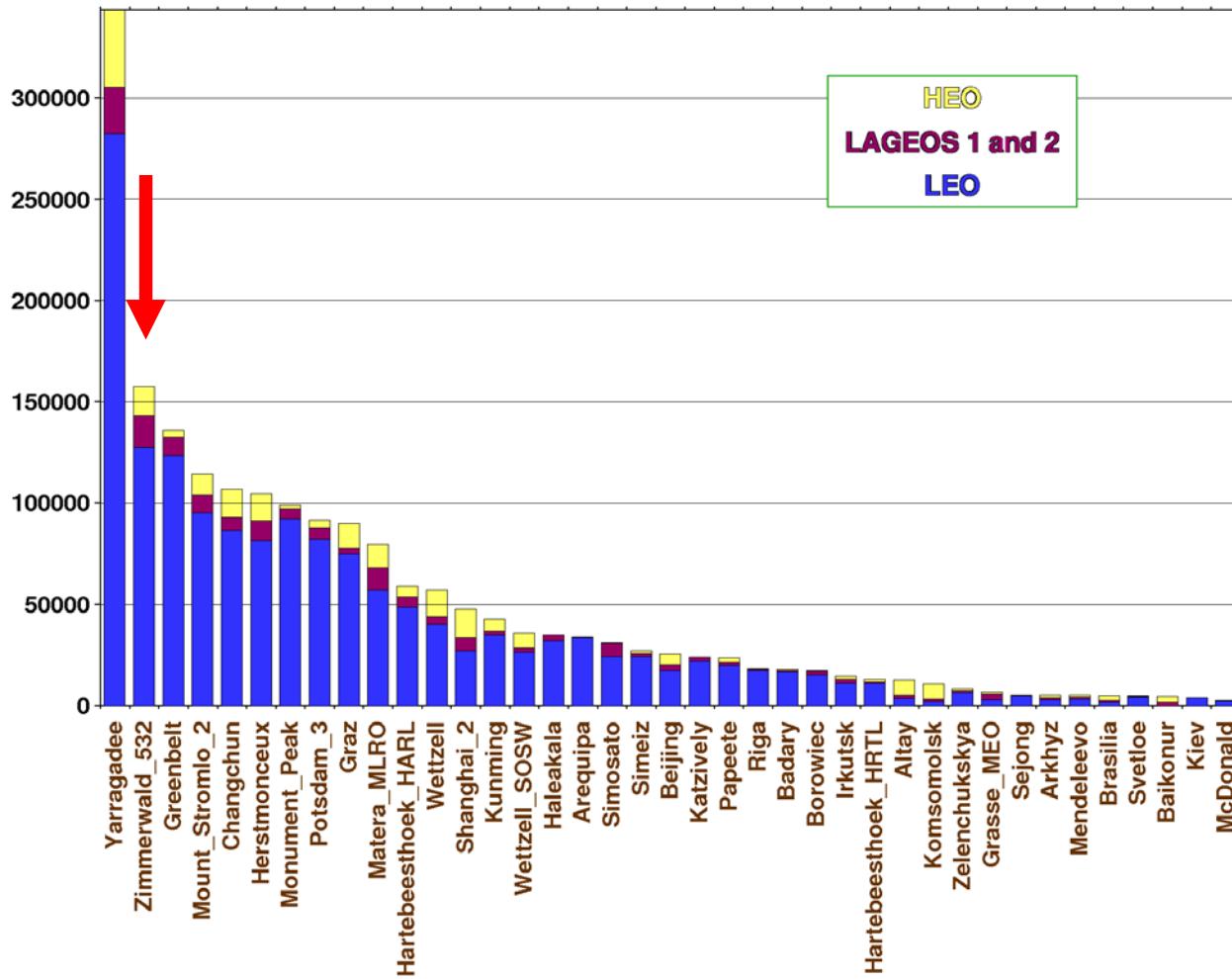
minutes of data
from October 1, 2017 through September 30, 2018



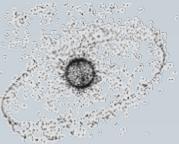
20181002

ILRS Station Performance

total normal points
from October 1, 2017 through September 30, 2018



20181002

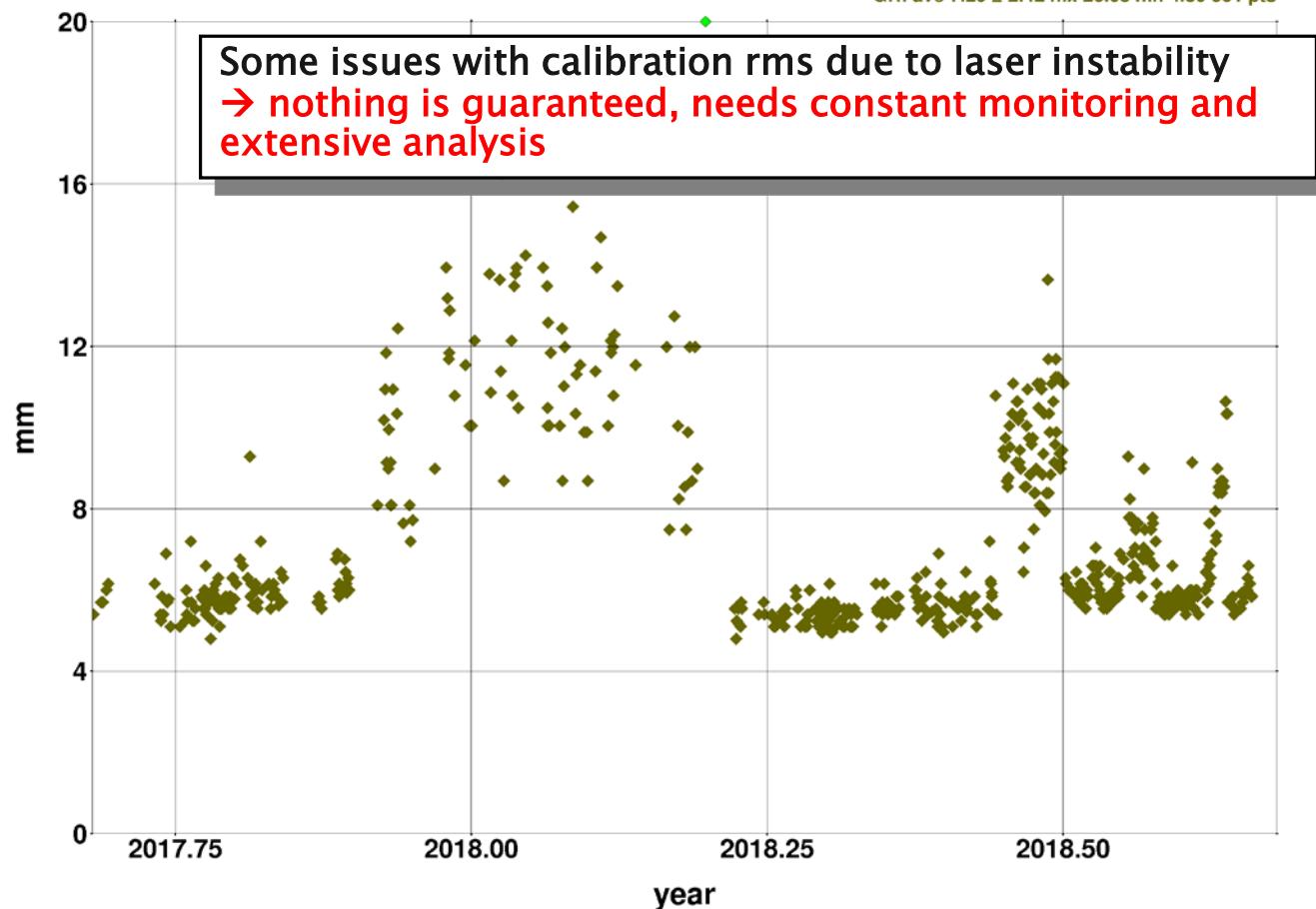


LAGEOS Calibration rms

Zimmerwald, Switzerland 7810

pass LAGEOS calibration rms

GR: ave 7.20 ± 2.42 mx 26.08 mn 4.80 561 pts



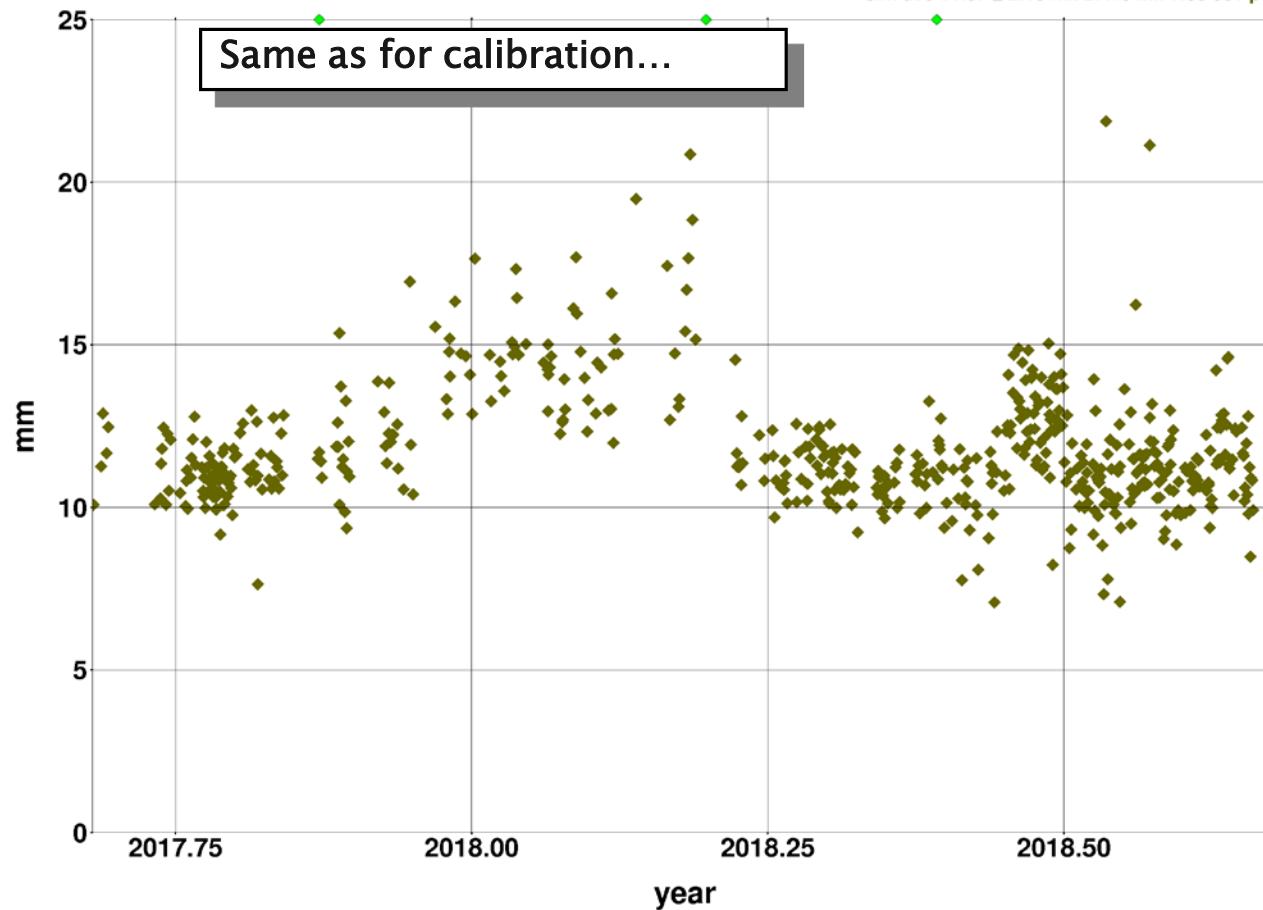
20180904 18:45

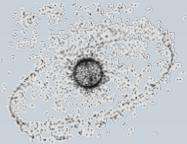


LAGEOS NPT rms

Zimmerwald, Switzerland 7810
pass average LAGEOS normal point rms

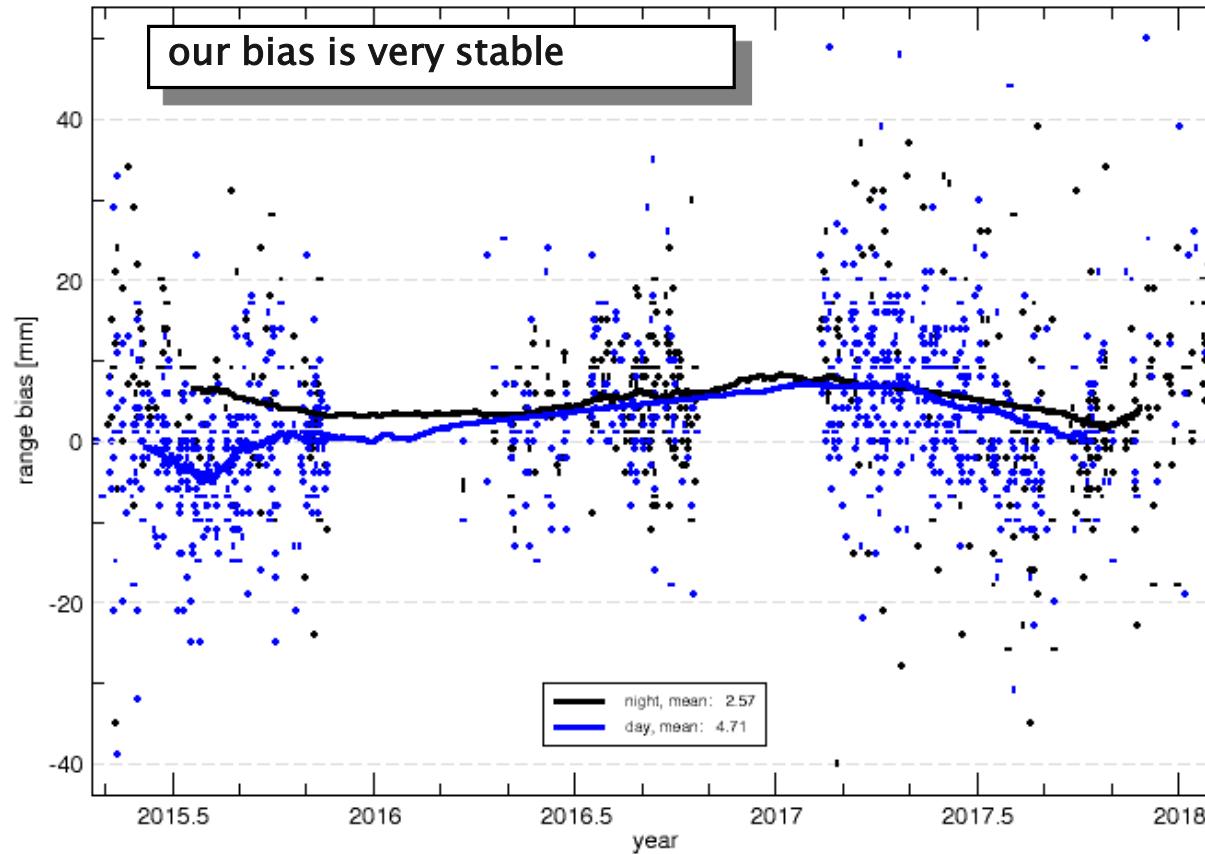
GR: ave 11.87 ± 2.19 mx 27.43 mn 7.08 561 pts



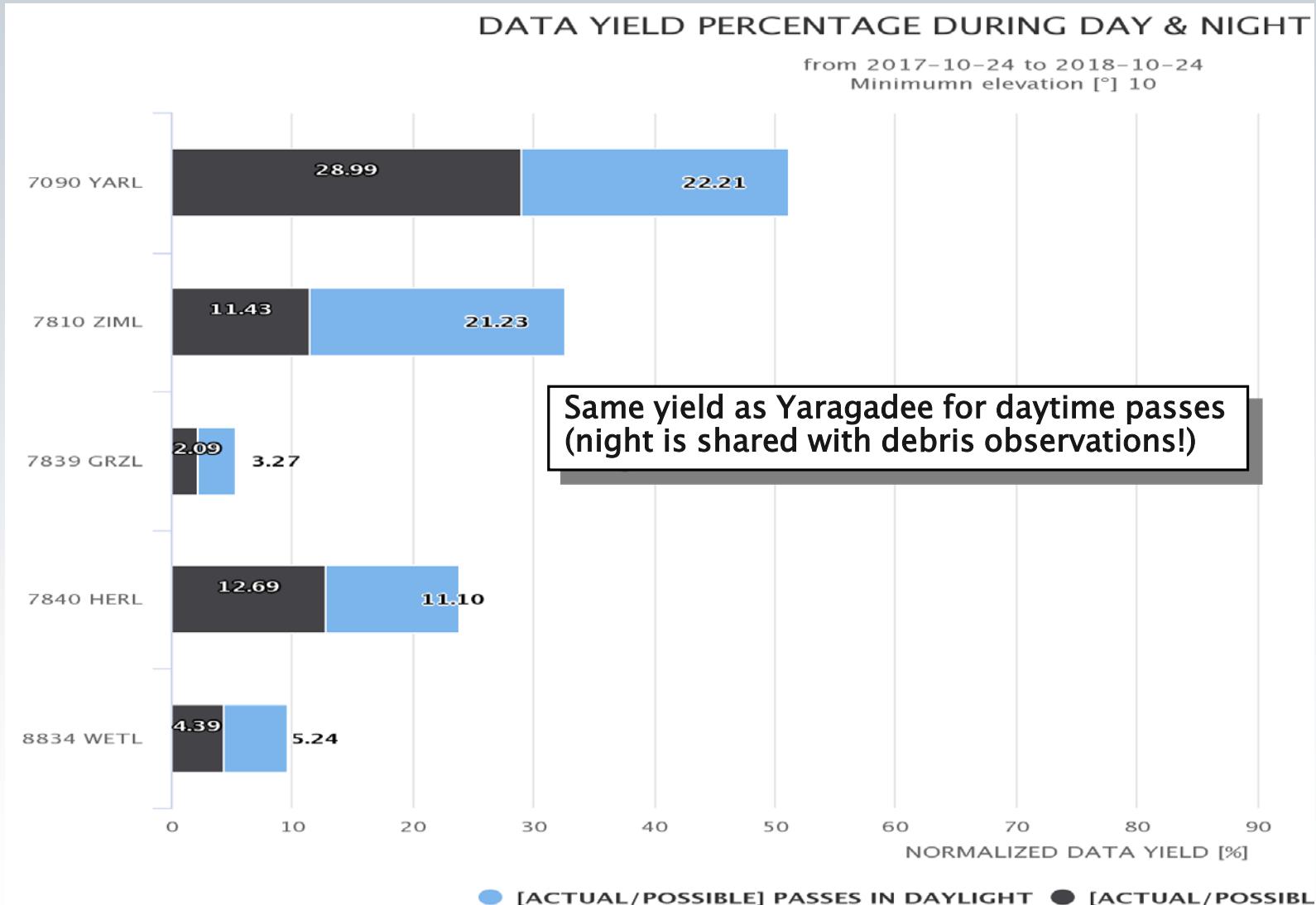


LAGEOS Bias Analysis

Zimmerwald (7810) bias analysis: one per pass and running average of 300 passes
LAGEOS-1 2015-2018 day vs. night passes



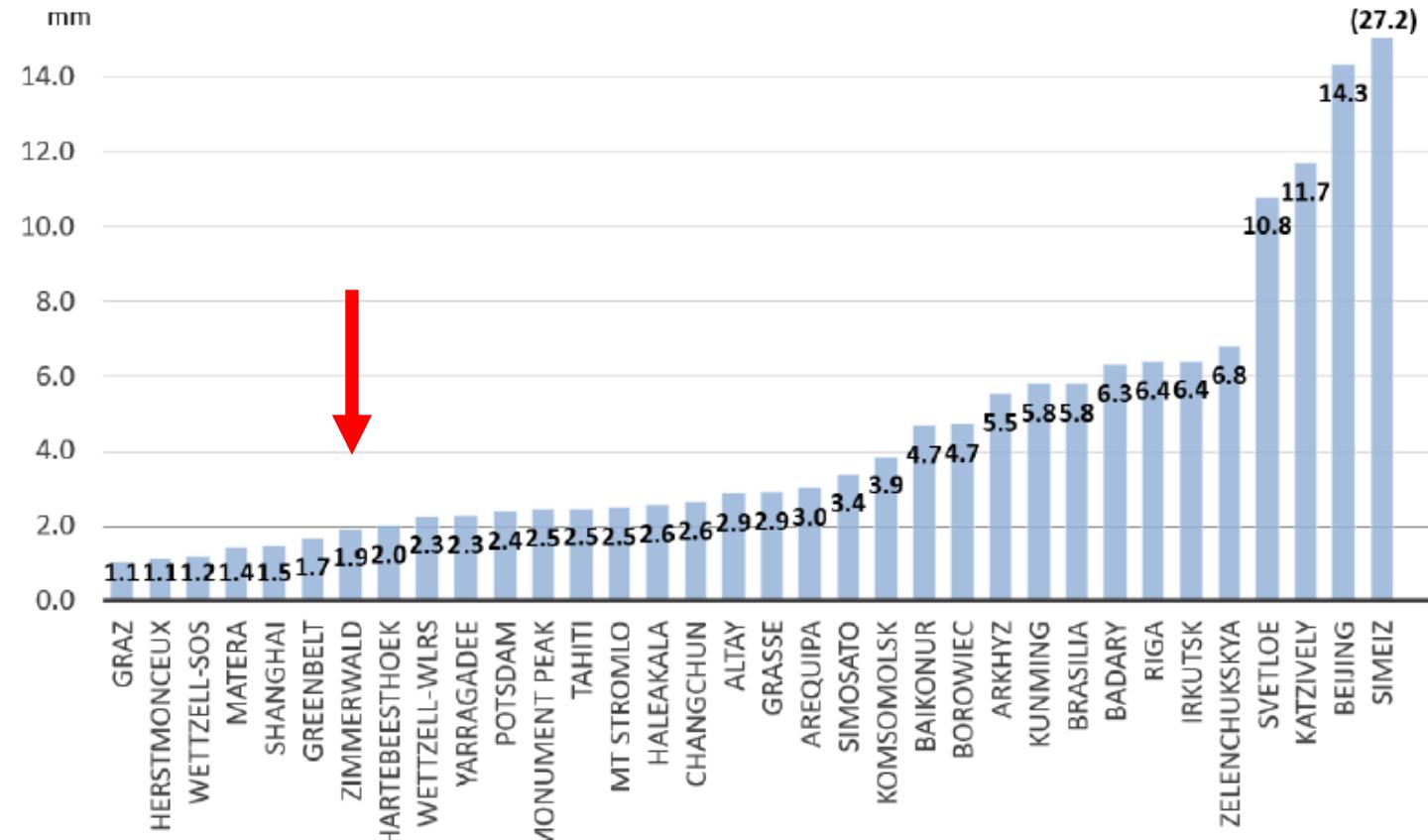
LAGEOS Day/Night Yield



ILRS NP RMS

1 year (July 2016-June 2017), LAG1+LAG2.
RB only or RB+TB smoothing applied for POD (c5++) post-fit residuals.

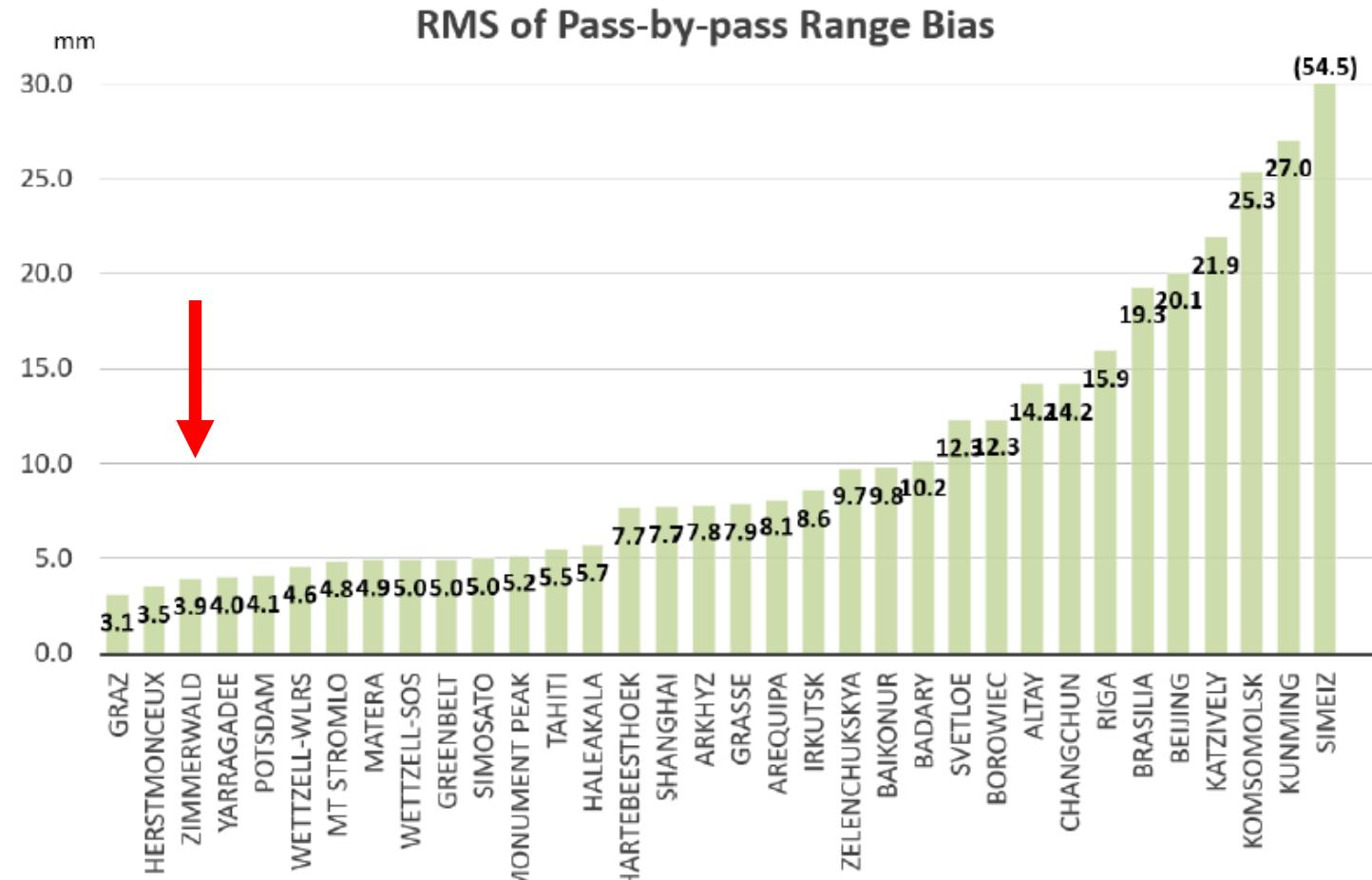
Mean "NP RMS" (pass smoothing applied)



ILRS Range Bias

1 year (July 2016-June 2017), LAG1+LAG2.

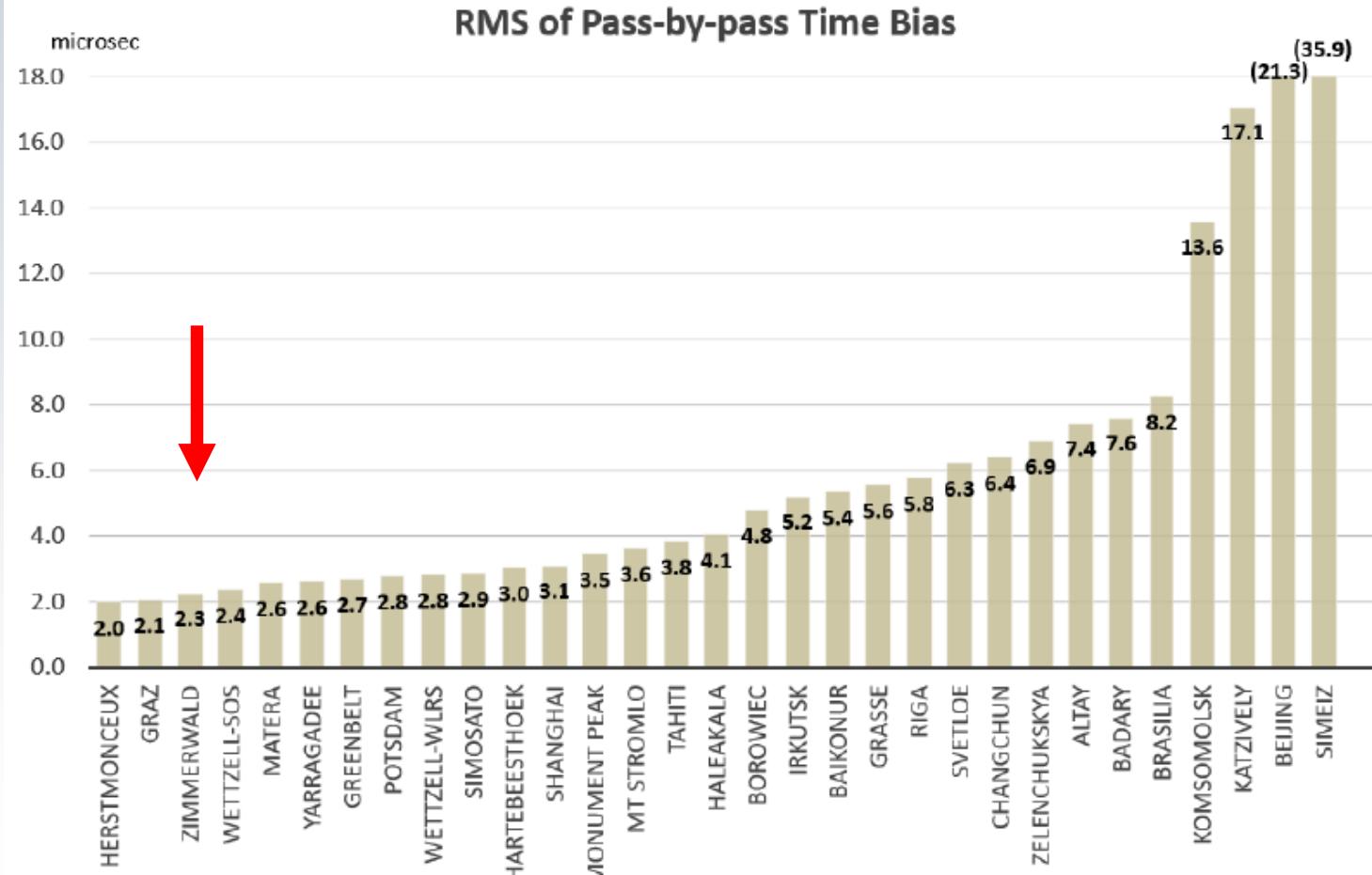
POD (c5++): station pos solved for. U-Strasbg atm+hyd loading applied.

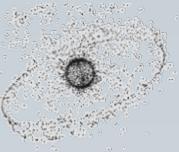


ILRS Time Bias

1 year (July 2016-June 2017), LAG1+LAG2.

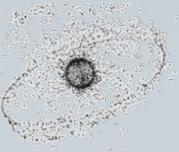
POD (c5++): station pos solved for. U-Strasbg atm+hyd loading applied.





SLR Operations

- **ILRS pass statistics**
 - again #1 in the northern hemisphere
 - very good performance
 - station range bias often at ≤ 1 mm, location of error source narrowed
- **Special Satellites/Restricted Tracking**
 - Sentinel-3A/B on a routine basis (Low Energy)
 - ICESat-2 upcoming (GoNoGo)
 - decommissioned satellites: # of debris targets increasing
- **System**
 - laser, electronics and mechanics perform reliably
 - mechanics: expect some refurbishments after 20 years of operation for dome



SLR Operations

- **Definition/evaluation of new laser**
 - 100Hz/kHz...? (quantum jump of technology not in sight)
 - two lasers? debris SLR on new 0.8m telescope?
 - new targets (nano, debris satellites) might affect choice
- **European Laser Time Transfer project (ELT) (ACES experiment on ISS)**
 - Trimble GPS–Receiver 1PPS 15 ns wrt UTC integrated
 - determination of internal calibration delays on-going



SLR Operations

- **Tracking camera / stare and chase**
 - find target in full frame image → correct telescope pointing (automatically!) → track satellite with laser
- **Quantum experiments**
 - telescope optics in the infrared fits requirements for entangled photons experiment with IAP (Hasler Stiftung project)
 - first to calibration target, later to a satellite?





Stare and Chase

?? Unknown Language Help

NightCam @ 38.8%

File Zoom Zimlas

Acquisition

Start Stop Exposure Time (s) Bit Depth Shutter Mode Readout Freq.

0.2 16-bit rolling 200 MHz

Rendering

Auto Scale Black Point 6 9

Range White Point Binning 1x1

Midtones

TCP Corr. Az/° Corr. El/° AutoImage 96 AutoImage 97

0.0 0.0

Send Correction

Resume Search Loop Save Images

Reset Correction Save Single (F8)

Ping 130.92.25.136

Min. #Hits Min. Interval/s

1 0.05

Save continuously

HWIST=33.90 AZIST=294.89 DRIST=40.00 HKKORR=937.98 HXKORR=-41.98 AMKORR=764.86 AXKORR=69.70 IHITPC=23 MMODEL=40

PGPLOT Window 11

Change Filter Tolerance: Option Q

0 10 0 -10

0 10 20 30 40 50 60 70 80 90 100

Name Date modified Type

Topex 18.10.2018 19:27 File folder

prefix20181018-210557-816.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210558-206.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210558-596.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210558-976.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210559-356.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210559-746.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210600-136.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210600-516.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210600-896.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210601-296.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210601-696.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210602-096.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210602-486.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210602-866.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210603-246.fit 18.10.2018 21:06 Maxim DL Image

prefix20181018-210603-636.fit 18.10.2018 21:06 Maxim DL Image

Cooling status: Stabilised / Sensor temperature: -30.3 °C G:\NightCam\Win64\Debug\images\20181018\

Tracking camera image before correcting telescope pointing

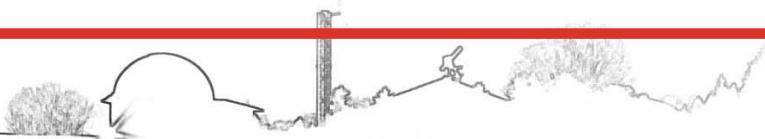
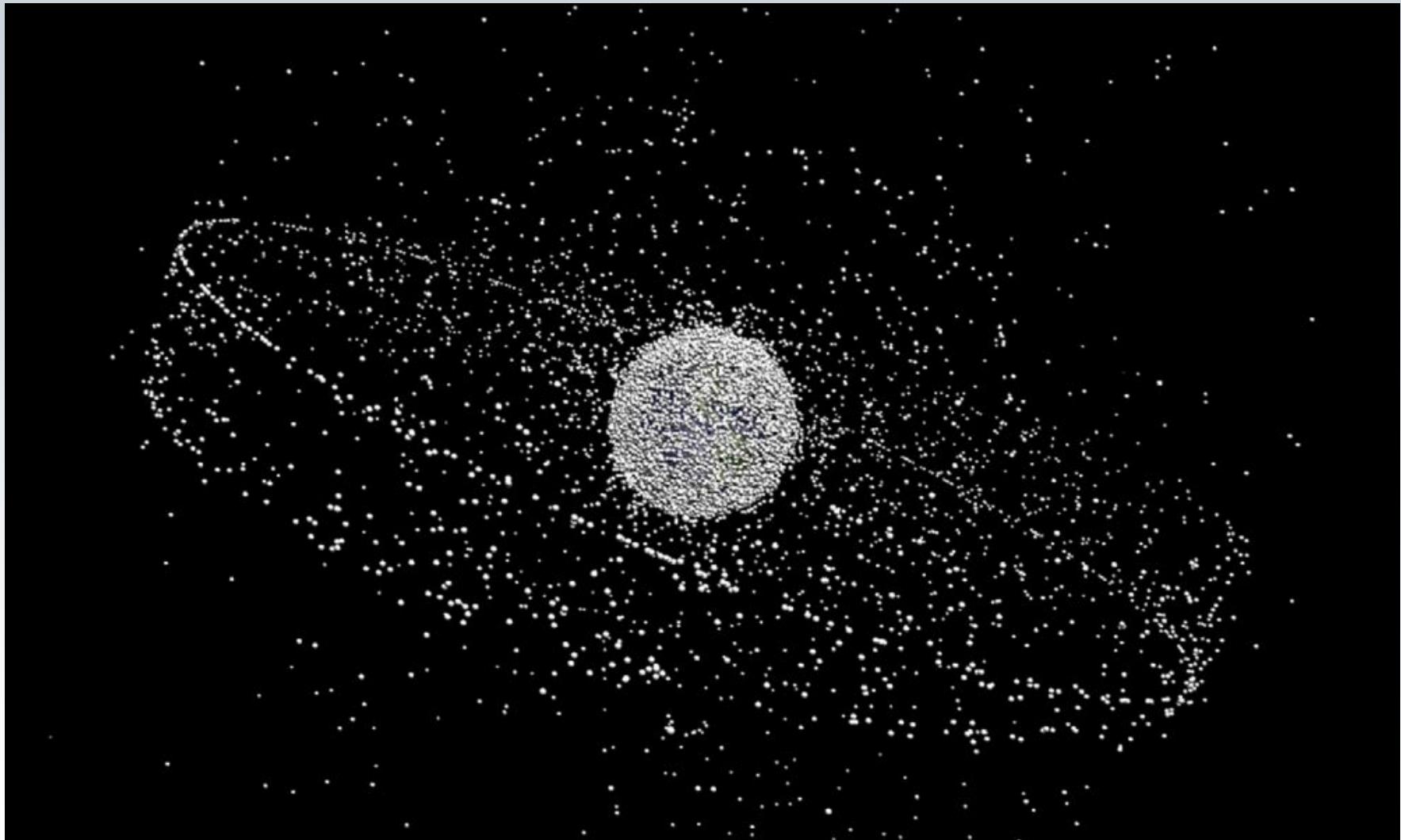
Returns after automatic pointing correction (TOPEX)!

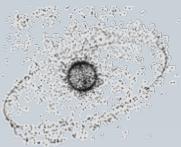
The screenshot shows the NightCam software interface. On the left, a large window displays a tracking camera image with a red crosshair indicating the target. Below it, a smaller window titled 'PGPLOT Window 11' shows a scatter plot of data points. To the right, a file list shows numerous FITS files from October 18, 2018, all named 'prefix20181018...' followed by a sequence of digits. A callout box highlights the tracking camera image with the text 'Tracking camera image before correcting telescope pointing'. Another callout box highlights the file list with the text 'Returns after automatic pointing correction (TOPEX)!'.



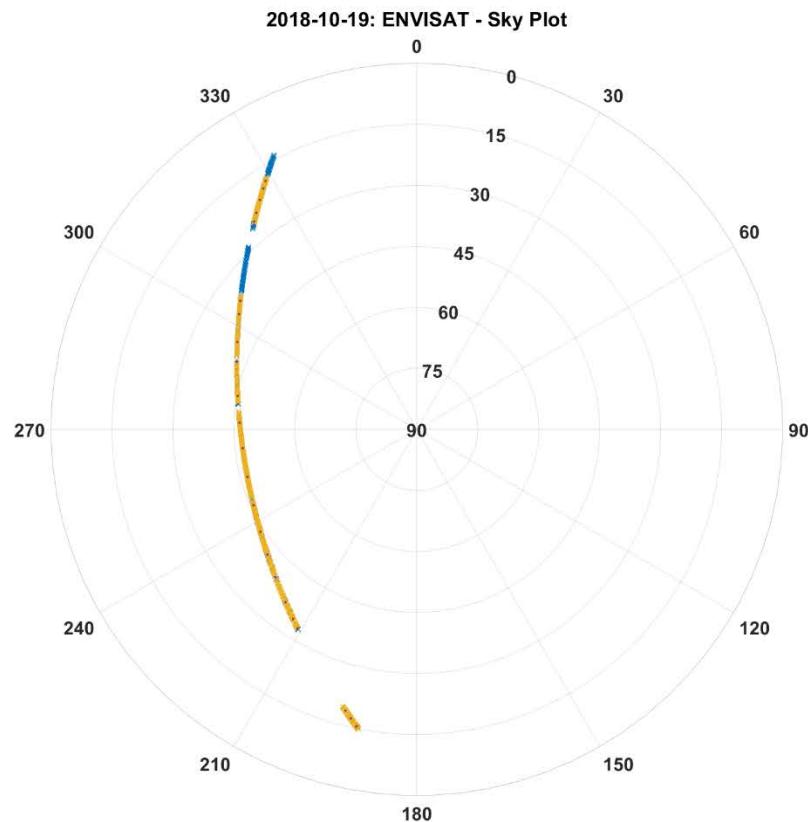
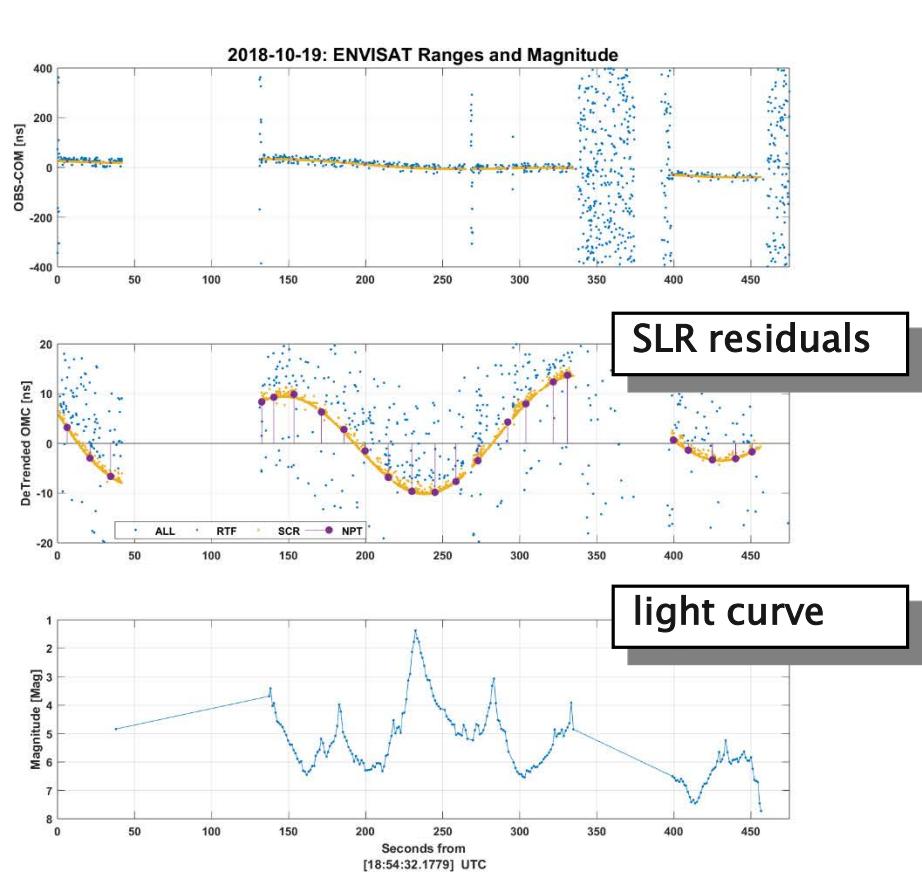


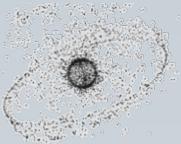
Space Debris



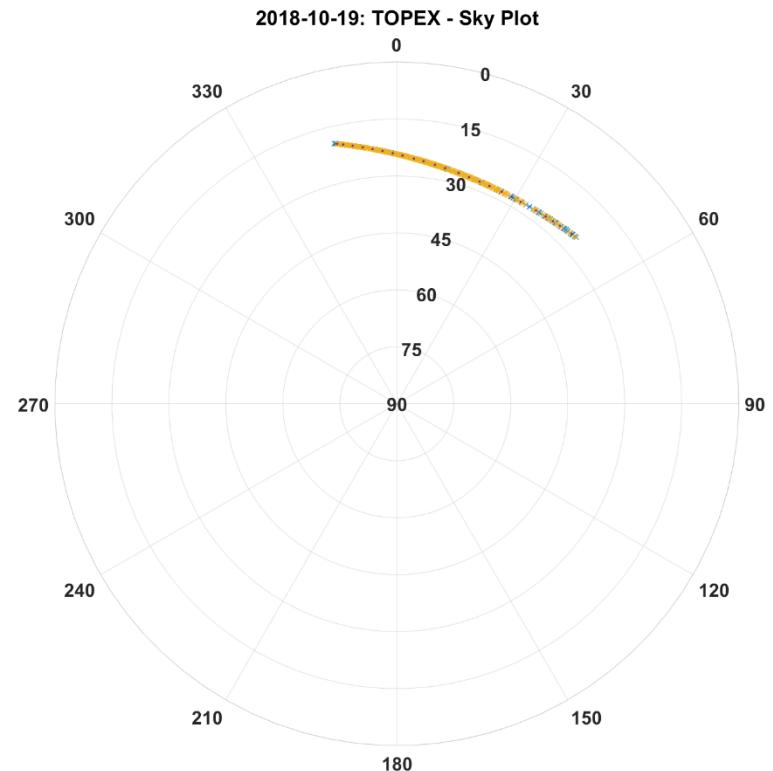
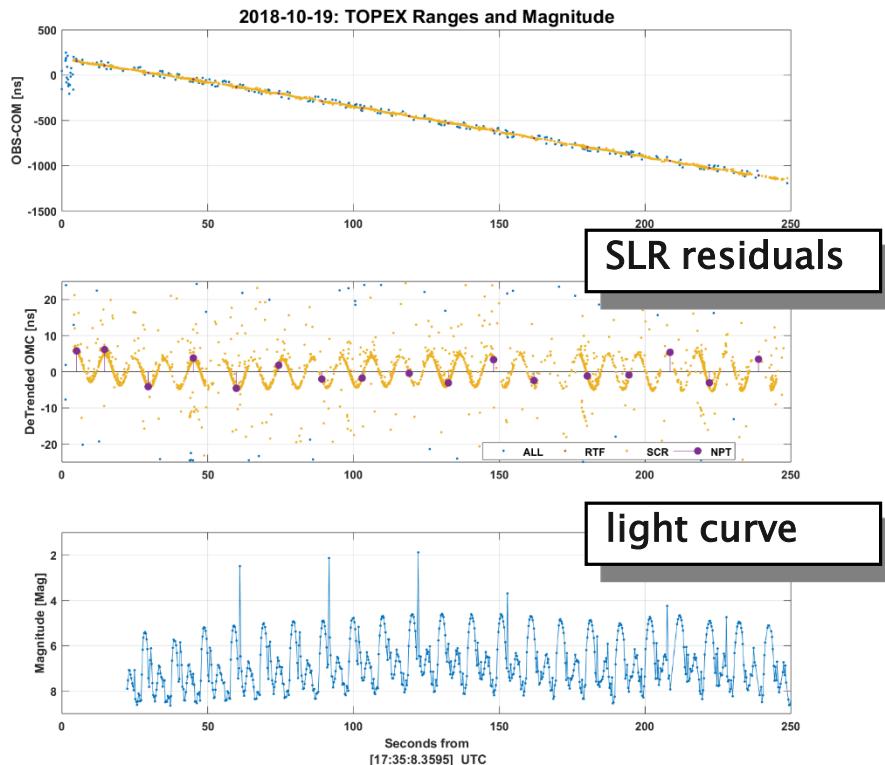


Simultaneous Light Curves and SLR Measurements – ENVISAT





Simultaneous Light Curves and SLR Measurements – TOPEX



Simultaneous Light Curves and SLR Measurements – GLONASS (decommissioned)

