GRACE gravity field determination with the Celestial Mechanics Approach at AIUB

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GRACE Science Team Meeting

11./12. Nov. 2010

Potsdam

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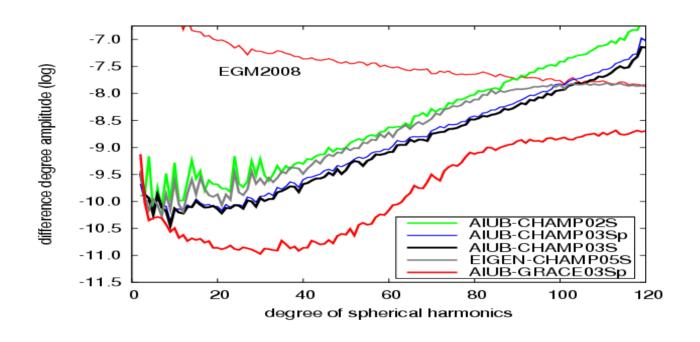
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CHAMP: static 8y field AIUB-CHAMP03S



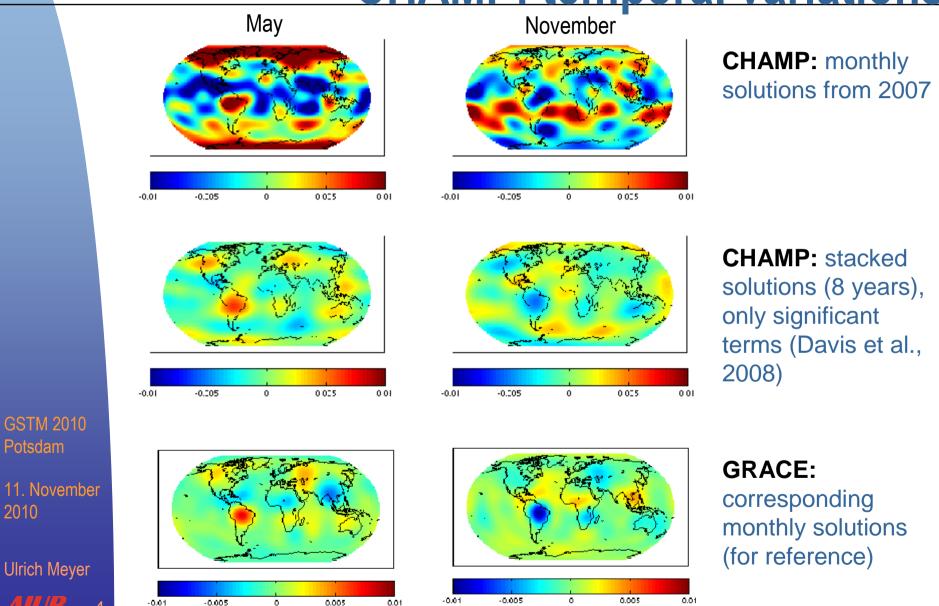
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CHAMP: temporal variations



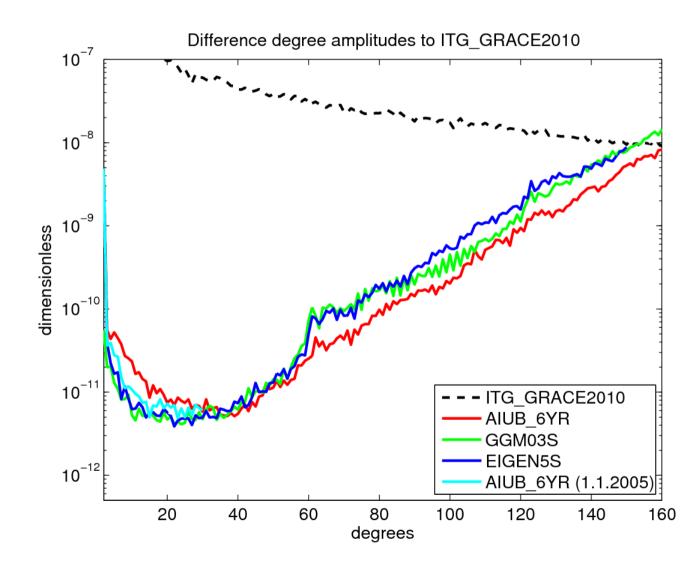
GRACE: static 6y field AIUB_6YR



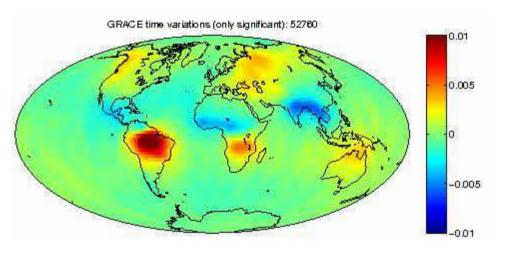
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GRACE: periodic time variations



significant terms (1-yearly, ½-yearly)

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Celestial Mechanics Approach

Key features:

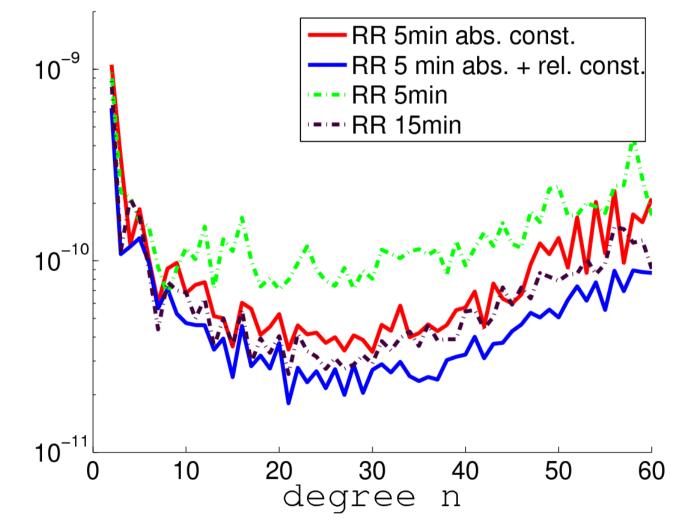
- Kinematic orbits (from GPS) as pseudoobservations (efficiency)
- Reduced dynamic orbits (GPS & K-Band) with stochastic pulses (flexibility)
- Intelligent way to compute variational equations (efficiency)
- NEQ-manipulation: relative weighting in combination, parameter pre-elimination, accumulation (flexibility and efficiency)

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Pulses: spacing and constraints



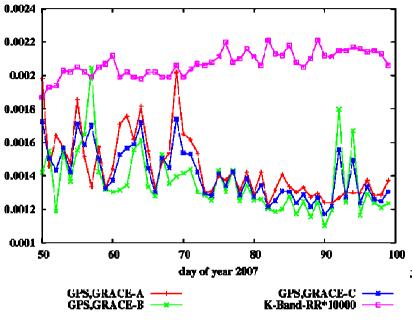
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GPS via K-Band: RMS and rel. weighting



K-Band-only solution:

- solved for difference $\frac{1}{2}(\mathbf{r}_1 \mathbf{r}_2)$
- mean value $\frac{1}{2}(\mathbf{r}_1 + \mathbf{r}_2)$ slightly constrained to kin. positions (from GPS)

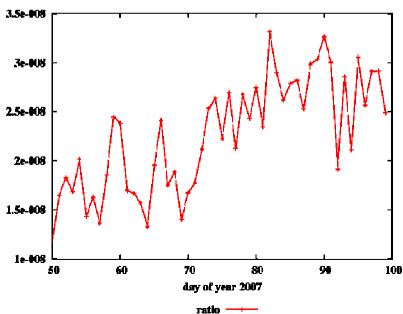
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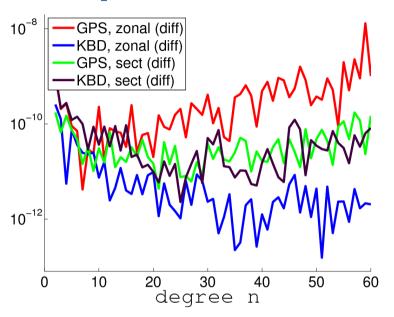
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Theoretical weighting factor $\sigma_{KBD}^{2}/\sigma_{GPS}^{2} = 2...3 *10^{-8} s^{-2}$ (better agreement with other GRACE-fields is achieved using $1*10^{-10} s^{-2}$, i.e. downweighting GPS)



Impact of GPS on combined solution



Impact of GPS mainly on very low degree and close to sectorial coefficients.

Combined minus K-Band only Coefficients

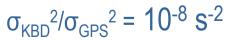
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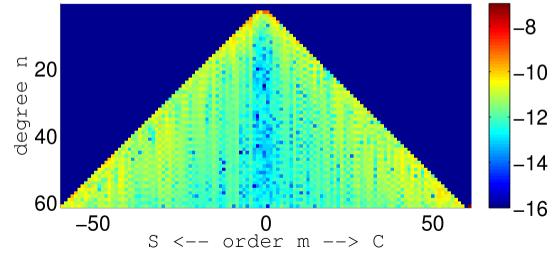
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Range-Differences and Correlation

correla	ited
	correla

RR	Range-Rate	RRC
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R ≈ RC

RR
$$\approx$$
 RD = (R(t₂) - R(t₁)) / Δ t

$$P_R = k * I$$

$$P_{RD} = ((\partial RD/\partial R)^T P_R \partial RD/\partial R)^{-1}$$

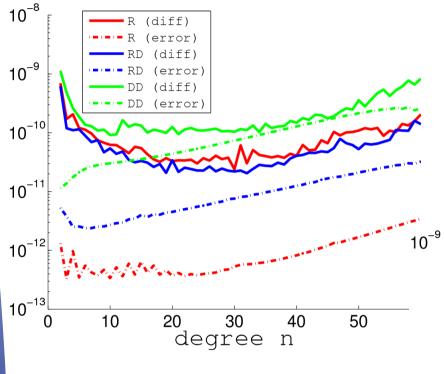
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Range, Range-Diff., Double-Diff.



Uncorrelated RD-solution outperforms R-solution (contradictory to formal errors).

R 30min (diff)

R 5min const. (diff)

Introducing frequent (5 min.) 11. November constrained pulses results in 2010 a competitive R-solution.

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