

# Bernese GNSS Software Version 5.4

**Tutorial** 

Processing Example Introductory Course Terminal Session

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# 1 Introduction to the Example Campaign

# 1.1 Stations in the Example Campaign

Data from seventeen European stations of the International GNSS Service (IGS) network and from the EUREF Permanent Network (EPN) were selected for the example campaign. They are listed in Table 1.1. The locations of these stations are given in Figure 1.1.

The observations for these stations are available for six days. Two days in year 2019 (day of year 044 and 045), two in 2020 (day of year 179 and 180), and two in 2021 (days 095 and 096). The observations are distributed in a mixture of RINEX 2 and RINEX 3 formats with related short and long filenames. Within the campaign area we use a Bernese specific naming convention for observation and navigation RINEX files that is neutral with respect to the original very different filenaming. The observation files contain also measurements from different GNSS, indicated in Table 1.1 by the RINEX system characters:

G GPS Global Positioning System R **GLONASS** Russian Global Navigation Satellite System Ε Galileo European Galileo navigation satellite system  $\mathbf{C}$ BDS BeiDou Navigation Satellite System J QZSS Quasi-Zenith Satellite System **NAVIC** Navigation with Indian Constellation

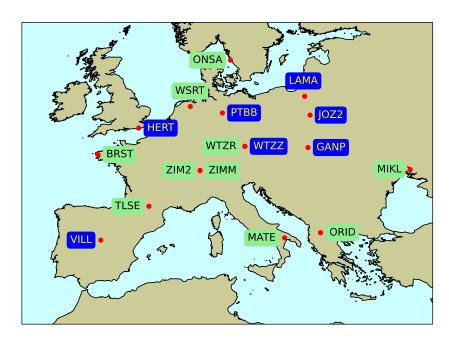


Figure 1.1: Stations used in example campaign (stations indicated by light boxes with coordinates given in the IGS 20 reference frame)

**Table 1.1:** List of stations used for the example campaign including receiver and antenna type as well as the antenna height.

	e antenna neight.		D . 1		Α
Station name	Location		Receiver type Antenna type	Radome	Antenna height
BRST00FRA 10004M004	Brest, France		TRIMBLE ALLOY TRM57971.00	NONE	(GREC) 2.0431 m
GANPOOSVK 11515M001	Ganovce, Slovakia		TRIMBLE ALLOY TRM59800.00	SCIS	$\begin{array}{c} (GRECJI) \\ 0.3830\mathrm{m} \end{array}$
HERTOOGBR 13212M010	Hailsham, United Kingdom		LEICA GRX1200G LEIAT504GG	GPRO NONE	$\begin{array}{c} (GR) \\ 0.0000\mathrm{m} \end{array}$
J0Z200P0L 12204M002	Jozefoslaw, Poland		LEICA GRX1200G LEIAT504GG :TRIMBLE NETR9 TRM59800.00	GPRO NONE NONE	(GRCJ) 0.0000 m (GRE) 0.0000 m
LAMA00POL 12209M001	Olsztyn, Poland		LEICA GRX1200+ LEIAT504GG	GNSS LEIS	$(GR) \\ 0.0600  m$
MATEOOITA 12734M008	Matera, Italy		LEICA GR30 LEIAR20	NONE	$\begin{array}{c} (GREC) \\ 0.1010\mathrm{m} \end{array}$
MIKLOOUKR 12335M001	Mykolaiv, Ukraine		LEICA GR10 LEIAR10	NONE	$^{\rm (GRE)}_{\rm 0.0237m}$
ONSA00SWE 10402M004	Onsala, Sweden		JAVAD TRE_G3TH AOAD/M_B :SEPT POLARX5TR AOAD/M_B	OSOD	(GR) 0.9950 m (GREC) 0.9950 m
ORIDOOMKD 15601M001	Ohrid, Republic of North Macedonia		LEICA GRX1200G LEIAT504GG :LEICA GR30 LEIAR25.R4	GPRO LEIS LEIT	(GR) 0.0640 m (GREC) 0.0000 m
PTBB00DEU 14234M001	Braunschweig, Germany		ASHTECH Z-XII3 ASH700936E :SEPT POLARX5TR LEIAR25.R4	SNOW	(G) 0.0562 m (GRECI) 0.0590 m
TLSE00FRA 10003M009	Toulouse, France		TRIMBLE NETR9 TRM59800.00	NONE	$\begin{array}{c} (GREC) \\ 1.0530\mathrm{m} \end{array}$
VILL00ESP 13406M001	Villafranca, Spain		SEPT POLARX4 SEPCHOKE_MC :SEPT POLARX5 SEPCHOKE_B3E6	NONE NONE	(GRECI) 0.0937 m (GRECI) 0.0937 m
WSRT00NLD 13506M005	Westerbork, The Netherlands		SEPT POLARX5 AOAD/M_T	DUTD	$\begin{array}{c} (GRECJ) \\ 0.3888\mathrm{m} \end{array}$
WTZR00DEU 14201M010	Kötzting, Germany		LEICA GR50 LEIAR25.R3	LEIT	$\begin{array}{c} (GRECI) \\ 0.0710\mathrm{m} \end{array}$
WTZZ00DEU 14201M014	Kötzting, Germany		JAVAD TRE_3 DE LEIAR25.R3	LTA LEIT	$\begin{array}{c} (GRECI) \\ 0.2840\mathrm{m} \end{array}$
ZIM200CHE 14001M008	Zimmerwald, Switze	erland	TRIMBLE NETR9 TRM59800.00	NONE	$_{0.0000\mathrm{m}}^{\mathrm{(GREC)}}$
ZIMMOOCHE 14001M004	Zimmerwald, Switze	erland	TRIMBLE NETR9 TRM29659.00	NONE	(G) 0.0000 m

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In the terminal sessions you will analyze the data in order to obtain a velocity field based on products computed by Center for Orbit Determination in Europe (CODE) for the IGS repro3 campaign – the contribution to the ITRF 2020. For nine of these stations, coordinates and velocities are given in the IGS 20 reference frame, an IGS–specific realization of the ITRF 2020 (see https://itrf.ign.fr/en/solutions/ITRF2020). Even though the scale of the ITRF 2020 has slightly changed with respect to the repro3 series due to the combination with the other space-geodetic techniques, the orbit and Earth orientation parameters (EOP) products can be used together with the IGS 20 reference frame. For Precise Point Positioning (PPP) applications (in particular with the ambiguity resolution) this inconsistency will, however, degrade the results. A related update of the product series is under consideration.

Between the selected days in 2019 and 2020 the receiver in ONSA and the full equipment at stations JOZ2, ORID, PTBB, and VILL was changed. The receiver type, the antenna type, and the antenna height are provided in Table 1.1.

Ideally, consistent (ROBOT) calibrations for each GNSS that should be used. However, those calibrations are not always available. The antenna calibrations available for the example network and related to the IGS 20 antenna model are shown in Table 1.2. For station ONSA we have the situation that for the antenna type/radome combination even

**Table 1.2:** List of antenna/radome combinations used in the example campaign together with the available antenna calibration values in IGS20 model.

Antenna type		Type of ca	libration	used at stations
AOAD/M_B	OSOD	ADOPTED from NONE	for GPS	ONSAOOSWE
		ADOPTED from GPS	for other systems	
AOAD/M_T	DUTD	ROBOT	for GPS and GLO	WSRTOONLD
	a	ADOPTED from GPS	for other systems	DEDD 0.00 PW (0.010)
ASH700936E	SNOW	ROBOT ADOPTED from GPS	for GPS for other systems	PTBB00DEU (2019)
LEIAR10	NONE	ROBOT	for all systems	MIKLOOUKR
			v	
LEIAR20	NONE	ROBOT	for all systems	MATEOOITA
LEIAR25.R3	LEIT	ROBOT	for all systems	WTZROODEU, WTZZOODEU
LEIAR25.R4	LEIT	ROBOT	for all systems	ORIDOOMKD (2020/21), PTBB00DEU (2020/21)
LEIAT504GG	LEIS	ROBOT	for GPS and GLO	LAMAOOPOL,
				ORIDOOMKD (2019)
		ADOPTED from GPS	for other systems	
LEIAT504GG	NONE	ROBOT	for all systems	HERTOOGBR,
SEPCHOKE_B3E	C NONE	ROBOT	for all systems	JOZ200POL (2019) VILL00ESP (2020/21)
			v	, , ,
SEPCHOKE_MC	NONE	ROBOT ADOPTED from GPS	for GPS and GLO for other systems	VILLOOESP (2019)
TDM000F0 00	NONE		v	ZIMMOOGUE
TRM29659.00	NONE	ROBOT	for all systems	ZIMMOOCHE
TRM57971.00	NONE	ROBOT	for all systems	BRST00FRA
TRM59800.00	NONE	ROBOT	for all systems	JOZ200POL (2020/21),
TRM59800.00	SCIS	ROBOT	for all systems	TLSEOOFRA, ZIM2OOCHE GANPOOSVK

no calibration is available. The values from the same antenna type without radome are used as a proxy.

The distances between occupied locations in the network are between 200 and  $1000\,\mathrm{km}$ . Two locations (Zimmerwald and Kötzting) are occupied by receiver/antenna pairs defining separate stations each: in Zimmerwald, the distance between ZIMM and ZIM2 is only  $19\,\mathrm{m}$ . In Kötzting, WTZR and WTZZ are separated by less than  $2\,\mathrm{m}$  — these are short baselines.

The receivers used at the stations MATE, ONSA, and PTBB are connected to H-Maser clocks, VILL to a Cesium clock. In 2020 and 2021 both receivers in Kötzting (WTZR and WTZZ) were connected to the same H-Maser (EFOS 18); in 2019 WTZZ is running on the internal receiver clock only.

# 1.2 Directory Structure

The data belonging to this example campaign is included in the distribution of the *Bernese GNSS Software*. Therefore, you may also use this document to generate solutions from the example dataset to train yourself in the use of the *Bernese GNSS Software* outside the environment of the *Bernese Introductory Course*.

There are three areas relevant for the data processing (in the environment of the Bernese Introductory Course they are all located in the \${HOME}/GPSDATA directory):

- \${D}: The DATAPOOL area is intended as an interface where all external files can be deposited after their download. It can be used by several processing campaigns.
- \${P}: The CAMPAIGN54 directory contains all processing campaigns for the Version 5.4 of the Bernese GNSS Software. In the Bernese Introductory Course environment all groups use \${P}/INTRO in their \${HOME} directory.
- \${S}: The SAVEDISK area serves as a product database where the result files from different processes/projects can be collected and archived. Before you start processing, only reference files (\*.\*\_REF) obtained with the example BPE from the distribution are available here.

### 1.2.1 The DATAPOOL Directory Structure (\${D})

#### Motivation for the DATAPOOL area

The idea of the DATAPOOL area is to place local copies of external files somewhere on your filesystem. It has several advantages compared to downloading the data each time when starting the processing:

- The files are downloaded only once, even if they are used for several campaigns.
- The data download can be organized with a set of scripts running independently from the Bernese GNSS Software environment, scheduled by the expected availability of the external files to download.
- The processing itself becomes independent from the availability of external data sources.

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#### Structure and content of the DATAPOOL area

The DATAPOOL area contains several subdirectories taking into account the different potential sources of files and their formats:

#### GNSS observation data:

#### RINEX-directory:

The data of GNSS stations are provided in Receiver INdependent EXchange format (RINEX) format. The directory may contain observation (Hatanaka-compressed) and navigation files. The files in this area may even be UNIX-or gzip-compressed. RINEX file versions 2, 3 or 4 are supported and may be introduced with the related filenaming conventions. If several versions of RINEX files for the same station are available, the script (RNX\_COP) decides on the priority.

These RINEX files are "originary" files that are not changed during the processing.

The RINEX files can be downloaded from international data centers. Project-specific files are copied into this area. If you mix the station lists from different projects, it is recommended to keep the four-character IDs of all stations in the RINEX file names unique (even if the remaining characters of the nine-character ID is different). The reason is that a mixture with four-character IDs from files using the RINEX 2 naming convention may lead to unpredictable effects.

#### HOURLY-directory:

The same as the RINEX directory but dedicated to hourly RINEX data used for near real-time applications.

Note: not all stations in this example provide hourly RINEX files.

#### GNSS products:

Orbits, EOP, satellite clock corrections and biases are basic external information for a GNSS analysis. We propose to use the ID of the source also to create related subdirectories:

- COD CODE operational final solution, currently containing GPS and GLONASS.
- COR CODE operational rapid solution, containing GPS, GLONASS and Galileo.
- CODORO3FIN CODE contribution to IGS repro3, containing GPS, GLONASS, Galileo.
- CODOMGXFIN CODE experimental Multi-GNSS Extension (MGEX) solution, containing GPS, GLONASS, Galileo, BeiDou, and QZSS.
- IGS combined IGS products from the final series, containing only GPS. GLONASS may be added from separate files (named IGL)

The files are provided at the download area from CODE (http://www.aiub.unibe.ch/download/CODE/). Instructions for downloading of combined IGS products as well as selected analysis center results are available at https://files.igs.org/pub/product/readme.txt

Note that not all analysis centers/solution series provide clock corrections and consistent bias products.

The files are either named with the GPS-week and the day of the week (apart from files containing information for the entire week, e.g., ERP, or the processing summaries). Alternatively also the new naming convention with long filenames indicating the beginning of the dataset with year, day of year, hour and minute together with the length of the dataset in the file is supported. Unix- or gzip-compressions may be applied to these files.

The IGS provides GPS and GLONASS orbits only in separate files (IGS/IGL-series from the final product line) stemming from independent combination procedures with different contributing analysis centers. Nevertheless, they are consistent enough to merge both files together as the first step of the processing. CODE as well as other analysis centers of the IGS (e.g., ESA or GFZ) contribute fully combined multi-GNSS solutions to the IGS final (and ultra-rapid) product line. The contributions of these groups are already combined GPS/GLONASS orbit files.

#### BSW54 :

In this directory we have placed files containing external input information in Bernese specific formats. The files are neutral with respect to the data you are going to process. Typically these are ionosphere maps. If the external GNSS product source does not provide biases also Bernese formatted bias files (OSB) may be placed here. Please take care on the consistency in this case. These files can be downloaded from http://www.aiub.unibe.ch/download/CODE/ or http://www.aiub.unibe.ch/download/BSWUSER54/ areas.

#### REF54:

Here we propose to collect files in Bernese format which are useful for several campaigns.

- At first these are the series of files related to the reference frame (e.g., IGB14 or IGS20) that may be updated via the download area http://www.aiub.unibe.ch/download/BSWUSER54/REF
- Typically, also the sets of files related to the project-specific station selection are prepared in this area. This allows the usage of these files in various campaigns for data processing. Such files are typically: station coordinates (CRD) and velocities (VEL), station information (STA) and equipment conflict (CRX) files, abbreviation (ABB), cluster (CLU) and tectonic plate (PLD) tables, as well as the ocean and atmosphere tidal loading correction tables (BLQ and ATL, respectively).
- The antenna phase center variation (PCV) are expanded according to the used antennas. The supported systems are extracted from the observation selection file (SEL). Both are sitting in the campaigns GEN directory for processing. If these files are prepared project-specific they can be taken from this REF54 directory, alternatively one file serving for all projects can be prepared and placed in the \${CONFIG} folder.
- The directory is also the source of the SINEX and IONEX skeleton files to be used to generate the various result files.

#### **LEO-data** (e.g., GRCC or GRCD):

These directories are intended to host files which are necessary for Low Earth Orbiter (LEO) data processing. For each LEO satellite a separate folder is foreseen.

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RINEX observation files (version 2 or 3/4) are stored in the subdirectory RINEX (of the LEO directory). The corresponding attitude files are placed in the subdirectory ATTIT.

Both locations can be addressed with PCF variables "V\_RNXDIR" and "V\_ATTDIR". Also the names of the files in these directories are flexible and can be specified via the variables "V\_RNXFIL" and "V\_ATTFIL" respectively.

These files are needed to run the example BPE on LEO orbit determination (LEOPOD. PCF). They are not used in the example during the *Bernese Introductory Course*.

#### SLR\_data:

The Satellite Laser Ranging (SLR) data is provided either in the traditional quicklook normal point or the current consolidated range data format. The directory contains the normal point files downloaded from the International Laser Ranging Service (ILRS) data centers.

These files are needed to run the example BPE on orbit validation using SLR observations (SLRVAL.PCF). They are not used during the *Bernese Introductory Course*.

#### Miscellaneous directories:

The files in these directories are not necessary for executing the example BPEs as they are provided with the software. They indicate additional files which might be useful when processing your own data with adjusted processing setups.

#### StationLogs:

This directory contains the site log files (e.g., from https://files.igs.org/pub/station/log) providing the information on the equipment history in order to be filled into the station information file to check the header of the RINEX observation files before processing.

#### VMF3:

The grids for the Vienna Mapping Function (VMF) (Version 3) are located in a separate directory. They are not used for the examples but it shall indicate that for other types of files other directories may be created.

#### GFZloading:

Grid files with crustal deformation corrections from the non-tidal loading models as provided by GFZ Potsdam for atmosphere, ocean and hydrology non-tidal loading effects.

All files and meta-information related to the 17 stations selected for the example campaign are already in this DATAPOOL—area (\${D}) after installing the *Bernese GNSS Software*. GNSS orbit information is available from CODE (legacy, rapid and MGEX) and IGS (directories \${D}/COD, \${D}/COR, \${D}/CODOMGXFIN or \${D}/IGS, respectively).

#### 1.2.2 The Campaign-Directory Structure

### Putting data from the DATAPOOL into the campaign

When running an automated processing using the BPE there is a script at the beginning of the process which copies the data from the DATAPOOL—area into the campaign. If you

are going to process data manually you first have to copy the necessary files into the campaign and decompress them if necessary using standard utilities (uncompress, gunzip<sup>1</sup>, or CRZ2RNX for RINEX-files).

# Content of the campaign area to process the example

All files needed to process the data according to this tutorial are already copied into the campaign area. If you want to follow the example outside the *Bernese Introductory Course* environment you have to put the following files at the correct places in the campaign directory structure.

\${P}/INTRO/ATM/	HOI_20190440.ION HOI_20201800.ION	HOI_2019045 HOI_2021095		HOI_20201790.ION HOI_20210960.ION	
4 (D) (TVMDQ (DF= )	-			-	
\${P}/INTRO/BPE/					
\${P}/INTRO/GEN/		ANTENNA_I2O.PCV SESSIONS.SES	I14.ATX SINEX_INTR	I20.ATX	
\${P}/INTRO/GRD/	VMF3_20190440.GRD VMF3_20201800.GRD	VMF3_201904 VMF3_202109		VMF3_20201790.GRD VMF3_20210960.GRD	
\${P}/INTRO/OBS/					
\${P}/INTRO/ORB/	COD_20190440.PRE	COD_2019044	O.TEP		
ψ (1 ), 1M1M3, 3M2,	COD_20190450.PRE	COD_2019045			
	COD_20201790.PRE	COD_2020179	O.IEP	COD_20201790.BIA	
	COD_20201800.PRE	COD_2020180		COD_20201800.BIA	
	COD_20210950.PRE	COD_2021095	O.IEP	COD_20210950.BIA	
	COD_20210960.PRE	COD_2021096	O.IEP	COD_20210960.BIA	
	IGS_20190440.PRE	IGS_2019045		IGS_20201790.PRE	
	IGS_20201800.PRE	IGS_2021095		IGS_20210960.PRE	
	IGS_20190440.IEP	IGS_2019045		IGS_20201790.IEP	
	IGS_20201800.IEP	IGS_2021095		IGS_20210960.IEP	
	IGL_20190440.PRE	IGL_2019045		IGL_20201790.PRE	
	IGL_20201800.PRE	IGL_2021095		IGL_20210960.PRE	
	BIA_20190440.0SB BIA_20201800.0SB	BIA_2019045 BIA_2021095		BIA_20201790.0SB BIA_20210960.0SB	
	B1A_20201000.03B	BIA_2021095	0.055	BIA_20210900.USB	
\${P}/INTRO/ORX/					
\${P}/INTRO/OUT/	COD_20190440.CLK	COD_2019045	O.CLK	COD_20201790.CLK	
	COD_20201800.CLK	COD_2021095	O.CLK	COD_20210960.CLK	
	IGS_20190440.CLK	IGS_2019045	O.CLK	IGS_20201790.CLK	
	IGS_20201800.CLK	IGS_2021095	O.CLK	IGS_20210960.CLK	
\${P}/INTRO/RAW/	BRST00FRA_20190440.	RXO BRSTOOFRA_2	0190450.RX0	BRST00FRA_20201790.RX	0
	BRST00FRA_20201800.	RXO BRSTOOFRA_2	0210950.RXO	BRST00FRA_20210960.RX	0
	GANPOOSVK_20190440.	_		GANP00SVK_20201790.RX	
	GANPOOSVK_20201800.			GANPOOSVK_20210960.RX	
	HERT00XYZ_20190440.I			HERTOOXYZ_20201790.RX	
	HERTOOXYZ_20201800.I			HERT00XYZ_20210960.RX	
	JOZ200XYZ_20190440.I		0190450.RX0	J0Z200POL_20201790.RX	
	JOZ200POL_20201800.I LAMA00XYZ_20190440.I		0210950.RX0	JOZ200POL_20210960.RX LAMA00XYZ_20201790.RX	
	LAMA00XYZ_20201800.1			LAMA00XYZ_20201730.RX	
	MATEOOITA_20190440.I			MATEOOITA_20201790.RX	
	MATEOOITA_20201800.I			MATEOOITA_20210960.RX	
	MIKLOOUKR_20190440.			MIKLOOUKR_20201790.RX	
	MIKL00UKR_20201800.			MIKL00UKR_20210960.RX	
	ONSA00XYZ_20190440.		0190450.RXO	ONSA00SWE_20201790.RX	
	ONSA00SWE_20201800.	RXO ONSAOOSWE_2	0210950.RXO	ONSA00SWE_20210960.RX	
	ORID00XYZ_20190440.		0190450.RXO	ORID00MKD_20201790.RX	
	ORIDOOMKD_20201800.		0210950.RX0	ORIDOOMKD_20210960.RX	
	PTBB00XYZ_20190440.I			PTBB00DEU_20201790.RX	
	PTBB00DEU_20201800.I			PTBB00DEU_20210960.RX	
	TLSE00FRA_20190440.I			TLSE00FRA_20201790.RX	
	TLSE00FRA_20201800.I			TLSE00FRA_20210960.RX	
	VILLOOESP_20190440.I			VILLOOESP_20201790.RX	
	VILL00ESP_20201800.F WSRT00NLD_20190440.F			VILLOOESP_20210960.RX WSRTOONLD_20201790.RX	
	WSRT00NLD_20130440.1		0210950.RX0	WSRT00NLD_20210960.RX	
	5.0.1 0 0 2.5 _ 2.0 2.0 1.0 0 0 . 1				-

<sup>&</sup>lt;sup>1</sup>These tools are also available for WINDOWS-platforms, see www.gzip.org. Note, that gunzip can also be used to uncompress UNIX-compressed files with the extension .2.

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```
WTZR00DEU_20201790.RX0
                    WTZR00DEU_20190440.RX0
                                               WTZR00DEU_20190450.RX0
                   WTZROODEU 20201800.RX0
                   WTZZ00DEU_20190440.RX0
                                                                          WTZZ00DEU_20201790.RX0
                                               WTZZ00DEU 20190450.RX0
                   WTZZ00DEU_20201800.RX0
                                               WTZZ00DEU_20210950.RX0
                                                                          WTZZ00DEU_20210960.RX0
                   ZIM200CHE_20190440.RX0
                                               ZIM200CHE 20190450.RX0
                                                                          ZIM200CHE 20201790.RX0
                   ZIM200CHE_20201800.RX0
                                               ZIM200CHE_20210950.RX0
                                                                          ZIM200CHE_20210960.RX0
                    ZIMMOOCHE_20190440.RX0
                                               ZIMM00CHE_20190450.RX0
                                                                          ZIMM00CHE_20201790.RX0
                   ZIMMOOCHE 20201800.RX0
                                               ZIMMOOCHE 20210950.RX0
                                                                          ZIMMOOCHE 20210960.RX0
${P}/INTRO/SOL/
${P}/INTRO/STA/
                   EXAMPLE.CRD
                                       EXAMPLE. VEL
                                                           EXAMPLE.STA
                                                                               EXAMPLE.ABB
                   EXAMPLE.BLQ
                                       EXAMPLE.ATL
                                                           EXAMPLE.CLU
                                                                               EXAMPLE.PLD
                   IGB14_R.CRD
                                       IGB14_R.VEL
                                                           IGB14.PSD
                    IGB14.FIX
                                       IGB14.SIG
                   IGB20_R.CRD
                                       IGB20_R.VEL
                                                           IGB20.PSD
                    IGB20.FIX
                                       IGB20.SIG
```

The directories \${P}/INTRO/MODEL/ and \${P}/INTRO/CONFIG/ contains copies of the files from the \${MODEL}- and \${CONFIG}-directories, which are used by the processing programs. If you want to view these files, please use those in your campaign and not the ones in the \${MDOEL}- and \${CONFIG}-directory to prevent potential interferences with your colleagues.

\${P}/INTRO/MDOEL/	CONST.BSW	EGM2008_SMALL.GRV	IAU2000R06.NUT	IERS2010XY.SUB
	OT_FES2004.TID	s1_s2_def_ce.dat	TIDE2000.TPO	
\${P}/INTRO/CONFIG/		DATUM.BSW	GPSUTC.BSW	POLOFF.POL
	SAT_2019.CRX	SAT_2020.CRX	SAT_2021.CRX	
	SATELLIT_I14.SAT	SATELLIT_I20.SAT		

# 1.2.3 Input Files for the Processing Examples

#### Atmosphere files ATM

The input files in this directory are global ionosphere models in the Bernese format obtained from the IGS processing at CODE. They will be used to support the phase ambiguity resolution with the QIF strategy and to enable the higher order ionosphere (HOI) corrections.

#### General files GEN

In this directory campaign-specific configuration files are located:

- SESSIONS.SES: Session table
  Initial versions for daily and hourly processing are available in \${PAN}.
- OBSERV.SEL: Priority list for selecting the observation/signal types for processing Initial versions as they are used for the processing at CODE can be downloaded from BSW aftp.
- ANTENNA\_I14.PCV or ANTENNA\_I20.PCV: Phase center eccentricities and variations Updates are needed when introducing new antenna corrections or new antenna/radome combinations. An Initial version can be downloaded from BSW aftp and updated with ATX2PCV.

• I14.ATX or I20.ATX: Antenna correction model in ANTenna correction EXchange format (ANTEX)

The corrections are related to IGS 14 and IGS 20, respectively.

• SINEX header information

The metadata in the SINEX header has to be adapted to your institution, It is applied to the general SINEX as well as troposphere and bias SINEX.

The skeleton of this file is available at \${DOC}/EXAMPLE\_SINEX.SKL; versions prepared for the distributed processing examples are copied from \${D}/REF54 into the campaign:

SINEX.RNX2SNX for RNX2SNX example SINEX.IONDET for IONDET example

• IONEX header information

Adapt IONEX header for your institution

The skeleton of this file is available at \${DOC}/EXAMPLE\_IONEX.SKL; versions prepared for the distributed processing examples are copied from \${D}/REF54 into the campaign:

IONEX.IONDET for IONDET example

#### Grid files GRD

In this directory the grid files \*.GRD are collected. To apply, e.g., the VMF troposphere model (a priori information from European Centre for Medium-Range Weather Forecasts (ECMWF) and Vienna mapping function) you need a grid with the necessary coefficients.

#### Orbit files ORB

The precise orbits in the files \*.PRE are usually the final products from CODE analysis center containing Global Positioning System (GPS), Russian Global Navigation Satellite System (GLONASS), and potentially European Galileo navigation satellite system (Galileo) orbits from a rigorous multi–GNSS analysis. Alternatively also the combined final products from the IGS can be used. They do not contain orbits for the GLONASS satellites. The combined GLONASS satellite orbits from the IGS are available in IGL–files. Both precise orbit files need to be merged for a multi–GNSS analysis. The corresponding EOP are given in weekly files with the extension \*.IEP (take care on full consistency with the orbit product).

Furthermore, the directory contains code and phase bias values either in the international Bias SINEX format with the extension BIA or the Bernese internal format for Observation-specific signal biasess (OSBs) (extension OSB).

#### Clock RINEX files OUT

The clock RINEX files are located in the OUT-directory. They are consistent with the GNSS orbits, EOP, and potentially the phase bias products in the ORB-directory. They contain station and satellite clock corrections with at least 5 minutes sampling — there

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are also files from the IGS or some of the analysis centers (ACs) providing satellite clock corrections with a sampling of 30 seconds.

#### RINEX files RAW

The raw data are given in RINEX format. In the Bernese internal naming scheme, we combine the nine character long station identifier with the four digit year and the session. The extension is RXO for observation and RXN for navigation RINEX files. In case of RINEX 2 files, the four characters are just extended by the default naming OOXYZ.

#### Station files STA

The coordinates and velocities of the stations given in the IGS realization of the reference frame ITRF 2020 are available in the files IGS20\_R.CRD and IGS20\_R.VEL. For some stations also corrections for Post Seismic Deformation (PSD) need to be applied that are provided in the file IGS20.PSD. The IGS core stations are listed in IGS20.FIX. This file will be used to define the geodetic datum when estimating station coordinates. The files for the previously published IGS realization (based on ITRF 2014) are available as well: IGB14\_R.CRD, IGB14\_R.VEL, IGB14.PSD and IGB14.FIX. You can browse all these files with a text editor or with the menu ("Menu>Campaign>Edit station files").

For all stations that have unknown coordinates in the IGS 20 reference frame a PPP using the example BPE (PPP.PCF) for day 044 of year 2019 has been executed. For our EXAMPLE-project a resulting coordinate file EXAMPLE.CRD has been generated. It contains all IGS core sites (copied from file IGS20\_R.CRD) and the PPP results for the remaining stations. The epoch of the coordinates is January 01, 2015. The corresponding velocity file EXAMPLE.VEL contains the velocities for the core sites (copied from file IGS20\_R.VEL) completed by the NNR-NUVEL1A velocities for the other stations. The assignment of stations to tectonic plates is given in the file EXAMPLE.PLD.

To make sure that you process the data in the Bernese GNSS Software with correct station information (station name, receiver type, antenna type, antenna height, etc.) the file EXAMPLE.STA is used to verify the RINEX header information. The reason to use this file has to be seen in the context that some antenna heights or receiver/antenna types in the RINEX files may not be correct or may be measured to a different antenna reference point. Similarly, the marker (station) names in the RINEX files may differ from the names we want to use in the processing. The antenna types have to correspond to those in the file \${P}/INTRO/GEN/ANTENNA\_I2O.PCV to ensure that the correct phase center offsets and variations are used.

For each station name unique four— and two—character abbreviations to construct the names for the Bernese observation files need to be defined in the file EXAMPLE.ABB. It was automatically generated by the PPP—example BPE. If you want to process big networks, the baselines need to be divided into clusters to speed up the processing. For that purpose each station has to be assigned to a region by a cluster number in the file EXAMPLE.CLU.

**Table 1.3:** List of global model and configuration files to be used in the Bernese programs for the processing example.

Model files located in \${C}/GLOBAL/MODEL

Filename	Content	Modification	Update from
CONST.BSW	All constants used in the	No	BSW aftp
	Bernese GNSS Software		
IAU2000R06.NUT	Nutation model	No	_
	coefficients		
IERS2010XY.SUB	Subdaily pole model	No	_
	coefficients		
OT_FES2004.TID	Ocean tides coefficients	No	_
TIDE2000.TPO	Solid Earth tides	No	_
	coefficients		
EGM2008_SMALL.	Earth potential coefficients	No	_
GRV			
	(reduced version, sufficient		
	for GNSS and LEO orbit		
	determination)		
s1_s2_def_ce.	S1/S2 atmospheric tidal	No	_
dat	loading coefficients		

Configuration files located in \${C}/GLOBAL/CONFIG

Filename	Content	Modification	Update from
DATUM.BSW	Definition of geodetic	Introducing new reference	BSW aftp
	datum	ellipsoid	
GPSUTC.BSW	Leap seconds	When a new leap second is	BSW aftp
		announced by the IERS	
POLOFF.POL	Pole offset coefficients	Introducing new values	
		from IERS annual report	
		(until 1997)	
SATELLIT_I14.	Satellite information file	New launched satellites	BSW aftp
SAT or SATELLIT_			
I20.SAT			
BOXWING.MAC	Definition of orbit model	New launched satellites	BSW aftp
	parameters		
SAT_\$Y+0.CRX	Satellite problems	Satellite maneuvers, bad	BSW aftp
		data,	

The last files to be mentioned in this directory are EXAMPLE.BLQ and EXAMPLE.ATL. They respectively provide the coefficients for the ocean and atmospheric tidal loading of the stations. They should at least be applied in the final run of GPSEST.

#### Global model and configuration files MODEL and CONFIG

These general model and configuration files contain information that is neither user— nor campaign—specific. They are accessed by all users, and changes in these files will affect processing for everyone. Consequently, these files have a central location in either \${C}/GLOBAL/MODEL or \${C}/GLOBAL/CONFIG directory. Table 1.3 shows the list of these files necessary for the processing example. It also shows which files need to be updated from time to time by downloading them from the anonymous ftp—server of AIUB (http://www.aiub.unibe.ch/download/BSWUSER54/).

The antenna corrections have to be applied consistently to the reference frame. Since we are going to use the IGS 20 coordinates and velocities to define the datum for our solution,

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the related antenna correction model has to be applied: IGS20.ATX. The related values are available in the SATELLIT\_I20.SAT and ANTENNA\_I20.PCV files. The first of these two files is located in the central CONFIG directory the latter one is contained in the campaign-specific GEN directory because it should contain only antenna/radome combinations that are processed in this campaign. Disregarding the two locations it is essential to use these two files consistently. For the ITRF 2014 related datum, the files are available as well: SATELLIT\_I14.SAT and ANTENNA\_I14.PCV.

Each Bernese processing program has its own panel for general files. Make sure that you use the correct files listed in Table 1.3.

# 1.2.4 The SAVEDISK Directory Structure (\${S})

#### Motivation for the SAVEDISK area

When processing GNSS data, a lot of files from various processing steps will populate your campaign directories. The main result files from the data analysis are collected in the SAVEDISK area. This area is intended as long-term archive for your result files.

Because the result files are stored in the SAVEDISK area, you can easily clean up your campaign area without loosing important files. Please keep in mind that the computing performance decreases if you have several thousands of files in a directory.

#### Structure and content of the SAVEDISK area

We propose to build subdirectories in the SAVEDISK area for each of your projects. If these projects collect data over several years, yearly subdirectories are recommended. It is also practical to use further subdirectories like ATM, ORB, OUT, SOL, STA to distribute the files and to get shorter listings if you are looking for a file. For reproducibility and documentation of the obtained results it might be a good idea to add the used configuration and model files to a GEN directory.

The SAVEDISK area contains after its installation a directory structure according to the description above. Each example BPE is assumed as a project. Therefore, you will find on the top level of the SAVEDISK the directories PPP, RNS2SNX, CLKDET, IONDET, LEOPOD, and SLRVAL (related to the different example BPEs).

In each of these directories you will find several files ending with \_REF. They are generated by running the example BPEs on the system at Astronomical Institute of the University of Bern (AIUB). Even though this tutorial and of the RNX2SNX example BPE are both aiming on processing the same dataset from a regional network, the results will not be identical since there are some differences in the processing strategies and selected options.

# 1.2.5 Preparing the environment

For the *Bernese Introductory Course* the three data areas are prepared in such a way that the processing of the example data can immediately start. Some of the data are copied from the DATAPOOL-area into the CAMPAIGN-area and decompressed. An archive file containing the related files is made available that can be installed by

```
cd $D
cd ..
tar -xvzf _where_ever_it_is_/TUTORIAL.tgz
```

(\_where\_ever\_it\_is\_ should be replaced by the location of this archive file on your system)

On Windows platforms you have to copy the three subdirectories of the archive to the related locations using the Windows File Explorer.

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# 2 Terminal Session: Monday

Today's terminal session is to:

- 1. become familiar with the UNIX environment, the menu of the Bernese GNSS Software, and the example campaign,
- 2. verify the campaign setup done for you (see sections 2.2 and 2.3, and also the handout for the terminal sessions),
- 3. generate the a priori coordinates for all 4 days using COOVEL (see Section 2.5), and
- 4. start to prepare pole and orbit information according to chapter 3.

#### 2.1 Start the Menu

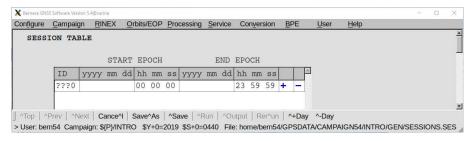
Start the menu program using the command  $G^1$ .

Navigate through the submenus to become familiar with the structure of the menu. Read the general help (available at "Menu>Help>General") to get an overview on the usage of the menu program of the Bernese GNSS Software.

For the terminal session in the *Bernese Introductory Course*, the campaign setup has already been done for each user. Please check that the campaign name in the statusbar of the Bernese Menu is set correctly to your campaign (i.e., Campaign  $\{P\}/INTRO\}$ ) and that the current session is set to the first session (i.e., Y+0=2019, S+0=0440). If this is not the case, please contact the staff in the terminal room.

#### 2.2 Select Current Session

Select "Menu>Campaign>Edit session table" to check the session table. It is recommended to use the wildcard string ???0 for the "List of sessions" in panel "SESSION TABLE". The panel below shows the session definition for a typical permanent campaign with 24-hours sessions. The setup of the session table is a very important task when you prepare a campaign. Please read the corresponding online help carefully.



<sup>&</sup>lt;sup>1</sup>At the exercise terminals the Bernese environment is loaded automatically during login. At home you have to source the file \${C}/LOADGPS.setvar on UNIX—platforms either manually or during login.

Save the session table (press the `Save button) and open the "Date Selection Dialog" in the "Menu>Configure>Set session/compute date" in order to define the current session:



# 2.3 Campaign Setup

Usually, a new campaign must be added to the campaign list ("Menu>Campaign>Edit list of campaigns") first and select it as the active campaign ("Menu>Campaign>Select active campaign"), before the directory structure can be created ("Menu>Campaign>Create new campaign"). In the Bernese Introductory Course environment this should already have been done for your campaign, but please verify that.

In the Bernese Introductory Course environment the selected campaign should be  $\P$  INTRO. In order to become familiar with the campaign structure, you can inspect your campaign directory and inspect the contents using the command line (using cd for changing directories and ls to create directory listings) or using a filemanager (e.g., midnight commander mc).

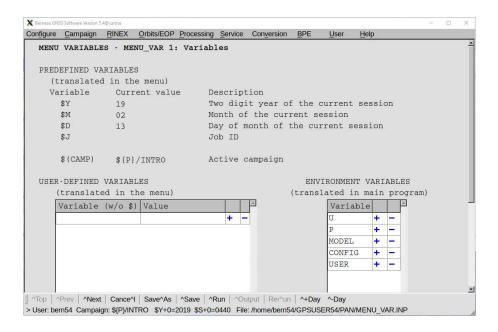
#### 2.4 Menu Variables

When processing GNSS data, it is often necessary to repeat a program run several times with only slightly different option settings. A typical example would be the processing of several sessions of data. The names of observation files change from session to session because the session number is typically a part of the filename. It would be very cumbersome to repeat all the runs selecting the correct files manually every time. For the BPE an automation is mandatory. For such cases the Bernese menu system provides a powerful tool: the so-called menu variables. The menu variables are defined in the user-specific menu input file \${U}/PAN/MENU\_VAR.INP that is accessible through "Menu >Configure>Menu yariables". Three kinds of menu variables are available: predefined variables (also called menu time variables), user-defined variables, and system environment variables.

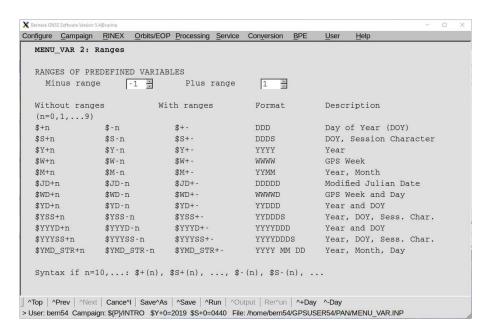
The use of system environment variables is necessary to generate the complete path to the files used in the Bernese GNSS Software. The campaign data are located in the directory \${P}/INTRO=\${HOME}/GPSDATA/CAMPAIGN54/INTRO. The user-dependent files can be found at \${U}=\${HOME}/GPSUSER54 — note that \${HOME} may have been already translated into the name of your home-directory. The temporary user files are

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saved in \${T}=/scratch/local/bern54 (change bern54 to your user name). Finally, the campaign-independent model and configuration files reside in \${MODEL}=\${C}/GLOBAL/MODEL and \${CONFIG}=\${C}/GLOBAL/CONFIG, respectively.



The predefined variables provide a set of time strings assigned to the current session. From the second panel of the menu variables you get an overview on the available variables and their usage:



Be aware that the variable \$S+1 refers to the next session. Because we are using a session table for daily processing it also corresponds to the next day.

These variables are automatically translated by the menu upon saving the panel or running the program. We recommend to make use of them in the input panels (e.g. for filename specification).

# 2.5 Generate A Priori Coordinates

As stated before the a priori coordinates generated from the PPP processing example BPE refer to the epoch January 01, 2015. The first step is to extrapolate the coordinates to the epoch that is currently processed. In the recent ITRF-solutions also Post Seismic Deformation (PSD) corrections have to be applied when the epoch of the coordinate sets are changed. They are provided in the input field "PSD corrections (since ITRF14)". Of course this feature is also included in the IGS 20 frame – the IGS-specific realization of the ITRF 2020.

Coordinate extrapolation is the task of the program COOVEL. Open the program input panel in "Menu> $\underline{S}$ ervice> $\underline{C}$ oordinate tools> $\underline{E}$  $\underline{x}$ trapolate coordinates":

X Bernese GNSS Software Version 5.4@carina	
Configure Campaign RINEX Orbits/EOP Processing Service Conversion BPE User Help	
EXTRAPOLATE COORDINATES - COOVEL 1: Filenames	
GENERAL FILES	
Show all general files	
INPUT FILES	
Input coordinate file EXAMPLE CRD	
Input velocity file EXAMPLE VEL	
PSD correction (since ITRF14) IGS20 PSD	
Annual/semiannual corrections TVC	
REFERENCE EPOCH   yyyy mm dd   hh mm ss	
RESULT FILE	
Output coordinate file APR_\$YYYSS+0 CRD	
Approx. velocities at ref. epoch VEL	
Stations without PSD corrections FIX	
GENERAL OUTPUT FILES	
Program output  wse COOVEL.Lnn or COOVEL	OUT
Error messages merged to program output or ERROR	MSG
TITLE EXAMPLE: Session \$Y+0-\$S+0: Coordinate propagation	
Top   ^Prev   ^Next   Cance^1   Save^As   ^Save   ^Run   ^Output   Rer^un   ^+Day ^-Day   > User: bem54   Campaign: \${P}/INTRO \$Y+0=2019 \$S+0=0440 File: /home/bem54/GPSUSER54/PAN/COOVEL.INP	

"Reference epoch: date" \$YMD\_STR+0  $\rightarrow$  2019 02 13 "Output coordinate file" APR\_\$YYYSS+0  $\rightarrow$  APR\_20190440 "Title line" Session \$Y+0-\$S+0:  $\rightarrow$  Session 2019-0440:

Start the program with the <code>^Run-button</code>. The program generates an output file <code>COOVEL.L??</code> in the directory f(P)/INTRO/OUT. This file may be browsed using the <code>^Output-button</code> or with <code>"Menu>Service>Browse program output"</code>. It should look like

Bernese GNSS Software, Version 5.4

Program : COOVEL
Purpose : Extrapolate coordinates

Campaign : \${P}/INTRO

Default session: 0440 year 2019
Date : 13-Jan-2024 15:36:36
User name : bern54

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```
EXAMPLE: Session 2019-0440: Coordinate propagation
INPUT AND OUTPUT FILENAMES
                        : ${CONFIG}/DATUM.BSW
Geodetic datum
Input coordinate file
                                 : ${P}/INTRO/STA/EXAMPLE.CRD
                                : ${P}/INTRO/STA/EXAMPLE.VEL
 Input velocity file
Output coordinate file
                                 : ${P}/INTRO/STA/APR_20190440.CRD
Approx. velocities at ref. epoch:
PSD corrections (since ITRF14) : ${P}/INTRO/STA/IGS20.PSD Annual/semiannual corrections (s: ---
Stations without PSD corrections:
                                : ${P}/INTRO/OUT/COOVEL.LOO
Program output
                                 : ${U}/WORK/ERROR.MSG
Error message
Session table
                                : ${P}/INTRO/GEN/SESSIONS.SES
REFERENCE EPOCH:
                          2015-01-01 00:00:00
INTERPOLATION FACTOR:
                          -4.1177275838466807
>>> CPU/Real time for pgm "COOVEL": 0:00:00.134 / 0:00:00.134
>>> Program finished successfully
```

The header area of the program output is standardized for all programs of the *Bernese GNSS Software*, Version 5.4. Furthermore each program has a title line that should characterize the program run. It is printed to the program output and to most of the result files. Many program output files furthermore provide a list of input and output files that have been used or generated.

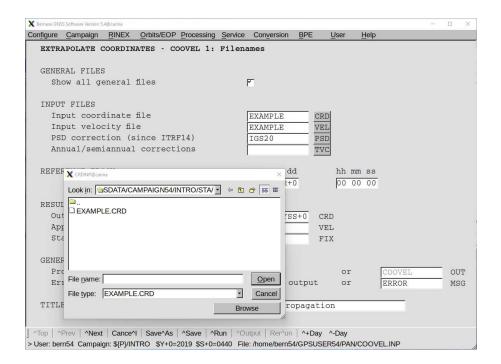
The last two lines of the above example appear also in each program output of the *Bernese GNSS Software*, Version 5.4. It reports the processing time and the status successful or with error.

The result of the COOVEL run is an a priori coordinate file (\${P}/INTRO/STA/APR\_20190440.CRD) containing the positions of the sites to be processed for the epoch of the current session (the lines for the other stations are ignored in the processing):

EXAM	EXAMPLE: Session 2019-0440: Coordinate propagation						
LOCA	LOCAL GEODETIC DATUM: IGS20 EPOCH: 2019-02-13 00:00:00						
NUM	STATI	ION NAME	X (M)	Y (M)	Z (M)	FLAG	
SYST	EM						
43		10004M004		-332746.42891		IGS20	
104		11515M001	3929181.30600	1455236.94546	4793653.99294	PPP	
122	HERT	13212M010	4033460.75197	23538.05244	4924318.34740	PPP	
143	JOZ2	12204M002	3664880.34767	1409190.82194	5009618.55508	PPP	
166	LAMA	12209M001	3524522.69472	1329693.85317	5129846.42860	PPP	
184	MATE	12734M008	4641949.29729	1393045.69364	4133287.66716	IGS20	
198	MIKL	12335M001	3698553.70970	2308676.23252	4639769.62655	IGS20	
235	ONSA	10402M004	3370658.33932	711877.34331	5349787.09062	IGS20	
236	ORID	15601M001	4498451.47983	1708267.26377	4173591.98565	IGS20	
264	PTBB	14234M001	3844059.76112	709661.55883	5023129.64494	PPP	
329	TLSE	10003M009	4627851.66234	119640.28843	4372993.71461	IGS20	
359	VILL	13406M001	4849833.60409	-335048.75055	4116015.05395	PPP	
369	WSRT	13506M005	3828735.65242	443305.18368	5064884.85551	IGS20	
370	WTZR	14201M010	4075580.32457	931854.03929	4801568.27077	IGS20	
371	WTZZ	14201M014	4075579.21251	931853.36368	4801569.13639	PPP	
383	ZIM2	14001M008	4331299.68387	567537.57591	4633133.88020	IGS20	
385	ZIMM	14001M004	4331296.87301	567556.13266	4633134.09986	IGS20	

Have a look at the LOCAL GEODETIC DATUM: in the resulting coordinate file. It is set to IGS20 in this case. If you go back to your input file (e.g., pressing the Rer^un button)

you may open the dialog to select the "Input coordinate file" by pressing on the button next to the input field. Select now the file EXAMPLE.CRD and press the button Browse in order to open a window where the selected file is displayed.



Here you can see the LOCAL GEODETIC DATUM: is set to IGS20\_0. This difference is the indicator whether the PSD corrections have been applied or not. Coordinate files indicating that the PSD corrections have not been applied cannot be used for processing GNSS data. At the same time it is protecting you from applying the corrections twice. For that reason the execution of the program COOVEL for applying the PSD corrections is also essential even if none of the stations in your processing (as in our example) is affected by these corrections.

You can repeat all steps for the other sessions in the example campaign by changing the current session using the ^+Day or ^-Day to change a limited number of days (not sessions) or via "Menu>Configure>Set session/compute date". You can then use the Rer^un button to restart the program. No options need to be changed because of the consequent use of the menu time variables was made. Even if you are going to process only the first day (044, year 2019) of the example dataset during the terminal sessions, you will need prepared coordinate files for all days later on Thursday. That's why, this step should be executed for all six days: 044 and 045 of year 2019, 179 and 180 of year 2020 as well as 095 and 096 of year 2021.

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# 2.6 Session Goals

At the end of this session, you should have created the following files:

1. a priori coordinates in your campaign's STA directory: for six sessions APR\_ 20190440. CRD, APR\_ 20190450. CRD, APR\_ 20201790. CRD, APR\_ 20201800. CRD, APR\_ 20210950. CRD, and APR\_ 20210960. CRD

Until the end of today's terminal session you should start with preparing the pole and orbit information, see Chapter 3.

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# 3 Terminal Session: Pole and Orbit Preparation (Monday/Tuesday)

The terminal session on pole and orbit preparation is to:

- 1. generate the pole information file in the Bernese format (POLUPD)
- 2. generate the Bernese standard orbit files from CODE precise orbit files (ORBGEN)
- 3. extract satellite clock corrections for receiver clock synchronization (SATCLK2)
- 4. convert biases from external to internal format (BIA2OSB)

You should start with these tasks during Monday's terminal sessions and finish the processing during the terminal session on Tuesday.

# **Introductory Remark**

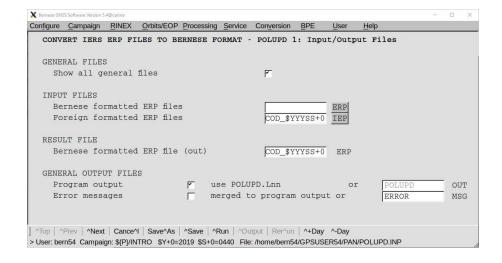
We recommend to use the final or reprocessed products from CODE because they contain consistent orbits for GPS, GLONASS, and Galileo (at least the reprocessing series). Alternatively, the CODE rapid series contains the orbits from all three systems since September 2019. All these series include all active GNSS satellites (even if they are unhealthy or during GPS satellite repositioning events) with the highest possible accuracy thanks to the three-day long-arc technology. Due to this choice you will get the best possible consistency between the external products and the software.

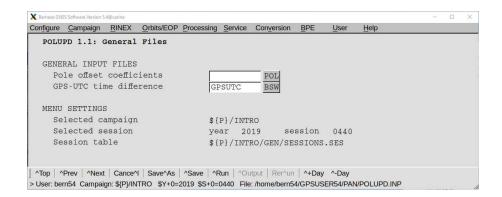
You may alternatively use the products from the IGS. Separate product files for GPS and GLONASS orbits (no combination of Galileo products is executed so far) exist from independent combination procedures that first need to be merged for a multi-GNSS processing. For most of the applications, merging the precise orbit files is sufficient — a tutorial on the procedure is given in Section 7.1 of this book. On the other hand, the consistency of the orbits can not be as good as that of CODE (or other analysis centers) following the strategy of the rigorous combined processing of GPS and GLONASS measurements for orbit determination.

Products for other systems are available in the MGEX solutions provided by CODE and other ACs. Note, that these are experimental products used, e.g., to test various orbit models or antenna calibration sets.

# 3.1 Prepare Pole Information

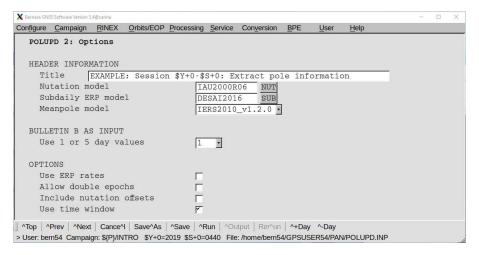
Together with the precise orbit files (PRE), a consistent set of Earth orientation parameters is provided in the ORB directory. Whereas the orbits are given in daily files the EOP may be available in weekly files for some product series (e.g., IGS). We have to convert the information from the International Earth Rotation and Reference Systems Service (IERS)/IGS standard format (file extension within the *Bernese GNSS Software* is IEP) into the internal Bernese EOP format (file extension within the *Bernese GNSS Software* is ERP). This is the task of the program POLUPD ("Menu>Orbits/EOP>Handle EOP files>Convert IERS to Bernese Format") which is also able to update the EOP records of an existing file.



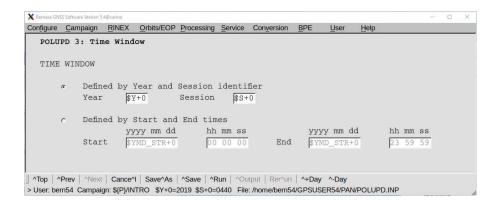


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Together with the nutation and subdaily pole model also the meanpole model is specified in this panel. They are reported in the resulting Bernese formatted pole file in order to check the consistent use.



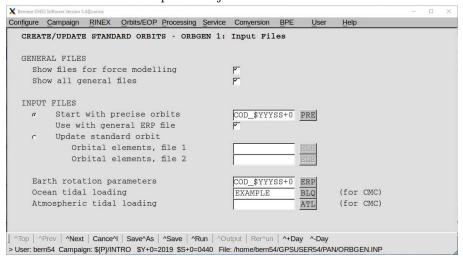
Take care on the option "Meanpole model". It has to be chosen consistently to the EOPs to be converted. The reprocessing series generated in preparation of ITRF 2020 use IERS2010\_v1.2.0. As soon as the ITRF 2020 respective the IGS 20 will be introduced in the operational processing of the IGS this "Meanpole model" has to be used also for the operational products. For earlier operational products the IERS2010XY has to be used. Another important remark is related to option "Use ERP rates": please check your input "Foreign formatted ERP files" whether they contain the midnight epochs requested to cover the entire period of the session (e.g., in the CODE products). In that case the checkbox should be deactivated. If you have only one record per day at noon in your input "Foreign formatted ERP files" (e.g., in IGS products) the rates from the file should be used, the option has to be activated.



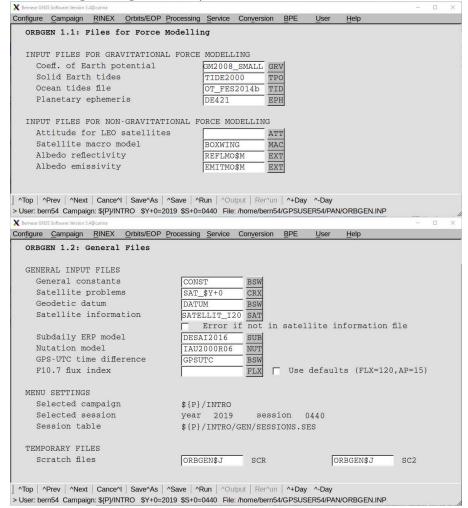
The last panel of the program POLUPD is an example for the specification of time windows in the *Bernese GNSS Software*, Version 5.4. Time windows can be defined by sessions (a single session or a range of sessions). Alternatively, a time window may be specified by a start and an end epoch. By entering either a start or an end epoch you may define only the beginning or the end of the time interval. We refer to the online help for more details.

#### 3.2 Generate Orbit Files

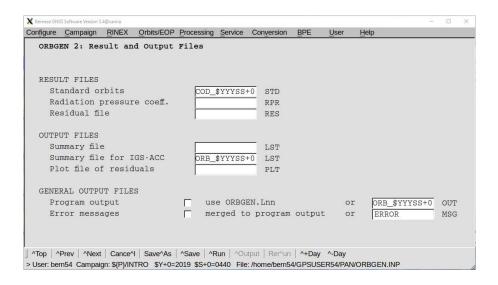
In this processing example we use only one program of the orbit part of the *Bernese GNSS Software*, namely ORBGEN ("Menu>Orbits/EOP>Create/update standard orbits"). It prepares the so-called standard orbits using the satellite positions in the precise orbit files as pseudo-observations for a least-squares adjustment.



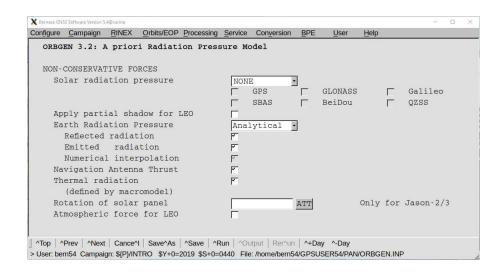
It is mandatory to consistently use the ERP file, the nutation, and the subdaily pole model together with the generated standard orbits in all subsequent processing steps/programs (otherwise a warning message is issued).



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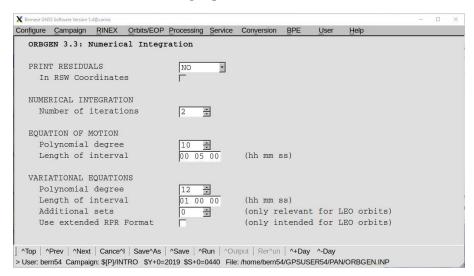


X Bernese GNSS Software Version 5.4@canina	×
Configure Campaign RINEX Orbits/EOP Processing Service Conversion BPE User Help	
ORBGEN 3.1: General Options	
TITLE EXAMPLE: Session \$Y+0-\$S+0: Standard orbit generation	
TIME FRAME, POTENTIAL AND TIDAL CORRECTIONS  Time frame  Earth potential degree  Ocean tides max degree  Apply CMC correction  Apply antenna offset	
SATELLITE SELECTION  P GPS P GLONASS P Galileo P SBAS P BeiDou P QZSS	
The satellites in the following list will be EXCLUDED Y Satellite list	
EXPERIMENTAL OPTIONS  True attitude modelling of satellites  Print beta angle and attitude switch info to STDOUT  Print Yaw angle to SYSOUT file	
7-Top   ^Prev   ^Next   Cance^1   Save^As   ^Save   ^Run   ^Output   Rer^un   ^+Day ^-Day   Ver: bern54   Campaign: \${P}/INTRO   \$Y+0=2019   S\$+0=0440   File: /home/bern54/GPSUSER54/PAN/ORBGEN.INP	

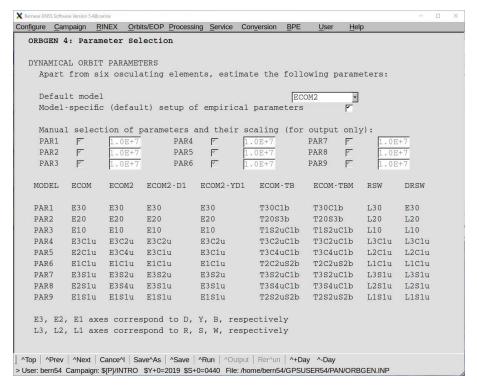


For the options "EQUATION OF MOTION: Length of interval" and "VARIATIONAL EQUATIONS: Length of interval" interval lengths of 00 05 00 and 01 00 00 are sufficient when processing GNSS orbits. Note, that stochastic pulses can only be placed between the integration intervals. When activating the option "Setup of stochastic pulses" later in panel "ORBGEN 5: Stochastic Pulses and Satellite Accuracy Codes" to a specific value (e.g., Orbit midnight) the intervals should allow a realization of this selection.

Shorter intervals will increase of the program run time.



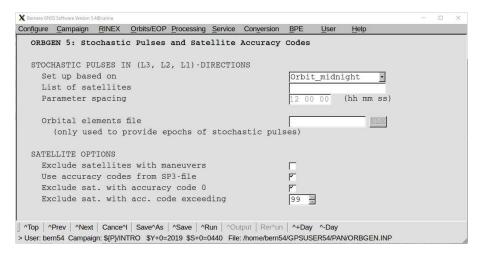
For the option "Number of iterations" 2 is sufficient for most of the cases. You may select a third iteration to check that the output does not change in the additional iteration.



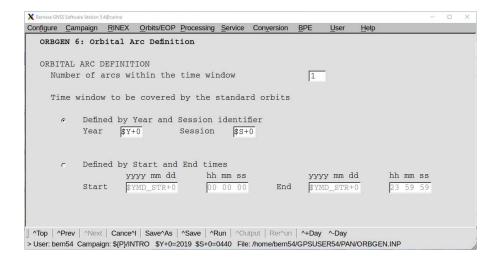
In this panel you see the available orbit models and which of the orbit parameters are used for which component in which coordinate system. We recommend to use the default setup as prepared in the \${CONFIG}/BOXWING.MAC ("Satellite macro model" in "ORBGEN 1.1: Files for Force Modelling").

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The stochastic pulses are set in the CODE reprocessing series – used during this exercise session – at Orbit midnight. For the operational final series from CODE this strategy is applied since GPS week 2112, June 2020. For the other products from CODE or other sources we recommend to change the option "Setup of stochastic pulses" to Spacing with 12 00 00 for "Interval for stochastic pulses".



If you process precise orbits from CODE it is generally not necessary to remove bad satellites, because CODE orbit files contain the correct accuracy codes. It might become necessary for orbit products from other sources (e.g., from the IGS).



The program produces an output file  $ORB_2019044.0UT$  which should look like

Close to the end of the output file you can find the reporting on the orbit model used for each of the satellite according to the "Default model for empirical forces" setup. Most of the GNSS satellites will be represented by the ECOM2 orbit model. Below you find another table that is reporting the parametrization of the selected orbit model.

	EMP MODEL											
EMP:	ECOM2		2	3	4	5	6	7	8	9	10	
EMP:	ECOM2	11	12	13	14	15	16	17	18	19	20	
EMP:	ECOM2	21	22	23	24	25	26	27	28	29	30	
EMP:	ECOM2	31	32	101	102	103	104	105	107	108	109	
EMP:	ECOM2	110	111	113	114	115	116	117	118	119	120	
EMP:	ECOM2	121	122	123	124	126	201	202	203	204	205	
	ECOM2	207	208	209	211	212	213	214	215	218	219	
EMP:	ECOM2	221	224	225	226	227	230	231	233	236		
····· )ESCF	RIPTION OF EMPI	RICAL MODE	 LS									
 DESCE	MODEL 1	 NAME	PAR II									
	MODEL	 NAME	PAR II									
 EMPP <i>I</i>	MODEL MODEL	 NAME	PAR II	1 E30					 F		т	
 EMPP <i>I</i>	MODEL 1 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20					 F F		T T	
EMPP <i>I</i> EMPP <i>I</i>	MODEL 1 AR: ECOM2 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20 3 E10	   				 F F F		T T T	
EMPP! EMPP! EMPP!	MODEL 1 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20	: : : :2u				 F F		T T	
EMPPA EMPPA EMPPA EMPPA	MODEL 1 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20 3 E10 4 E30	: : :2u :4u				F F F F		T T T T	
EMPPA EMPPA EMPPA EMPPA EMPPA	MODEL 1 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20 3 E10 4 E30 5 E30	: :2u :4u				F F F F F		T T T T T	
EMPP I EMPP I EMPP I EMPP I EMPP I EMPP I	MODEL 1 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2 AR: ECOM2	 NAME	PAR II	1 E30 2 E20 3 E10 4 E30 5 E30 6 E10	2u 4u 1u 2u				F F F F		T T T T F	

Inspecting for instance the program output for day 179 in the year 2020, you will find a series of Galileo satellites represented by the ECOM2-YD1 because of their beta angles (elevation angle of the Sun above the orbital plane) are below the limit 14 degrees. The related definition comes from the BOXWING.MAC file:

The most important part of the program output is the quality of the orbit representation by the model defined for the current program run:

RMS	ERRORS	AND MAX. RES	SIDUALS	ARC NUMBE	ER: 1			ITERATI	ON: 2
			QUADRA	TIC MEAN	OF 0-C	(M)	MAX. I	RESIDUAL	S (M)
SAT	#POS	RMS (M)	TOTAL	RADIAL	ALONG	OUT	RADIAL	ALONG	OUT
1	288	0.001	0.001	0.001	0.001	0.001	0.003	0.002	0.003
2	288	0.001	0.001	0.001	0.001	0.001	0.002	0.002	0.003
3	288	0.001	0.001	0.001	0.001	0.001	0.005	0.003	0.001
4	288	0.000	0.000	0.000	0.001	0.000	0.001	0.002	0.001
5	288	0.001	0.001	0.001	0.001	0.001	0.004	0.003	0.001
6	288	0.001	0.001	0.001	0.001	0.001	0.003	0.002	0.003
31	288	0.001	0.001	0.000	0.001	0.000	0.001	0.002	0.001
32	288	0.001	0.001	0.001	0.001	0.001	0.002	0.002	0.003

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101	288	0.001	0.001	0.002	0.001	0.001	0.007	0.004	0.003
102	288	0.001	0.002	0.002	0.001	0.001	0.006	0.004	0.003
103	288	0.001	0.001	0.002	0.001	0.001	0.009	0.004	0.004
124	288	0.001	0.001	0.001	0.001	0.001	0.002	0.002	0.004
126	288	0.001	0.001	0.000	0.001	0.000	0.001	0.002	0.001
201	288	0.001	0.001	0.001	0.001	0.001	0.004	0.003	0.004
202	288	0.001	0.001	0.001	0.001	0.001	0.006	0.003	0.004
124	288	0.001	0.001	0.001	0.001	0.001	0.002	0.002	0.004
126	288	0.001	0.001	0.000	0.001	0.000	0.001	0.002	0.001
201	288	0.001	0.001	0.001	0.001	0.001	0.004	0.003	0.004
202	288	0.001	0.001	0.001	0.001	0.001	0.006	0.003	0.004

These should be 1 mm (for older orbits it may also be 3...5 mm) if precise orbits from CODE were used together with the consistent EOP information (the actual RMS errors depend on the quality of the precise orbits, on the pole file used for the transformation between IERS Terrestrial Reference Frame (ITRF) and International Celestial Reference Frame (ICRF), and on the orbit model used in ORBGEN). Comparing the RMS error from the third and the fourth iteration you will see that three iterations are sufficient to produce precise standard orbits for GNSS satellites.

The file \${P}/INTRO/OUT/ORB\_20190440.LST summarizes the orbit fit RMS values in one table:

The output shows that 8 satellites are in eclipse (indicated above the satellite number by E for Earth or M for Moon).

A similar file may be generated using the orbit products from IGS (including the EOP information) following the procedure described in Section 7.1. As you can notice the RMS errors are slightly larger. It does not mean that the orbits from CODE are better than the IGS orbits. The orbit model in ORBGEN is just more consistent with the orbit model used at CODE:

```
EXAMPLE: Session 2019-0440: Standard orbit generation TIME FROM DAY : 3 GPS WEEK: 2040  TO \quad DAY : 4 \ GPS \ WEEK: 2040 
                                                                    14-JAN-24 05:22
ORBIT REPEATABILITY FROM A 1-DAY FIT THROUGH DAILY ORBIT SOLUTIONS (MM)
                          8 E /
# ECLIPSING SATELLITES:
                                  0 M ( 0 EM)
                                          Ε.
                                                   E. ...
ECL
                     Ε.
                                      9
                                                                           28 29
                                                                                    30 31 32
                           6
                                  8
                                          10 11 12 ... 24 25 26 27
DOY
             3
                     5
044
      2 2
              3 3
                     3 2
                               2
                                 3
                                      2
                                          3
                                               2
                                                    4 ...
                                                            2 3 3 2
                                                  4 ...
                               2 3 2 3
                                                            2 3 3 2
      2 2 3 3 3 2
```

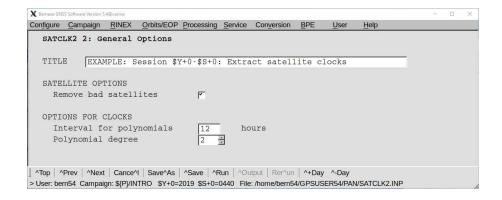
(Note the missing GLONASS and Galileo satellites in the IGS orbit product.)

In the example for day 045 of year 2019 a satellite 76 appears. The GPS—satellite 26 had a repositioning event at 14-Feb-2019 19:47:44 (see \${CONFIG}/SAT\_2019.CRX). The satellite is introduced in the processing with two independent arcs: one before (number 26) and one after (number 76) the event (you may verify this by the number of epochs available for each of these two satellite arcs).

# 3.3 Extract satellite clock corrections

As the next step we run SATCLK2 from "Menu>Orbits/EOP>Import satellite clocks>From precise orbit files" in order to generate a satellite clock file (CLK) from the precise orbit file (PRE). The clock file will be needed, e.g., in program CODSPP (see Section 4.2.1, to be discussed in the Tuesday's lectures on pre-processing).

X Bernese GN:	SS Software Version 5	.4@carina									<u> 25</u> 5		×
Configure	Campaign	RINEX	Orbits/EOP	Processing	Service	Conversion	BPE	User	Help				
EXTR.	ACT SATE	LLITE C	LOCKS -	SATCLK2	1: File	names							
	RAL FILES		files	F									
	T FILE ecise epl	nemeris	file	COD_	\$YYYSS+	0 PRE							
	LT FILE tellite (	clock fi	ile	COD_	\$YYYSS+	O CLK							
GENE	RAL OUTPU	JT FILE	S										
Pr	ogram ou	tput		₹ us	e SATCI	K2.Lnn		or		SATCLK2		OU	Т
Er	ror mess	ages		☐ me	rged to	program	output	or		ERROR		MS	G
						tput   Rer^un			SATCL	<2.INP			



Panel "SATCLK2 2: General Options" contains the options for extracting the satellite clock information. The clock values in the precise orbit file are sampled to 15 min. We interpolate with a "Polynomial degree" of 2 with an "Interval for polynomials [h]" of 12 hours. This is good enough for the receiver clock synchronization in CODSPP.

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A message like this is expected:

```
### FN O_SATCLK2:chkSufficientNbClkValues, PG SATCLK2 (13-Jan-2024 15:46:23)
    Satellite clock correction values missing
    Minimal number : 3
    Number in sample: 0
    Satellite : R26
    File name : ${P}/INTRO/ORB/COD_20190440.PRE

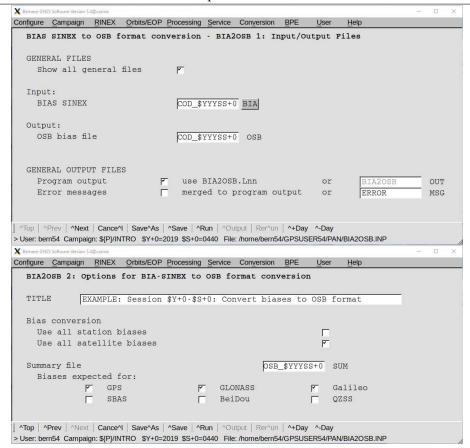
### FN O_SATCLK2:chkSufficientNbClkValues, PG SATCLK2 (13-Jan-2024 15:46:23)
    Satellite clock correction values missing
    Minimal number : 3
    Number in sample: 0
    Satellite : E01
    File name : ${P}/INTRO/ORB/COD_20190440.PRE
...
```

It indicates that the precise orbit files do not contain clock corrections for Galileo satellites. Consequently they are not available for the synchronization of the receiver clocks in CODSPP. It does not matter because it is sufficient to do this process based on the GPS satellite clocks only.

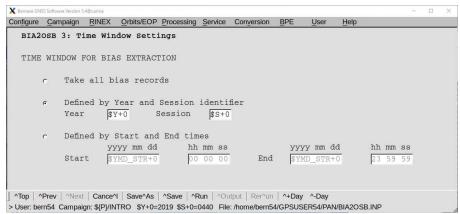
# 3.4 Convert Bias Corrections from Bias SINEX into Bernese bias Format

Ideally the clock corrections (either in a separate clock RINEX format or also in a precise orbit file) are provided together with a consistent set of code and potentially phase biases. Biases provided by a different source could potentially be less consistent.

In this context we run BIA2OSB from "Menu>Conversion>BIAS-SINEX to Bernese format". in order to extract biases from international Bias SINEX format file (BIA) into the Bernese formatted file with OSBs (OSB, both files are located in the ORB directory of the campaign).



We select the current session as the time window in order to extract the expected set of satellites if the input Bias SINEX file contains a longer interval than just the current session (like in our case) and within this interval a satellite change takes place:



If no bias information is provided we can use default code biases from the CODE processing (http://www.aiub.unibe.ch/download/BSWUSER54/ORB).

#### 3.5 Session Goals

At the end of this session, you should have created the following files:

- 1. Bernese pole file in the campaign's ORB directory: COD\_ 20190440. ERP,
- 2. Bernese standard orbit file in the ORB directory: COD\_ 20190440. STD,
- 3. Bernese satellite clock file in the ORB directory: COD\_ 20190440. CLK.
- 4. Biases in Bernese format file in the ORB directory: COD\_ 20190440. OSB.

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# 4 Terminal Session: Tuesday

Today's terminal session is to:

- 1. import the observations from the RINEX format into the Bernese format using RXOBV3 (section 4.1).
- 2. preprocess the Bernese observation files:
  - receiver clock synchronization (CODSPP, section 4.2.1)
  - baseline generation (SNGDIF, section 4.2.2)
  - preprocess baselines (MAUPRP, section 4.2.3)

#### 4.1 Importing the Observations

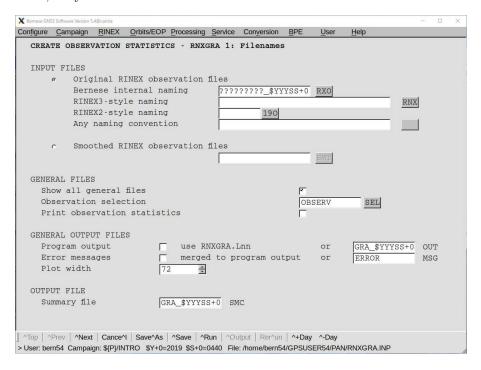
The campaign has been set up and all necessary files are available. The first part of processing consists of the transfer of the observations from RINEX to Bernese (binary) format.

#### 4.1.1 Inventory of the input RINEX files

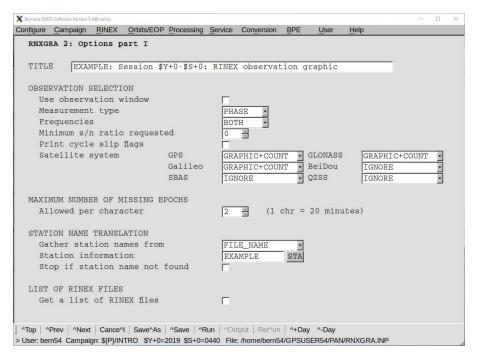
Before starting the data analysis, it is recommended to get an overview of the data availability, completeness, and a first impression on the data quality. For this purpose you may generate a pseudographic from the RINEX observation files using the program RNXGRA in "Menu>RINEX>RINEX utilities>Create observation statistics".

Note that this step is not mandatory and can also be skipped during the tutorial exercise.

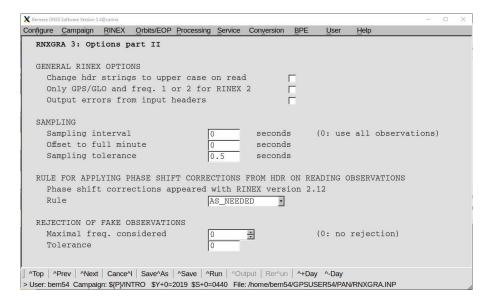
You have to select the RINEX observation files for option "Original RINEX observation files". All three naming schemes of the files are supported by the user interface: "Bernese internal naming", "RINEX3-style naming", and "RINEX2-style naming". If there are files specified in more than one field, they are cumulated.



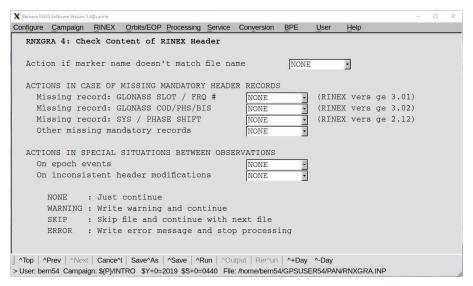
As only GPS, GLONASS, and Galileo satellite orbits are processed in this example, only those systems are selected



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All validations are disabled in this step because they can also be carried out when importing the data by the program RXOBV3 described in Section 4.1.2.



From the processing you get a summary in \${P}/INTRO/OUT/GRA\_20190440. SMC:

```
DATE : 2019-02-13
PHASE OBSERVATIONS
BOTH FREQUENCIES
    |*************
                                   11237631 -
                                                               -1*
      -17****************
                                           -3221 -
              -4555555557777777777
  4 1
                                                      111 -
                                               -125*******82-
  5 I
                 -3************
              -19**********
                                       -138********
  6 |
    -17*****
             -1***************
                                                     -2221-
                                   -5******41-
 10 | ******731 -
                                                        -3******
231 | -366776621 -
                             --368999****9999******99871-
      -13577888999999998874--
                                    -118999999999998651-
233 I
                         -3676632-
                                                   -3789999******
                                 12
                                                                24
NUMBER OF SATELLITES INCLUDED IN DATA FILES:
```

The summary reports the number of stations tracking each satellite during a certain period of the day. If more than 9 stations are in view a \*-character is written

The second part shows how many satellites observed by each station during the course of the day (again the \*-character is indicating a number > 9). This part is shown separately for each selected satellite system.

```
DATE : 2019-02-13
PHASE OBSERVATIONS
BOTH FREQUENCIES
SATELLITE SYSTEM: GPS
   |9*************9*999**9998998878*9**9*8899**9*9997888778888999999
GANP
HERT
   JOZ2
   LAMA
   1*9******9****9998888*998887679877887*9*997888999998867886776877899**89
MATE | ***************999898899888889989888889989*9888*978777889889899*9
   MIKL
   ******************************
ONSA
ORID |*****************998889887877888**99899***99899*97887888*888899**
PTBB
  67*****9*****88898789*99987868986787799987878999998877777768757899*98
   TLSE
  |9******99***9***9887888878656878989987898899866668768799999
VILL
            ---9----6***999*9****
          ****9***989978**88898***9***999998*99*7789898*9*
          *********9**89**98898***9***99998*9997779898*9***
7.TM2
  1*************************
```

From this output it can be concluded that station WSRT seems to have some tracking problems in the first half of the day. It is therefore recommended to have a particular look at the results from this station in the different processing steps. In addition, the stations JOZ2 and WTZR show short outages of one hour. This typically caused by data management issues at the station (related to the extraction of the hourly RINEX files).

Similar tables for the other systems are available as well.

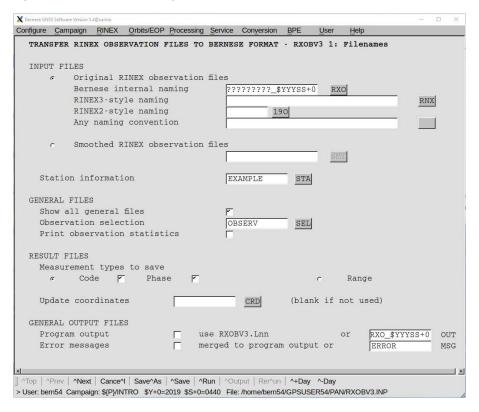
The program output file \${P}/INTRO/OUT/GRA\_20190440.OUT provides even more detailed information on each individual RINEX file.

#### 4.1.2 Converting the Observations from RINEX into Bernese format

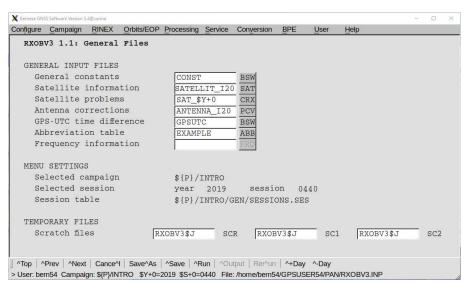
Importing the RINEX observation files is the task of the program RXOBV3 in "Menu>RINEX >Import RINEX to Bernese format>Observation files" (we do not use the RINEX navigation files in this processing example).

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The principle of specifying the input RINEX files is the same as for all RINEX related programs (description see page 36):

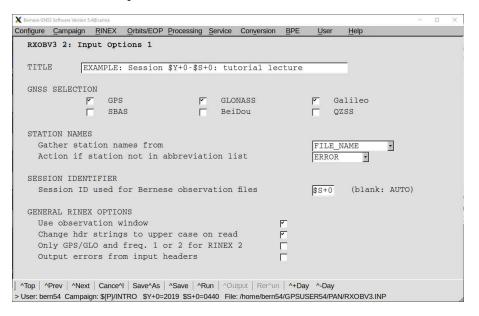


The next panel specifies the general input files. The "Satellite information" and "Antenna phase center corrections" files should consistently refer to the used reference frame even if the corrections itself are not yet needed for the import step.

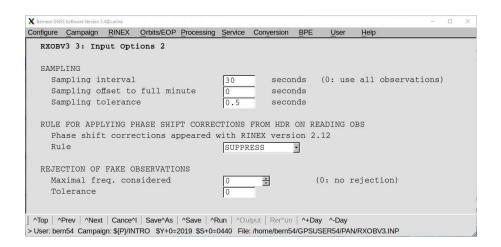


Other panels allow to select the data to be imported and to specify a few parameters for the Bernese observation header files.

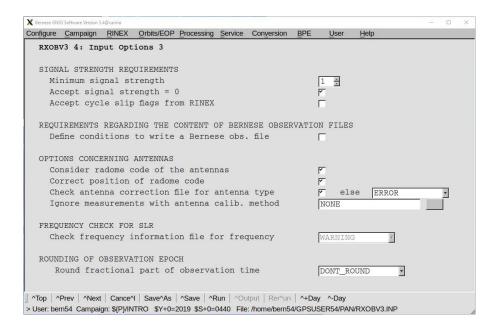
Because the orbits contain only the GPS-, GLONASS, and Galileo-satellites, activate the related checkboxes in the option "GNSS SELECTION".



X Bernese GNS	S Software Version :	5,4@carina									<u> </u>	×
Configure	Campaign	RINEX	Orbits/EOP	Processing	Service	Conversion	BPE	User	Help			
RXOB	v3 2.1:	Observa	tion Wind	low								
OBSE	RVATION	WINDOW										
	e Defir Year	nacco ne <del>s</del> , c	Year and Y+0 S	Session Session	identi: \$S+0							
	c Defir	essent and	Start and yyy mm dd		es mm ss		ΥΥ	yy mm do	ì	hh mm	SS	
	Star		MD_STR+0	,	00 00	End		MD_STR+(		23 59	59	
2						tput   Rer^un						
> User: ber	n54 Campa	ign: \${P}/IN	TRO \$Y+0=2	2019 \$S+0=	0440 File	/home/bern54	/GPSUS	ER54/PAN	RXOBV	3.INP		,

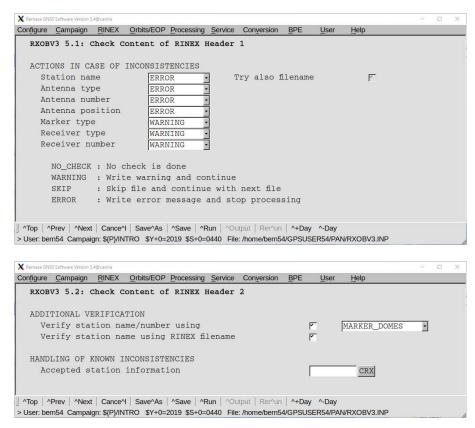


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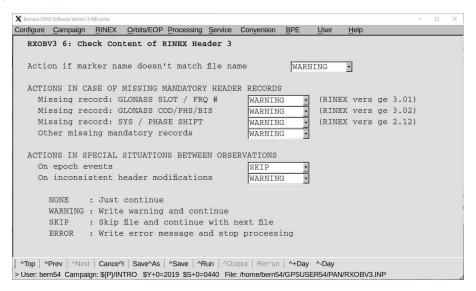


When switching for instance the option "Ignore measurements with antenna calib. method" to ADOPTED from GPS the observations from those GNSS are not imported where no antenna calibration is available for (see Table 1.2). It may also make sense to completely exclude stations where no calibration for the antenna with the specific radome is available (ADOPTED from NONE). In the case of the example dataset this would affect the station ONSA.

The next two panels are only displayed, if you specified a station information file in "RXOBV3 1: Filenames". They allow you to configure the RINEX header information verification:



The last panel allows for checking the compliance of the input RINEX file with the format definition.



Start the program with the 'Run-button.

Various warning messages appear. In an automated processing they are sorted according to a priority and become better readable. In this case they are ordered just by the processing order. Most of them are not critical and do just inform about assumptions made by the program RXOBV3 based on the RINEX format description:

```
### SR SAVMEA: Receiver clock offset in RINEX file found.

It is assumed, that the observations in the
RINEX file are not corrected by applying the
receiver clock offset. Please check the
keyword "RCV CLOCK OFFS APPL" in the RINEX file
header!
RINEX FILENAME: ...RO/RAW/BRST00FRA_20190440.RX0
```

or actions taken on the observations that do not influence the processing results:

```
### SR SAVMEA: JUMP INTRODUCED INTO PHASES

DUE TO FORMAT OVERFLOW

RINEX FILE : ...RO/RAW/HERTOOXYZ_20190440.RXO

FREQUENCY : L1

SATELLITE :122

JUMP (CYCLES): 579337875.0

EPOCH :2019-02-13 18:10:00
```

In some cases where there are several (inconsistent) header sequences in the RINEX files containing different information (e.g., related to the APPROX POSITION XYZ record):

```
### SR O_RNXBASE:wrtHdrEntryChkErrsToLfnErr, PG RXOBV3
    Inconsistency between the following two RINEX headers detected:
    Filename (header to update):
        ${P}/INTRO/RAW/JOZ200XYZ_20190440.RXO
    Filename (new header):
        ${P}/INTRO/RAW/JOZ200XYZ_20190440.RXO
    Geocentric approx. ant. marker pos. (APPROX POSITION XYZ)
        in header to update: '(3.6648784825E+06,1.4091900878E+06,5.0096170459E+06)'
        in new header: '(3.6648768963E+06,1.4091892139E+06,5.0096167523E+06)'
```

The amount of these massages may be managed by the setting in panel "RXOBV3 6: Check Content of RINEX Header 3".

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More relevant are messages indicating observations were removed due to an entry in the "Satellite problems" file. In the example, station ONSA provides data for the two GLONASS satellites R12 and R06 (satellite system R for GLONASS in RINEX whereas satellite numbers between 100 and 199 are in use within the *Bernese GNSS Software*) which are removed because of an entry in the "Satellite problems" file specified. in panel "RXOBV3 1.1: General Files".

```
### SR SAVMEA: PROBLEM FOR SATELLITE:
               INDICATED IN SATCRUX : ${CONFIG}/SAT_2019.CRX
               PROBLEM.
                                    : BAD PHASE+CODE
               REQUESTED ACTION
                                   : OBS. REMOVED
                                   : 2016-11-22 00:00:00 2019-06-19 23:59:59
               TIME WINDOW
                                   : ...RO/RAW/ONSAOOXYZ_20190440.RXO
               IN RINEX FILE
### SR SAVMEA: PROBLEM FOR SATELLITE:
                                      106
               INDICATED IN SATCRUX : ${CONFIG}/SAT_2019.CRX
               PROBLEM
                                    : BAD PHASE+CODE
               REQUESTED ACTION
                                    : OBS. REMOVED
                                    : 2018-04-08 00:00:00 2099-12-31 23:59:59
               TIME WINDOW
               IN RINEX FILE
                                     ...RO/RAW/ONSAOOXYZ_20190440.RXO
```

The program produces an output file RXO\_20190440.0UT in the directory \${P}/INTRO/OUT (resp. corresponding filenames for the other sessions). This file may be browsed using the ^Output button or with "Menu>Service>Browse program output". After echoing the input options, the file provides an overview of the observation intervals), on the station information records in the RINEX observation file header and on the values that are used for the processing in the Bernese GNSS Software. In addition some observation statistics are available. In the last section you may check the completeness of the Bernese observation files by the available number of epochs:

TARI.	E OF INPUT AND OUTPUT FILE	NAMES:						
				#observ	7.	#sat	ell.	
Num	Rinex file name	Bernese code header file	#epo	#C1/P1				GAL
		Bernese code observ. file						
		Bernese phase header file	#epo	#L1	#L2	GPS	GLO	GAL
		Bernese phase observ. file	•					
1	BRST00FRA_20190440.RX0	\${P}/INTRO/OBS/BRSTO440.CZH	2880	75780	74934	31	22	24
		\${P}/INTRO/OBS/BRST0440.CZO \${P}/INTRO/OBS/BRST0440.PZH	0000	75654	74004	0.4	22	0.4
		\${P}/INTRO/OBS/BRST0440.PZ0	2880	75654	74864	31	22	24
		\$\frac{1}{1}\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\\						
2	GANPOOSVK_20190440.RXO	\${P}/INTRO/OBS/GANPO440.CZH	2880	77299	76553	32	22	24
_		\${P}/INTRO/OBS/GANPO440.CZO	2000			02		
		\${P}/INTRO/OBS/GANPO440.PZH	2880	77040	76362	32	22	24
		\${P}/INTRO/OBS/GANPO440.PZO						
3	HERT00XYZ_20190440.RX0		2880	52454	51981	31	22	0
		\${P}/INTRO/OBS/HERT0440.CZO						
		\${P}/INTRO/OBS/HERTO440.PZH	2880	52454	51970	31	22	0
		\${P}/INTRO/OBS/HERTO440.PZO						
16	ZIM200CHE_20190440.RX0	\${P}/INTRO/OBS/ZIM20440.CZH	2880	79238	78472	32	22	24
		\${P}/INTRO/OBS/ZIM20440.CZO	2000	.0200		02		
		\${P}/INTRO/OBS/ZIM20440.PZH	2880	78966	78274	32	22	24
		\${P}/INTRO/OBS/ZIM20440.PZO						
17	ZIMMOOCHE_20190440.RXO	\${P}/INTRO/OBS/ZIMMO440.CZH	2880	28964	28819	32	0	0
		\${P}/INTRO/OBS/ZIMMO440.CZO						
		\${P}/INTRO/OBS/ZIMMO440.PZH	2880	28844	28729	32	0	0
		\${P}/INTRO/OBS/ZIMMO440.PZO						

If epochs or satellites are missing for some RINEX files you may check this with the RINEX observation graphic from program RNXGRA ("Menu>RINEX>RINEX utilities>Create observation statistics"). In February 2019 32 GPS, 22 GLONASS, and 24 Galileo satellites were active.

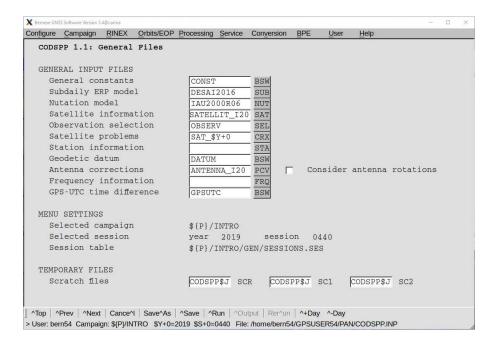
At the end of the RXOBV3 output there is a table of RINEX files rejected during the import for some reasons.

## 4.2 Data Preprocessing (I)

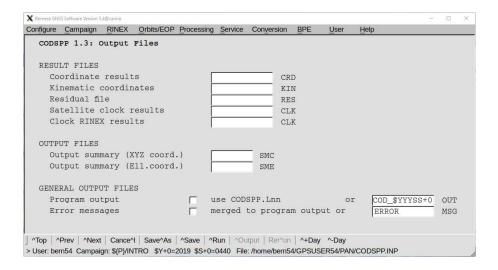
#### 4.2.1 Receiver Clock Synchronization

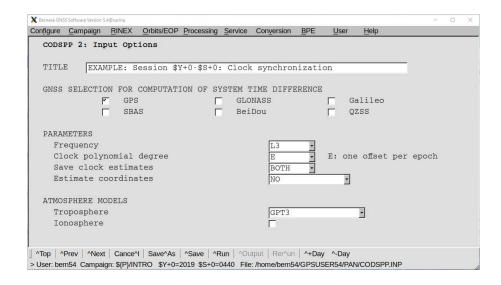
Now we are ready to enter the processing part of the *Bernese GNSS Software*. We have to run three programs for this example. The first program is called CODSPP ("Menu>Processing >Code-based clock synchronization"). Its main task is to compute the receiver clock corrections with respect to the GPS system time.

X Bernese	GNSS Software Version 5	.4@carina										×
Configu	e <u>C</u> ampaign	RINEX	Orbits/EOP	Processing	Service	Conversion	BPE	User	<u>H</u> elp			
	DE-BASED CI		NCHRONIZ#	ATION - (	CODSPP	1: Filen	ames					
	Show all ge	eneral	files	F								
INI	UT FILES											
	Code obser	vation	files		????	\$S+0 (	CZH					
	Standard of	rbits			COD_\$	YYYSS+0	STD					
	Earth rota	-				YYYSS+0						
	Satellite (			ns	-	YYYSS+0	CONTRACTOR OF THE PARTY OF THE					
	Code bias	input i	ile		COD_\$	YYYSS+0	DSB					
	Station co	ordinat	es		APR_\$	YYYSS+0	CRD					
	Kinematic (					I	KIN					
***	Observation	Contract Section					sos					
1	Station eco	centrio	cities		1	<u>I</u>	ECC					
1	Estimated	troposp	here				TRP					
	/MF grid fi	le					GRD					
1	LEO files			П								
^Top	^Prev   ^Next	Cance/	Save^As	^Save   ^F	Run   ^Oı	utput   Rer^ur		^-Day				
> User:	pern54 Campai	gn: \${P}/IN	TRO \$Y+0=	2019 \$S+0=	0440 File	e: /home/bern5	4/GPSUS	SER54/PAN	VCODSPP.	INP		/

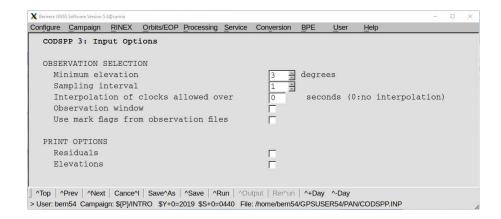


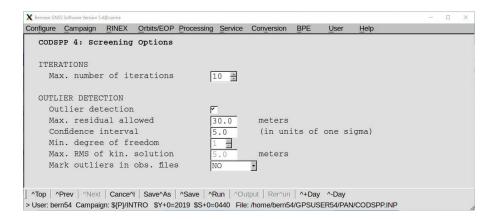
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We already have geocentric coordinates of good quality available for the sites from the PPP example BPE. Therefore, the option "Estimate coordinates" may be set to NO. The most important option in this CODSPP run is "Save clock estimates". It has to be set to BOTH in order to write the estimated receiver clock corrections into both, the code observation files and the phase observation files.





#### CODSPP produces the following output:

AY OF YEAR	:	44											
				FROM	1				TO				
BSERVATIONS	:		2019-02	2-13 (	00:00:	0.00	2	2019-02	2-13	23:59:3	0.00		
EQUESTED WINDOW	:												
EASUREMENT INTERV			EC										
AMPLING RATE													
ROCESSED FREQUENC													
LEVATION LIMIT	:	3 D	EG										
				SPHERE	3			OSPHER	RE				
TMOSPHERE MODELS	:		GPT3				NON	ΙE					
TATISTICS FOR SAT				PS									
ATELLITE NUMBER BSERVATIONS IN FI	:	1	2	3	5	6	7	8	9	10	11		TOTA
SED OBSERVATIONS IN FI													
AD OBSERVATIONS (	%):	2.3	2.0	1.0	2.0	0.2	1.5	3.6	0.4	1.8	1.8		1
MS ERROR (M)													
ESULTS:													
BSERVATIONS IN FI	LE:	29	017										
SED OBSERVATIONS AD OBSERVATIONS	:	28	520										
AD OBSERVATIONS	:		1.71	%									
MS OF UNIT WEIGHT UMBER OF ITERATIO	:		1.17	M									
UMBER OF ITERATIO	NS:		2										
TATION COORDINATE													
OCAL GEODETIC DAT	UM:	IGS20		Δ PRT(	ד א ר		NEW		NEW	_ A PRT	OR T	RMS	FRROR
BRST 10004M004	Х			1231162	2.42	42	31162.4	12		- A PRI 0.00		0.	00
(MARKER)				-332746	6.43	-3	32746.4	13		0.00		0 .	
	Z						45131.0			0.00		0 .	.00
	HEIG	нт		65	5.82		65.8	32		0.00		0 .	00
	LATI LONG	TUDE ITUDE	- 48	3 22 49 1 29 47	9.784 7.729	48 :	22 49.7 29 47.7	784 729	0 0	0.000		0.	0000
LOCK PARAMETERS:													
FFSET FOR REFEREN	CE EP	OCH:			-0.00	000000	19 SEC	;					
LOCK OFFSETS STOR													
TOOK OTTOBIO DION	עד עם	CODE	· I HADE	ואיםטעט	VIII OIA								

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```
SUMMARY OF BAD OBSERVATIONS
 MAXIMUM RESIDUAL DIFFERENCE ALLOWED :
                                            30.00 M
 CONFIDENCE INTERVAL OF F*SIGMA WITH F:
                                             5.00
NUMBER OF BAD OBSERVATION PIECES
                                               28
                             TYP SAT
 NUMB FIL STATION
                                             FROM
                                                                    TO
                                                                                 #EPO
       2 GANP 11515M001
                             OUT
                                  4 2019 02 13 04 54 00 2019 02 13 04 54 30
                             OUT
       8 ONSA 10402M004
                                 30
                                      2019 02 13 00 16 30
                                                            2019 02 13 00 16 30
          ONSA 10402M004
                             OUT
                                      2019 02 13 00 35 30
                                                            2019 02 13 00 35 30
       8
                                  15
       8 ONSA 10402M004
                             OUT 17
                                      2019 02 13 01 49 00
                                                            2019 02 13 01 49 00
          ONSA 10402M004
                             OUT
                                      2019 02 13 05
                                                            2019 02 13 05 17
                                 25
   6
       8 ONSA 10402M004
                             OUT
                                      2019 02 13 06 02 00
                                                            2019 02 13 06 02 00
                             OUT
   18
      13 WSRT 13506M005
                                  23
                                      2019 02 13 05 59 30
                                                            2019 02 13 05 59 30
                                      2019 02 13 06 13 30
   19
          WSRT 13506M005
                             OUT
                                                            2019 02 13 06 14
      13
   20
           WSRT 13506M005
                             OUT
                                      2019 02 13 07 31 00
                                                            2019 02 13 07 32 00
      13
                                           02 13 07 35 30
      13
          WSRT 13506M005
                             OUT
                                      2019
                                                            2019 02 13 07 35
   22
          WSRT 13506M005
                             OUT
                                      2019 02 13 08 16 00
                                                            2019 02 13 08 16
       13
   23
      13
          WSRT 13506M005
                             OUT
                                 17
                                      2019 02 13 14 14 00
                                                            2019 02 13 14 14
   24
                                      2019 02 13 14 28 30
      13
         WSRT 13506M005
                             OUT
                                  12
                                                            2019 02 13 14 32 00
                                                                                    8
                             OUT
   25
      13
          WSRT 13506M005
                                 32
                                      2019 02 13 14 28 30
                                                            2019 02 13 14 32 00
                                                                                    8
          WSRT 13506M005
                             OUT
                                 24
                                      2019 02 13 14 29 00
                                                            2019 02 13 14 32 00
      13
           WSRT 13506M005
                             OUT
                                      2019 02 13 14 29 00
                                                            2019 02 13 14 32 00
   28 13 WSRT 13506M005
                             OUT 24 2019 02 13 14 34 30
                                                           2019 02 13 14 35 30
>>> CPU/Real time for pgm "CODSPP": 0:01:13.209 / 0:01:13.232
>>> Program finished successfully
```

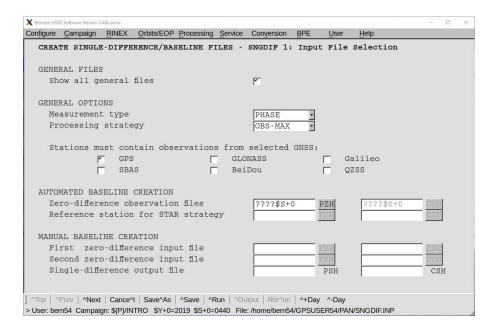
The most important message in the output file is CLOCK OFFSETS STORED IN CODE+PHASE OBSERVATION FILES. This indicates that the receiver clock corrections computed by CODSPP are stored in code and phase observation files.

The a posteriori RMS error (for each zero difference file processed) should be checked in the CODSPP output file. A value of about 20–30 m is normal if Selective Availability (SA) — artificial degradation of the satellite clock accuracy — is on (before May 2000). Without SA a value of about 3 m is expected if P–code measurements are available (this is the case for the time interval of the processing example). However, even worse code measurements would still be sufficiently accurate to compute the receiver clock corrections with the necessary accuracy of  $1\,\mu s$ .

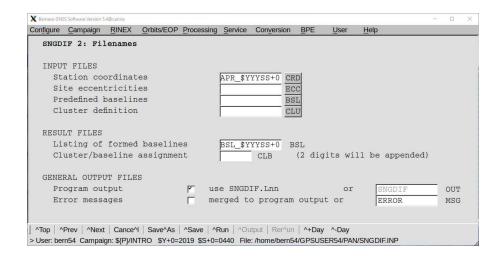
You may use the extraction program CODXTR ("Menu>Processing>Program output extraction>Code-based clock synchronization") to generate a short summary from the CODSPP program output:

#### 4.2.2 Form Baselines

The second preprocessing program is called SNGDIF and may be activated in "Menu >Processing>Create baseline files". SNGDIF creates the single differences and stores them into single-difference observation files. We use the strategy OBS-MAX for PHASE observation files

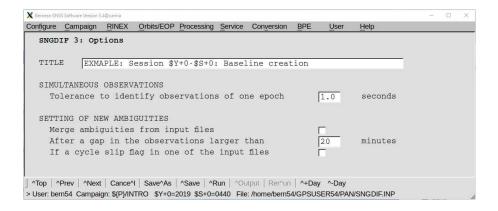


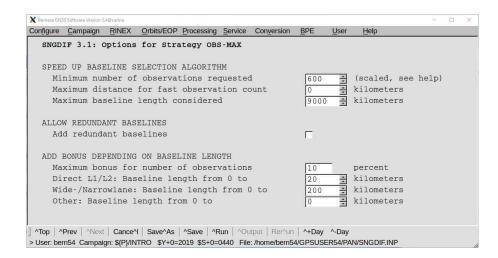
By activating the option "Stations must contain observ. from GPS" you demand that stations included in the baseline creation must contain at least GPS measurements. This option can be used to exclude those datasets containing no GPS observations.



The result file specified in option "Listing of formed baselines" stores the selected baseline configuration for further use.

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The output of SNGDIF echoes the zero difference files used and the single difference files created. The first table confirms that all stations provide at least data from GPS. For that reason all stations are included in the baseline creation.

NUM	HEADER FILE NAMES	STATION NAME	#SAT	SYS	REMARK
1	\${P}/INTRO/OBS/BRST0440.PZH	BRST 10004M004	77	GRE	included
2	\${P}/INTRO/OBS/GANPO440.PZH	GANP 11515M001	78	GRE	included
3	\${P}/INTRO/OBS/HERTO440.PZH	HERT 13212M010	53	GR	included
4	\${P}/INTRO/OBS/JOZ20440.PZH	JOZ2 12204M002	53	GR	included
5	\${P}/INTRO/OBS/LAMAO440.PZH	LAMA 12209M001	53	GR	included
6	\${P}/INTRO/OBS/MATE0440.PZH	MATE 12734M008	75	GRE	included
7	\${P}/INTRO/OBS/MIKLO440.PZH	MIKL 12335M001	75	GRE	included
8	\${P}/INTRO/OBS/ONSA0440.PZH	ONSA 10402M004	54	GR	included
9	\${P}/INTRO/OBS/ORIDO440.PZH	ORID 15601M001	53	GR	included
10	\${P}/INTRO/OBS/PTBB0440.PZH	PTBB 14234M001	32	G	included
11	\${P}/INTRO/OBS/TLSE0440.PZH	TLSE 10003M009	77	GRE	included
12	\${P}/INTRO/OBS/VILLO440.PZH	VILL 13406M001	79	GRE	included
13	\${P}/INTRO/OBS/WSRT0440.PZH	WSRT 13506M005	74	GRE	included
14	\${P}/INTRO/OBS/WTZRO440.PZH	WTZR 14201M010	77	GRE	included
15	\${P}/INTRO/OBS/WTZZO440.PZH	WTZZ 14201M014	79	GRE	included
16	\${P}/INTRO/OBS/ZIM20440.PZH	ZIM2 14001M008	78	GRE	included
17	\${P}/INTRO/OBS/ZIMMO440.PZH	ZIMM 14001M004	32	G	included

The creation of the following 16 baseline files from 17 zero difference observation files is reported:

```
SNGDIF: INPUT AND OUTPUT OBSERVATION FILE NAMES
O-DIF. HEADER FILE NAMES (INPUT) O-DIF. OBS. FILE NAMES (INPUT)
${P}/INTRO/OBS/BRST0440.PZH
                                   ${P}/INTRO/OBS/BRST0440.PZO
                                   ${P}/INTRO/OBS/TLSE0440.PZO
${P}/INTRO/OBS/TLSE0440.PZH
${P}/INTRO/OBS/GANPO440.PZH
                                   ${P}/INTRO/OBS/GANPO440.PZO
                                                                         3
${P}/INTRO/OBS/MIKLO440.PZH
                                   ${P}/INTRO/OBS/MIKL0440.PZO
                                                                         4
                                   ${P}/INTRO/OBS/GANPO440.PZO
${P}/INTRO/OBS/GANPO440.PZH
                                                                         5
${P}/INTRO/OBS/WTZZO440.PZH
                                   ${P}/INTRO/OBS/WTZZO440.PZO
                                                                         6
1-DIF. HEADER FILE NAMES (OUT)
                                   1-DIF. OBS. FILE NAMES (OUT)
                                                                       NR1 NR2 STAT.
${P}/INTRO/OBS/BRTL0440.PSH
                                   ${P}/INTRO/OBS/BRTL0440.PSO
                                                                                 OΚ
${P}/INTRO/OBS/GAMIO440.PSH
                                   ${P}/INTRO/OBS/GAMI0440.PSO
                                                                                 OK
${P}/INTRO/OBS/GAWZ0440.PSH
                                   ${P}/INTRO/OBS/GAWZ0440.PSO
                                                                                 OK
${P}/INTRO/OBS/HETLO440.PSH
                                   ${P}/INTRO/OBS/HETL0440.PSO
                                                                                  OK
${P}/INTRO/OBS/JOON0440.PSH
                                   ${P}/INTRO/OBS/JOON0440.PSO
                                                                         9
                                                                            10
                                                                                 OK
${P}/INTRO/OBS/LAONO440.PSH
                                   ${P}/INTRO/OBS/LAONO440.PSO
                                                                        11
                                                                            12
                                                                                 OK
${P}/INTRO/OBS/MAORO440.PSH
                                   ${P}/INTRO/OBS/MAORO440.PSO
                                                                        13
                                                                            14
                                                                                 OK
${P}/INTRO/OBS/MAZIO440.PSH
                                   ${P}/INTRO/OBS/MAZIO440.PSO
                                                                        15
                                                                            16
                                                                                 OK
${P}/INTRO/OBS/ONPTO440.PSH
                                   ${P}/INTRO/OBS/ONPTO440.PSO
                                                                        17
                                                                            18
                                                                                  OK
${P}/INTRO/OBS/ONWZO440.PSH
                                   ${P}/INTRO/OBS/ONWZ0440.PSO
                                                                        19
                                                                            20
${P}/INTRO/OBS/TLZI0440.PSH
                                   ${P}/INTRO/OBS/TLZI0440.PSO
                                                                        21
                                                                            22
                                                                                  OK
${P}/INTRO/OBS/VIWZO440.PSH
${P}/INTRO/OBS/WSWZO440.PSH
                                   ${P}/INTRO/OBS/VIWZ0440.PS0
${P}/INTRO/OBS/WSWZ0440.PS0
                                                                        23
25
                                                                            24
                                                                                  ΠK
                                                                            26 WARNG
${P}/INTRO/OBS/WTWZ0440.PSH
                                   ${P}/INTRO/OBS/WTWZ0440.PSO
                                                                        27
                                                                            28
                                                                                 OK
${P}/INTRO/OBS/WZZIO440.PSH
                                   ${P}/INTRO/OBS/WZZI0440.PSO
                                                                        29
                                                                            30
                                                                                  OK
${P}/INTRO/OBS/ZIZMO440.PSH
                                   ${P}/INTRO/OBS/ZIZMO440.PSO
```

For the baseline WSWZ (WSRT-WTZZ) it is indicated that warning messages related to this baseline are issued that can be displayed via "Menu>Service>Browse error message":

```
### PG SNGDIF: JUMP INTRODUCED TO AVOID FORMAT OVERFLOW
FREQUENCY: 1
SATELLITE: 1
EPOCH NUMBER: 299
JUMP (CYCLES): 872366831.
FILE: ${P}/INTRO/OBS/WSWT0440.PSH
...
```

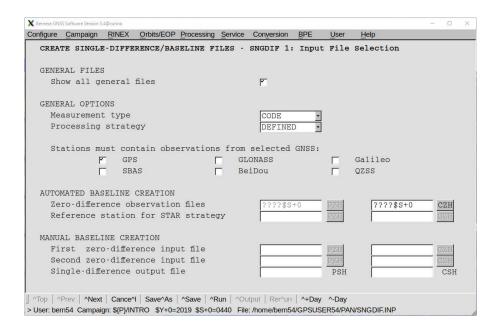
These jumps do not influence the obtained results.

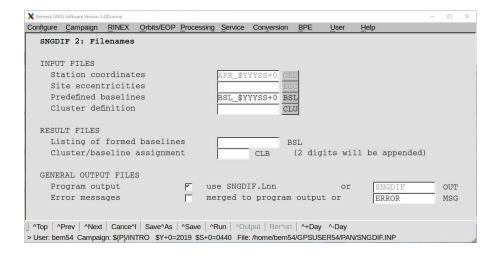
If the strategy OBS-MAX was selected all possible pairs of zero difference files are listed with the corresponding criterion value. The baselines belonging to the created network configuration are labeled with OK.

```
1 BRST 10004M004
                              11515M001
                                            #SAT:
                                                       CRIT .:
                                                                 32293
                                                       CRIT.:
  2 BRST 10004M004
                         HERT 13212M010
                                           #SAT:
                                                   53
                                                                 24582
                                                       CRIT.:
  3 BRST 10004M004
                        JUZ2 12204M002
                                           #SAT:
                                                   53
                                                                 22972
  4 BRST 10004M004
                       - LAMA 12209M001
                                           #SAT:
                                                   53
                                                       CRIT.:
                                                                 22225
                       - PTBB 14234M001
  9 BRST 10004M004
                                            #SAT:
                                                   31
                                                       CRIT.:
                                                                 11945
 10 BRST 10004M004
                      - TLSE 10003M009
                                            #SAT:
                                                   77
                                                       CRIT.:
                                                                 35538
                                                       CRIT.:
 11 BRST 10004M004
                       - VILL 13406M001
                                           #SAT:
                                                   77
                                                                 31995
 12 BRST 10004M004
                       - WSRT 13506M005
                                           #SAT:
                                                   73
                                                       CRIT.:
                                                                 24936
123 VILL 13406M001
                       - WTZR 14201M010
                                                       CRIT.:
                                                                 30079
                                           #SAT:
                                                   77
                      - WTZZ 14201M014
124 VILL 13406M001
                                                   79
                                                       CRIT .:
                                           #SAT:
                                                                 32284
                                                                         OK
125 VILL 13406M001
                                            #SAT:
                        ZIM2 14001M008
                                                   78
                                                       CRIT.:
                                                                 32108
126 VILL 13406M001
                         ZIMM 14001M004
                                           #SAT:
                                                   32
                                                       CRIT.:
                                                                 12797
                                                       CRIT.:
127 WSRT 13506M005
                        WTZR 14201M010
                                           #SAT:
                                                   73
                                                                 25785
128 WSRT 13506M005
                       - WTZZ 14201M014
                                           #SAT:
                                                   74
                                                       CRIT .:
                                                                 26483
                                                                         OK
129 WSRT 13506M005
                       - ZIM2 14001M008
                                                   74
                                                       CRIT.:
                                           #SAT:
                                                                 26102
130 WSRT 13506M005
                        ZIMM 14001M004
                                           #SAT:
                                                   32
                                                       CRIT .:
                                                                 10829
131 WTZR 14201M010
                                                   77
                         WTZZ
                              14201M014
                                            #SAT:
                                                       CRIT .:
                                                                 45323
132 WTZR
         14201M010
                                            #SAT:
                         ZIM2
                              14001M008
                                                       CRIT.:
                                                                 36323
                                                       CRIT.:
133 WTZR 14201M010
                        ZIMM 14001M004
                                           #SAT:
                                                   31
                                                                 13183
                                                       CRIT.:
134 WTZZ 14201M014
                        ZIM2 14001M008
                                           #SAT:
                                                   78
                                                                 38261
                                                                         OK
135 WTZZ 14201M014
                         ZIMM
                              14001M004
                                            #SAT:
                                                   32
                                                       CRIT.:
                                                                 14212
136 ZIM2 14001M008
                         ZIMM 14001M004
                                                       CRIT.:
                                                                         OK
                                            #SAT:
                                                                 17346
```

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The baseline file stored in option "Listing of formed baselines" can be used for various purposes, e.g., to create code single difference observation files for the same baselines as the corresponding phase files. One application is the ambiguity resolution using the Melbourne-Wübbena linear combination (e.g., in the tomorrows terminal session). For this purpose the file introduced to option "Predefined baselines" when processing code measurements.

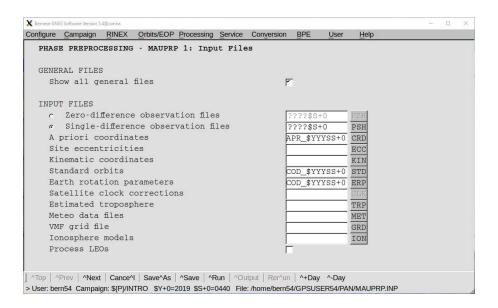




The other options do not matter because we use a predefined baseline configuration.

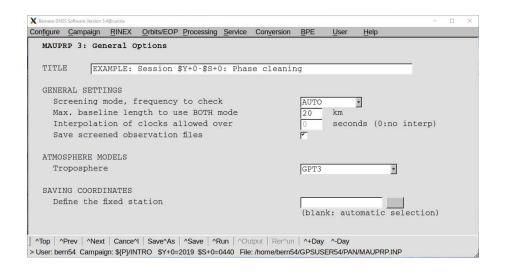
#### 4.2.3 Preprocessing of the Phase Baseline Files

The main task of the program MAUPRP is the cycle—slip detection and correction. It is started using "Menu>Processing>Phase preprocessing".

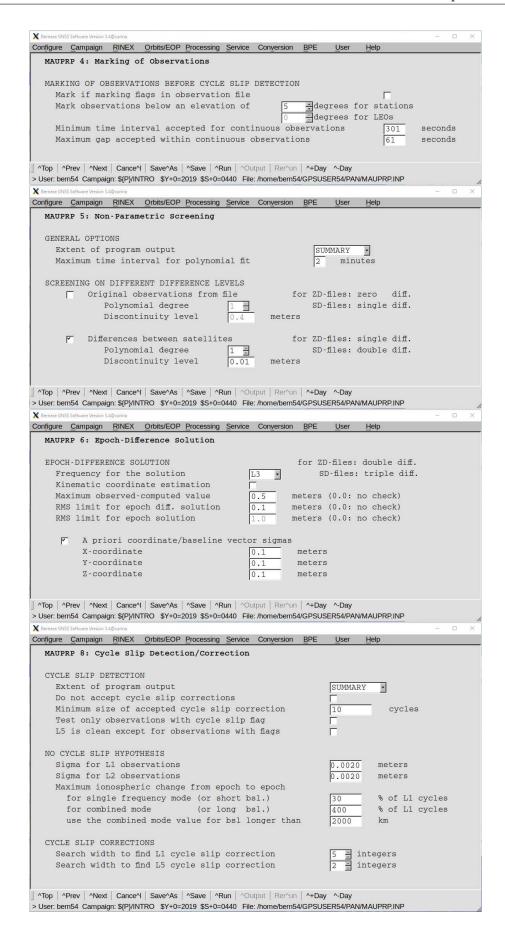


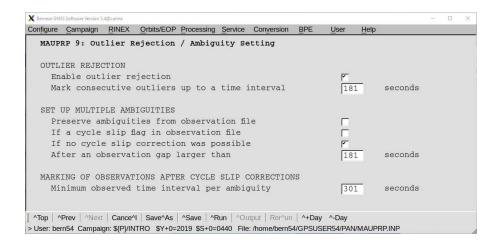
In the next input panel "MAUPRP 2: Output Files" you only need to specify the "Program output", e.g., to MPR\_\$YYYSS+0.





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The output of the program MAUPRP is discussed in detail in the lecture session. The software manual contains a detailed description, too.

You can find here the results of the adjustment of the input parameters for the maximum accepted change of the ionosphere from one epoch to the next, which is computed according to the baseline length (AUTO in option "Screening mode, frequency to check", panel "MAUPRP 3: General Options").

The most important item to check is the epoch difference solution:

The epoch difference solution is used as the reference for the data screening. For a successful phase preprocessing the RMS OF EPOCH DIFF. SOLUTION has to be below  $2\,\mathrm{cm}$ . The estimates for the coordinates in the epoch difference solution are expected to be smaller than about  $0.5\,\mathrm{m}$ .

It should be pointed out that it is not necessary to run the program MAUPRP more than once for each baseline. However, it is mandatory to run MAUPRP again if you (for whatever reason) have to re-create the baselines with program SNGDIF.

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You might get some warning messages regarding too large O-C (i.e., observed minus computed) values on certain baselines for certain epochs. The corresponding observations get flagged, and will not disturb the processing.

You can use the extraction program MPRXTR ("Menu>Processing>Program output extraction>Phase preprocessing") to generate a short summary of the MAUPRP output. The file you have specified in "MAUPRP station summary file" looks like this:

						****											
SESS	FIL	OK?	ST1	ST2	L(KM)	#OBS.	RMS	DX	DY	DZ	#SL	#DL	#MA	MAXL3	SLIP		
 0440	1	OK	BRST	TLSE	707	60983	13	 -57	20	-62	9	1522	146	24	11		
0440	2	OK	GANP	MIKL	897	59647	11	-96	-31	-171	12	927	108	47	11		
0440	3	OK	GANP	WTZZ	544	65887	10	164	7	149	4	873	112	12	31		
0440	4	OK	HERT	TLSE	816	42085	12	-192	19	-282	11	1106	117	45	11		
0440	5	OK	JOZ2	ONSA	830	43494	13	-19	25	-75	6	1049	129	34	13		
0440	6	OK	LAMA	ONSA	674	43514	12	-14	17	-83	35	809	135	42	11		
0440	7	OK	MATE	ORID	349	42135	9	2	17	22	29	735	103	44	11		
0440	8	OK	MATE	ZIM2	1014	57683	11	214	56	207	24	1390	136	44	11		
0440	9	OK	ONSA	PTBB	575	21952	13	120	35	164	63	265	38	50	13		
0440	10	OK	ONSA	WTZZ	920	44738	12	180	-3	214	3	717	83	25	11		
0440	11	OK	TLSE	ZIM2	597	62587	13	210	-28	203	6	1364	134	43	13		
0440	12	OK	VILL	WTZZ	1635	58894	10	237	4	250	5	712	73	16	17		
0440	13	OK	WSRT	WTZZ	607	40007	10	-14	24	-40	5	4134	404	49	12	<	Αl
0440	14	OK	WTZR	WTZZ	0	67195	11	-10	-17	39	4	563	116	0	0		
0440	15	OK	WTZZ			66640	12	-85	-3	-167	4	1406	105	30	11		
0440	16	OK	ZIM2	ZIMM	0	24300	13	19	10	16	7	245	49	0	0		
 Гоt:	16				665	67195	13	237	4	250	63	4134	404	50	0		

Note that in the bottom line the maximum values for each column are reported to show the "worst case".

The indicators show that there is a baseline with exceptionally many AMBiguities and marked observations. This is the baseline containing the station WSRT where data quality issues have already been detected in the RNXGRA summary output.

## 4.3 Daily Goals

At the end of today's session, you should have created the following files:

- 1. Bernese formatted zero difference observation files in your campaign's OBS directory: BRST0440. CZH, BRST0440. PZH, BRST0440. CZO, BRST0440. PZO, ... (for all stations).
- 2. Single difference files (baseline files) in the OBS directory: BRTL0440. PSH, BRTL0440. PSO, GAMI0440. PSH, GAMI0440. PSO,... for all baselines,
- 3. you should also have verified the outputs of these programs: ORBGEN, RXOBV3, CODSPP, SNGDIF, and MAUPRP

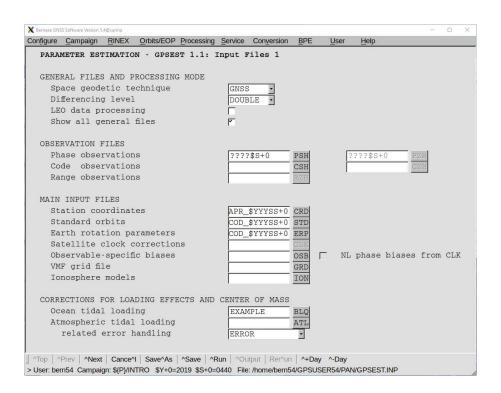
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# 5 Terminal Session: Wednesday

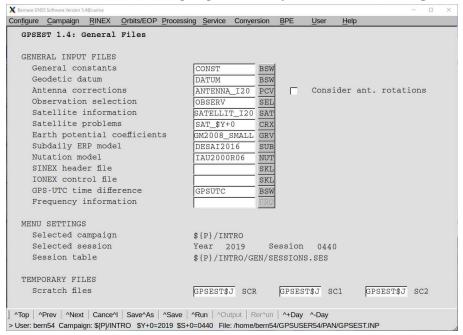
Today's terminal session is to:
1. perform a residual screening (GPSEST, RESRMS, SATMRK),
2. generate a first estimation for coordinates and troposphere parameters (GPSEST),
3. resolve the double difference ambiguities (GPSEST).

## 5.1 Data Preprocessing (II)

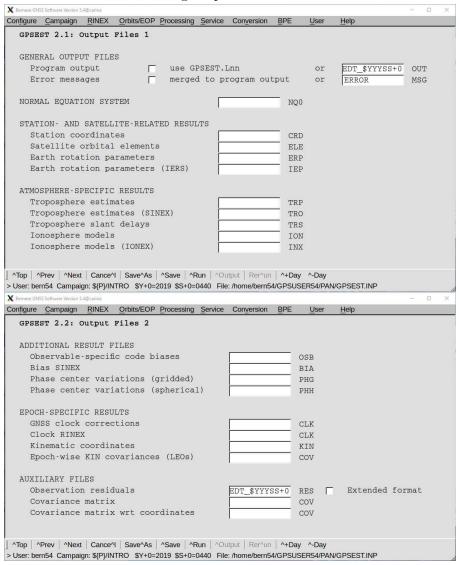
The main parameter estimation based on a least–squares adjustment is the task of program GPSEST. It is a good idea to start GPSEST first in the session mode and to produce a  $L_3$  solution (ionosphere–free linear combination) with real–valued ambiguities. We do not expect any final results from this run but we want to check the quality of data and save the residuals after the least–squares adjustment. The program is available via "Menu>Processing >Parameter estimation". We use the following options:



No files are selected in the second input panel. Verify the model and configuration files:

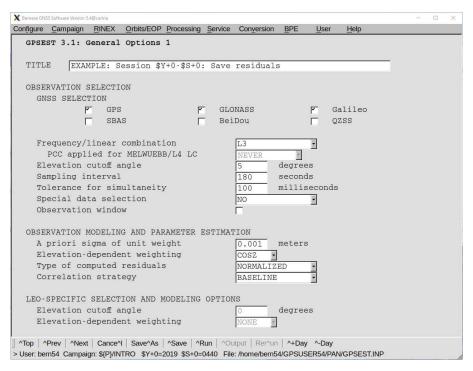


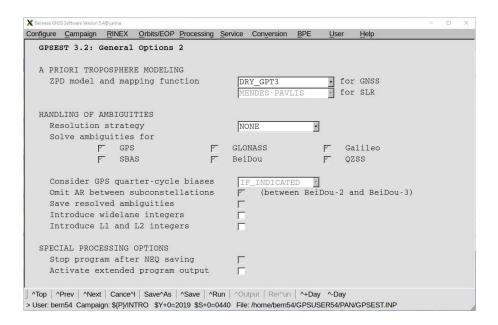
We recommend to select the following output files:



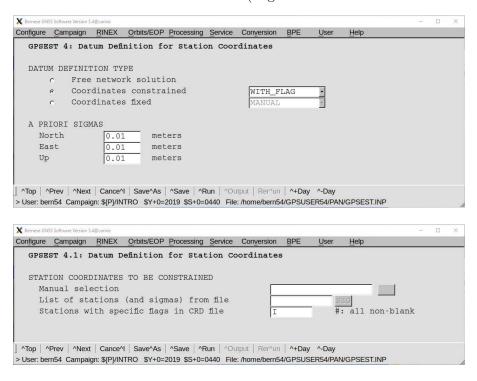
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This run is intended to screen the post–fit residuals for outliers. Later on, for the ambiguity resolution, all observations are needed without down–sampling the data. To run the program GPSEST for the network with 17 stations and the observations to more than 75 satellites without reducing the data sampling rate takes easily 10 minutes or more. For that reason we are forced to sample the data, e.g., down to 3 minutes — we will see in the next step how to solve this discrepancy.

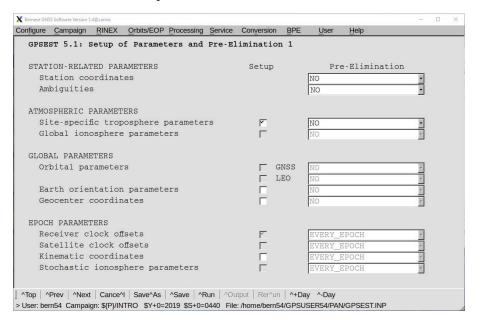




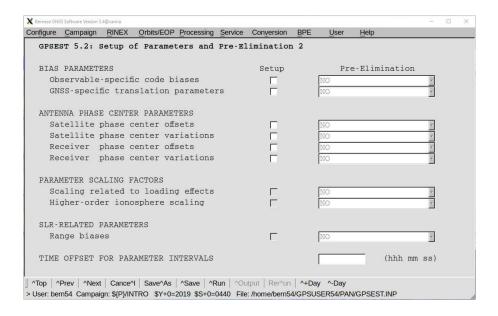
We want to put loose constraints on the station coordinates that are available from the IGS realization of ITRF 2020 reference frame (flag I like IGS20 in the coordinate file).



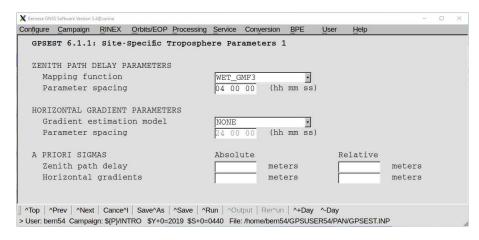
No parameters (not even ambiguity parameters) can be pre-eliminated if residuals should be written into the residual output file:



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A 4 hour resolution in time for the troposphere parameters is sufficient for this purpose:



X Bernese GNS	S Software Version 5.	4@carina									20	×
Configure	Campaign	RINEX	Orbits/EOP	Processing	Service	Conversion	BPE	User	<u>H</u> elp			
GPSE	ST 6.1.2:	Site-	Specific 7	roposph	ere Par	ameters 2	:					
STAT	IONS TO E	E EXCL	UDED FROM	I TROPOSI	PHERE E	STIMATION	Ī					
St	ation sel	Lection	1	1	IONE	•						
1000	ation lis		file			FIX						
Mai	nual sele	ection										
STAT	IONS WITH	SPECI	AL A PRIC	ORI SIGM	AS							
St	ation sel	Lection	L	Ī	IONE	•						
St	ation lis	st from	file			SIG						
Ma	nual sele	ection										
SPEC	IAL A PRI	ORI SI	GMAS	А	bsolute	2		Rela	tive			
Ze	nith path	n delay	1	Γ		meters	5			meters		
Но	rizontal	gradie	ents	Γ		meters	3			meters		
^Top   ^	Prev   ^Next	Cance <sup>2</sup>	I Save^As	^Save   ^F	Run   ^Ou	tput   Rer^un	^+Da	y ^-Day				
User: ber	n54 Campaig	gn: \${P}/IN	TRO \$Y+0=	2019 \$S+0=	0440 File	/home/bern54	1/GPSU	SER54/PAN	GPSES	T.INP		

The program output of GPSEST summarizes all important input options, input data, and reports the estimated results. An important information in the output file is the a posteriori RMS error:

```
Statistics:
Total number of authentic observations
                                                     133349
Total number of pseudo-observations
                                                          46
Total number of explicit parameters
                                                       2909
Total number of implicit parameters
                                                          0
Total number of observations
                                                     133395
Total number of adjusted parameters Degree of freedom (DOF)  \begin{tabular}{ll} \hline \end{tabular} 
                                                       2909
                                                     130486
A posteriori RMS of unit weight
                                                  0.001041 m
Chi **2/DOF
                                                       1.08
Total number of observation files
                                                         269
{\tt Total\ number\ of\ unobserved\ ambiguities}
Total number of stations
                                                         17
```

An a posteriori RMS error of about 1.0...1.5 mm is expected if elevation—dependent weighting is used. A significant higher RMS error indicates that either your data stems from low—quality receivers, that the data was collected under extremely bad conditions, or that the preprocessing step (MAUPRP and CODSPP) was not successfully performed.

Below this part the program output reports the results of the parameter estimation in a standardized format for all parameter types:

Sol	Stati	on name	Тур	Correction	Estimated value	RMS error	A priori value	 Ab
1	BRST	10004M004	X	0.01152	4231162.42917	0.00303	4231162.41765	 # C
1	BRST	10004M004	Y	-0.00335	-332746.43226	0.00282	-332746.42891	 # (
1	BRST	10004M004	Z	0.00415	4745131.07383	0.00300	4745131.06968	 #(
1	TLSE	10003M009	Х	0.00633 -0.00418 -0.00281	4627851.66867	0.00299	4627851.66234	 #(
1	TLSE	10003M009	Y	-0.00418	119640.28425	0.00276	119640.28843	 #(
1	TLSE	10003M009	Z	-0.00281	4372993.71180	0.00295	4372993.71461	 #(
1	GANP	11515M001	X	-0.00393	3929181.30207	0.00302	3929181.30600	 #(
1	GANP	11515M001		0.00394	1455236.94940	0.00278	1455236.94546	 # (
1	GANP	11515M001		0.03959 0.00239	4793654.03253	0.00296	4793653.99294	 # (
1	MIKL	12335M001	Х	0.00239	3698553.71209	0.00305	3698553.70970	 #(
		cific tropos						
	ion i		Тур		Estimated value	RMS error	A priori value	 Αì
tat	ion i				Estimated value		A priori value	 
tat	ion 1	name	Тур	Correction	2.42203			 #1
tat SRST	ion 1	name 	Typ U	Correction 	2.42203	0.00081	2.31184	 #1
tat RST RST	ion 1	name  04M004 04M004	Typ U U	Correction 0.11019 0.09745	2.42203 2.40934	0.00081 0.00061	2.31184 2.31190	 #1
Stat BRST BRST BRST	ion i	name  04 M004 04 M004	Typ U U U	Correction 0.11019 0.09745 0.07787	2.42203 2.40934 2.38976	0.00081 0.00061 0.00063	2.31184 2.31190 2.31188	 #: #:
tat RST RST RST RST	ion i	name  04 M004 04 M004 04 M004	Typ  U U U U	Correction 0.11019 0.09745 0.07787 0.08644	2.42203 2.40934 2.38976 2.39831	0.00081 0.00061 0.00063 0.00063	2.31184 2.31190 2.31188 2.31187	 #1 #1 #1 #1
Stat BRST BRST BRST BRST BRST	ion 1000 1000 1000 1000 1000 1000 1000 10	name  04 M004 04 M004 04 M004 04 M004 04 M004	Typ U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075	2.42203 2.40934 2.38976 2.39831 2.39261	0.00081 0.00061 0.00063 0.00063 0.00061	2.31184 2.31190 2.31188 2.31187 2.31186	 #: #: #: #:
Stat BRST BRST BRST BRST BRST	ion 1 1000 1000 1000 1000 1000	name 04M004 04M004 04M004 04M004 04M004 04M004	Typ  U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06701	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886	0.00081 0.00061 0.00063 0.00063 0.00061 0.00067	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185	 # # # # # # # # # # # # # # # # # # #
Stat BRST BRST BRST BRST BRST BRST	ion i 1000 1000 1000 1000 1000 1000	1ame 04M004 04M004 04M004 04M004 04M004 04M004 04M004	Typ  U U U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942	0.00081 0.00061 0.00063 0.00063 0.00061 0.00067 0.00098	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185 2.31177	  #1 #1 #1 #1 #1
RST RST RST RST RST RST RST	ion i 1000 1000 1000 1000 1000 1000 1000 1151	name	Typ  U U U U U U U U U U U U U U U U U	0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765 0.04676	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866	0.00081 0.00061 0.00063 0.00063 0.00061 0.00067 0.00098	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185 2.31177 2.12190	 #3 #3 #3 #3 #4 #3 #3 #3
Stat BRST BRST BRST BRST BRST BRST BRST BRST	ion 1 1000 1000 1000 1000 1151 1151 1151	name	Typ  U U U U U U U U U U U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06761 0.06765 0.04676 0.04605	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866 2.16796	0.00081 0.00063 0.00063 0.00063 0.00067 0.00067 0.00048	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185 2.31177 2.12190 2.12192	 #5 #5 #5 #5 #5 #5 #5 #5
tat BRST BRST BRST BRST BRST BRST BRST BRST	ion 1 1000 1000 1000 1151 1151 1151	name	Typ  U U U U U U U U U U U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765 0.04676 0.04605 0.07752	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866 2.16796 2.19943	0.00081 0.00061 0.00063 0.00063 0.00067 0.00067 0.00098 0.00067	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185 2.31177 2.12190 2.12192 2.12191	  #11 #12 #13 #13 #13 #14 #15
tat RST RST RST RST RST RST ANF ANF	ion i 1000 1000 1000 1000 1151 1151 1151 11	1ame 14M004 14M004 14M004 14M004 14M004 14M004 14M004 15M001 15M001	Typ  U U U U U U U U U U U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765 0.04676 0.04605 0.07752 0.08049	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866 2.16796 2.19943 2.20239	0.00081 0.00061 0.00063 0.00063 0.00067 0.00067 0.00098 0.00067 0.00048 0.00045	2.31184 2.31190 2.31188 2.31187 2.31186 2.31185 2.31177 2.12190 2.12192 2.12191 2.12191	  #10 #10 #10 #10 #10 #10 #10 #10
tat BRST BRST BRST BRST BRST BRST BRST BRST	ion i 1000 1000 1000 1000 1000 1151 1151 11	1ame 14M004 14M004 14M004 14M004 14M004 14M004 14M004 15M001 15M001 15M001 15M001	Typ  U U U U U U U U U U U U U U U U U U	Correction 0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765 0.04676 0.04605 0.07752 0.08049 0.06357	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866 2.16796 2.19943 2.20239 2.18548	0.00081 0.00061 0.00063 0.00063 0.00067 0.00098 0.00067 0.00048 0.00047	2.31184 2.31190 2.31188 2.31187 2.31185 2.31177 2.12190 2.12192 2.12191 2.12191 2.12191	  #10 #10 #10 #10 #10 #10 #10 #10
Stat BRSI BRSI BRSI BRSI BRSI BRSI GANF GANF GANF	ion i 1000 1000 1000 1000 1000 1000 1000 1151 1151 1151 1151	1ame 14M004 14M004 14M004 14M004 14M004 14M004 15M001 15M001 15M001 15M001 15M001 15M001	Typ  U U U U U U U U U U U U U U U U U U	0.11019 0.09745 0.07787 0.08644 0.08075 0.06701 0.06765 0.04676 0.04605 0.07752 0.08049 0.06357 0.06543	2.42203 2.40934 2.38976 2.39831 2.39261 2.37886 2.37942 2.16866 2.16796 2.19943 2.20239 2.18548 2.18733	0.00081 0.00063 0.00063 0.00067 0.00067 0.00048 0.00045 0.00045 0.00049	2.31184 2.31187 2.31186 2.31185 2.31187 2.31186 2.31187 2.12190 2.12192 2.12191 2.12191 2.12191 2.12190 2.12190	 ######################################

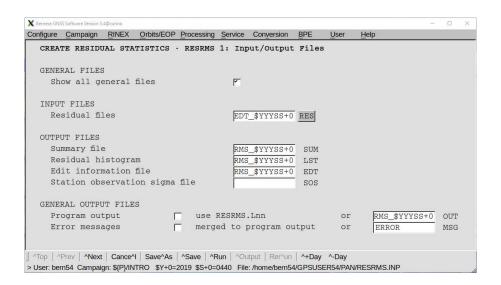
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You should check the improvement for the a priori coordinates. If all stations get approximately the same improvement in the order of decimeters, very likely the datum definition failed. Check that you have really selected datum stations.

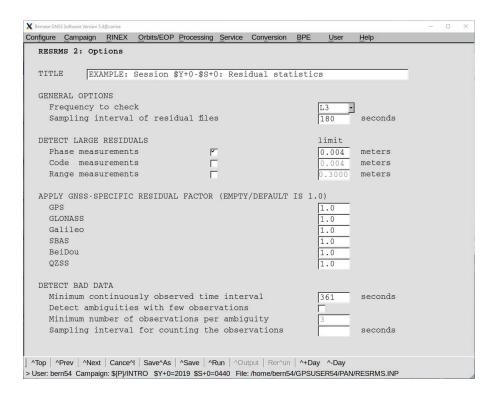
After running the program a warning message pops up indicating that the normalization of residuals based on the uncertainty of the estimated parameters (NORMALIZED in option "Type of computed residuals") fails. Some parameters related to the station WSRT (troposphere and ambiguities) show comparable high formal errors. In that case, no normalization of the residuals takes place. They ending as high residuals in the related result file and the related observations will be detected and removed in the subsequent steps.

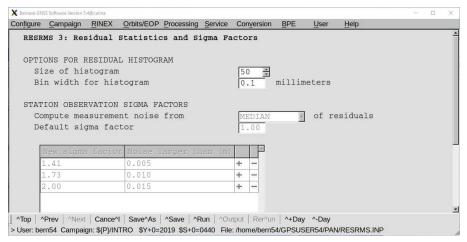
If the residuals have been stored in the binary residual file (specified in "GPSEST 2.2: Output Files 2") it is possible to have a look at the residuals (program REDISP, "Menu>Service>Residual files>Display residual file").

To screen the residuals automatically use the program RESRMS in "Menu> $\underline{S}$ ervice> $\underline{R}$ esidual files >Create residual statistics".

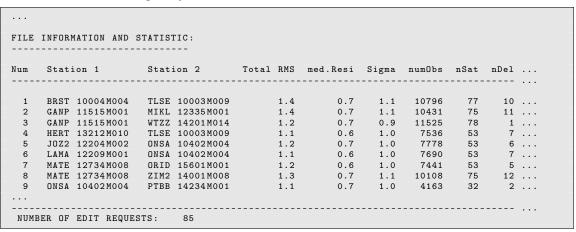


The sampling interval you have previously introduced in option "Sampling interval" in program GPSEST has to be repeated here. RESRMS makes the assumption that the observations between two outliers in the sampled residual file are also bad.



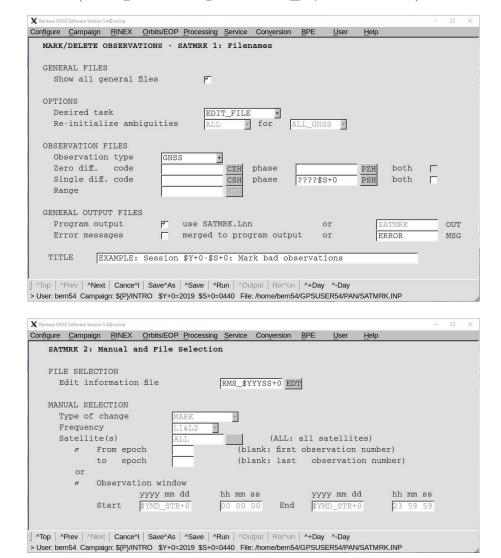


The program output of RESRMS ( $\P$ /INTRO/OUT/RMS\_2019044.OUT) provides a nice overview on the data quality.

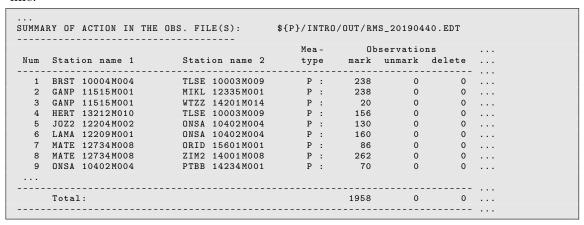


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In addition, files containing a summary table (\${P}/INTRO/OUT/RMS\_20190440.SUM) and a histogram (\${P}/INTRO/OUT/RMS\_20190440.LST) of the residuals are available. The most important result file for the data screening is the "Edit information file" (\${P}/INTRO/OUT/RMS\_20190440.EDT), which may be used by the program SATMRK to mark outliers in the observation files ("Menu>Service>Bernese observation files>Mark/delete observations"):

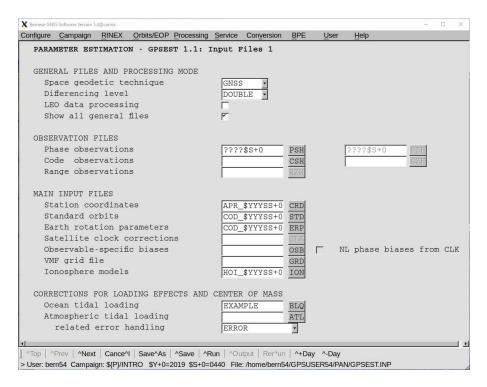


The program output from SATMRK reports the number of marked observations per baseline:



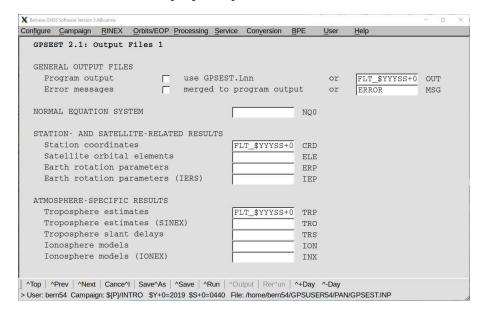
#### 5.2 Produce a First Network Solution

After screening the observations for outliers we can generate an ionosphere–free  $(L_3)$  solution with unresolved ambiguities. A detailed discussion on the troposphere/ionosphere modeling will be given in a dedicated lecture tomorrow. The input options are very similar to the previous preprocessing step. There are only a few differences shown in the following panels:



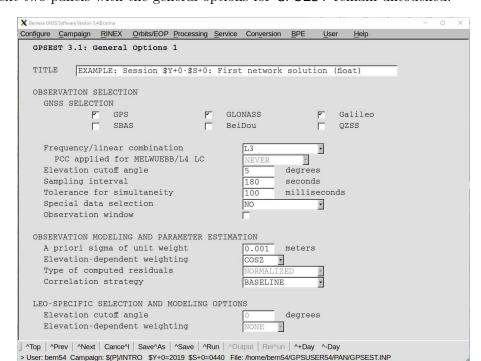
The file in the input field "lonosphere models" enables the HOI–corrections.

We store the coordinates and troposphere parameters into files to be re-introduced later:

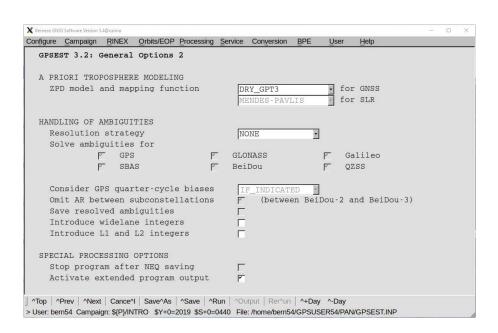


In the subsequent panel you should remove the output filename for the "Residuals" because we do not need the residuals from this run.

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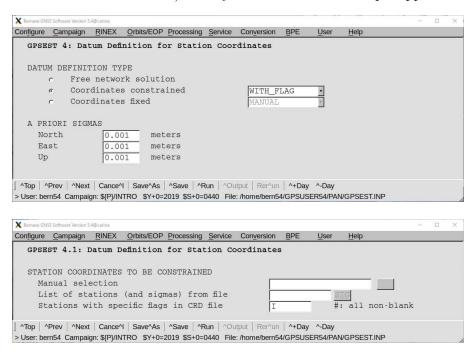


The next two panels with the general options for GPSEST remain untouched:

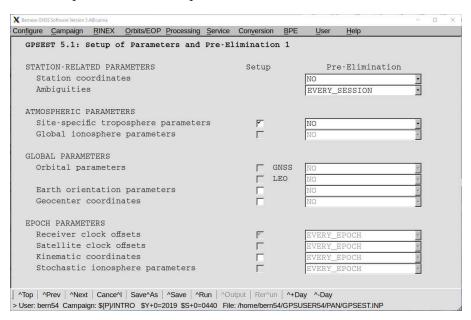


Nevertheless, it might be an interesting idea to enable the option "Activate extended program output". Two more panels appear where much more information on the data and results can be asked for. In this example we just enable the two options for "Printing: Extended summary concerning coordinates" and "Printing: Extended summary concerning troposphere". You are invited to discover these options further in order to adjust it to your needs.

To heavily constrain the coordinates of the IGS core sites is not the best way to realize the geodetic datum for a solution. The program ADDNEQ2 offers more sophisticated options (e.g., minimum constraint solution). Today we will follow this simple approach:

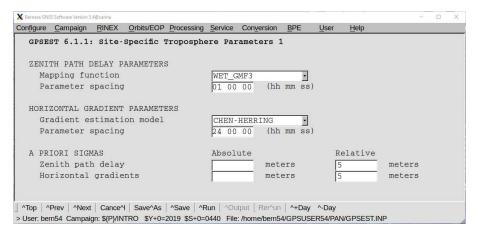


Since we do not store residual files in this run, ambiguity parameters may be pre-eliminated from the normal equation before the parameters are estimated:



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The estimation of troposphere parameters is mandatory for a campaign of this type. We increase the number of estimated parameters (e.g., to 24 instead of 6 parameters per station and session). In addition, it is recommended to set up troposphere gradient parameters. In order to avoid a format overflow in the "Troposphere estimates" output file that may happen if a troposphere parameter is estimated based on very few observations concentrated at one end of the interval of parameter validity, a loose relative constraint (e.g., sigmas of 5 meter) may help.



In the first part of the output generated by program GPSEST, the selected options are echoed. The result part starts with some statistics on the parameters and the observations:

```
SUMMARY OF RESULTS
Number of parameters:
                               Adjusted Expl. Impl. (elim)* Del. Ref. Sing. No obs
Parameter type
Station coordinates / velocities
                                      51 51
2739 0
493 493
                                                                 0 0
0 0
0 0
                                                     2739 (bfst)
                                                                             40
                                                                            0 0
Site-specific troposphere parameters
                                                     0
Total number
                                      3243 544
                                                    2699
                                                                   0
                                                                      0
                                                                           40
                                                                                  0
*Abbreviation pre-elimination (elim): before stacking (bfst), after stacking (afst),
                                   epoch-wise (epow), after inversion (afin)
```

Then the a posteriori RMS error and the results of the initial least–squares adjustment are given

```
Statistics:
Total number of authentic observations
                                             133235
Total number of pseudo-observations
Total number of explicit parameters
                                                544
                                              2699
Total number of implicit parameters
                                             133265
Total number of observations
Total number of adjusted parameters
                                               3243
Degree of freedom (DOF)
                                             130022
A posteriori RMS of unit weight
                                           0.000987 m
Chi * * 2 / DOF
                                               0.97
 . . .
```

Below you find the extended output of the results for coordinates and troposphere parameters at the end of the file that we have asked for in panel "GPSEST 3.2.1.1: Extended Program Output Options":

Reference epoch: 20	019-02-13	11:59:45				
Station name	Тур	A priori value	Estimated value	Correction	RMS error	• •
BRST 10004M004	X	4231162.41765	4231162.41830	0.00065	0.00063	• • •
	Y	-332746.42891	-332746.42771	0.00120	0.00056	
	Z	4745131.06968	4745131.07283	0.00315	0.00060	
	U	65.81939	65.82211	0.00272	0.00076	
	N	48.3804957	48.3804957	0.00168	0.00042	
	E	-4.4965914	-4.4965914	0.00125	0.00056	
TLSE 10003M009	Х	4627851.66234	4627851.66205	-0.00029	0.00051	
	Y	119640.28843	119640.28820	-0.00023	0.00046	
	Z	4372993.71461	4372993.71376	-0.00085	0.00051	
	U	207.18324	207.18244	-0.00080	0.00062	
	N	43.5606968	43.5606968	-0.00041	0.00037	
	E	1.4808939	1.4808938	-0.00022	0.00046	
GANP 11515M001	Х	3929181.30600	3929181.29822	-0.00778	0.00061	
	Y	1455236.94546	1455236.95378	0.00832	0.00053	
	Z	4793653.99294	4793654.03586	0.04292	0.00063	
	U	746.00460	746.03412	0.02952	0.00078	
	N	49.0347149	49.0347152	0.03144	0.00041	
	E	20.3229395	20.3229396	0.01046	0.00051	
MIKL 12335M001	Х	3698553.70970	3698553.70973	0.00003	0.00064	
	Y	2308676.23252	2308676.23206	-0.00046	0.00060	
	Z	4639769.62655	4639769.62475	-0.00180	0.00066	
	U	93.92954	93.92808	-0.00146	0.00081	
	N	46.9727862	46.9727862	-0.00107	0.00045	
	E	31.9728450	31.9728450	-0.00040	0.00059	

Reference elevati	ion angle o	f gradien	t terms :	45.0 de	grees		
Minimum elevation			:		grees		
Mapping factor at	t minimum e	levation	angle :	11.4			
	9		( )	DWG			
			(m)		errors (		• • •
Station name	North	East	Zenith	North	East	Zenith	
DDGT 40004W004	0.00044		0.44007		0.00044	0.00450	• • •
BRST 10004M004	0.00044	0.00003	0.11297		0.00011		
BRST 10004M004	0.00043	0.00003	0.10486	0.00008	0.00010	0.00105	• • •
BRST 10004M004	0.00043	0.00003	0.10191	0.00008	0.00009	0.00113	• • •
BRST 10004M004	0.00043	0.00002	0.10370	0.00007	0.00009	0.00108	• • •
BRST 10004M004	0.00042	0.00002	0.09834	0.00007	0.00008	0.00104	• • •
BRST 10004M004		-0.00007	0.06907	0.00007	0.00009	0.00106	
BRST 10004M004	0.00034	-0.00007	0.06841	0.00008	0.00010	0.00113	
BRST 10004M004	0.00034	-0.00008	0.07125	0.00008	0.00011	0.00116	
BRST 10004M004	0.00033	-0.00008	0.06842	0.00009	0.00011	0.00212	
GANP 11515M001	-0.00001	-0.00088	0.04855	0.00006	0.00006	0.00110	
GANP 11515M001	-0.00003	-0.00087	0.04251	0.00006	0.00006	0.00082	
GANP 11515M001	-0.00005	-0.00086	0.04326	0.00005	0.00006	0.00081	
GANP 11515M001	-0.00007	-0.00085	0.04354	0.00005	0.00005	0.00076	
GANP 11515M001	-0.00009	-0.00083	0.04544	0.00005	0.00005	0.00085	
GANP 11515M001		-0.00082	0.05617	0.00004	0.00005	0.00073	

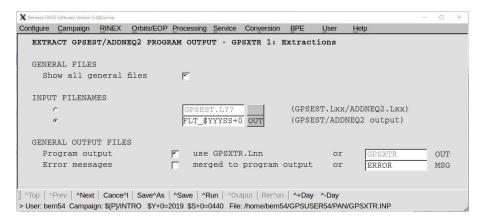
Because outliers have been removed in the previous step, the obtained a posteriori RMS error should decrease (at least not increase). If this is not the case, it is likely that the observations and the heavily constrained a priori coordinates are inconsistent.

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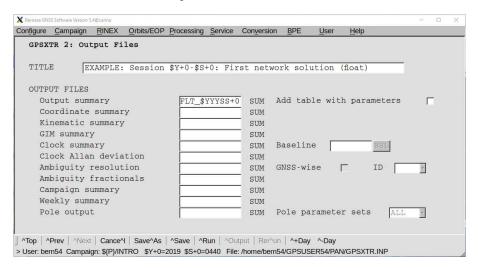
## 5.2.1 Program Output Extraction

The program output from the program GPSEST is quite long and contains many information that help to identify potential problems in the processing. On the other hand, it would be smart to extract the key parameters into a short summary that can for instance be included in a comprehensive processing protocol. The program GPSXTR (accessible via "Menu>Processing>Program output extraction>Parameter estimation/stacking") provides such a summary.

Select the program output(s) to be analyzed:



and select the name of the summary file



The short extraction from the full program output in \${P}/INTRO/OUT/FLT\_20190440.OUT reads like

```
FLT_20190440.0UT Rms: 0.99 , # fil.: 16 , # obs.: 133265 , # par.: 3243 (DOY: 2019-044)
```

Also other styles of output are available, check the online help for some examples. Please note, that some summaries only work on specific program output files containing the expected parameter types.

## 5.3 Ambiguity Resolution

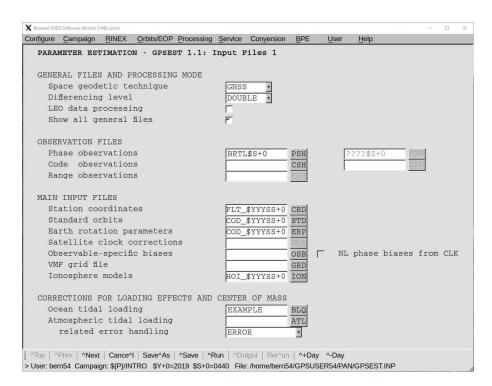
To resolve the ambiguities, we process the baselines separately one by one using the Quasi–Ionosphere–Free (QIF) strategy. This baseline processing mode is necessary because of the tremendous number of parameters. The attempt to resolve the ambiguities in a session solution might require too much CPU and memory to be feasible (several iterations with inversions of the full normal equation (NEQ) are necessary).

## 5.3.1 Ambiguity Resolution: Quasi-lonosphere-Free (QIF)

The complete list of baseline observation files of a session (e.g., session 0440 of year 2019) can be generated by listing all phase single–difference header files in the campaign's observation directory of your campaign:

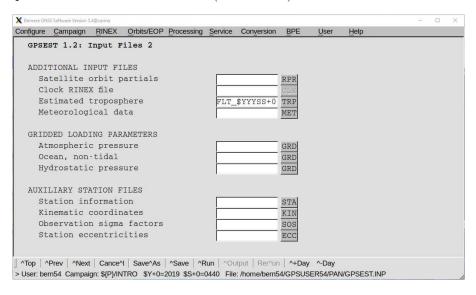
```
bern52@carina: > ls ${P}/INTRO/OBS/????0440.PSH
${P}/INTRO/OBS/BRTL0440.PSH
${P}/INTRO/OBS/GAMIO440.PSH
${P}/INTRO/OBS/GAWZO440.PSH
${P}/INTRO/OBS/HETLO440.PSH
${P}/INTRO/OBS/JOONO440.PSH
${P}/INTRO/OBS/LAONO440.PSH
${P}/INTRO/OBS/MAORO440.PSH
${P}/INTRO/OBS/MAZIO440.PSH
${P}/INTRO/OBS/ONPTO440.PSH
${P}/INTRO/OBS/ONWZO440.PSH
${P}/INTRO/OBS/TLZI0440.PSH
${P}/INTRO/OBS/VIWZO440.PSH
${P}/INTRO/OBS/WSWZ0440.PSH
${P}/INTRO/OBS/WTWZ0440.PSH
${P}/INTRO/OBS/WZZI0440.PSH
${P}/INTRO/OBS/ZIZMO440.PSH
```

The first baseline for this session is from BRST to TLSE with the observation filename BRTL0440. Using the menu time variables this name is specified as BRTL\$S+0. The following options are used for the ambiguity resolution step:

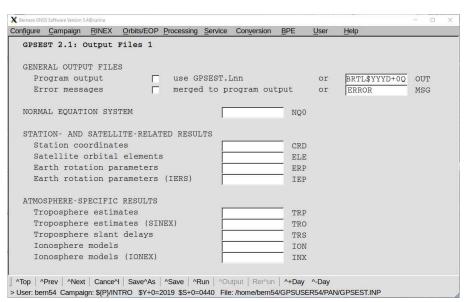


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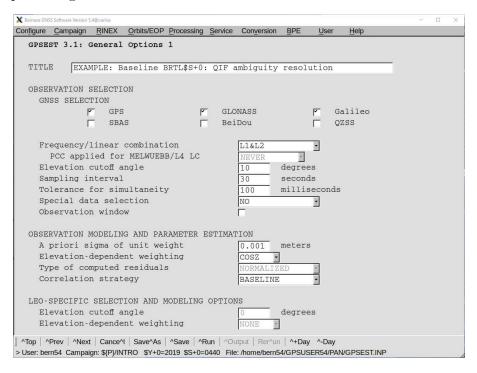
Only one baseline file is selected. Coordinates and troposphere estimates are introduced from the previous first network solution (Section 5.2).



Specify a baseline specific output to prevent overwriting in subsequent runs: BRTL\$YYYD+0Q. The Q at the end shall indicate that it is the output from the QIF–strategy.

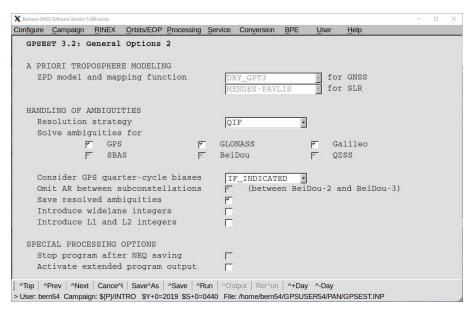


Because the QIF–ambiguity resolution strategy is very sensitive to the formal errors of the ambiguity parameters we have to include all measurements with the full sampling of 30 s into the processing.



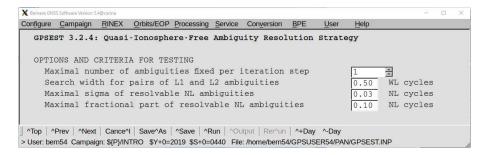
The selection of a "ZPD model and mapping function (GNSS)" is disabled because a troposphere file has been introduced in panel "GPSEST 1.2: Input Files 2". The program uses the troposphere model from this input file and allows no other selection for consistency reasons.

In the subsequent panel the "Resolution strategy" is chosen. Please, do not forget to store the resolved integer ambiguities in your observation file (mark checkbox "Save resolved ambiguities").

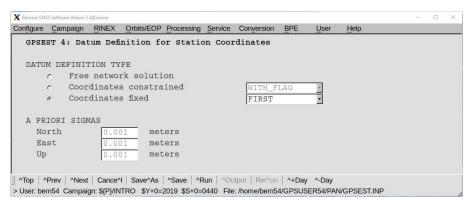


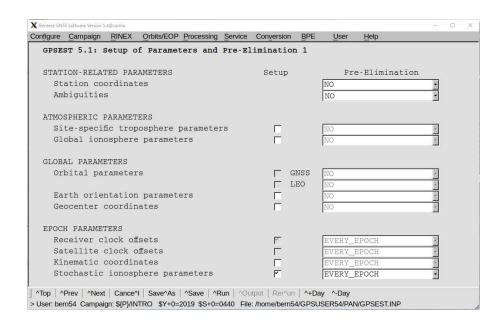
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In case of ambiguity resolution including GLONASS, only one ambiguity per iteration can be resolved. The program will adjust the setting automatically issuing a warning message.

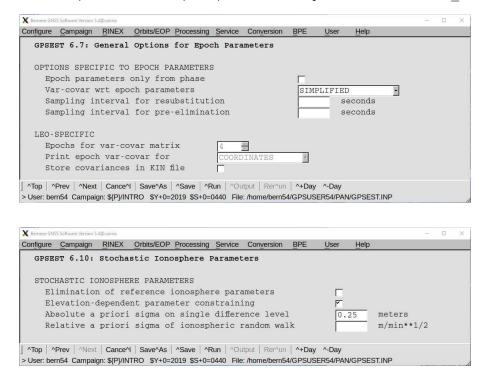


Even if we introduce the coordinate solution from the previous network solution from Section 5.2, we fix only the coordinates of the first station and estimate those from the second one with each iteration of ambiguity resolution. In this way, the geometry can react to the resolved ambiguities.

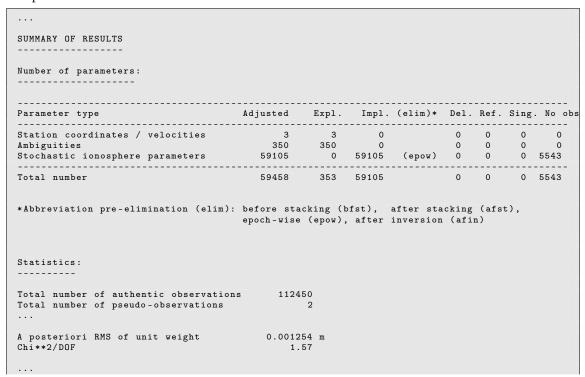




An additional panel with options specific to epoch–parameters is displayed now because the "Parameter Setup: stochastic ionosphere parameters" are pre-eliminated EVERY EPOCH.



After reporting input options and input data for the current run of GPSEST, the results are presented in two parts. The first part refers to the solution where the ambiguities are estimated as real values whereas the second part reports the results after resolving the ambiguity parameters to integer values. The real-valued estimates for the ambiguities may be found below the Station coordinates and velocities section of the program output:



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	tation			Тур		rect	ion	Estim				error				
	LSE 10			X Y Z		0.00	 110	462								
1 T	LSE 1	0003M	009	Y	(	0.000	010	11	9640	.28830	0.0	00021	119640	28820		# C
1 T	LSE 1	0003M	009	Z	-	0.000	092	437	2993	71284	0.0	00018 00021 00016	4372993	3.71376		# C
MBTG	UITIES	3 :														
						D.I		NCE								
MBI	FILE	SAT	S EPO	L EPO	FRO		EFERE CLU		CLU	AMBI	GUITY	7 RMS	TOTAI.	AMBIGU		
				704				470	0.4.7		0.4			00770		
1	1	1	1	701	1			172	217		0.11					
2	1	1	2830	2880	1	1	2	172	217			0.85		32629.50		
3	1	7	720	1326	1	1	4	172 172	217			7 0.12		14156.73		
4 5	1 1	7 8	2646 1	2825 401	1	1	5	172	217 217			7 0.47		14121.63		
6	1	8	2576	2581				172	217			0.13		01549.57		
6 7	1		2576	2880	1 1			172	217			3.04 5 0.16		)1598.90 )1598.25		
8	1	10	2588	2880	1			172	217			1 0.20		10600.94		
9		10	1548	1718	1		11		217			1 0.20		40558.06		
10	1	10	1725	1852	1	1	12		217			1 0.35		40558.00 40553.86		
11	1	10	2561	2825	1	1		172	217		5.07			40673.93		
12	1	10	2832	2880	1	1	15		217			7 0.30		40673.33 40674.73		
13	_	11	1	606	1	1	16		217			9 0.11		16706.31		
14	1	11	2791	2825	1	1		172	217		-0.42			16725.58		
15		11	2849	2880	1	1	19		217		3.13			16724.13		
16	_	16	1	78	1	1	20		217		-1.69			04795.69		
17	1	16	2307	2842	1	1		172	217		0.19			04862.81		
18		16	2849	2880	1	1	23		217			0.55		04861.15		
19	1	18	1	561	1	1	24	172	217			0.12		19999.28		
20	1	18	2689	2842	1	1	25	172	217		-1.60	0.31	134	19919.40	)	
21	1	18	2849	2880	1	1	26	172	217		3.24	0.42	134	19919.24	1	
22	1	20	1	106	1	1	27	172	217		-1.43	0.45	-13	39284.43	3	
23	1	20	1424	1432	1	1	29	172	217		-2.66			39505.66	3	
24	1		1439	1453	1	1	30		217			1.48		39498.11		
25	1	20	1460	1779	1	1		172	217		-2.99			39500.99		
26	1	20	2465	2842	1	1	32		217		-0.08			39436.08		
27	1		2849	2880	1	1	33		217			0.56		39435.42		
28	1	22	19	751	1			172	217		3.07			58169.93		
29	1	26	2167	2825	1	1	36		217		-3.61			10511.61		
30	1	27	1	272	1			172	217			0.16		99410.68		
31		27	2411	2825				172	217		1.20			99408.80		
32	1	27	2832	2880	1			172	217			0.26		99407.48		
33	1	30	836	1486	1	1	41	172	217		-1.23	0.12	35	72427.77		

If GLONASS data are processed, single–difference (instead of double–difference) ambiguities are resolved and no — REFERENCE — as in case of, e.g., a GPS–only solution appears.

In the next part of the output the result of the QIF ambiguity resolution algorithm is given:

```
AMBIGUITY RESOLUTION:

STRATEGY: QUASI-IONOSPHERE-FREE AMBIGUITY RESOLUTION (QIF)

AMBIGUITY RESOLUTION ITERATION: 1

BEST INT. CORRECTIONS IN CYCLES

FILE AM1 CL1 #AM1 AM2 CL2 #AM2 L1 L2 L1 L2 L5 L3 RMS(L3) SA1 SA2

1 154 188 1 175 221 1 -4 0 -0.40 -0.54 0.138 0.007 0.003 211 225
```

	GUITY	Y RE	SOLUT	ION	ITER	ATION:	:	2							
FILE	AM1	CL1	# A M 1	AM2	CL2	#AM2	BEST L1	INT. L2	COR L1	RECTIONS L2	IN CYC	LES L3	RMS(L3)	SA1	SA2
				456	400				0.00	0.00	0.000	0.000	0.000	001	
1	66	79	1	156	190								0.003		
AMBI	GUITY	Y RE	SOLUT	ION	ITER	ATION:									
FILE	AM1	CL1	# A M 1			#AM2	L1	L2	L1		L5	L3	RMS(L3)		
	70												0.003		
													0.003		
	GUITY	Y RE	SOLUT	ION	ITER	ATION:	15	5							
										DECETIONS					
FILE	AM1	CL1	#AM1	AM2	CL2	#AM2	L1	L2	L1	RECTIONS L2	L5	LES L3	RMS(L3)	SA1	SA2
													0.004		
						ATION:									
FILE	AM1	CL1	# AM1	AM2	CL2	#AM2	BEST L1	INT. L2	COR L1	RECTIONS L2	IN CYC	LES L3	RMS(L3)	SA1	SA2
1	28	34	2	98	118	1	4	3	-0.00	-0.00	0.001	0.000	0.003	22	3
AMBI						 ATION:									
FILE	AM1	CL1	# AM1	AM2	CL2					RECTIONS L2			RMS(L3)	SA1	SA2
1	51	62	1	113	136	1	8	5	-0.45	-0.58	0.130	0.004	0.005	114	110
						 ATION:									
C11 C	A M 1	CI 1	# A M 4	АМО	CI O					RECTIONS			DWG(I3)	C A 1	g A O
	AM1		# AM1	AM2		#AM2		L2		L2	гр	L3	RMS(L3)	5A1	5A2
1	26	32	1	173	219	12	-8	-8	-0.79	-1.00	0.211	-0.044	0.005	20	25
AMBI						ATION:									
										RECTIONS					
FILE										L2			RMS(L3)	SA1	SA2
				101	121	1	1	0	0.78	1.00	-0.217	0.021	0.005	107	103
	44	55	1												
1															
1	GUITY	r RE	SOLUT	 ION	ITER	ATION:	5:	1							
1	GUITY	r RE	SOLUT	 ION	ITER	ATION:	5:	1							
1  AMBI( 	GUITY	RE	SOLUT	ION 	ITER	#AM2	BEST L1	1  INT. L2	COR L1	RECTIONS	IN CYC	LES L3	RMS(L3)	SA1	SA2
1 AMBIO	AM1	Y RE	SOLUT	ION  AM2	CL2	# AM2	5: BEST L1	1  INT. L2	COR L1	RECTIONS L2	IN CYC	LES L3	RMS(L3)	SA1	SA2

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The individual iteration steps are first described (we specified that only one ambiguity may be resolved within each iteration step — see panel "GPSEST 3.2.4: Quasi-lonosphere-Free Ambiguity Resolution Strategy"). The following information is listed for each resolved double—difference ambiguity:

FILE file number (1 in our case; we process one baseline only),

AM1 first ambiguity number (single-difference level),

CL1 corresponding ambiguity cluster,

#AM1 number of ambiguities belonging to the same cluster,

#### AM2, CL2, #AM2

similar information for the second ambiguity.

#### BEST INT. L1, L2

are the integer corrections to the a priori values (a priori values are computed using the a priori coordinates and may be rather inaccurate).

#### CORRECTIONS IN CYCLES

for carriers L1 and L2 gives the information about the fractional parts of the  $L_1$  and  $L_2$  ambiguities. The CORRECTIONS IN CYCLES L5 and L3 are of greater interest. The value L5 represents the ionosphere–induced bias expressed in  $L_5$  cycles. These values may not be greater than the maximum value specified in panel "GPSEST 3.2.4: Quasi-lonosphere-Free Ambiguity Resolution Strategy" (option "Search width for pairs of L1 and L2 ambiguities"). RMS(L3) is the criterion according to which the ambiguities are sorted. Ambiguities with  $L_3$  RMS errors larger than the value specified in the program input panel (in our example 0.03) will not be resolved.

#### SA1, SA2

first and second satellite number related to the ambiguities.

The first ambiguities are resolved for Galileo; with iteration AMBIGUITY RESOLUTION ITERATION: 15 the first GPS satellites follow. In ITERATION: 48, 50 and 51 there are examples for resolving pairs of ambiguities from the satellite using the same frequency. Later there are even examples for path—to—path ambiguity resolution, e.g., ITERATION: 65 or 76, indicated by the same satellite number for SA1 and SA2.

The following table summarizes the results of the ambiguity resolution:

						R	EFERE	NCE					
AMBI	FILE	SAT.	S_EPO	L_EPO	FRQ	WLF	CLU	AMBI	CLU	AMBIGUITY	RMS	TOTAL AMBIGU.	DL/L
1	1	1	1	701	1	1	1	28	34	-3		1432773.	0.00000
2	1	1	2830	2880	1	1	2	173	219	-4		1432622.	0.00000
3	1	7	720	1326	1	1	4	126	157	-2		2614159.	0.00000
4	1	7	2646	2825	1	1	5	173	219	-11		2614112.	0.00000
5	1	8	1	401	1	1	7	125	155	-4		1201547.	0.00000
6	1	8	2576	2581	1	1	8	172	217	-3.32	3.14	1201596.68	
7	1	8	2588	2880	1	1	9	173	219	-10		1201590.	0.00000
8	1	10	1	240	1	1	10	141	174	2		-940597.	0.00000
9	1	10	1548	1718	1	1	11	173	219	-7		-940567.	0.00000
10	1	10	1725	1852	1	1	12	172	217	-1.17	0.46	-940556.17	
11	1	10	2561	2825	1	1	14	141	174	9		-940670.	0.00000
12	1	10	2832	2880	1	1	15	173	219	-5		-940683.	0.00000
13	1	11	1	606	1	1	16	19	24	0		2346708.	0.00000
14	1	11	2791	2825	1	1	18	172	217	-2.67	0.97	2346723.33	
15	1	11	2849	2880	1	1	19	173	219	-5		2346716.	0.00000
16	1	16	1	78	1	1	20	173	219	-11		-2504805.	0.00000

17	1	16	2307	2842	1	1	22	31	39	-1		-2504864.	0.00000
18	1	16	2849	2880	1	1	23	143	176	1		-2504864.	0.00000
19	1	18	1	561	1	1	24	98	118	-1		1350000.	0.00000
20	1	18	2689	2842	1	1	25	36	46	-1		1349920.	0.00000
21	1	18	2849	2880	1	1	26	173	219	-5		1349911.	0.00000
22	1	20	1	106	1	1	27	143	176	-3		-139286.	0.00000
23	1	20	1424	1432	1	1	29	172	217	-4.97	2.49	-139507.97	
24	1	20	1439	1453	1	1	30	172	217	3.57	1.53	-139500.43	
25	1	20	1460	1779	1	1	31	173	219	-12		-139510.	0.00000
26	1	20	2465	2842	1	1	32	173	219	-8		-139444.	0.00000
27	1	20	2849	2880	1	1	33	143	176	2		-139437.	0.00000
28	1	22	19	751	1	1	34	98	118	4		-1558169.	0.00000
29	1	26	2167	2825	1	1	36	165	208	-10		-3210518.	0.00000
30	1	27	1	272	1	1	38	141	174	2		-899407.	0.00000
31	1	27	2411	2825	1	1	39	141	174	5		-899405.	0.00000
32	1	27	2832	2880	1	1	40	143	176	1		-899410.	0.00000
33	1	30	836	1486	1	1	41	173	219	-10		3572419.	0.00000

The ambiguities for which an RMS is specified could not be resolved (these ambiguities will be treated as real values by all subsequent program runs). In case of GLONASS, only ambiguities with the same channel number are resolved in Version 5.4 of  $Bernese\ GNSS\ Software$ .

Ambiguity resolution has an influence on other parameters. Therefore, the results of the ambiguity–fixed solution are given in Part 2 of the output:

SUMMARY OF RESULTS								
Number of parameters:								
Parameter type		Expl.	Impl.	 (elim)*	Del.	 Ref.	Sing	 . No ob
Station coordinates / velocities Ambiguities Stochastic ionosphere parameters	3 86 59105	3 84 0	0 0 59105	(epow)	0 0 0	0 2 0	0 0 0	0 0 5543
Total number	59194	87	59105		0	0	0	5543
*Abbreviation pre-elimination (elim	): before star epoch-wise						t),	
Statistics:								
Total number of authentic observati Total number of pseudo-observations								
Total number of explicit parameters Total number of implicit parameters		39 )5						
Total number of observations Total number of adjusted parameters Degree of freedom (DOF)	1124 591 532	94						
A posteriori RMS of unit weight Chi**2/DOF	0.00129							
Total number of observation files Total number of unobserved ambiguit Total number of stations	ies	1 92 1						
Station coordinates and velocities:								
Sol Station name Typ Correct								
1 TLSE 10003M009 X -0.0 1 TLSE 10003M009 Y 0.0 1 TLSE 10003M009 Z -0.0	00063 46278 00013 1196 00057 43729	51.66142 10.28832 93.71319	0.00 0.00 0.00	014 4 007 014 4	162785 11964 137299	1.662 0.288 3.713	05 19 76	#CRD #CRD #CRD

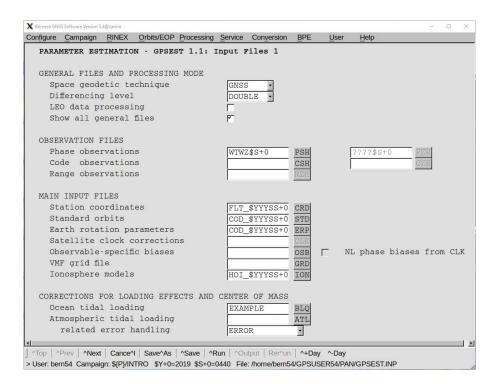
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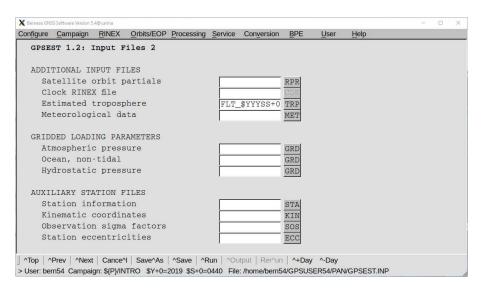
You may see from the output that from a total of 350 ambiguities 266 ambiguities could be resolved (compare part 1 AMBIGUITIES with part 2 AMBIGUITIES). Note that these numbers include reference ambiguities for each GNSS, GLONASS frequency number and frequency.

## 5.3.2 Ambiguity Resolution: Short Baselines

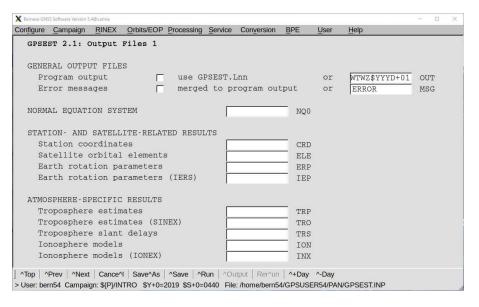
There are two very short baselines in the network where a direct ambiguity resolution for the  $L_1$  and  $L_2$  signal is possible applying the sigma–strategy.

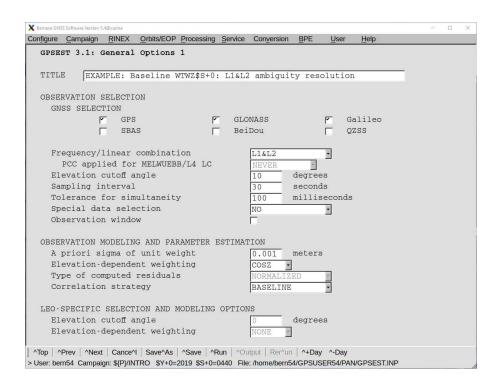
The ultra–short baseline in Kötzting is between WTZR and WTZZ (WTWZ0440.PSH). The GPSEST input panels should look like follows:



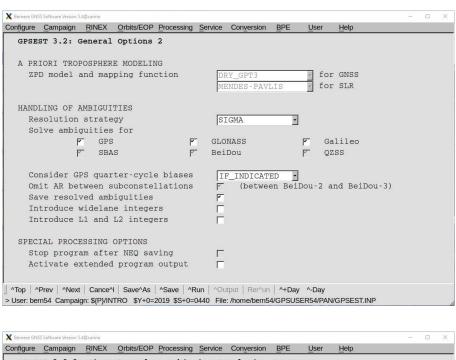


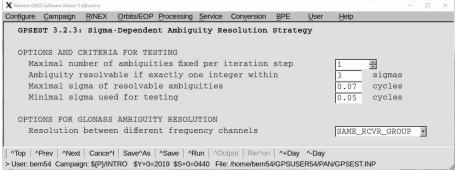
The program output name is again related to the name of the baseline but contains an identifier 1 at the end to distinguish the files from the output files of the QIF-strategy:

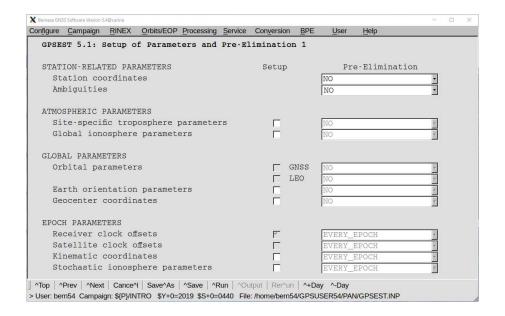




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The structure of the program output is the same as it has extensively been described in the previous section for the QIF ambiguity resolution strategy. It starts with PART 1 for the solution before the ambiguity resolution. Here are the corresponding parameter statistics:

```
SUMMARY OF RESULTS
Number of parameters:
                                        Adjusted Expl. Impl. (elim)* Del. Ref. Sing. No obs
Parameter type
                                       3 3 0 0 0 0 0 0 3 344 344 0 0 0 0 0 0
Station coordinates / velocities
Total number
                                                347 347 0
                                                                                  0 0 0 0
*Abbreviation pre-elimination (elim): before stacking (bfst), after stacking (afst), epoch-wise (epow), after inversion (afin)
Statistics:
Total number of authentic observations
Total number of pseudo-observations
Total number of explicit parameters 347
Total number of implicit parameters 0

Total number of observations 118378
Total number of adjusted parameters 347
Total number of freedom (DOF) 118031
A posteriori RMS of unit weight
                                            0.001128 m
Chi**2/DOF
                                                   1.27
Total number of observation files
Total number of unobserved ambiguities
                                                       42
Total number of stations
```

After the ambiguity resolution the same statistics is provided in PART 2:

```
SUMMARY OF RESULTS
Number of parameters:
                             Adjusted Expl. Impl. (elim)* Del. Ref. Sing. No obs
Parameter type
Station coordinates / velocities 3 3 0 0 0 0 0 Ambiguities 56 52 0 0 4 0 0
          ------
                                          59 55 0
Total number
*Abbreviation pre-elimination (elim): before stacking (bfst), after stacking (afst),
                                     epoch-wise (epow), after inversion (afin)
Statistics:
Total number of authentic observations 118374
Total number of authentic observations

Total number of pseudo-observations

Total number of explicit parameters

Total number of implicit parameters
Total number of observations 118378
Total number of adjusted parameters 61
118317
Degree of freedom (DOF)
A posteriori RMS of unit weight 0.001326 m
Chi**2/DOF
                                            1.76
Total number of observation files
Total number of unobserved ambiguities 42
Total number of stations
                                                1
```

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From the number of ambiguity parameters it can be computed that 288 out of 344 ambiguities from GPS, GLONASS, and Galileo have been resolved to their integer numbers. Please remind, that at least 6 ambiguity parameters must remain real values because one reference ambiguity per GNSS and frequency are needed. For the Code Division Multiple Access (CDMA) systems, GPS and Galileo they are reported as — Reference —. Because the two receivers belong to different groups regarding "GLONASS amb. resol. between different frequencies" only the ambiguities between the same frequency numbers have been resolved. In addition, depending on the receiver type, not all ambiguities for GPS are allowed to be resolved too, to prevent problems with the quarter–cycle bias between the P- and C-code signal (see option "Consider GPS quarter-cycle biases" in panel "GPSEST 3.2: General Options 2" and the lecture on ambiguity resolution).

## 5.3.3 Ambiguity Resolution: BPE

Admittedly, it is cumbersome to process the baselines "manually" one after the other — you have sixteen baselines per session for this small example campaign. On Thursday you will have a lecture on automation of the data processing using the BPE.

In the example BPE RNX2SNX.PCF the sequence for the ambiguity resolution is included. For this tutorial lecture a small part of this BPE is extracted into a separate TUTORIAL.PCF BPE. The process control file (PCF) is located in the directory \${U}-PCF. If you are not in the Bernese Introductory Course environment, you have to use the RNX2SNX.PCF instead of the TUTORIAL.PCF and skip the scripts that are not needed. You also need to adjust the variables in the lower section (a brief description is given in the panel description of the RUNBPE) program).

The TUTORIAL.PCF consists of two parts: the first part defines the scripts and the option directories where the program's input files are taken from. In addition the CPU-specification and waiting conditions are defined to keep the correct order for the execution of the scripts. The keyword PARALLEL indicates a special execution mode for the scripts, e.g., GNSAMB\_P with PID 412 or GNSQIF\_P with PID 432. They may run in parallel for each individual baseline. The preparatory scripts, e.g., GNSAMBAP or GNSQIFAP define the list of baselines to be processed. The second part defines the so called BPE— or PCF—variables that can be used within the scripts or in the input fields of the menu. A detailed introduction to the BPE will be the topic of a lecture on Thursday.

The TUTORIAL.PCF is responsible for the following tasks:

#### 401 SATMRK:

All previously resolved ambiguities are re–initialized to start for all files from unresolved ambiguities (otherwise the interpretation of the statistic of resolved ambiguities may become difficult).

#### 411 GNSAMBAP and 412 GNSAMB\_P:

Applies the SIGMA ambiguity resolution strategy using the Melbourne-Wübbena linear combination to all baselines longer than 200 km of the example in a baseline—by—baseline mode, where several baselines may be processed at the same time in parallel.

#### 421 GNSL53AP and 422 GNSL53\_P:

Applies the SIGMA ambiguity resolution strategy using the wide- and narrow-lane

```
# TUTORIAL . PCF
# Purpose:
# Run the ambiguity resolution for one session in the <<Terminal Session>>
# of the Bernese Software Introductory Course
            : R. Dach
: 27-Aug-2022
# Created
# Changes
PID SCRIPT OPT_DIR PARAMETERS
# Resolve phase ambiguities
401 SATMRK
                          CPU = ANY
                R2S_GEN
411 GNSAMBAP R2S_AMB
                            CPU=ANY; WAIT=401
                          CPU=ANY; WAIT=411; PARALLEL=411
CPU=ANY; WAIT=412
CPU=ANY; WAIT=421; PARALLEL=421
CPU=ANY; WAIT=422
CPU=ANY; WAIT=431; PARALLEL=431
412 GNSAMB_P R2S_AMB
421 GNSL53AP R2S_L53
422 GNSL53_P R2S_L53
431 GNSQIFAP R2S_QIF
432 GNSQIF_P R2S_QIF
441 GNSL12AP R2S_L12
                            CPU=ANY; WAIT=432
442 GNSL12_P
443 AMBXTR
                R2S_L12 CPU=ANY; WAIT=441; PARALLEL=441
R2S_AMB CPU=ANY; WAIT=442
# End of BPE
899 DUMMY
                NO_OPT
                            CPU=ANY; WAIT=443
VARIABLE DEFAULT
                           PARAMETERS
# General and model files:
V_SUBMOD = DESAI2016;
                             {\tt DESCRIPTION=Subdaily\ pole\ model}
V_NUTMOD = IAU2000R06;
                            DESCRIPTION=Nutation model
                             DESCRIPTION = Antenna phase center (PCV) model DESCRIPTION = PCV information file
V_PCV
         = I20;
DESCRIPTION=Satellite information file
V_ORB
         = COD;
                             DESCRIPTION=Orbit/ERP, CLK, bias information
\# Reference frame and station related files:
                            DESCRIPTION=Merged CRD/VEL filename
V_STAINF = EXAMPLE;
                             DESCRIPTION=_Station information file
V_BLQINF = EXAMPLE;
                            DESCRIPTION=_BLQ file name, CMC corrections
V ATLINF = EXAMPLE;
                             DESCRIPTION=_ATL file name, CMC corrections
V_HOIFIL = HOI_$YYYSS+0; DESCRIPTION=_Tonosphere model for higher order iono V_OSBFIL = COD_$YYYSS+0; DESCRIPTION=_OSB-file with GNSS satellite biases
# Data selection:
# Data selection:
V_OBSSEL = OBSERV;
                             DESCRIPTION = Observation selection file
                             DESCRIPTION=Select the GNSS (max. supported: GRECJ)
V_SATSYS = GRE;
# Ambiguity resolution:
V_GNSSAR = GRE;
                             DESCRIPTION=Select the GNSS for amb. resolu. (GECJ)
V_BL_AMB = 6000;
                             DESCRIPTION=Maximum baseline length for MW/L3 AR
V_BL_QIF = 2000;
                             DESCRIPTION=Maximum baseline length for QIF AR
V_BL_L53 = 200;
                             DESCRIPTION=Maximum baseline length for L5/L3 AR
                             DESCRIPTION=Maximum baseline length for L1&L2 AR
V BL L12 = 20:
# Other solution IDs used in the BPE:
V_APR
         = APR;
                             DESCRIPTION=A priori information
                             DESCRIPTION=Preliminary (ambiguity-float) results
# DO NOT USE V_D, V_J, V_M, V_Y VARIABLES!
# (they are used already by the menu)
```

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linear combinations to all baselines shorter than  $200\,\mathrm{km}$  and longer than  $20\,\mathrm{km}$ . In this example we don't have such a baseline, this section will be skipped automatically.

#### 431 GNSQIFAP and 432 GNSQIF\_P:

Applies the QIF ambiguity resolution strategy to all baselines of the example in a baseline—by—baseline mode, where several baselines may be processed at the same time in parallel. The resolved ambiguities from the previous steps are introduced and, consequently the algorithm only applies to the remaining unresolved ambiguities.

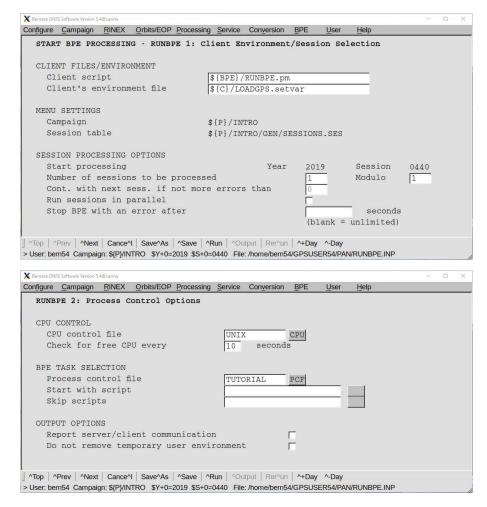
#### 441 GNSL12AP and 442 GNSL12\_P:

Applies the SIGMA ambiguity resolution strategy directly to the original observations on the  $L_1$  and  $L_2$  frequency for both short baselines in the example network: Kötzting (WTZR and WTZZ) and Zimmerwald (ZIM2 and ZIMM). These scripts also run in a baseline—by—baseline mode, allowing for a parallel processing.

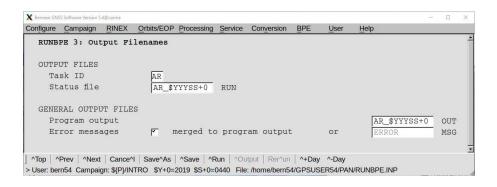
#### 443 AMBXTR :

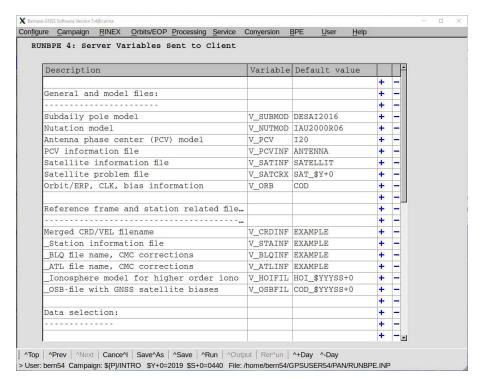
Creates a series of summary files on the success rate of the ambiguity resolution process in the various steps.

In the *Bernese Introductory Course* environment the TUTORIAL.PCF BPE can be started for one session (e.g., day 044 of year 2019) using "Menu>BPE>Start BPE processing":



If you follow the tutorial outside from the environment of the *Bernese Introductory Course* you can select RNX2SNX in option "Process control file" and skip all scripts apart from the range between PID 401 SATMRK to PID 442 GNSL12\_P (option "Skip scripts").





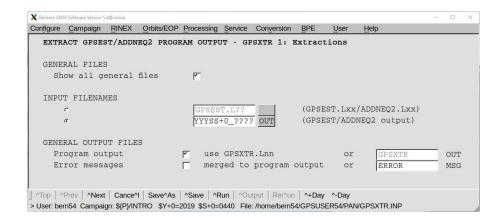
Run the BPE for the current session. If the BPE stops with an error you can inspect the files \${P}/INTRO/BPE/AR190440\_...LOG belonging to your current session. These files report for instance if an input file is missing. This might be the case if you did not follow the naming convention proposed in the tutorial. In that case you have to copy the file from your naming to the expected one.

## 5.3.4 Ambiguity Resolution: Summary

For each observation file a corresponding program output file is generated. To get an overview on the success rate of the various ambiguity resolution steps a one-line summary for each baseline (and optionally for each satellite system) is generated by the BPE in the last step (443 AMBXTR).

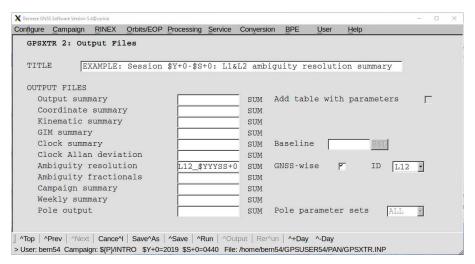
The program GPSXTR ("Menu>Processing>Program output extraction>Parameter estimation/stacking") was used in several configurations. Here we have a look at the extraction of the last step of the ambiguity resolution, direct use of the SIGMA-strategy on  $L_1$  and  $L_2$  frequencies (solution indicator L12):

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The entry for option "GPSEST/ADDNEQ output files" may be incomplete in the screenshot. It should be L12 \$YYYSS+0 ???? (you may check with bottom right to the input field).

All program output files related to the L12 ambiguity resolution method have been specified with program output filenames fitting in the shape \${P}/INTRO/OUT/L12\_\$YYYSS+0\_????..OUT and can, therefore, easily be selected in the input field "GPSEST/ADDNEQ output files".



The resulting protocol in the file \${P}/INTRO/OUT/L12\_20190440.SUM looks like:

File		Sta1	Sta2	Length (km)								
				(KIII)								
WTWZ	0440	WTZR	WTZZ	0.002							#AR_L12	
			WTZZ	0.002					62.5		——————————————————————————————————————	
WTWZ	0440	WTZR	WTZZ	0.002	90	1.1	0	1.3	100.0	Е	#AR_L12	
WTWZ	0440	WTZR	WTZZ	0.002	332	1.1	44	1.3	86.7	GRE	#AR_L12	
ZIZM	0440	ZIM2	ZIMM	0.019	112	1.2	4	1.3	96.4	G	#AR_L12	
Tot:	2			0.010	242	1.2	6	1.3	97.5	G	#AR_L12	
Tot:				0.002					62.5			
Tot:				0.002							#AR_L12	
Tot:	2			0.010	444	1.2	48	1.3	89.2	GRE	#AR_L12	
Legen	d:											
QTR	cons	sider	GPS q	uater cycle							(never)	A (always)
						I	(indi	cated)	)	S	(selected)	
FRQ	GLOI	NASS	resolu	tion between	fre	q.: -	(not	applio	cable)	N	(never)	A (always)
						T	(same	rcvr	. type	) M	(same rcvr. model)	

If you compare the number of ambiguity parameters in the GPSEST program output with the number of ambiguities in the GPSXTR summary files, the number of reference ambiguities that need to be kept unresolved are considered.

Strategy: direct  $L_1/L_2$ :

The solution from the BPE are identical only to one of the manually processed solutions. You can compare the following two files in \${P}/INTRO/OUT for each strategy (e.g., with tkdiff):

#### Filenames from

Manual processing processed by the BPE WTWZ20190441.0UT L12\_20190440\_WTWZ.0UT

The QIF-based ambiguity resolution in the automated processing started from the resolved ambiguities based on the Melbourne-Wübbena strategy whereas in the manual mode the QIF-strategy was applied to the plain files.

Let's have a look at the complete summary file generated by the script in the BPE available in \${P}/INTRO/OUT/AMB\_20190440.SUM. The first part is reporting about the wide-lane/narrow-lane ambiguity resolution starting from the Melbourne-Wübbena linear combination:

ile	Sta1	Sta2	Length	Bef	ore	Aft	er	Res	Sys		
			(km)	#Amb	(mm)	#Amb	(mm)	(%)			
			707.462								#AR
TL0440	BRST	TLSE	707.462	46	1.4	15	1.5	67.4	Е		#AR
RTL0440	BRST	TLSE	707.462 707.462	115	1.4	30	1.5	73.9	G E		#AR_V
AMI0440	GANP	MIKL	897.345 897.345 897.345	51	0.8	4	0.8	92.2	G		#AR_W
AMI0440	GANP	MIKL	897.345	28	0.8	5	0.8	82.1	Е		#AR_W
AMIO440	GANP	MIKL	897.345	79	0.8	9	0.8	88.6	G E		#AR_WI
AWZ0440	GANP	WTZZ	543.530	56	1.0	9	1.1	83.9	G		#AR_WI
WZ0440	GANP	WTZZ	543.530	40	1.0	2	1.1	95.0	E		#AR_WI
WZ0440	GANP	WTZZ	543.530 543.530 543.530	96	1.0	11	1.1	88.5	G E		#AR_WI
_											
ZI0440	WTZZ	ZIM2	475.909	54	1.1	6	1.2	88.9	G		#AR_WI
ZZI0440	WTZZ	ZIM2	475.909	37	1.1	2	1.2	94.6	E		#AR_WI
ZI0440	WTZZ	ZIM2	475.909 475.909 475.909	91	1.1	8	1.2	91.2	G E		#AR_WI
t: 14			760.073	957	1.3	274	1.4	71.4	G		#AR_WI
t: 8			809.713	269	1.1	39	1.2	85.5	E		#AR_WI
: 14			760.073	1226	1.3	313	1 4	7/ F	0 E		#AR WI
Code-Base	d Na:	====== rrowla: ======	======= ne (NL) Amb	===== iguity =====	Reso	lution	:==== . (<60 :====	===== 00 km	==== ) =====		
Code-Base ====== File	d Nam	=====: rrowlar =====: Sta2	======= ne (NL) Amb ====================================	===== iguity ====== Bef #Amb	Reso ===== ore (mm)	lution ====== Aft #Amb	er (mm)	===== 00 km ====== Res (%)	===== ) ===== Sys	==== ====	=====
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
ode-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
ode-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
ode-Base	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys	  	======================================
de - Base ======= ile  RTL0440 RTL0440 AMI0440 AMI0440 AMI0440 AWZ0440 AWZ0440	d Name ====:	rrowlan	Length	===== iguity ===== Bef #Amb	Reso re (mm)	lution ====== Aft #Amb	er (mm)	======================================	===== ) ===== Sys		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
ode-Base ======= ile  RTL0440 RTL0440 AMI0440 AMI0440 AMI0440 AWZ0440 AWZ0440	d Name = = = = = = = = = = = = = = = = = = =	TLSE TLSE TLSE MIKL MIKL MIKL WTZZ WTZZ	Length (km)  707.462 707.462 707.462 897.345 897.345 543.530 543.530 543.530	Bef #Amb 69 46 115 52 37 89 56 40 96	Reso  ore (mm) 1.1 1.1 1.1 1.1 1.1 0.8 0.8	Aft #Amb 15 15 30 5 14 19 10 2	er (mm)  1.2 1.2 1.1 1.1 1.0 1.0	Res (%) 78.3 67.4 73.9 90.4 62.2 78.7 82.1 95.0 87.5	Sys GEGEGEGEGE		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
ode-Base ======= ile  RTL0440 RTL0440 RTL0440 AMI0440 AMI0440 AWZ0440 AWZ0440	d Name = = = = = = = = = = = = = = = = = = =	TLSE TLSE TLSE MIKL MIKL MIKL WTZZ WTZZ	Length (km)  707.462 707.462 707.462 897.345 897.345 543.530 543.530 543.530	Bef #Amb 69 46 115 52 37 89 56 40 96	Reso  ore (mm) 1.1 1.1 1.1 1.1 1.1 0.8 0.8	Aft #Amb 15 15 30 5 14 19 10 2	er (mm)  1.2 1.2 1.1 1.1 1.0 1.0	Res (%) 78.3 67.4 73.9 90.4 62.2 78.7 82.1 95.0 87.5	Sys GEGEGEGEGE		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
ode-Base ======= ile  RTL0440 RTL0440 RTL0440 AMI0440 AMI0440 AWZ0440 AWZ0440	d Name = = = = = = = = = = = = = = = = = = =	TLSE TLSE TLSE MIKL MIKL MIKL WTZZ WTZZ	Length (km)  707.462 707.462 707.462 897.345 897.345 543.530 543.530 543.530	Bef #Amb 69 46 115 52 37 89 56 40 96	Reso  ore (mm) 1.1 1.1 1.1 1.1 1.1 0.8 0.8	Aft #Amb 15 15 30 5 14 19 10 2	er (mm)  1.2 1.2 1.1 1.1 1.0 1.0	Res (%) 78.3 67.4 73.9 90.4 62.2 78.7 82.1 95.0 87.5	Sys GEGEGEGEGE		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
ode - Base ======= ile  RTL0440 RTL0440 AMI0440 AMI0440 AMI0440 AWZ0440 AWZ0440 AWZ0440  ZZI0440 ZZI0440	d Na: ====: Sta1 BRST BRST BRST GANP GANP GANP GANP GANP GANP WTZZ WTZZ	Sta2  TLSE TLSE TLSE MIKL MIKL MIKL WIZZ WIZZ WIZZ ZIM2 ZIM2 ZIM2	Length (km)	======iguity ====================================	Reso ===== ore (mm)  1.1 1.1 1.1 1.1 1.1 0.8 0.8 0.8	#Amb Aft #Amb 15 15 15 19 10 2 12	er (mm)  1.2 1.2 1.1 1.1 1.0 1.0 1.1 1.1	======================================	Sys GEGEGEGEGEGEGEGEGEGEGEGEGEGEGEGEGEGEGE		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
ode-Base ======= ile ==========================	d Name = = = = = = = = = = = = = = = = = = =	TLSE TLSE TLSE TLSE MIKL MIKL WTZZ WTZZ WTZZ ZIM2 ZIM2 ZIM2	Length (km)  707.462 707.462 707.462 897.345 897.345 543.530 543.530 543.530 475.909 475.909	=====iguity ====== Bef #Amb 69 46 115 52 37 89 56 40 96 54 37 91	===== Reso ===== ore (mm) 1.1 1.1 1.1 1.1 1.1 0.8 0.8 0.9 0.9 0.9	#Amb  Aft #Amb  15 15 30 5 14 19 10 2 12	er (<60 cmm)  1.2 1.2 1.1 1.1 1.0 1.0 1.0 1.0 1.1 1.1 1.1 1.1	======================================	Sys  G		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI
de-Base  1e  1TL0440  TTL0440  TTL0440  MI0440  MI0440  MI0440  MI0440  MZ0440  MZ0440  MZ0440  1ZI0440  2ZI0440	d Name = = = = = = = = = = = = = = = = = = =	TLSE TLSE TLSE TLSE MIKL MIKL WTZZ WTZZ WTZZ ZIM2 ZIM2 ZIM2	Length (km)	=====iguity ====== Bef #Amb 69 46 115 52 37 89 56 40 96 54 37 91	===== Reso ===== ore (mm) 1.1 1.1 1.1 1.1 1.1 0.8 0.8 0.9 0.9 0.9	#Amb  Aft #Amb  15 15 30 5 14 19 10 2 12	er (<60 cmm)  1.2 1.2 1.1 1.1 1.0 1.0 1.0 1.0 1.1 1.1 1.1 1.1	======================================	Sys  G		#AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI #AR_NI

In particular for the Melbourne-Wübbena linear combination we observe a higher resolution rate for Galileo than for GPS. The reason is the lower noise level of code measurements obtained from the Galileo system. GLONASS is not included because it is not supported for the Melbourne-Wübbena linear combination with Version 5.4 of *Bernese GNSS Software*.

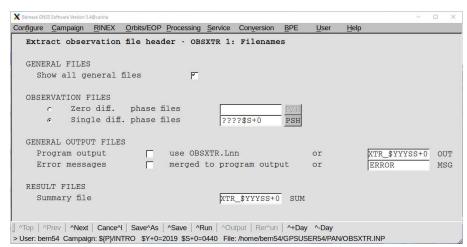
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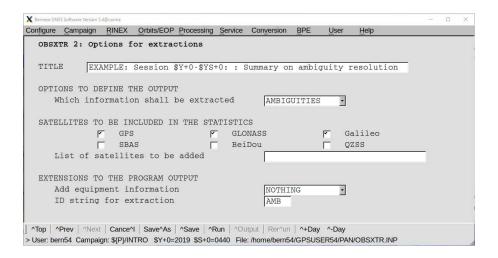
The next block of the ambiguity resolution summary is empty because no baselines for the related lengths are available in the example dataset. The next section is related to the QIF-strategy which is only applied to those ambiguities that are not resolved so far. For this reason the ambiguity resolution rate for GPS and Galileo are lower than it might be expected:

		e (QIF) Am	hianit	_						
				ty Res	olutio	on (<2	000 ki	n)		
			=====							
Sta1	Sta2	Length	Bet	fore	Aft	er	Res	Sys		
		(km)	#Amb	(mm)	#Amb	(mm)	(%)			
										#AR_QIF
										#AR_QIF
										#AR_QIF
										#AR_QIF
						1.2	60.0	G		#AR_QIF
O GANP	MIKL	897.345	96	1.2	38	1.2	60.4	R		#AR_QIF
O GANP	MIKL	897.345	28	1.2	6	1.2	78.6	Е		#AR_QIF
O GANP	MIKL	897.345	134	1.2	48	1.2	64.2	GRE		#AR_QIF
O GANP	WTZZ	543.530	20	1.0	14	1.0	30.0	G		#AR_QIF
O GANP	WTZZ	543.530	100	1.0	34	1.0	66.0	R		#AR_QIF
O GANP	WTZZ	543.530	4	1.0	2	1.0	50.0	E		#AR_QIF
O GANP	WTZZ	543.530	124	1.0	50	1.0	59.7	GRE		#AR_QIF
O WTZZ	ZIM2	475.909	12	1.1	10	1.2	16.7	G		#AR_QIF
O WTZZ	ZIM2	475.909	94	1.1	28	1.2	70.2	R		#AR_QIF
O WTZZ	ZIM2	475.909	4	1.1	4	1.2	0.0	E		#AR_QIF
O WTZZ	ZIM2	475.909	110	1.1	42	1.2	61.8	GRE		#AR_QIF
4		760.073	622	1.1	468	1.1	24.8	G		#AR_QIF
3		774.297	1532	1.1	714	1.1	53.4	R		#AR_QIF
8		809.713	148	1.1	56	1.2	62.2	Е		#AR_QIF
4		760.073	2302	1.1	1238	1.1	46.2	GRE		#AR_QIF
	O BRST O BRST O BRST O GANP O GANP O GANP O GANP O GANP O GANP O WTZZ O WTZZ	3 8	(km)  0 BRST TLSE 707.462 0 BRST TLSE 707.462 0 BRST TLSE 707.462 0 BRST TLSE 707.462 0 GANP MIKL 897.345 0 GANP MIKL 897.345 0 GANP MIKL 897.345 0 GANP MIKL 897.345 0 GANP WIZZ 543.530 0 GANP WTZZ 543.530 0 WTZZ ZIM2 475.909	(km) #Amb  0 BRST TLSE 707.462 30 0 BRST TLSE 707.462 114 0 BRST TLSE 707.462 174 0 BRST TLSE 707.462 174 0 GANP MIKL 897.345 10 0 GANP MIKL 897.345 96 0 GANP MIKL 897.345 28 0 GANP MIKL 897.345 134 0 GANP WTZZ 543.530 20 0 GANP WTZZ 543.530 100 0 GANP WTZZ 543.530 120 0 GANP WTZZ 543.530 124 0 GANP WTZZ 543.530 124 0 GANP WTZZ 543.530 124 0 WTZZ ZIM2 475.909 12 0 WTZZ ZIM2 475.909 94 0 WTZZ ZIM2 475.909 110	(km) #Amb (mm)  0 BRST TLSE 707.462 30 1.3  0 BRST TLSE 707.462 114 1.3  0 BRST TLSE 707.462 30 1.3  0 BRST TLSE 707.462 174 1.3  0 BRST TLSE 707.462 174 1.3  0 GANP MIKL 897.345 10 1.2  0 GANP MIKL 897.345 28 1.2  0 GANP MIKL 897.345 28 1.2  0 GANP MIKL 897.345 134 1.2  0 GANP WTZZ 543.530 20 1.0  0 GANP WTZZ 543.530 100 1.0  0 GANP WTZZ 543.530 4 1.0  0 GANP WTZZ 543.530 124 1.0  0 WTZZ ZIM2 475.909 12 1.1  0 WTZZ ZIM2 475.909 4 1.1  0 WTZZ ZIM2 475.909 110 1.1  4 760.073 622 1.1  3 774.297 1532 1.1  8 809.713 148 1.1	(km)         #Amb (mm)         #Amb           0         BRST TLSE         707.462         30         1.3         28           0         BRST TLSE         707.462         114         1.3         36           0         BRST TLSE         707.462         174         1.3         78           0         BRST TLSE         707.462         174         1.3         78           0         GANP MIKL         897.345         10         1.2         4           0         GANP MIKL         897.345         28         1.2         6           0         GANP MIKL         897.345         134         1.2         48           0         GANP MIKL         897.345         134         1.2         48           0         GANP MIKL         897.345         134         1.2         48           0         GANP WIZZ         543.530         20         1.0         14           0         GANP WIZZ         543.530         4         1.0         2           0         GANP WIZZ         543.530         124         1.0         50           0         WIZZ         ZIM2         475.909         12         1.1	(km)         #Amb (mm)         #Amb (mm)           0         BRST TLSE         707.462         30         1.3         28         1.3           0         BRST TLSE         707.462         114         1.3         36         1.3           0         BRST TLSE         707.462         174         1.3         78         1.3           0         BRST TLSE         707.462         174         1.3         78         1.3           0         GANP MIKL         897.345         10         1.2         4         1.2           0         GANP MIKL         897.345         96         1.2         38         1.2           0         GANP MIKL         897.345         28         1.2         6         1.2           0         GANP MIKL         897.345         134         1.2         48         1.2           0         GANP WIZZ         543.530         20         1.0         14         1.0           0         GANP WIZZ         543.530         100         1.0         34         1.0           0         GANP WIZZ         543.530         124         1.0         50         1.0           0         WIZZ	(km)         #Amb         (mm)         #Amb         (mm)         (%)           0         BRST         TLSE         707.462         30         1.3         28         1.3         6.7           0         BRST         TLSE         707.462         114         1.3         36         1.3         68.4           0         BRST         TLSE         707.462         130         1.3         14         1.3         53.3           0         BRST         TLSE         707.462         174         1.3         78         1.3         55.2           0         GANP         MIKL         897.345         10         1.2         4         1.2         60.0           0         GANP         MIKL         897.345         96         1.2         38         1.2         60.4           0         GANP         MIKL         897.345         134         1.2         48         1.2         64.2           0         GANP         MITZZ         543.530         20         1.0         14         1.0         30.0           0         GANP         WTZZ         543.530         100         1.0         34         1.0         66.0	(km) #Amb (mm) #Amb (mm) (%)  0 BRST TLSE 707.462 30 1.3 28 1.3 6.7 G  0 BRST TLSE 707.462 114 1.3 36 1.3 68.4 R  0 BRST TLSE 707.462 30 1.3 14 1.3 53.3 E  0 BRST TLSE 707.462 174 1.3 78 1.3 55.2 GRE  0 GANP MIKL 897.345 10 1.2 4 1.2 60.0 G  0 GANP MIKL 897.345 96 1.2 38 1.2 60.4 R  0 GANP MIKL 897.345 28 1.2 6 1.2 78.6 E  0 GANP MIKL 897.345 134 1.2 48 1.2 64.2 GRE  0 GANP MIKL 897.345 134 1.2 48 1.2 64.2 GRE  0 GANP WTZZ 543.530 20 1.0 14 1.0 30.0 G  0 GANP WTZZ 543.530 100 1.0 34 1.0 66.0 R  0 GANP WTZZ 543.530 4 1.0 2 1.0 50.0 E  0 GANP WTZZ 543.530 124 1.0 50 1.0 59.7 GRE  0 WTZZ ZIM2 475.909 12 1.1 10 1.2 16.7 G  0 WTZZ ZIM2 475.909 14 1.1 28 1.2 70.2 R  0 WTZZ ZIM2 475.909 110 1.1 42 1.2 61.8 GRE  4 760.073 622 1.1 468 1.1 24.8 G  3 774.297 1532 1.1 714 1.1 53.4 R  8 809.713 148 1.1 56 1.2 62.2 E	(km)         #Amb         (mm)         #Amb         (mm)         (%)            0         BRST         TLSE         707.462         30         1.3         28         1.3         6.7         G            0         BRST         TLSE         707.462         114         1.3         36         1.3         68.4         R            0         BRST         TLSE         707.462         174         1.3         78         1.3         55.2         GRE            0         GANP         MIKL         897.345         10         1.2         4         1.2         60.0         G            0         GANP         MIKL         897.345         96         1.2         38         1.2         60.4         R            0         GANP         MIKL         897.345         96         1.2         38         1.2         60.4         R            0         GANP         MIKL         897.345         134         1.2         48         1.2         64.2         GRE            0         GANP         WTZZ         543.530         20

The next section in the ambiguity resolution summary is related to the direct ambiguity resolution applying the SIGMA-strategy directly on the  $L_1$  and  $L_2$  observations for very short baselines. The same summary results as we have obtained above manually.

This multiple-step ambiguity resolution makes it difficult to get an overall overview on the success-rate of the ambiguity resolution. Such an overview can be provided by the program OBSXTR ("Menu>Service>Bernese observation files>Extract statistics from header file") that evaluates the header of the phase observation files on the relation of resolved ambiguities





The program output and a satellite-wise summary are attached to the ambiguity resolution summary  $\PP/INTRO/OUT/AMB_20190440.SUM$  as well:

					numb.	ambiguity	res	solution	sta	tistic		
File												
1		TLSE	10003 M009	GPS	31	70	15	78 6%	15	78 69	 (	± Δ Ν
1	BRST 10001M004	TLSE	1000311000	GIO	22	58	19	67 2%	19	67 29		# A N
1	BRST 10001M004	TLSE	1000311000	GAT.	24	47	8	83 0%	8	83 0%		# A 1
1	BRST 10001M004	TLSE	1000311000	tot	77	175	42	76.0%	42	76 0%		# A 1
2	GAND 11515M001	MIKI	12335 MOO1	CPS	31	53	3	94 3%	3	94 39	,	# A 1
2	CAND 11515M001	MIKL	12335 MOO1	מוט	22	40	20	54.5%	20	59.37	 ,	# A I
2	CAND 11515M001	MIKL	12335 MOO1	CAI	22	38	20	89.5%	20	80 5%	 '	# A 1
2	GAND 11515M001	MIKL	12335 1001	GAL	75	140	07	09.5%	07	00.07	 ,	# A
3	CAND 11515M001	UT77	14201 M014	CDC	20	140 E7	21	00.1%	21	00.17	,	# A
3	GAND 11515M001	WIZZ	142011014	GFD	00	51	10	64.7%	10	64.79	,	# A.
	GAND 11515MUU1	WIZZ	14201 M 0 1 4	GLU	22	51	10	04.7%	10	04.77	,	# A
3	GANP 11515M001	WIZZ	14201M014	GAL	24	41	2	95.1%	2	95.1/	,	# A
3	GANP 11515M001	WTZZ	14201M014	tot	78	149	28	81.2%	28	81.27		# A.
 15	UT77 1/201M01/	7 T M O	1/001 M008	CDG	3.0	5.5	6	80 1%	6	80 19	,	# A 1
15	WT77 14201H014	7 T M 2	140011000	GLD	22	48	15	68 84	15	68 89	,	# A
15	WIZZ 14201H014	7 T M O	1400111000	CAT	24	20	10	00.0%	10	00.07	,	# A
15	WIZZ 14201M014	ZIMO	140011008	GAL	70	30	3	92.1%	3	92.1/	,	# A
	WIZZ 14201M014	ZIMZ	14001 M000	tot	70	141	24	03.0%	24	03.0/	,	# A
16	ZIM2 14001M008	ZIMM	14001M004	GPS	32	57	3	94.7%	3	94.77	,	# A
16	WTZZ 14201M014 WTZZ 14201M014 WTZZ 14201M014 WTZZ 14201M014 ZIM2 14201M008 ZIM2 14001M008	ZIMM	14001M004	tot	32	57	3	94.7%	3	94.77		# A
	files do contair			GPS	32	1130 846	266	76.5%	266	76.5%	 	
	files do contair			GLO	23	846	403	52.4%	403	52.4%	, , , , , , , , , , , , , , , , , , ,	
	files do contair	- 1		GAT.	24	367 2343	47	87.2%	47	87.29		
	files do contain						- '	0 270		0,		
 >> CP1	 U/Real time for p <sub>{</sub>	sm "OBSX	TR": 0:00				716	69.4%	716	69.47	6	
>> CPI >> Pro		gm "OBSX ccessful	TR": 0:00 ly .ution Sta	: :00.023 / ======== tistics	0:00:	 00.023	===:			69.4%	<b></b>	
>> CPI >> Pro ===== Satel	U/Real time for pg ogram finished sud =========== lite-wise Ambiguit	gm "OBSX ccessful	TR": 0:00 ly 	:00.023 /	/ 0:00:	00.023	===:			69.4%		
>> CPN >> Pro ====== Satel: ===== RN	U/Real time for pg ogram finished sud  ===================================	gm "OBSX ccessful ====== y Resol ====== 2 ambigu #/clu	TR": 0:00 ly	:00.023 /	/ 0:00:	00.023	===:			69.4%	6	
>> CPI >> Pr( ===== Satel: ===== RN	U/Real time for pg ogram finished suc  ===================================	gm "OBSX ccessful ======= gy Resol ======= 2 ambigu #/clu	TR": 0:00 ly  Lution Sta  Lities   rel	:00.023 /	0:00:	00.023	===:			69.4%	6	
>> CPI >> Pr( ===== Satel: ===== RN	U/Real time for pg ogram finished suc  ===================================	gm "OBSX ccessful ======= gy Resol ======= 2 ambigu #/clu	TR": 0:00 ly  Lution Sta  Lities   rel	:00.023 /	0:00:	00.023	===:			69.4%		
>> CPN >> Pro ====== Satel: ====== RN	U/Real time for pg ogram finished sud  ===================================	gm "OBSX ccessful ====== y Resol ====== 2 ambigu #/clu  7.7 12.2	TR": 0:00 ly ution Sta ities   rel   65.2   80.6	:00.023 / tistics L5 an solved	/ 0:00: / 0:00: // 0:00: // 0:00: // 0:00: // 0:00: // 0:00: // 0:00:	00.023  ies rel 39.8 66.3	===:			69.4%		
>> CPN >> Pro ====== Satel: ====== RN	U/Real time for pg ogram finished sud  ===================================	gm "OBSX ccessful ====== y Resol ====== 2 ambigu #/clu  7.7 12.2	TR": 0:00 ly ution Sta ities   rel   65.2   80.6	:00.023 / tistics L5 an solved	/ 0:00: / 0:00: // 0:00: // 0:00: // 0:00: // 0:00: // 0:00: // 0:00:	00.023  ies rel 39.8 66.3	===:			69.4%	6	
>> CPP >> Pro ====== Satel: ====================================	U/Real time for pg ogram finished suc  ===================================	gm "OBS) ccessful ====== cy Resol ====== 2 ambigu #/clu	TR": 0:00 ly	:00.023 / ======= tistics ====== Solved 16 29 15	7 0:00: 	00.023  ========  ies     rel 39.8 66.3 41.5 0.0	===:			69.4%		
>> CPN >> Pro Satel1 =====: RN    +- 1   2   3   4   5	U/Real time for pg ogram finished suc  ===================================	gm "OBS) ccessful ====== cy Resol ====== 2 ambigu #/clu	TR": 0:00 ly	:00.023 / ======= tistics ====== Solved 16 29 15	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6	===:			69.47		
>> CPT >> Pro ====================================	U/Real time for pg ogram finished suc  ===================================	gm "OBS) ccessful ====== cy Resol ====== 2 ambigu #/clu	TR": 0:00 ly	:00.023 / ======= tistics ====== Solved 16 29 15	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7	===:			69.4%		
>>> CPP >>> Pr ===== Satel: ====== 1	Mmb   L1 & L2 total   solved   33   21   6   0   35   30   44   34   31   27	gm "OBS% ccessful "y Resol "y	TR": 0:00 ly 	:00.023 / tistics L5 am solved 16 29 15 0 24 28 19	/ 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7	===:			69.47		
>>> CPN >>> Pro ====== Satel: ====================================	U/Real time for pg ogram finished suc  ===================================	gm "OBS% ccessful "y Resol "y	TR": 0:00 ly 	:00.023 / tistics L5 am solved 16 29 15 0 24 28 19	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7	===:			69.47		
>> CPI >> Pri =====: Satel: =====: N      +- 1   2   3   4   5   6   7   8	U/Real time for pg ogram finished sud sud sud sud sud sud sud sud sud su	gm "OBS) ccessful =======  y Resol ======  2 ambigu #/clu  7.7 12.2 8.8 0.0 20.2 14.3 13.9 12.6	TR": 0:00 ly	:00.023 / tistics tistics L5 an solved 29 15 0 24 28 19 34	"0:00: ""0:00: """"""""""""""""""""""""	00.023  ===================================	===:			69.47		
>> CPI >> Pri ===== Satel: ===== RN      +- 1   2   3   4   5   6   7   8   	Mmb   L1 & L2 total   solved   Si   Si   Si   Si   Si   Si   Si   S	gm "OBS) ccessful yy Resol 2 ambigu #/clu 7.7 12.2 8.8 0.0 20.2 14.3 13.9 12.6	TR": 0:00 ly  Lution Sta  Lution Sta  rel    65.2    80.6    56.4    0.0    81.5    71.9    80.9    78.9	:00.023 /  tistics  L5 an solved  16 29 15 0 24 28 19 34	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7 63.9	===:			69.47		
	### April	gm "OBS) ccessful yy Resol 2 ambigu #/clu 7.7 12.2 8.8 0.0 20.2 14.3 13.9 12.6 16.5 18.5	TR": 0:00 ly  Lution Sta  Lution Sta  rel    65.2    80.6    56.4    0.0    81.5    71.9    80.9    78.9	:00.023 /  tistics  L5 an solved  16 29 15 0 24 28 19 34	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7 63.9	===:			69.47		
>> CPT >> Pro >> Pro	### April	gm "OBSY ccessful "Sy Resolution "A To	TR": 0:00 ly  Lution Sta  Lution Sta  rel    65.2    80.6    56.4    0.0    81.5    71.9    80.9    78.9	:00.023 /  tistics  L5 an solved  16 29 15 0 24 28 19 34	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7 63.9	===:			69.47		
>> CPT >> Pr >> Pr	### April	gm "OBSX ccessful "Sy Resolution "A To	TR": 0:00 ly  Lution Sta  Lution Sta  rel    65.2    80.6    56.4    0.0    81.5    71.9    80.9    78.9	:00.023 /  tistics  L5 an solved  16 29 15 0 24 28 19 34	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7 63.9	===:			69.47		
	### April	gm "OBSY ccessful "Sy Resolution "A To	TR": 0:00 ly	:00.023 / ====================================	7 0:00: 	00.023  ies rel 39.8 66.3 41.5 0.0 64.6 59.7 56.7 63.9	===:			69.47		

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101	50 l	41	3.8	60.3	0	0.0	0.0	
102	55	45	3.2	56.4	0	0.0	0.0	
103	36 I	22	3.4	43.1	0	0.0	0.0	
104	30 I	17	4.1	42.9	0	0.0	0.0	
105	28 I	16	3.9	42.3	0	0.0	0.0	
107	30	25	3.3	58.3	0	0.0	0.0	
122	24 I	18	3.3	52.3	0	0.0	0.0	
123	39	28	3.8	52.7	0	0.0	0.0	
124	45 I	30	4.0	49.9	0	0.0	0.0	
126	2	2	2.0	50.0	0	0.0	0.0	
201	21	18	18.2	81.0	14	16.0	62.5	
202	12	12	26.9	96.3	10	20.3	79.2	
203	16	15	15.7	87.8	11	17.5	64.8	
204	17	17	29.0	96.5	15	16.4	82.9	
230	17	16	36.3	91.5	12	17.6	66.6	
231	16	13	9.0	72.2	8	11.0	45.5	
233	21	20	14.3	88.6	13	18.1	58.5	
236	9	9	24.7	95.9	8	18.8	84.2	
+-	+-			+				
GPS	1130	928	14.5	76.5	732	14.9	60.4	
GLO	846 I	616	3.6	52.4	0	0.0	0.0	
GAL I	367	342	15.5	87.2	242	20.2	62.7	

Please note, that the program OBSXTR does only evaluate the header of the observation files after the ambiguity resolution, whereas the program GPSXTR inspects the program output files from the specific GPSEST run, in which the ambiguity resolution is executed. Due to this the program GPSXTR has access to more information related to reference ambiguities due to user settings that are considered in the statistics. This detailed information is not available to OBSXTR. As a consequence the statistics provided by these two programs will slightly differ.

# 5.4 Daily Goals

At the end of today's session, you should have:

- used GPSEST for residual screening, created files: EDT\_20190440. OUT, EDT\_20190440. RES in your campaign's OUT directory,
- 2. screened the residual files from the above run using RESRMS: created files RMS\_20190440.SUM, RMS\_20190440.LST, RMS\_20190440.EDT, and RMS\_20190440.OUT,
- 3. used SATMRK to mark the identified outliers,
- 4. used GPSEST for a first coordinate and troposphere estimation, created files: FLT\_20190440. CRD and FLT\_20190440. TRP,
- 5. used GPSEST for QIF ambiguity resolution, created files: BRTL2019044Q.
  OUT,
- 6. used GPSEST for direct SIGMA ambiguity resolution, created files: WTWZ20190441.OUT,
- 7. apply the ambiguity resolution to all baselines running a BPE,
- 8. used GPSXTR to create a summary of the ambiguity resolution, created file: QIF\_20190440. SUM and L12\_20190440. SUM.

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# 6 Terminal Session: Thursday

Finish the work of yesterday by resolving the ambiguities for all baselines (day 044 year 2019) if not yet done so far.

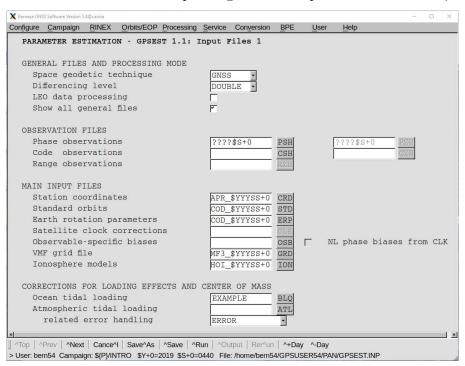
Today's terminal session is to:

- 1. compute a final network solution of the day (GPSEST),
- 2. check the coordinates of the fiducial sites (ADDNEQ2, HELMR1),
- 3. check the daily repeatability (COMPAR2),
- 4. recompute the final solution and generate reduced size normal equation files (ADDNEQ2),
- 5. compute velocities (ADDNEQ2),

for the current session. Compare the final coordinate results of the daily solutions (which are already processed and available).

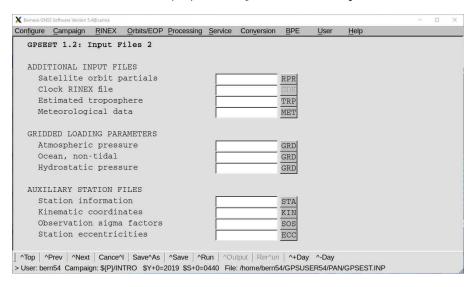
## 6.1 Final Network Solution

The resolved ambiguities may be introduced from the Bernese observation files into the final network solution. To start the program GPSEST in session mode you have to select all single difference files of the corresponding session. In panel "GPSEST 1.1: Input Files 1":

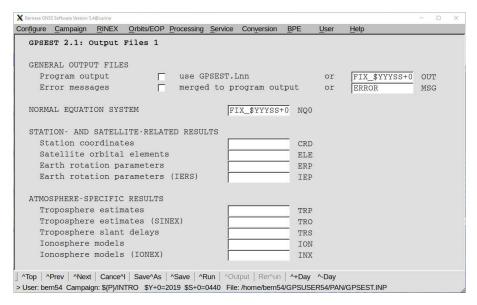


If the grid files for the coefficients of the VMF are available (downloaded from https://vmf.geo.tuwien.ac.at/trop\_products/GRID/1x1/VMF3/VMF3\_OP/ and the five grid files of the day are concatenated), we can introduce VMF3\_\$YYYSS+0 into panel "GPSEST 1.1: Input Files 1".

We do not introduce "Estimated troposphere" anymore from the previous solutions.

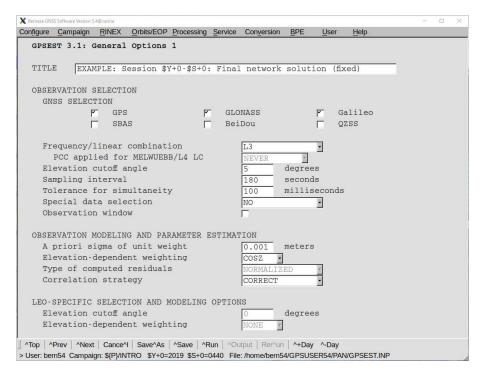


In panel "GPSEST 2.1: Output Files 1" we request the normal equation file as the only output file.

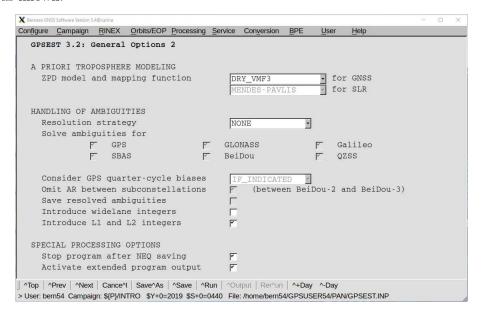


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For the final run of GPSEST we consider the correlations between the observations correctly:

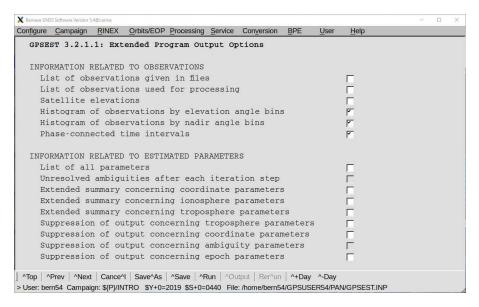


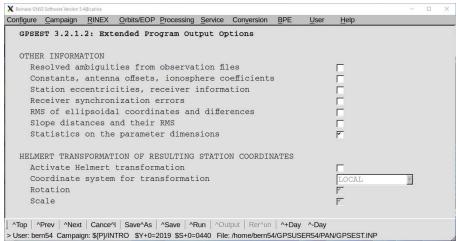
If the VMF grid files are available we can switch now from DRY\_GPT3 to DRY\_VMF3 to use the VMF instead of the Global Mapping Function (GMF). In the *Bernese Introductory Course* environment these files are available. Remember that you need to specify the grid files with the coefficients in the input field "VMF grid file" in panel "GPSEST 1.1: Input Files 1". Ambiguities which have been resolved in the previous runs of program GPSEST are introduced as known.



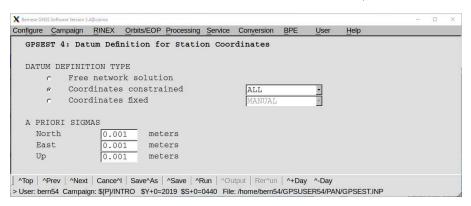
The checkbox in option "Stop program after NEQ saving" reduces the task of GPSEST to setting up the NEQ but not solving it. In particular in case of bigger networks this may save a lot of computing time because the solution for the session will be computed later on by ADDNEQ2 anyhow.

Since this is the final run of GPSEST, it is worthwhile to add some additional information about the observation files into the program output. This is useful if you archive the program output of this run together with the observation files and the resulting normal equation files.



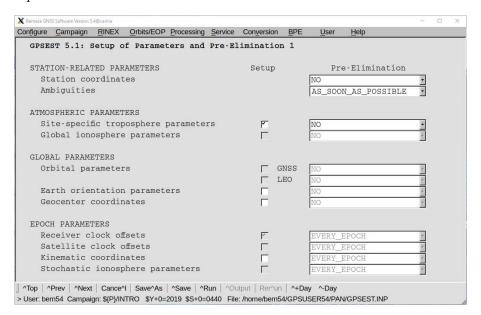


We do not fix any stations on their a priori position, i.e., the coordinates of all stations will be estimated. This retains the flexibility for later changes in the realization of the reference frame (station constraints) with program ADDNEQ2. Because no solution in GPSEST is computed you can select here all types of datum definition apart from "Coordinates fixed" (the normal equations are always stored without any constraints):

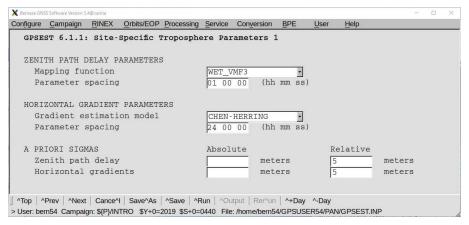


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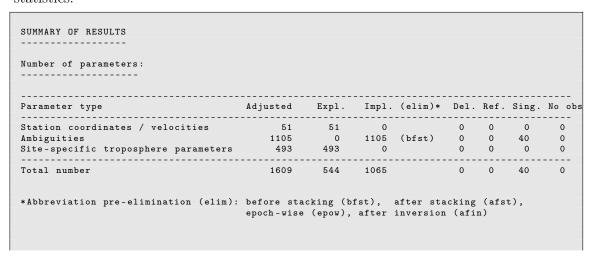
The remaining unresolved ambiguities are pre–eliminated. In addition we may set up additional parameters of interest.



The selection of the mapping function has to be consistent with the selection of the troposphere model in "ZPD model and mapping function (GNSS)" in panel "GPSEST 3.2: General Options 2".



The output of the GPSEST contains only the input parameter and ends with the parameter statistics:



```
Statistics:
------

Total number of authentic observations 133235
Total number of pseudo-observations 0
Total number of observations 133235

Total number of explicit parameters 544
Total number of implicit parameters 1065
Total number of parameters 1609

Total number of observation files 16
Total number of stations 17
Total number of satellites 0

>>> CPU/Real time for pgm "GPSEST": 0:01:54.383 / 0:01:54.399
>>> Program finished successfully
```

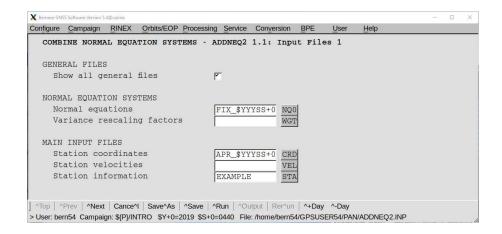
After running GPSEST in session mode the normal equation file  $FIX_20190440.NQ0$  should be available in the directory  $\P$ INTRO/SOL.

In the environment of the *Bernese Introductory Course* these files are provided in the archive \${S}/RNX2SNX/2019/SOL/ \${S}/RNX2SNX/2020/SOL/, and \${S}/RNX2SNX/2021/SOL/ respectively. Copy the files of the additional five days into your campaign. Following files should be now available in the directory \${P}/INTRO/SOL

```
FIX_20190440.NQ0, FIX_20190450.NQ0, FIX_20201790.NQ0, FIX_20201800.NQ0, and FIX_20210950.NQ0, FIX_20210960.NQ0.
```

## 6.2 Check the Coordinates of the Fiducial Sites

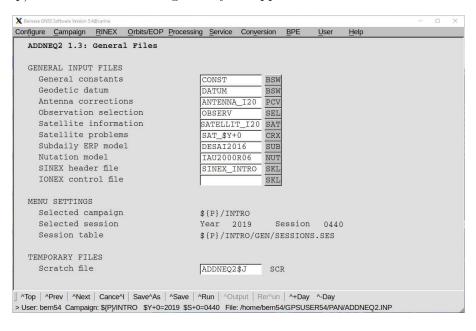
To check the consistency of our network solution with respect to the coordinates available in the IGS 20 reference frame we generate a minimum constraint solution for the network using program ADDNEQ2 ("Menu>Processing>Combine normal equation systems") with the following options:



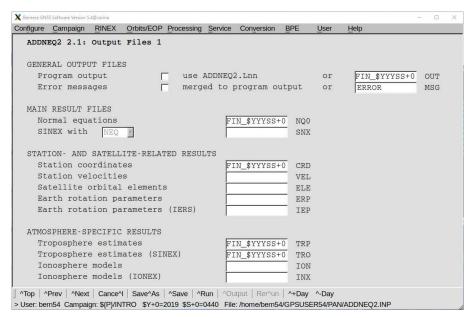
No further input files in the next panel are needed.

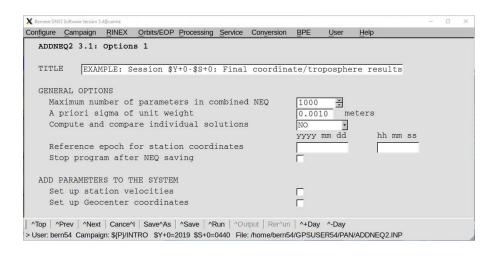
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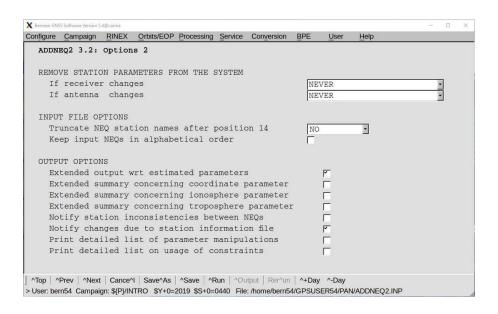
If we intend to store a troposphere Solution INdependent EXchange format (SINEX) file to make the troposphere estimates also available in the international format, we need to add the "SINEX header file". Please adjust the content (in particular the institution and authorship) of this file before using it for your applications.



The troposphere estimates are now written in the Bernese internal (option "Troposphere estimates") and the international format (option "Troposphere estimates (SINEX)"). If there is no need to store both, you may also limit yourself to one of them or even to not store the troposphere results at all.



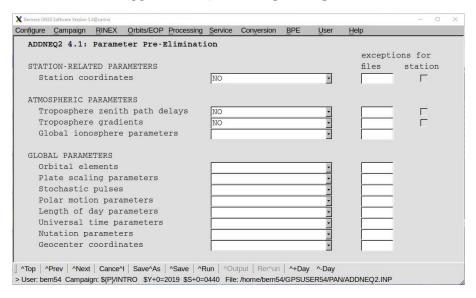




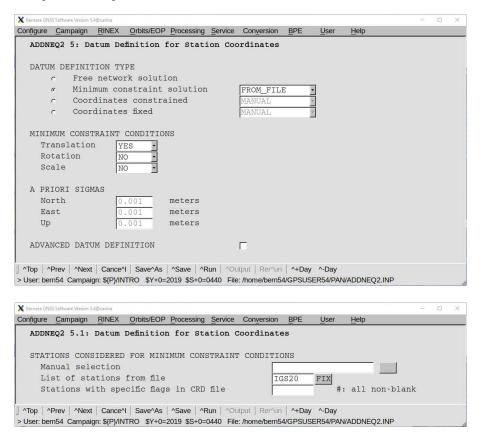
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In the following three panels, all parameter types supported by ADDNEQ2 are listed. You may specify whether a parameter shall be pre-eliminated or not. An empty entry means that the parameter is not expected in the input NEQ files.

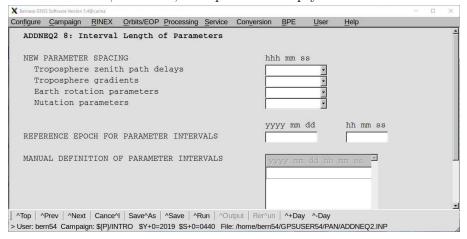
Please note that an automated preselection is not possible for technical reasons. If a parameter with an empty input field is detected in the input NEQ files, the program will stop with an error. In the opposite case, a warning message is issued.



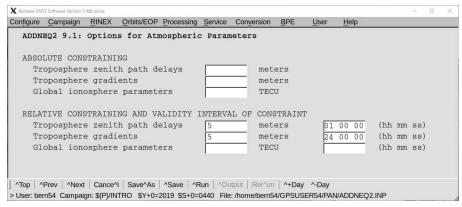
For the validation of the datum stations we choose the "Minimum constraint solution". Because it is a regional network, we only apply the no–translation condition. The other conditions are sufficiently defined by the satellite orbits.



The following panel allows to change the parameter spacing. We do not need this feature at the moment and leave, therefore, all input fields empty.

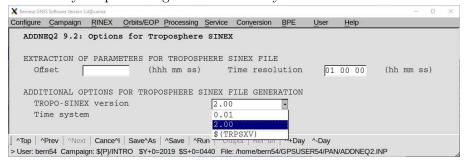


The relative constraints applied to the vertical troposphere and gradient parameters are so loose that they do not affect the solution. As in GPSEST, they simply prevent a format overflow in the output troposphere file in case of very weakly observed parameters due to gaps in the observation scenario.



If you have selected troposphere output in SINEX format you may select the spacing of the sampling of the reported records. Please note that this has no impact on the parametrization of the troposphere parameters.

Furthermore you can decide whether the traditional (0.01) or the latest (2.0) version of the troposphere SINEX format shall be written. In the latter case the nine character station IDs are taken from the related column in PART II of the "Station information" file you have selected in panel "ADDNEQ2 1.1: Input Files 1". Please make sure that you provide for each station in your processing the related entry.



For option "Time system for TROPO-SINEX" we select GPS because the processing takes place in the GPS time system.

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The ADDNEQ2 program output starts with some information about the parameters contained in the input NQO file(s). The input options for the program run follow. An important part is the statistics for the current ADDNEQ2 solution with the same layout you know already from GPSEST:

```
. . .
SUMMARY OF RESULTS
Number of parameters:
Parameter type
                                         Adjusted Expl. Impl. (elim)* Del. Ref. Sing. No obs

      Station coordinates / velocities
      51
      51
      0
      0
      0
      0
      0

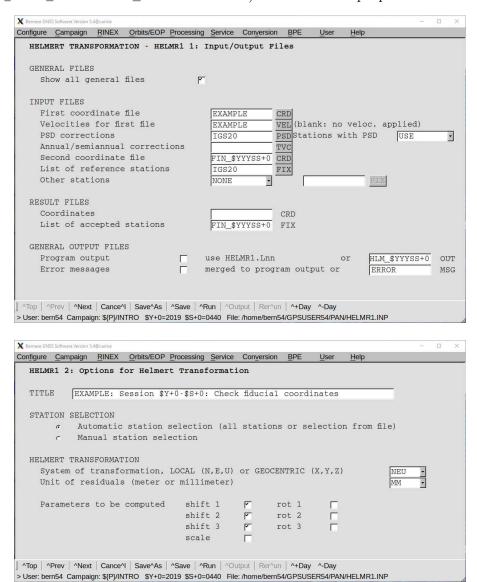
      Site-specific troposphere parameters
      493
      493
      0
      0
      0
      0
      0

Previously pre-eliminated parameters
                                             1065 1065
                                               1609 544 1065
Total number
*Abbreviation pre-elimination (elim): before stacking (bfst), after stacking (afst),
                                           epoch-wise (epow), after inversion (afin)
Statistics:
Total number of authentic observations
                                                 133235
Total number of pseudo-observations
Total number of explicit parameters
                                                    544
                                                  1065
Total number of implicit parameters
Total number of observations
Total number of adjusted parameters
                                               133680
Degree of freedom (DOF)
                                                132071
A posteriori RMS of unit weight
                                               0.001089 m
Chi **2/DOF
                                                   1.19
Total number of observation files
```

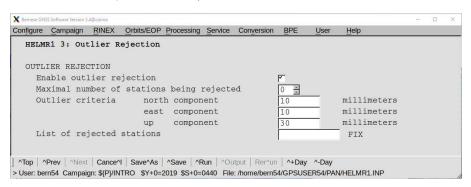
Below this part the program output reports the results of the parameter estimation in the same standardized format for all parameter types as in the program GPSEST:

Julion Coolainate	s and velo	CIUICD.				
Sol Station name	Тур	Correction	Estimated value	RMS error	A priori value	 Abb
1 BRST 10004M004	. Х	0.00152	4231162.41917	0.00047	4231162.41765	 #CF
1 BRST 10004M004	Y	0.00316	-332746.42575	0.00015	-332746.42891	 #CI
1 BRST 10004M004	Z	0.00309	4745131.07277	0.00047	4745131.06968	 #CI
1 GANP 11515M001	X	-0.00773	3929181.29827	0.00041	3929181.30600	 #CI
1 GANP 11515M001	Y	0.00706	1455236.95252	0.00019	1455236.94546	 #CI
1 GANP 11515M001	Z	0.04306	4793654.03600	0.00045	4793653.99294	 # C1
1 HERT 13212M010						
	X	-0.01012	4033460.74185	0.00058	4033460.75197	 # C.
I HERT 13212M010 . Site-specific trop			4033460.74185	0.00058	4033460.75197	 #C1
Site-specific trop	osphere pa	rameters:		RMS error	A priori value	
lite-specific trop	osphere pa Typ	rameters:		RMS error		 Abl
	osphere pa Typ	rameters:  Correction  0.00009	Estimated value	RMS error	A priori value	Abl  #Tl
Site-specific trop Station name SRST 10004M004 SRST 10004M004	Typ	rameters:  Correction  0.00009	Estimated value	RMS error	A priori value 0.00000	 Abi  #T:
Site-specific trop Station name STATE 10004M004 STRST 10004M004 STRST 10004M004	Typ  N E	rameters: Correction 0.00009 -0.00001	Estimated value 0.00009 -0.00001	RMS error 0.00005 0.00008	A priori value  0.00000 0.00000	 Abi  #T: #T:
Site-specific trop Station name BRST 10004M004	Typ  N E U	rameters: Correction 0.00009 -0.00001 0.07659	Estimated value	RMS error  0.00005 0.00008 0.00093	A priori value  0.00000 0.00000 2.35173	 Abl #Tl #Tl #Tl #Tl

The coordinate solution for the session (\${P}/INTRO/STA/FIN\_20190440.CRD) may be compared with the a priori coordinates of the IGS core sites. The program HELMR1 ("Menu>Service>Coordinate tools>Helmert transformation") is used for this purpose:



The last panel allows even for an automated quality control and manages a potential update of the list of stations used for the datum definition (option "List of accepted stations" in panel "HELMR1 1: Input/Output Files"):



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For our example we get the following output.

```
______
 Bernese GNSS Software, Version 5.4
Program : HELMR1
Purpose : Helmert Transformation
               : ${P}/INTRO
Campaign
Default session: 0440 year 2019
Date : 13-Jan-2024 18:31:56
User name : bern54
EXAMPLE: Session 2019-0440: Check fiducial coordinates
FILE 1: EXAMPLE.CRD: IGS20: coordinate list
FILE 2: FIN_20190440.CRD: EXAMPLE: Session 2019-0440: Final coordinate/troposph
LOCAL GEODETIC DATUM: IGS20
 RESIDUALS IN LOCAL SYSTEM (NORTH, EAST, UP)
 LIST OF REMOVED STATIONS:
OUTLIER CRITERIA:
                                     10.00
                                                10.00
                                                          30.00
| NUM | NAME | FLG | RESIDUALS IN MILLIMETERS | |
   OVERALL RMS/IQR/MAX(3D)
                                                  3.07
                                       3.71
                                                            10.27
                                                                        ORID ...
| ALL | RMS / COMPONENT | | 2.18 | 1.84 | 5.76 | | | ALL | IQR | | | 2.06 | 2.49 | 8.34 | | | ALL | MEAN | | 0.00 | 0.00 | 0.00 | | ALL | MEDIAN | | -0.25 | -0.18 | 1.37 | | ALL | MIN | | -4.87 | -2.80 | -9.50 | | ALL | MAX | | | 3.56 | 3.48 | 7.58 |
                                                   3.07
        OVERALL RMS/IQR/MAX(3D)
                                        3.71
                                                            10.27
                                                                         ORID ...
 NUMBER OF PARAMETERS :
NUMBER OF STATIONS : 10
NUMBER OF COORDINATES : 30
 RMS OF TRANSFORMATION :
                            3.71 MM
BARYCENTER COORDINATES:
LATITUDE :
LONGITUDE :
                            47 48 37.88
                            11 25 53.36
HEIGHT
                                  -64.817 KM
PARAMETERS:
TRANSLATION IN N : TRANSLATION IN E : TRANSLATION IN U :
                                   0.02 +- 1.17
0.03 +- 1.17
-0.00 +- 1.17
                                                          MM
                                  -0.00
NUMBER OF ITERATIONS :
 ACCEPTED STATIONS
REJECTED STATIONS : VERIFIED STATIONS :
>>> CPU/Real time for pgm "HELMR1": 0:00:00.135 / 0:00:00.135
>>> Program finished successfully
```

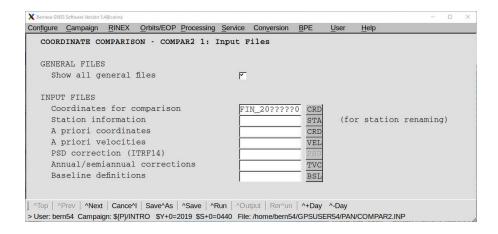
We can conclude that no problems concerning the stations used for the datum definition were detected.

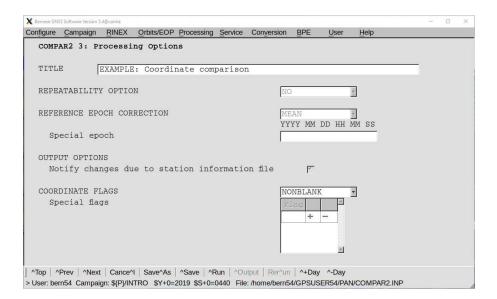
If there were problems, the ADDNEQ2 run needs to be repeated with the problematic station either removed from the file \${P}/INTRO/STA/IGS20.FIX (e.g., by using the result file \${P}/INTRO/STA/FIN\_20190440.FIX) or with manual selection of the stations used for the datum definition in panel "ADDNEQ2 5.1: Datum Definition for Station Coordinates".

In order to check the repeatability of the coordinate solutions for all six days (or make at least sure to have four days available), repeat the above steps for the remaining days.

## 6.3 Check the Daily Repeatability

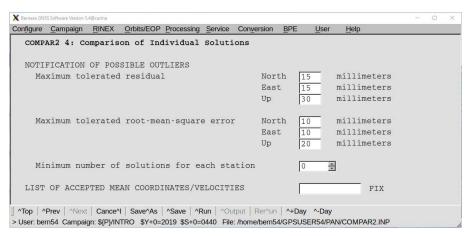
If the minimum constraint solutions of the six sessions are available, the repeatability of the coordinate solutions may be checked using the program COMPAR2 ("Menu>Service >Coordinate tools>Coordinate comparison").





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In the last panel you can define thresholds for detecting stations with exceptionally poor coordinate repeatability.



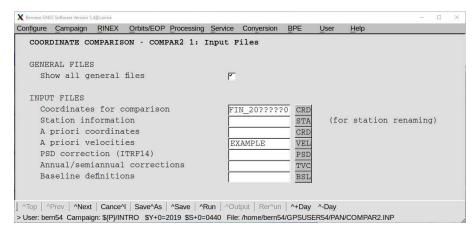
The program computes the arithmetic mean for all station coordinates. The difference of each individual coordinate set to this mean value and the overall RMS are reported in the following section of the program output:

Compa	arison of i	ndivi	dual sol	utions:					
BRST	10004M004	N	15.82			3.09			
BRST	10004M004	E	14.91		-17.54		1.30	15.85	16.15
BRST	10004M004	U	2.96	4.28	3.07	-0.83	-2.16	-2.97	-1.38
GANP	11515M001	N	12.89	-15.43	-15.25	3.04	1.20	13.32	13.13
GANP	11515M001		20.64					19.25	
GANP	11515M001	U	2.41	-0.85	-2.46	-0.94	-1.64	3.78	2.11
HERT	13212M010	N	16.23	-19.80	-19.60	4.40	2.98	15.95	16.07
HERT	13212M010	E	15.93	-17.77	-18.15	-0.08	0.69	17.86	17.45
HERT	13212M010	U	3.04	-17.77 1.09	-0.74	3.16	3.24	-2.60	-4.15
J0Z2	12204M002	N	13.92	-17.39	-16.58	4.37	2.55	13.55	13.49
J0Z2	12204M002		23.96						
J0Z2	12204M002	U	1.61	-0.88	1.03	2.13	0.44	-0.21	-2.51
LAMA	12209M001	N	14.72	-18.18	-18.12	4.97	3.47	13.78	14.08
	12209M001	Е	20.49	-24.67	-25.36	5.26	4.86		
	12209M001		3.21	-4.52	-0.68	-1.70	0.72	5.11	1.08
		-							
MATE	12734M008	N	18.44	-22.42	-21.76	2.79	4.00	18.65	18.74
	12734M008	E		-28.00					
			2.94				-4.75		
	12.01000	·	2.01	2.20	0.00	2.00	11.0	0.0.	2.0.
MTKI.	12335M001	N	11.88	-14.43	-13.46	1.21	1.69	12.49	12.49
	12335M001		24.79					23.55	
				-3.64					
	1200001	·	0.2.	0.01	2.01		0.22	1.00	0.00
ONSA	10402M004	N	15.83	-19.32	-19 47	4.89	3 40	15.27	15 24
				-20.05				15.88	
	10402M004		3.06					4.33	
0.1.0.11	10102001	·	0.00	0.00	0.11	0.10		1.00	2.00
ORTD	15601M001	N	8 56	-9.69	-9 64	-0.36	0.79	0 18	9 72
	15601M001		25.27		_31 42	5.09	6.42	24.31	25 52
				7.10		-2 93	_2 79		
OILLD	1000111001	U	7.11	7.10	2.50	-2.00	-2.13	-5.10	1.27
ртрр	14234M001	N	18.04	-22.12	22 32	1 71	E 11	17 11	17 10
	14234M001								
	14234M001	U	2 37	-21.78 -0.88	2 10	2.51	2.20	3 00	0.77
FIDD	1423411001	U	2.01	-0.00	-2.10	-2.00	2.13	3.00	-0.22
TICE	10003M009	N	15 70	-18.79	_10 /0	3 60	3 42	15.89	15 37
	10003M009	F	15.79	-10.79	-13.43	2.91	2.76	18.63	
				-0.16					
ILDE	1000311009	U	1.13	-0.10	-0.19	-0.73	-0.30	-0.55	2.31
VIII	13406M001	N	15.17	-18.25	18 20	2 47	3 6F	14.85	15 56
	13406M001		16.41		-10.29		2.28		
	13406M001		5.13	7 02	1 70	-4.54	4.20	0.00	-3.35
VILL	13400 MOOI	U	5.13	1.23	4.70	-4.54	-4.94	0.69	-3.35
uan	425068065		46.00	04.64	00 57	4 26	4 00	46.05	45.00
		N	16.92	-21.04 -20.10	-20.57	4.38	4.98	16.95	15.30
	13506M005	E	16.56	-20.10	-19.77	3.01	3.62	15.75	17.50
WSRT	13506M005	U	5.60	-5.75	-3.93	8.59	5.05	-2.85	-1.10

WTZR	14201M010	N	12.14	-10.37	-10.65	11.09	9.93			
WTZR	14201M010	E	15.86	-13.64	-13.83	14.29	13.18			
WTZR	14201M010	U	2.71	-2.73	-1.54	3.36	0.91			
WTZZ	14201M014	N	15.38	-18.60	-18.58	4.08	2.52	14.68	15.91	
WTZZ	14201M014	E	19.81	-23.69	-23.77	4.39	2.85	19.86	20.36	
WTZZ	14201M014	U	3.58	-4.46	-4.57	2.30	1.83	3.63	1.27	
ZIM2	14001M008	N	15.20	-18.20	-18.03	2.60	2.49	14.69	16.46	
ZIM2	14001M008	E	18.98	-23.00	-22.86	4.05	3.93	19.49	18.40	
ZIM2	14001M008	U	1.87	-2.03	-2.07	-0.44	0.76	2.67	1.10	
ZIMM	14001M004	N	15.74	-18.83	-19.01	3.20	2.94	15.27	16.44	
ZIMM	14001M004	Ε	18.69	-23.03	-22.56	4.43	4.73	18.97	17.47	
ZIMM	14001M004	U	2.35	-1.34	-2.58	-0.65	-0.11	4.32	0.37	

In the section Notification of detected outliers all stations from the first (in 2019) and last (in 2021) pair of solutions get listed. While interpreting this output, keep in mind that the the six columns refer to different epochs. The difference between these epochs is about one year. Obviously, station velocities need to be estimated (this will be done in the next step described in Section 6.5).

If reliable velocities for all stations are available they can be introduced:



You may check the influence on the repeatability on your own. Please be reminded, that for the ITRF 2020 (IGS 20) reference frame the linear station velocities are not sufficient. You may see the effect of the PSD corrections in the repeatability. This is not the case for this example because none of the stations is affected by an earthquake and, therefore, no PSD corrections have to be considered.

This output may be used for quality assessment. Stations with a problem in one or more sessions can be identified and excluded from the final ADDNEQ2 solution by adding them to section TYPE 003: STATION PROBLEMS in the station information file (\${P}/INTRO/STA/EXAMPLE.STA). All parameters of these stations will be pre-eliminated before the normal equations are stacked and, therefore, also before the solution is computed.

## 6.4 Compute the Reduced Solution of the Sessions

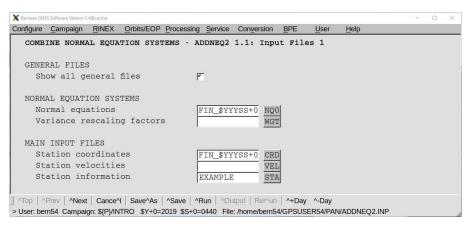
If one or more stations have to be excluded from the session solution or if the datum definition of the solution is still not acceptable, the final solution of the session has to be re-computed by repeating the ADDNEQ2 from Section 6.2. Finally, the result files for the final solution of the session are:

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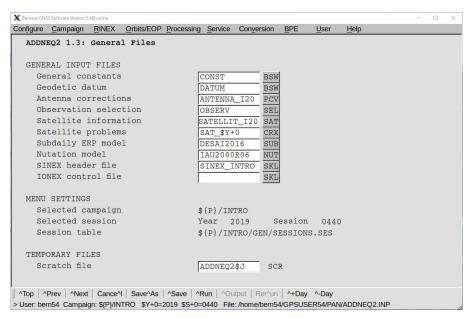
```
${P}/INTRO/SOL/FIN_$YYYSS+0.NQO,
${P}/INTRO/STA/FIN_$YYYSS+0.CRD,
${P}/INTRO/ATM/FIN_$YYYSS+0.TRP, and potentially
${P}/INTRO/ATM/FIN_$YYYSS+0.TRO.
```

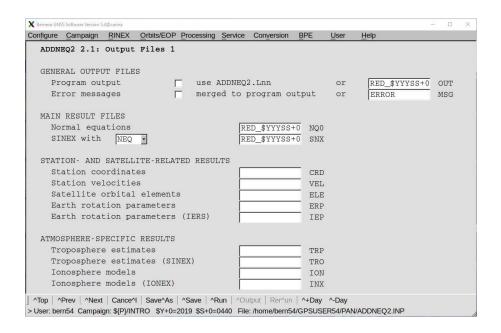
It is preferable for the velocity estimation to have smaller normal equation files containing only the coordinate parameters for each session. In addition, we generate a coordinate SINEX file (in NEQ representation) as the final solution of the day, so the troposphere parameters have to be pre-eliminated before the solution is computed. We introduce the station coordinates (\${P}/INTRO/STA/FIN\_\$YYYSS+0.CRD) obtained with the minimum constraint solution in the previous run of ADDNEQ2 and constrain the solution to these coordinates.

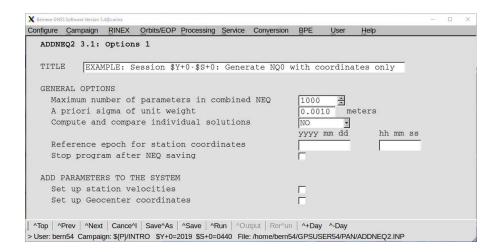
To generate these reduced NQO files and the SINEX file, the execution of ADDNEQ2 has to be repeated with the following changes in the input options:



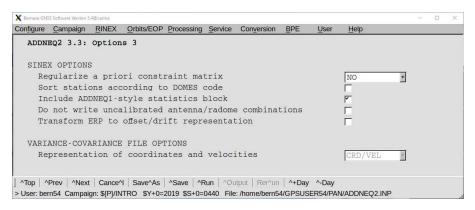
Please note that the output SINEX file includes again the skeleton from \${P}/INTRO/GEN/SINEX\_INTRO (Copied from \${D}/REF54 directory into the campaign) – the same file we used as skeleton for the troposphere SINEX generation.





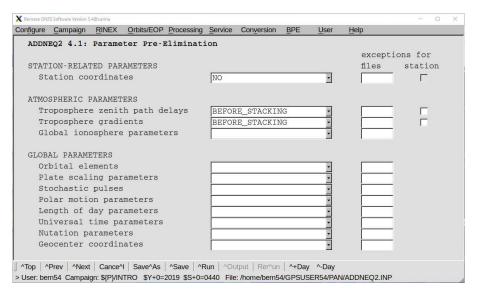


Because you are storing a SINEX file in NEQ representation (see option "Content of SINEX") no regularization is necessary.

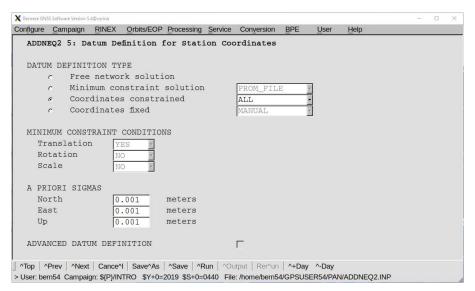


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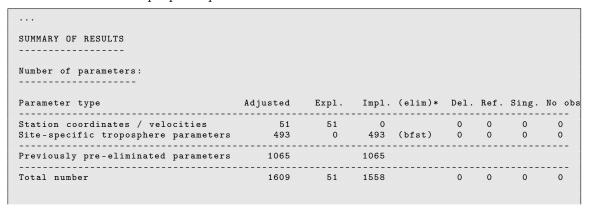
The troposphere parameters are pre-eliminated:



To keep the a priori and estimated sets of coordinates in the resulting SINEX file consistent, we introduce the coordinate solution of the session in "Station coordinates" in "ADDNEQ2 1.1: Input Files 1" and constrain all coordinate parameters to these values.



The normal equation file (\${P}/INTRO/SOL/RED\_20190440.NQ0) contains only the station coordinate parameters. The following section of the program output documents the pre-elimination of the troposphere parameters:



```
Statistics:
Total number of authentic observations
                                                  133235
Total number of pseudo-observations
                                                     493
Total number of explicit parameters Total number of implicit parameters
                                                      51
                                                    1558
Total number of observations
                                                  133728
Total number of adjusted parameters
                                                    1609
Degree of freedom (DOF)
                                                  132119
A posteriori RMS of unit weight
                                               0.001089 m
                                                    1.19
Total number of observation files
                                                      16
Total number of stations
```

You can also see that the number of explicit parameters in the NQO file was dramatically reduced (from 544 to 51). This is an advantage for the combination of a big number of normal equation files for the estimation of station velocities.

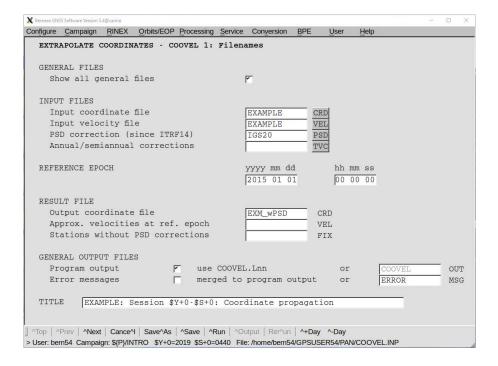
## 6.5 Velocity Estimation

### 6.5.1 Preparation for ITRF 2020/IGS 20 Velocity Estimation

This section can be skipped if no ITRF  $2020/IGS\,20$  reference frame is used as geodetic datum.

Because of the PSD corrections, the linear station velocity may not represent the actual station velocity and one has to prepare the station coordinate and velocity files before they can be used for the datum definition in the program ADDNEQ2. We have to compute the station coordinates at the reference epoch using the program COOVEL ("Menu>Service >Coordinate tools>Extrapolate coordinates"):

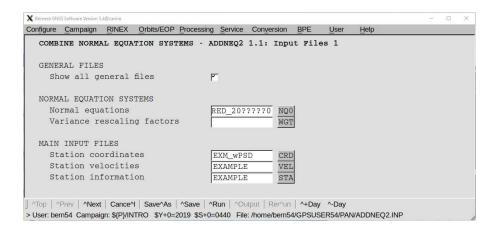
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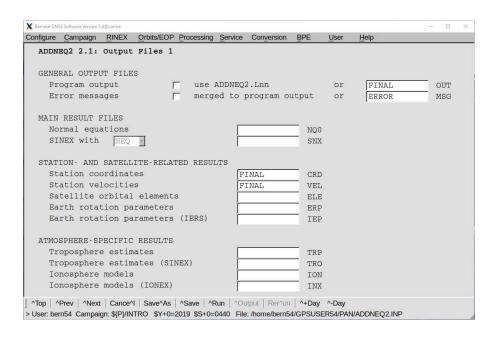


In our example none of the stations that shall be used for the datum definition is related to the PSD corrections. So, you can continue here. If you have another station selection where the PSD corrections become relevant the distributed BPE \${U}/PCF/ITRF.PCF may be helpful. In the related README file you will find further instructions. It can be accessed via "Menu>Help>Readme" (Section Example BPE).

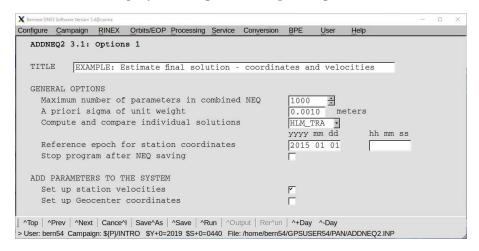
#### 6.5.2 Velocity Estimation Based on NEQ Files

The velocity estimation in program ADDNEQ2 is easy. Introduce the normal equation files containing only the station coordinate parameters. Independent from your success in following the example so far, you may copy the prepared files for all days (044/045 year 2019, 179/180 year 2020, and 095/096 year 2021) from \${S}/RNX2SNX/2019/SOL/RED\_2019\*NQ0, \${S}/RNX2SNX/2020/SOL/RED\_2020\*NQ0, and \${S}/RNX2SNX/2021/SOL/RED\_2021\*NQ0 into the \${P}/INTRO/SOL/ directory. The normal equation files have to cover a reasonable time interval to reliably estimate velocities (in this case about two years):





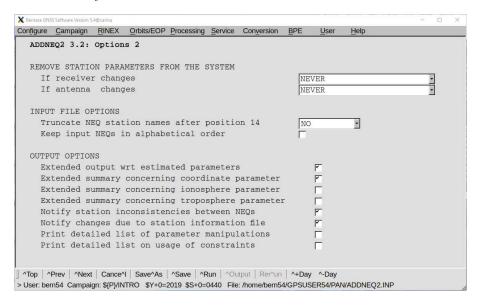
Station velocities are set up by marking the corresponding checkbox:



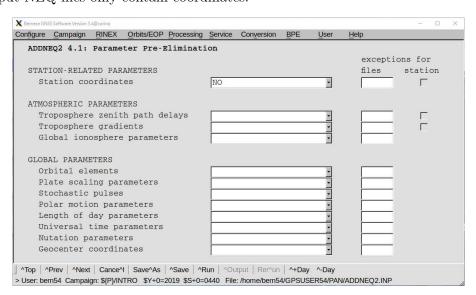
Furthermore, we check the repeatability of the daily solutions after the velocity estimation. The coordinates in the resulting file will refer to the epoch 2015 01 01.

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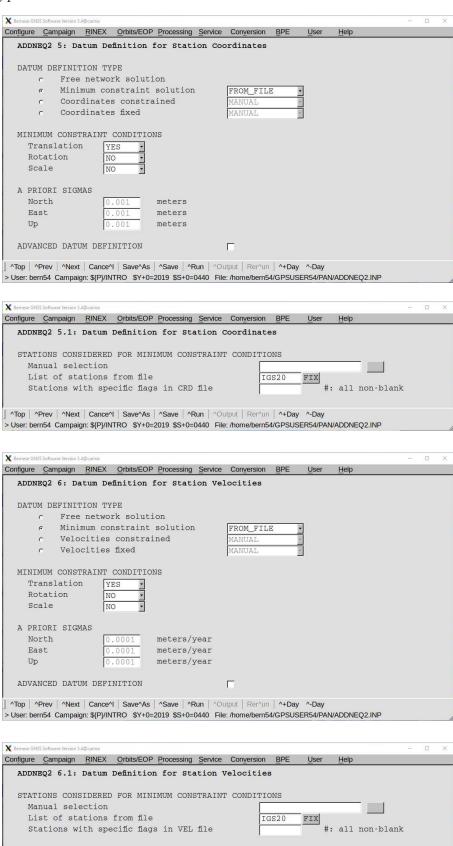
This panel is important for stations where an equipment change has taken place during the interval covered by this solution.



The input NEQ files only contain coordinates:



The realization of the geodetic datum is done separately for positions and velocities in the following panels:

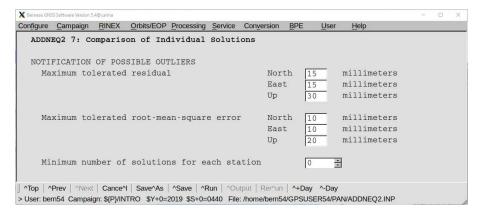


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^Top | ^Prev | ^Next | Cance^1 | Save^As | ^Save | ^Run | ^Output | Rer^un | ^+Day ^-Day

 $\verb| > User: bern54 Campaign: \$[P]/INTRO \$Y+0=2019 \$S+0=0440 File: /home/bern54/GPSUSER54/PAN/ADDNEQ2.INP| | File: /home/bern54/GPSUSER54/$ 

The following panel provides options to detect bad daily solutions based on the repeatability:



After the velocity estimation the repeatability of the coordinate solutions from the individual normal equations look like:

				#1	Observat:	ions		# P	arameters	
		DOF C								
1 0	.00109	132071 134591 147487 147425 138482 136210	1.19	133	3677	3		51	1558	
2 0	.00106	134591	1.13	136	080	3	5	51	1391	
3 0	.00148	147487	2.19	148	3986	3		51	1451	
4 0	.00145	147425	2.09	148	3917	3		51	1444	
5 0	.00110	138482	1.22	139	9733	3	4	18	1206	
6 0	.00111	136210	1.23	13	7543	3	4	18	1288	
omparison		vidual solu								
DGM 40004				0.04	0.07		0.04	0.44		
RST 10004	MOO4 N	0.69	0.38	0.24	-0.67	-0.91	0.94	0.11		
RST 10004	MOOA II	1.34 1.21	0.84	0.26	-1.89	-1.51	1.04	1.26		
ANP 11515	M001 N	0.82	0.28	0.34	-0.06	-1.59	0.69	0.37		
ANP 11515	M001 E	1.41	-0.66	-0.27	2.75	0.12	-1.17	-0.71		
ANP 11515	M001 U	0.82 1.41 1.23	1.29	-0.23	-0.88	-1.88	1.11	0.62		
ED# 42040	MO4O N	0.04	0.00	0.00	0.50	0.45	0.00	0.00		
EKI 13212	MOIO N	0.31	-0.03	0.03	0.52	-0.45	-0.00	-0.09		
ERI 13212 FRT 13919	MO10 E	2.28 2.97	-0.47	-2 19	3 80	-2.35 3.58	-2 20	-2 54		
								-2.04		
0Z2 12204	M002 N	0.57 2.41 1.64	-0.53	0.15	1.00	-0.49	-0.05	-0.30		
0Z2 12204	M002 E	2.41	-1.58	-1.22	4.34	0.83	-1.73	-1.57		
OZ2 12204	M002 U	1.64	-2.06	-0.05	2.51	0.57	-0.43	-1.56		
ΔΜΔ 12209	MOO1 N	0.78	-0.26	-0 34	1 45	0.28	-0.62	-0.54		
AMA 12209	M001 R	0.76	0.09	-0.34	0.71	0.26	-0.52	-0.46		
AMA 12209	M001 U	0.78 0.66 1.78	-1.52	2.40	-1.80	0.36	1.83	-1.04		
ATE 12734	M008 N	0.72 0.33	0.01	0.54	-1.38	0.12	0.37	0.48		
ATE 12734	M008 E	0.33	-0.10	0.35	-0.26	-0.40	0.42	-0.05		
ATE 12734	MOOR A	2.99	-2.29	3.31	1.87	-4.54	-0.12	2.08		
IKL 12335	M001 N	1.05	0.07	0.95	-1.59	-0.91	0.81	0.77		
IKL 12335	M001 E	1.21	-0.36	-0.66	1.55	1.51	-0.91	-1.10		
IKL 12335	M001 U	1.05 1.21 2.50	-1.21	0.01	-1.49	5.12	-0.69	-0.93		
NSA 10402	M004 N	0.56	-0.09	-0.38	1.08	0.00	-0.15	-0.51		
NSA 10402	M004 E	0.40 1.82	-U.35 2 19	-0.55	-0.29	0.00	1 36	0.44		
RID 15601	M001 N	1.50	0.91	0.88	-2.44	-1.00	0.84	1.42		
RID 15601	M001 E	1.50 0.92	0.60	-1.02	-0.45	1.42	-0.76	0.13		
RID 15601	M001 U	2.14	2.53	-1.50	-1.68	-1.85	-0.24	2.82		
TBB 14234	M001 N	0.96	-0.71	-1.04	0.38	1.50	-0.49	-0.64		
TBB 14234	M001 E	0.84 1.73	0.67	1.01	-0.55	-1.15	0.57	0.37		
TER 1/12/	M001 U	1 73	0.33	-0.81	-2.37	2.57	1.03	-1.01		

otal										14.96587	
Stati	ion coordina	ates			All			3.86	8857	14.96587	3
	ion coordina				Horizont	al				18.82593	2
Stati	ion coordina	ates			U/vertication	al		2.69	179	7.24576	1
Stati	ion coordina	ates			E			5.23	3596	27.41528	1
Stati	ion coordin	tes			N			3.19	946	10.23658	1
aran	meter type				Componen	t Sca	aling fac	tor wrt	RMS /	variance	D
/aria	ance-covari	ance	scaling f	actors							
IMM	14001M004	U	1.01	0.81	-0.33	-0.44	-0.29	1.75	-0.98		
	14001M004		0.83	-0.47	-0.07	0.44	1.17	0.38	-1.24		
	14001M004	N	0.50	0.32	0.00	-0.54	-0.42	-0.25	0.80		
11112	1400111008	U	0.30	-0.01	0.03	-0.21	0.04	0.11	-0.25		
	14001M008 14001M008		0.41		-0.01 0.05		0.29 0.54				
	14001M008	N			0.36		-0.73				
	14201M014	U	1.39				1.32				
	14201M014	E			0.03			0.17			
T77	14201M014	N	0.52	0 15	0.03	0 44	-0.75	-0 42	0.65		
TZR	14201M010	U	1.91	0.01	1.29	-0.24	-3.04				
	14201M010	E			-0.16						
ITZR	14201M010	N	0.37								
, DIL I	1550011005	U	0.40	-3.30	-1.00	0.00	4.13	-0.02	-2.00		
	13506M005 13506M005	E U	0.63 5.48	-0.12 -3.56			0.44 4.79	-0.83			
	13506M005	N	0.98		-0.11						
	13406M001	Ū			0.09		-4.01				
	13406M001	N E	0.89		0.07		-0.80				
	13406M001	N	0.60	0.02	0.07	1 02	0.55	0 00	0 63		
TLSE	10003M009	U	1.28	0.33	-0.19	-0.44	-0.51	-1.91	1.99		
	10003M009	E	0.62		0.37		-0.67				
	10003M009	N	0.32		-0.37			0.24	-0.32		

Below this table, bad daily solutions according to the settings in panel "ADDNEQ2 7: Comparison of Individual Solutions" are summarized (if there are any). In this example we have no additional section and, therefore, no outliers.

Directly below the Solution Statistics potential changes of the equipment are reported that might influence the velocity estimation. In our example, four stations are affected:

Stati	ion inconsi:	stencies:							
C+2+i	ion	First obs	enoch		noch		er type	Antenna type	
			еросп		. росп	116661	er type	• •	
JOZ2	12204M002	2019-02-13	00:00:00	2019-02-14	23:59:30	LEICA	GRX1200GGPRO	LEIAT504GG	
JOZ2	12204M002	2020-06-27	00:00:00	2021-04-06	23:59:30	TRIMBL	E NETR9	TRM59800.00	
ONSA	10402M004	2019-02-13	00:00:00	2019-02-14	23:59:30	JAVAD	TRE_G3TH DELTA	AOAD/M_B	
ONSA	10402M004	2020-06-27	00:00:00	2021-04-06	23:59:30	SEPT F	OLARX5TR	AOAD/M_B	
ORID	15601M001	2019-02-13	00:00:00	2019-02-14	23:59:30	LEICA	GRX1200GGPRO	LEIAT504GG	
ORID	15601M001	2020-06-27	00:00:00	2021-04-06	23:59:30	LEICA	GR30	LEIAR25.R4	٠.
PTBB	14234M001	2019-02-13	00:00:00	2019-02-14	23:59:30	ASHTEC	H Z-XII3T	ASH700936E	
PTBB	14234M001	2020-06-27	00:00:00	2021-04-06	23:59:30	SEPT F	OLARX5TR	LEIAR25.R4	
VILL	13406M001	2019-02-13	00:00:00	2019-02-14	23:59:30	SEPT F	OLARX4	SEPCHOKE_MC	
VILL	13406M001	2020-06-27	00:00:00	2021-04-06	23:59:30	SEPT F	OLARX5	SEPCHOKE_B3E6	6

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If you compare th	e velocities obtain	ned for the two	sites in Kötzting	(WTZR and WTZZ)
and Zimmerwald	(ZIM2 and ZIMM	) you will find	small differences:	

Station coordinate	s and vel	locities:				
Reference epoch: 2	015-01-01	00:00:00				
Station name	Тур	A priori value	Estimated value	Correction	RMS error	
TZR 14201M010	VX	-0.01593			0.00028	
	VY	0.01721		0.00044	0.00012	
	٧Z	0.01006	0.01349	0.00343	0.00031	
	1777	0.00004	0.00475	0.00470	0.00010	
	VU	-0.00004	0.00475	0.00479	0.00040	
	VN	0.01543			0.00013	
	VE	0.02033	0.02001	-0.00032	0.00010	
 NTZZ 14201M014	VX	-0.01593	-0.01426	0.00167	0.00020	
	VY	0.01721	0.01772	0.00107	0.00020	
	VZ	0.01721	0.01273	0.00267	0.00022	
	• •	0.02000	0.012.0	0.0020.	0.00022	
	VU	-0.00004	0.00312	0.00316	0.00029	
	VN	0.01543	0.01585	0.00043	0.00009	
	VE	0.02033	0.02045	0.00012	0.00007	
ZIM2 14001M008	VX	-0.01378		0.00092	0.00020	
	VY	0.01810	0.01806	-0.00004	0.00008	
	٧Z	0.01164	0.01179	0.00015	0.00021	
	VU	0 00076	0.00150	0.00073	0 00007	
	VU	0.00076 0.01621		0.00073 -0.00056	0.00027 0.00009	
	V N VE	0.01621	0.01565	-0.00056	0.00009	
	V E	0.01974	0.01957	-0.00016	0.00007	• • •
ZIMM 14001M004	VX	-0.01378	-0.01318	0.00060	0.00031	
	VY	0.01810	0.01769	-0.00041	0.00011	
	٧Z	0.01164	0.01222	0.00058	0.00031	
	VU	0.00076	0.00156	0.00080	0.00042	
	VN	0.01621	0.01621	0.00000	0.00014	
	VE	0.01974	0.01925	-0.00048	0.00010	

You can constrain the velocity estimates for the pairs of receivers at one location in the station information file. Copy the original station information file \${P}/INTRO/STA/EXAMPLE. STA and add the following lines to part TYPE 004: STATION COORDINATES AND VELOCITIES (ADDNEQ) of this copy.

TYPE 004: STATION	COORDINATES AND VELO	CITIES (ADDNEQ)				
		REL. CONSTR.	POSITION	RELATIVE	CONSTR.	VELOCITY
STATION NAME 1	STATION NAME 2	EAST	UP I	NORTH E	AST	UP
**********	**********	**.****	**.****	**.****	*.****	**.****
WTZR 14201M010	WTZZ 14201M014			0.00001	0.00001	0.00001
ZIM2 14001M008	ZIMM 14001M004			0.00001	0.00001	0.00001

(Pay attention to the number of blank lines before the next section starts.)

When introducing this information, the program ADDNEQ2 will issue the following message:

```
### SR AOPTNET: You are going to use relative constraints for station coordinates/velocities from station info file.

Please keep in mind that you will NOT constrain the estimated results but only the improvements of the apriori values.
```

If only the improvements (column Correction) for the velocities are constrained, you must make sure that also the a priori values (column A priori value) for the velocities are identical to obtain (column Estimated value) the same velocities for a group of stations.

You can verify this in the input velocity file \${P}/INTRO/STA/EXAMPLE.VEL:

370	WTZR 14201M010	-0.01593	0.01721	0.01006	IG20 EURA
371	WTZZ 14201M014	-0.01593	0.01721	0.01006	NNR EURA
383	ZIM2 14001M008	-0.01378	0.01810	0.01164	IG20 EURA
385	ZIMM 14001M004	-0.01378	0.01810	0.01164	IG20 EURA

If this is not the case, you should unify the a priori values.

The relative constraining of the velocity estimates is confirmed in the section of the input parameters (below the a priori coordinates and velocities) of the ADDNEQ2-program output:

```
...

Relative constraints between stations:

... relative constraints for velocities

Station names

... N (m/year) E (m/year) U (m/year)

WTZR 14201M010 WTZZ 14201M014 ... 0.00001 0.00001

ZIM2 14001M008 ZIMM 14001M004 ... 0.00001 0.00001
```

Introducing this modified station information file instead of the original one you will get the following estimates for the station velocities in Kötzting and Zimmerwald:

Station name	Тур					
		A priori value	Estimated value	Correction	RMS error	
VTZR 14201M010	VX	-0.01593	-0.01371	0.00222	0.00016	
	VY	0.01721	0.01768	0.00047	0.00007	
	VZ	0.01006	0.01300	0.00294	0.00018	
	VU	-0.00004		0.00371	0.00023	
	VN	0.01543			0.00008	
	VE	0.02033	0.02029	-0.00004	0.00006	
VTZZ 14201M014	VX	-0.01593	-0.01372	0.00221	0.00016	
	VY	0.01721		0.00047	0.00007	
	٧Z	0.01006	0.01301	0.00295	0.00018	
	VU	-0.00004		0.00371	0.00023	
	VN	0.01543				
	VE	0.02033	0.02029	-0.00004	0.00006	
ZIM2 14001M008	VX	-0.01378	-0.01283	0.00095	0.00016	
	VY	0.01810	0.01793	-0.00017	0.00006	
	٧Z	0.01164	0.01198	0.00034	0.00016	
	VU	0.00076	0.00164	0.00088	0.00021	
	VN	0.01621	0.01577	-0.00044	0.00007	
	VE	0.01974	0.01945	-0.00029	0.00005	
ZIMM 14001M004	VX	-0.01378		0.00095	0.00016	
	VY	0.01810		-0.00017	0.00006	
	٧Z	0.01164	0.01198	0.00034	0.00016	
	VU	0.00076	0.00164	0.00088	0.00021	
	VN	0.01621	0.01578	-0.00044	0.00007	

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The final results are contained in the files \${P}/INTRO/STA/FINAL.CRD

EXAM	PLE: E	stimate fina	al solution - co	ordinates and v	elocities 13	- JAN -24	18:48
T OCA		ETTC DATUM.	 IGS20	EDOCH . 201E	01 01 00.00.00		
LUCA	r GEOD	EIIC DAIUM:	16520	EPUCH: 2015-	01-01 00:00:00		
NUM	CTATT	ON NAME	X (M)	Y (M)	Z (M)	FLAG	
SYST		ON NAME	A (FI)	1 (11)	Z (H)	LLAG	
5151	En						
43	BRST	10004M004	4231162.47226	-332746.49379	4745131.03712	W	GRE
104	GANP	11515M001	3929181.36193	1455236.88350	4793653.99378	A	GRE
122	HERT	13212M010	4033460.79940	23537.97230	4924318.35818	A	GR
143	JOZ2	12204M002	3664880.40509	1409190.73252	5009618.56142	A	GRE
166	LAMA	12209M001	3524522.74510	1329693.77172	5129846.42519	A	GR
184	MATE	12734M008	4641949.37574	1393045.61381	4133287.60971	W	GRE
198	MIKL	12335M001	3698553.79492	2308676.15753	4639769.58706	W	GRE
235	ONSA	10402M004	3370658.41000	711877.28744	5349787.05070	W	GRE
236	ORID	15601M001	4498451.55310	1708267.17400	4173591.97666	W	GRE
264	PTBB	14234M001	3844059.81622	709661.47880	5023129.63889	A	GRE
329	TLSE	10003M009	4627851.70710	119640.21487	4372993.66294	W	GRE
359	VILL	13406M001	4849833.62039	-335048.82494	4116015.07867	A	GRE
369	WSRT	13506M005	3828735.71196	443305.12097	5064884.80160	W	GRE
370	WTZR	14201M010	4075580.37967	931853.96581	4801568.21514	W	GRE
371		14201M014	4075579.25914	931853.28128	4801569.12266	A	GRE
383		14001M008	4331299.73308	567537.50277	4633133.82755	W	GRE
385	ZIMM	14001M004	4331296.92165	567556.05878	4633134.04683	W	G

#### and \${P}/INTRO/STA/FINAL.VEL

EXAM	PLE: Estimate final	solution - coord	linates and velo	cities 13	- JAN - 24	18:48
LOCA	L GEODETIC DATUM: IC	 3S20				
NUM	STATION NAME	VX (M/Y)	VY (M/Y)	VZ (M/Y)	FLAG	PLATE
., 011	DIMITON NAME	VA (11/17)	VI (11/ I/	V2 (11, 1)	I Dire	1 11111
43	BRST 10004M004	-0.01300	0.01632	0.00849	W	EURA
104	GANP 11515M001	-0.01571	0.01686	0.00999	A	EURA
122	HERT 13212M010	-0.01396	0.01617	0.00944	A	EURA
143	J0Z2 12204M002	-0.01979	0.01848	0.00802	Α	EURA
166	LAMA 12209M001	-0.01757	0.01594	0.01088	A	EURA
184	MATE 12734M008	-0.01881	0.01920	0.01427	W	EURA
198	MIKL 12335M001	-0.02005	0.01765	0.00970	W	EURA
235	ONSA 10402M004	-0.01568	0.01386	0.01069	W	EURA
236	ORID 15601M001	-0.01745	0.02122	0.00401	W	EURA
264	PTBB 14234M001	-0.01706	0.01620	0.01195	A	EURA
329	TLSE 10003M009	-0.01135	0.01810	0.01210	W	EURA
359	VILL 13406M001	-0.01205	0.01769	0.00916	A	EURA
369	WSRT 13506M005	-0.01473	0.01548	0.01192	W	EURA
370	WTZR 14201M010	-0.01371	0.01768	0.01300	W	EURA
371	WTZZ 14201M014	-0.01372	0.01768	0.01301	A	EURA
383	ZIM2 14001M008	-0.01283	0.01793	0.01198	W	EURA
385	ZIMM 14001M004	-0.01283	0.01793	0.01198	W	EURA

## 6.6 Daily Goals

At the end of today's session, you should have:

- used GPSEST to compute a final solution of the day, created files: FIX\_ 20190440. OUT, FIX\_20190440. NQO,
- 2. checked the coordinates of the fiducial sites using ADDNEQ2 and HELMR1, created files: FIN\_ 20190440. CRD, FIN\_ 20190440. TRP, optionally also FIN\_ 20190440. TRO, FIN\_ 20190440. OUT, and HLM\_ 20190440. OUT these results should be available for all sessions,
- 3. used COMPAR2 to check the daily repeatabilities, created file:  ${\tt COMPAR.OUT}$ ,
- 4. used ADDNEQ2 to create a final session solution, and reduced size NEQs, created files: RED\_ 20190440. NQO and RED\_ 20190440. SNX,
- 5. if possible, used ADDNEQ2 for velocity estimation, created files: FINAL. CRD and FINAL. VEL.

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## 7 Additional Examples

In the previous terminal sessions you have estimated coordinates, velocities, and troposphere parameters. This is the standard application of the *Bernese GNSS Software* for most users.

If you have finished this work or if you want to follow more examples at home, this section of the document provides some suggestions to practice:

- generation of a combined GPS/GLONASS orbit from IGS legacy product series (see Section 7.1 on page 126),
- kinematic positioning for a station (see Section 7.2 on page 132), and
- simulation of GNSS observations (see Section 7.3 on page 139).

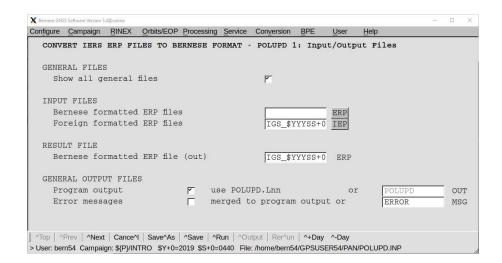
# 7.1 Preparing Combined GPS and GLONASS legacy IGS-Orbits

In this section the differences to the standard procedure using CODE products containing GPS and GLONASS orbits with respect to IGS products are demonstrated. The IGS uses independent combination procedures for GPS and GLONASS orbits resulting in two sets of precise orbit files. That's why the orbits for the two GNSS first need to be merged.

In contrast, CODE (and other AC) uses a rigorous combined multi–GNSS processing scheme, hence producing a single precise orbit file.

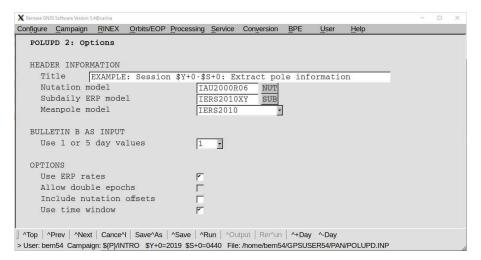
#### 7.1.1 Prepare Pole Information

For the IGS precise orbit files (PRE) the consistent EOP need to be available in the ORB directory (which is the case in the *Bernese Introductory Course* environment). As for the use of CODE products in Section 3.1, the EOP information has to be converted from the IERS/IGS standard format (file extension within the *Bernese GNSS Software* is IEP) to the internal Bernese EOP format (file extension within the *Bernese GNSS Software* is ERP). This is the task of the program POLUPD ("Menu>Orbits/EOP>Handle EOP files>Convert IERS to Bernese Format"). Simply specify IGS—related filenames.



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Because the input "Foreign formatted ERP files" is generated using the modelling of the operational processing within the IGS in that time for the "Subdaily ERP model" as well as the "Meanpole model" we have to switch it back to maintain the compatibility:



Because the IGS "Foreign formatted ERP files" contain only the noon epoch we have to enable the checkbox for option "Use ERP rates" in order to cover the full interval of the session in the resulting "Bernese formatted ERP file (out)".

#### The messages

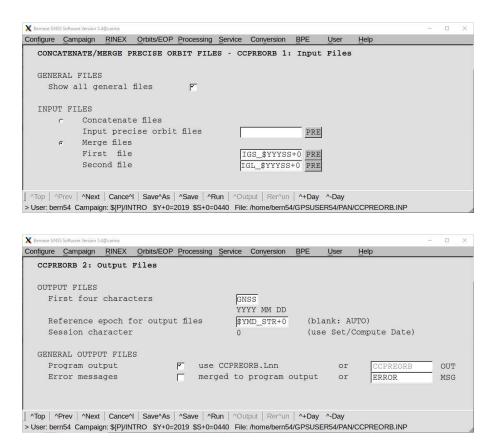
```
### PG POLUPD: NUTATION MODEL NOT SPECIFIED IN INPUT ERP FILE
USING NUTATION MODEL NAME : IAU2000R06

### PG POLUPD: SUBDAILY POLE MODEL NOT SPECIFIED IN INPUT ERP FILE
USING SUBDAILY POLE MODEL NAME : IERS2010
```

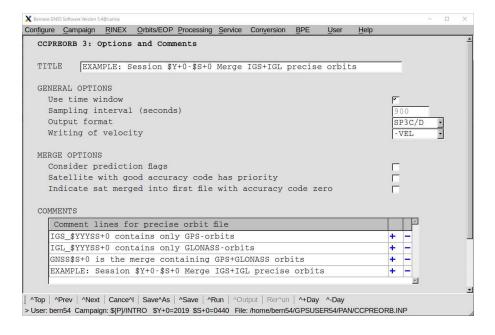
just inform you that the nutation and subdaily pole models from the files in the input panel are written to the output file because no Bernese formatted ERP file was used as input. This is different to importing the EOP from CODE products because here the information on the nutation and sub daily pole model is also available in the international format (with the extension IEP).

#### 7.1.2 Merging Precise Orbit Files

Before we can prepare the orbits from the IGS for a combined GPS+GLONASS processing we need to merge the two separate files IGS15941.PRE and IGL15941.PRE. This is the task of the program CCPREORB ("Menu> $\underline{O}$ rbits/EOP> $\underline{C}$ oncatenate/merge precise orbit files"):



The resulting filename consists of the solution identifier ("First four characters") and the session of the first epoch (if "Reference epoch for output files" is empty). The reference epoch may also be specified by the user as the above example shows. Using the input options from above panel we expect the result file named as \${P}/INTO/ORB/GNSSO440. PRE.

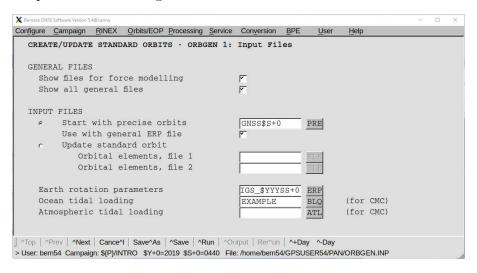


In the next panel you may specify the time window for which the satellite positions shall be included in the resulting precise orbit file.

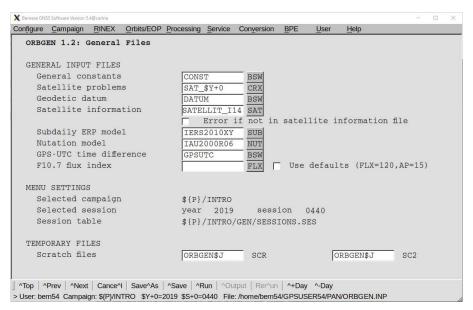
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#### 7.1.3 Generating Standard Orbit Files

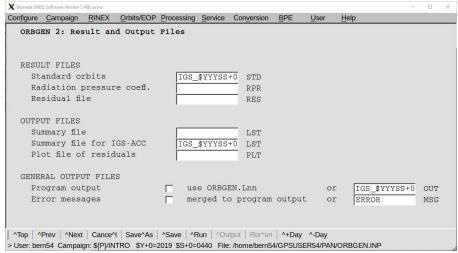
To generate the standard orbits (extension STD) from the merged precise orbits the program ORBGEN ("Menu>Orbits/EOP>Create/update standard orbits") has to be used. Introduce the merged IGS—related precise orbit file together with the consistent ERP file:



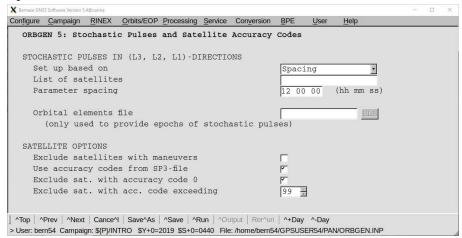
Adjust the entries for the "Satellite information" (including the correct antenna model based on IGS 14) and the "Subdaily ERP model" to be consistent to the introduced "Earth rotation parameters" file:



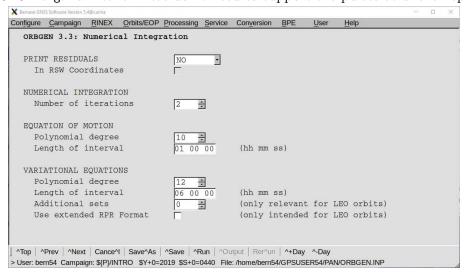
The name of the resulting orbit file should also be related to IGS:



In the general case it is preferable to setup the pulses just in an equidistant spacing by changing option "Setup of stochastic pulses" to Spacing. The interval in "Interval for stochastic pulses" is recommended to be relatively short in order to obtain a "reasonable" representation of the introduced positions in the precise orbit file. In case of IGS orbits one set of pulses at noon is sufficient.



Now you should go back to panel "ORBGEN 3.3: Numerical Integration" and increase the values for options "EQUATION OF MOTION: Length of interval" and "VARIATIONAL EQUATIONS: Length of interval". We don't need to support the pulses at uneven epoch.



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The other options can be used as given in Section 3.2. The resulting program output is expected to look like

RMS	ERRORS	AND MAX. R	ESIDUALS	ARC NUMBI	ER: 1			ITERATI	ON:
			QUADR	ATIC MEAN	OF 0-C	(M)	MAX. F	ESIDUAL	
SAT	#POS	RMS (M)	TOTAL	RADIAL	ALUNG	00T	RADIAL	ALUNG	OUT
1	96	0.002	0.002	0.002	0.002		0.006	0.005	0.00
2	96	0.002	0.002	0.002	0.002	0.001	0.008	0.005	0.00
3	96	0.003	0.003	0.004	0.002	0.003	0.013	0.007	0.00
4	96	0.004	0.003	0.005	0.003	0.002	0.011	0.007	0.00
5	96	0.003	0.003	0.005	0.002	0.001	0.020	0.007	0.00
6	96	0.002	0.002	0.002	0.002	0.002	0.005	0.004	0.00
7	96	0.002	0.002	0.001	0.001	0.002	0.003	0.003	0.00
8	96	0.003	0.003	0.003	0.003	0.002	0.006	0.007	0.00
9	96	0.002	0.002	0.002	0.001	0.002	0.003	0.003	0.00
10	96	0.003	0.003	0.004	0.002	0.002	0.015	0.006	0.00
11	96	0.003	0.002	0.003	0.003	0.002	0.006	0.007	0.00
12	96	0.004	0.004	0.005	0.003	0.002	0.018	0.007	0.00
13	96	0.001	0.001	0.001	0.001	0.001	0.003	0.003	0.00
		0.005	0 004				0 004	0 007	0 00
28	96	0.005	0.004	0.007	0.003	0.002	0.031	0.007	0.00
29	96	0.003	0.003	0.004	0.003	0.002	0.007	0.008	0.00
30	96	0.002	0.002	0.003	0.002	0.002	0.005	0.004	0.00
31	96	0.001	0.001	0.001	0.001	0.001	0.003	0.003	0.00
32	96	0.001	0.001	0.001	0.001	0.001	0.003	0.004	0.00
101	96	0.007	0.007	0.010	0.006	0.003	0.018	0.013	0.00
102	96	0.003	0.003	0.004	0.002	0.002	0.019	0.006	0.00
103	96	0.005	0.005	0.008	0.003	0.002	0.014	0.010	0.00
104	96	0.006	0.006	0.009	0.004	0.003	0.015	0.019	0.00
105	96	0.005	0.005	0.007	0.003	0.003	0.012	0.013	0.00
107	96	0.006	0.006	0.009	0.004	0.003	0.017	0.016	0.00
108	96	0.005	0.005	0.008	0.003	0.002	0.013	0.010	0.00
109	96	0.004	0.004	0.006	0.003	0.003	0.012	0.012	0.00
120	96	0.003	0.003	0.003	0.002	0.004	0.007	0.007	0.01
120	96	0.003	0.003	0.003	0.002	0.004	0.007	0.007	0.01
121	96	0.004	0.004	0.005	0.002	0.003	0.009	0.008	0.00
122	96	0.004	0.004	0.006	0.002	0.003	0.014	0.007	0.00
124	96	0.003	0.003	0.002	0.001	0.004	0.004	0.004	0.01
126	96	0.007	0.006	0.007	0.006	0.007	0.013	0.011	0.01

The RMS error for the orbit fit for precise IGS orbits should be below  $5\,\mathrm{mm}$  (for older orbits it may also achieve  $10...15\,\mathrm{mm}$ ).

The file \${P}/INTRO/OUT/IGS\_2019044.LST contains the same results as displayed on page 31 but contains also the GLONASS satellite orbits:

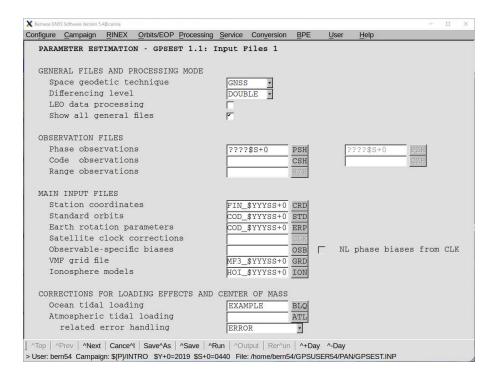
```
EXAMPLE: Session 2019-0440: Standard orbit generation
                                                                    14-JAN-24 05:22
TIME FROM DAY : 3 GPS WEEK: 2040
TO DAY : 4 GPS WEEK: 2040
ORBIT REPEATABILITY FROM A 1-DAY FIT THROUGH DAILY ORBIT SOLUTIONS (MM)
                                   0 M ( 0 EM)
                                  8 ...
                          6
                                          31 32 101 102 103 104 105 107 ... 122 123 124 126
                                                                        6 ...
                                  3 ...
         2
                     3
                          2
                              2
                                               1
                                                       3
                                                                    5
                                                                                        3 6
044
             3
                 3
                                           1
```

If you want to use these orbit files in the processing programs, you have to replace the CODE-related by the IGS-related filenames for the standard orbit and EOP files.

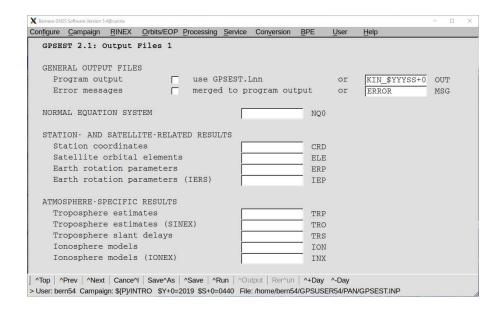
## 7.2 Kinematic Positioning

#### 7.2.1 Estimating Kinematic Positions in a Double-Difference Solution

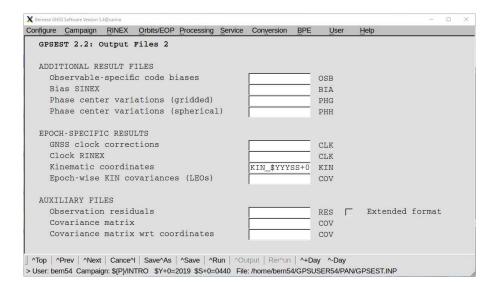
The example campaign contains no really roving stations. You can, however, define one of them to be kinematic (e.g., station GANP). Introduce the coordinates from the final solution (\${P}/INTRO/STA/FIN\_20190440.CRD) for all other sites.



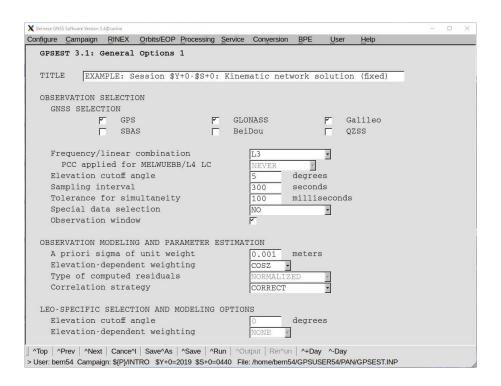
Remove the name of the resulting "Normal equations" file in panel "GPSEST 2.1: Output Files 1" if there is any entry in this input field. Store the kinematic coordinates in an output file ("Output kinematic coordinates" in panel "GPSEST 2.2: Output Files 2").

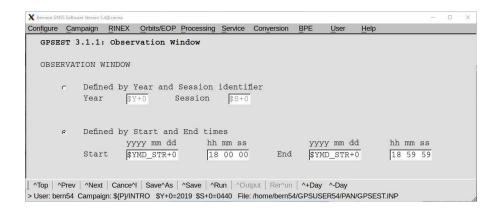


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Because the number of parameters for the kinematic positioning may become very large, we select only a short data interval of one hour for this kinematic positioning:

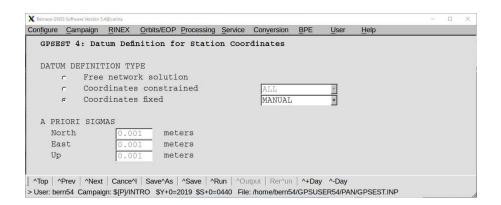




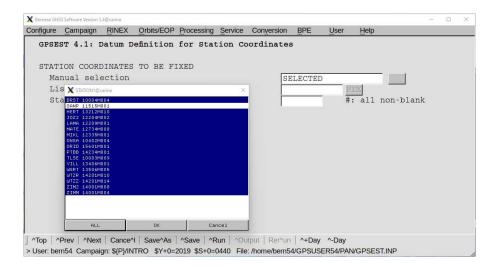
The option "Activate extended program output" may be disabled now:



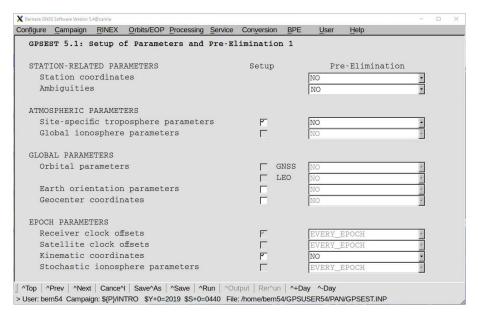
Fix all station coordinates apart from GANP in the panels "GPSEST 4: Datum Definition for Station Coordinates" (choose MANUAL in panel "GPSEST 4: Datum Definition for Station Coordinates" and select all stations except GANP in panel "GPSEST 4.2").



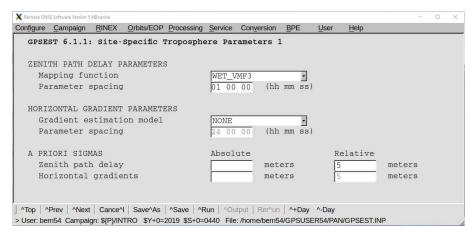
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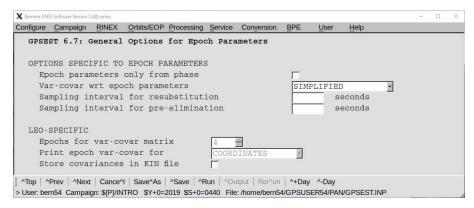
Enable the kinematic coordinates option without any pre-elimination in a first run:



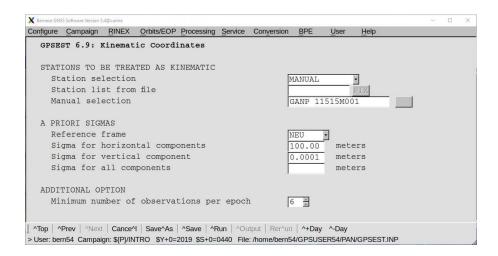
Because of the short analysis interval of only one hour we disable the estimation of troposphere gradients:



An additional panel for options related to epoch parameters is displayed where you can accept the default values:

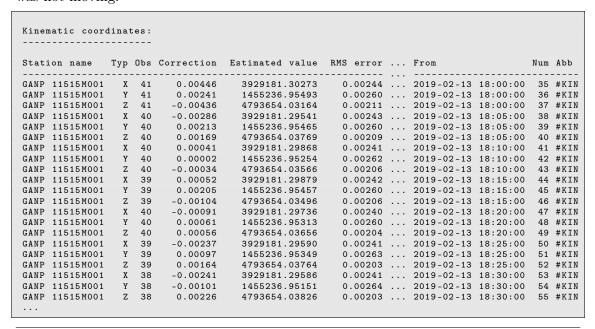


Let us assume only horizontal movements for this site:



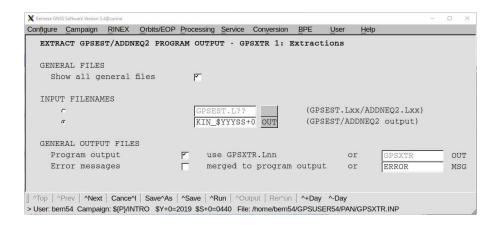
#### 7.2.2 Extracting the Program Output from a Kinematic Positioning

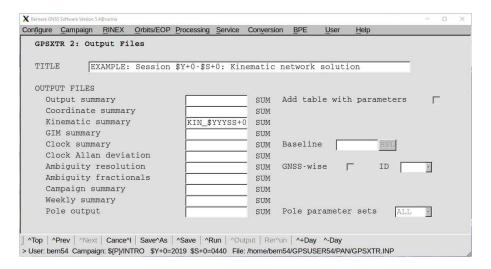
As expected you will get only small estimates for the kinematic coordinates since GANP was not moving:



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With the program GPSXTR ("Menu>Processing>Program output extraction>Parameter estimation/stacking") a comprehensive summary of the estimates for the kinematic solution can be extracted:





The resulting summary file looks like:

*STATION	GANP 11515M00	01		
*XEPO	12			
*INI	3929181.2983	1455236.9525	4793654.0360	
*EST	0.0000	0.0000	0.0000	
*DIFXYZ	0.0006	0.0005	-0.0006	
*DIFNEU	-0.0010	0.0002	-0.0000	
EPOCH	DN	DE	DU	
58527.75000	-0.0066	0.0007	-0.0000	
58527.75347	0.0026	0.0030	0.0000	
58527.75694	-0.0005	-0.0001	-0.0000	
58527.76042	-0.0016	0.0017	0.0000	
58527.76389	0.0009	0.0009	0.0000	
58527.76736	0.0025	0.0017	0.0000	
58527.77083	0.0035	-0.0001	-0.0000	
58527.77431	-0.0026	0.0014	0.0000	
58527.77778	-0.0035	-0.0018	-0.0000	
58527.78125	-0.0042	0.0032	0.0000	
58527.78472	-0.0004	-0.0033	-0.0000	
58527.78819	-0.0017	-0.0045	-0.0000	
 * A V G	-0.0010	0.0002	-0.0000	
*SIG	0.0030	0.0024	0.0000	
*RMS	0.0031	0.0023	0.0000	
*RMSTC	0.0011			

The different components of the summary are described in the online help.

#### 7.2.3 Further suggestions

- Introduce the result file with kinematic coordinates as an input file for another run of GPSEST. If the estimates become zero it is a confirmation that the file was correctly considered as the a priori kinematic positions for the station GANP.
- Use the pre-elimination EVERY\_EPOCH for the "Kinematic coordinates" (they are back—substituted by the program in order to get a solution also for those parameters). Compare the results with the first solution.
- Switch the "Var-covar wrt epoch parameters" in panel "GPSEST 6.7: General Options for Epoch Parameters" from SIMPLIFIED to CORRECT. Compare the results again with the first solution.
- Compute kinematic coordinates for the full day using the epoch—wise pre–elimination and back—substitution algorithm. To save computing power we recommend to sample the data to 300 s.
- Repeat the kinematic solution considering only one of the GNSSs at the time in panel "GPSEST 3.1: General Options 1").

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#### 7.3 Simulation of GNSS Observations

The Bernese GNSS Software provides the simulation tool GPSSIM ("Menu>Service>Generate simulated observation data"). It generates synthetic GNSS (i.e., GPS, GLONASS, Galileo, BDS, and QZSS) observations for terrestrial static or kinematic stations as well as for LEOs. Code and phase zero difference observation files can be created based on an observation scenario defined by

- GNSS satellite geometry given by a standard orbit and
- a set of receivers with positions from a coordinate file, kinematic positions, or a LEO standard orbit file.

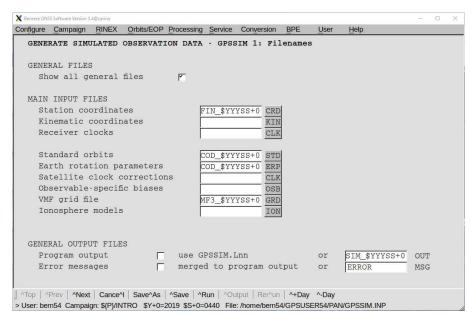
#### 7.3.1 Simulation of GNSS Observations

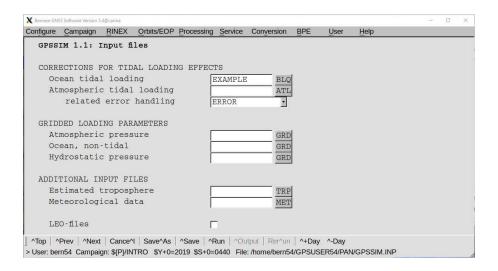
It is important that you remove all previously existing observation files for this session from the OBS directory of your campaign before you start to simulate observations. Otherwise you run into the danger of mixing your current set of simulated observations with other measurements:

```
bern54@carina:~ > rm ${P}/INTRO/OBS/????0440.CZ?
bern54@carina:~ > rm ${P}/INTRO/OBS/????0440.PZ?
```

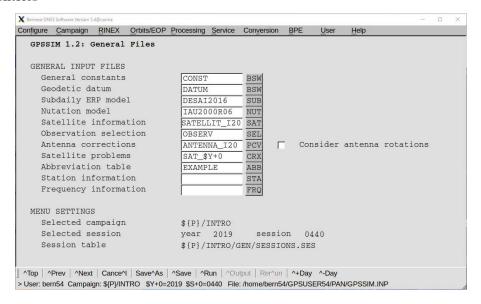
Please keep in mind that the observation files from your previous work in this campaign are lost due to this command. If you still need them, please copy them away.

The input files for the generation of synthetic GNSS observations with program GPSSIM ("Menu>Service>Generate simulated observation data") are defined in the first input panel:

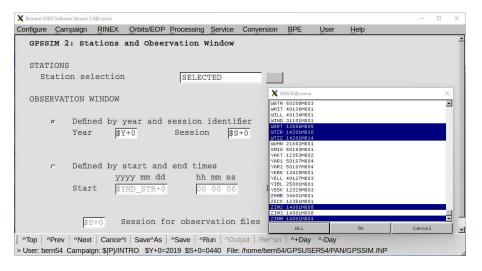




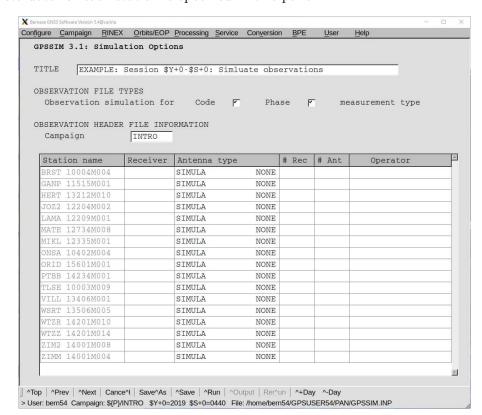
Select EXAMPLE.ABB in option "Abbreviation table". Take care on the consistency for the other entries



In the next panel, the interval for data simulation is defined and the list of stations selected from all sites in the input "Coordinate file":



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The meta data for each station is specified in this panel:

The "Receiver" for each station must start with SIMULA to indicate that this is a simulated station. In the input field you may only extend this string by a user input.

The selection of the systems for the simulation is given by the content of your standard orbit file. The signal types used for the two frequencies for each of these GNSS are given in the "Observation selection" file also defined in panel "GPSSIM 1.2: General Files":

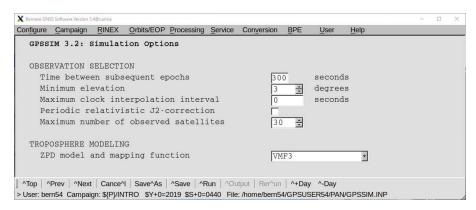
```
Receiver type
                         S/S O/F RINEX observation codes and their priority
STMULA
                         G
                               T.1
                                      1 P
                         G
                               L2
                                      2P
                         G
                               C1
                                      1 P
                         G
                               C2
                                      2P
                               C1
                                      1 P
                         R.
                               C2
                                      2P
                                      1 C
                         Ε
                               L1
                         Ε
                               L2
                                      5 Q
                               C1
                                      1 C
```

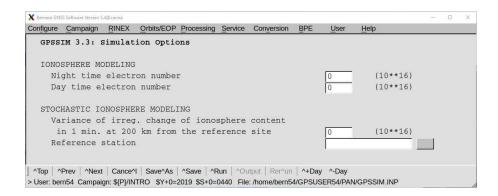
Note that each of these frequencies needs a corresponding antenna correction in the "Antenna corrections" file given in panel "GPSSIM 1.2: General Files". The simplest definition for this application is:

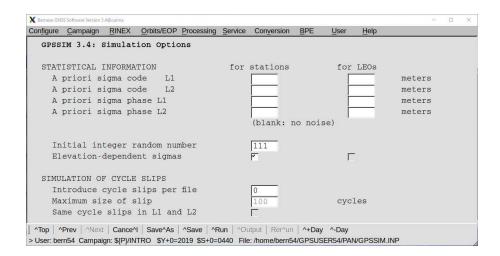
```
ANTENNA/RADOME TYPE NUMBER SYS FRQ TYP D(0) D(Z) D(A) M(Z)
                                   2
SIMULA
                NONE
                          0 G
                                       0
                                            0
                                                 0
                                                       0
                                                           90
                            R.
                                       0
                                            0
                                                 Ω
                                                       0
                                                           90
                             Ε
                                       0
                                                 0
                                                       0
                                                           90
```

```
NORTH MM EAST MM UP MM
                                    FACTOR
       *****.** ****.** ****.**
G01 0
          0.00
                    0.00
                             0.00
                                     0.100E+01
G02 0
          0.00
                    0.00
                              0.00
                                     0.100E+01
R01 0
          0.00
                    0.00
                              0.00
                                     0.100E+01
R02 0
          0.00
                    0.00
                              0.00
                                     0.100E+01
E01 0
          0.00
                    0.00
                              0.00
                                     0.100E+01
E05 0
          0.00
                    0.00
                              0.00
                                     0.100E+01
. . .
```

In the subsequent panels some basic characteristics, assumptions on the ionospheric conditions, and the noise level are introduced. Even cycle slips may be simulated.





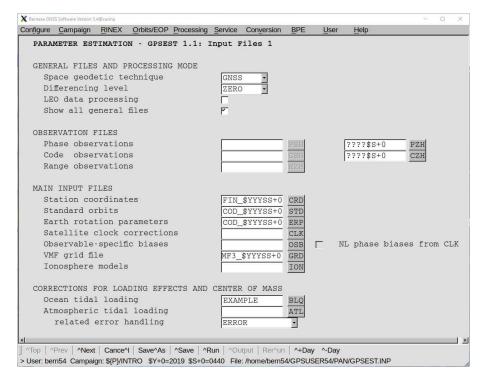


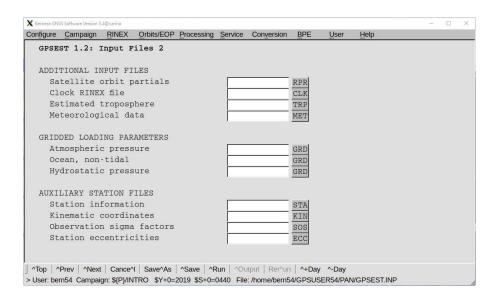
The observations are generated without noise to check the consistency with the processing program in the subsequent sections.

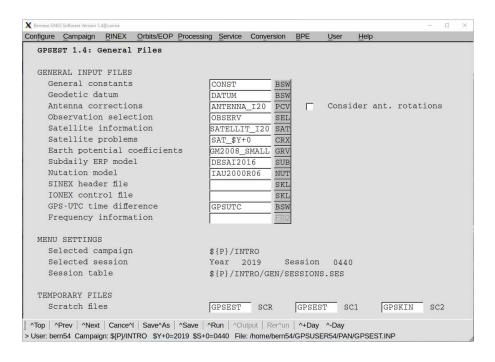
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## 7.3.2 Zero Difference Solution from Simulated GNSS Observations

The simulated observations can directly be introduced in program GPSEST for an analysis on zero difference level (if you have not simulated cycle slips). Please pay attention to the consistency of all input files with respect to the simulation:

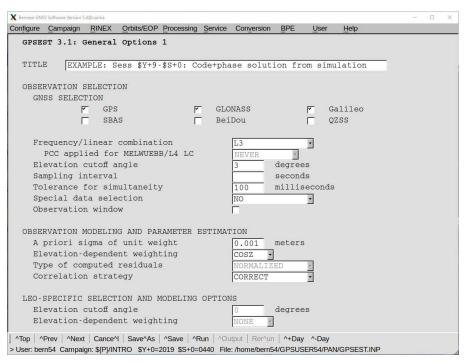




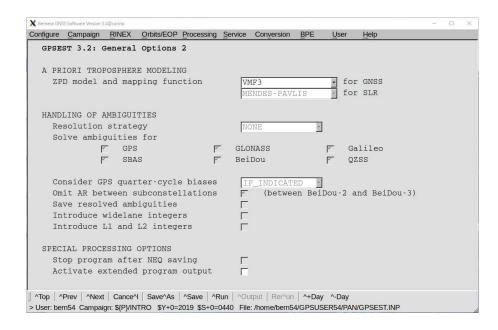


Result files can be specified to compare the results with the inputs for GPSSIM. A residual file might also be useful.

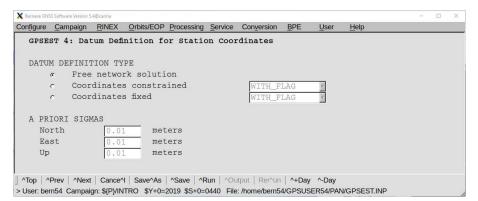
The processing models also have to be consistent with the simulation or to contain well defined differences which are the subject of your investigation:



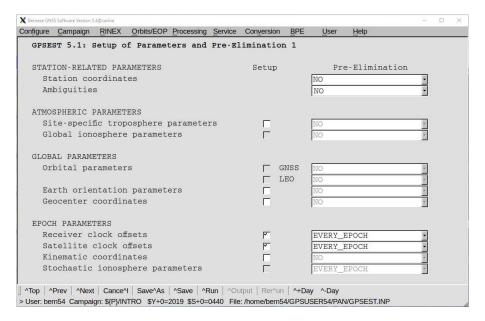
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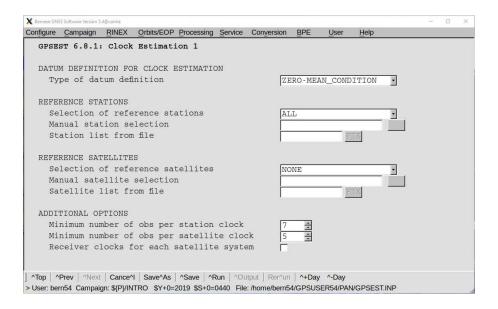


No constraints for datum definition are needed because these are noise–free simulated data:

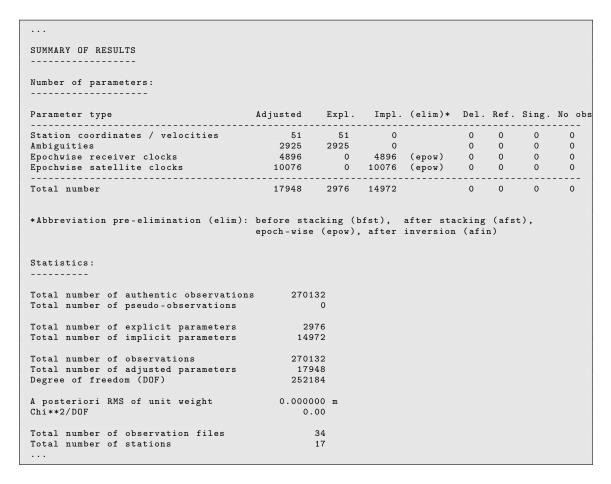


Only the receiver and satellite clocks are estimated:





The resulting program output file looks like the usual output from GPSEST but with perfect observations without any noise:



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No improvements for the station coordinates and other parameters are expected:

Stat	tion coordinates	and velo	cities:				
Sol	Station name	Тур	Correction	Estimated value	RMS error	A priori value	 Abb
1	BRST 10004M004	X	-0.00000	4231162.41917	0.00000	4231162.41917	 #CRD
1	BRST 10004M004	Y	0.00000	-332746.42575	0.00000	-332746.42575	 #CRD
1	BRST 10004M004	Z	-0.00000	4745131.07277	0.00000	4745131.07277	 #CRD
1	GANP 11515M001	Х	-0.00000	3929181.29827	0.00000	3929181.29827	 #CRD
1	GANP 11515M001	Y	0.00000	1455236.95252	0.00000	1455236.95252	 #CRD
1	GANP 11515M001	Z	-0.00000	4793654.03600	0.00000	4793654.03600	 #CRD
1	HERT 13212M010	Х	-0.00000	4033460.74185	0.00000	4033460.74185	 #CRD
1	HERT 13212M010	Y	0.00000	23538.04010	0.00000	23538.04010	 #CRD

The inter-system and inter-frequency biases have been assumed to be zero during the simulation (what is equivalent to any other constant number):

Station name	#obs	Correction	Estimated value	RMS error	From	Abb
	60 F		0.00000		2019-02-13 00:00:00	#RCK
GANP 11515M001	60 F		0.00000		2019-02-13 00:00:00	#RCK
HERT 13212M010	60 F		-0.00000		2019-02-13 00:00:00	#RCK
JOZ2 12204M002	56 F		-0.00000		2019-02-13 00:00:00	#RCF
LAMA 12209M001	58 F		-0.00000		2019-02-13 00:00:00	#RCF
MATE 12734M008	60 F		-0.00000		2019-02-13 00:00:00	#RCF
MIKL 12335M001	60 F		0.00000		2019-02-13 00:00:00	#RCF
ONSA 10402M004	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
ORID 15601M001	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
PTBB 14234M001	60 F	-0.00000	-0.00000	0.00000	2019-02-13 00:00:00	#RCF
TLSE 10003M009	60 F	-0.00000	-0.00000	0.00000	2019-02-13 00:00:00	#RCF
VILL 13406M001	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
WSRT 13506M005	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
WTZR 14201M010	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
WTZZ 14201M014	60 F	-0.0000	-0.00000	0.00000	2019-02-13 00:00:00	#RCI
ZIM2 14001M008	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF
ZIMM 14001M004	60 F	0.00000	0.00000	0.00000	2019-02-13 00:00:00	#RCF

The same holds for the epoch–wise clocks.

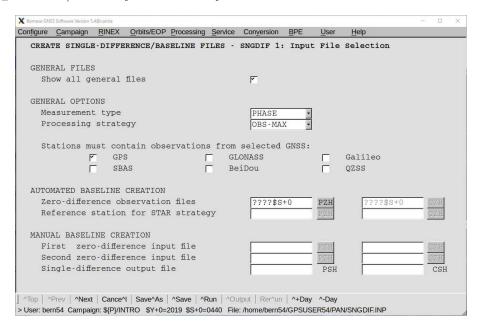
The residuals in the optional residual output file of such a dataset are well below the 1  $\mu$ m level.

## 7.3.3 Double-Difference Solution from Simulated GNSS Observations

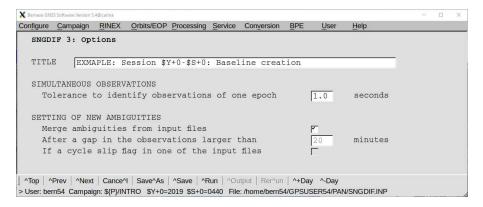
As in the beginning of the simulation, you should make sure that no other baseline observation files for the current session exist in the OBS-directory of your campaign to prevent any interferences and mixtures of simulated measurements with other ones.

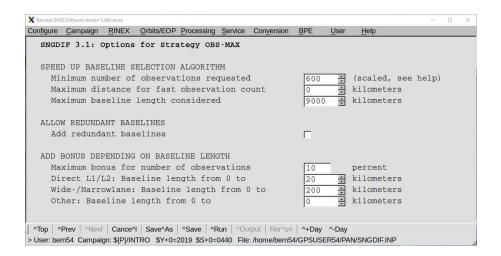
```
bern54@carina:~ > rm ${P}/INTRO/OBS/????0440.PS?
```

The simulated measurements can also be processed in the double-difference mode. In that case you have to start with forming baselines using the program SNGDIF ("Menu>Processing >Create baseline files") in nearly the same way as in Section 4.2.2 for real observations:



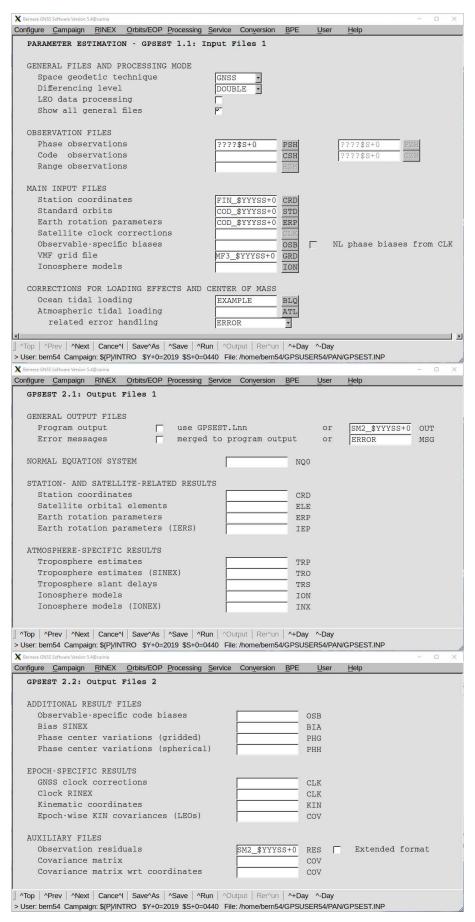
The main difference is that you should also keep all ambiguities from the zero difference in the baseline observation files, what is managed by checking the box for option "Merge ambiguities from input files":



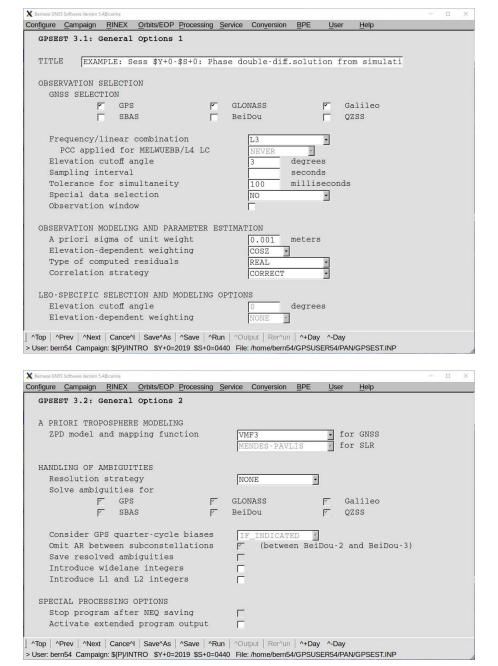


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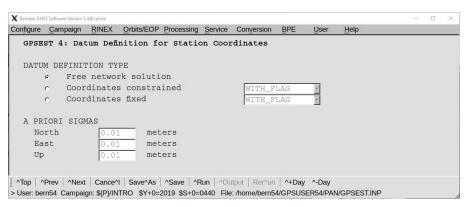
The resulting baseline files from simulated observations can now be analyzed with program GPSEST:



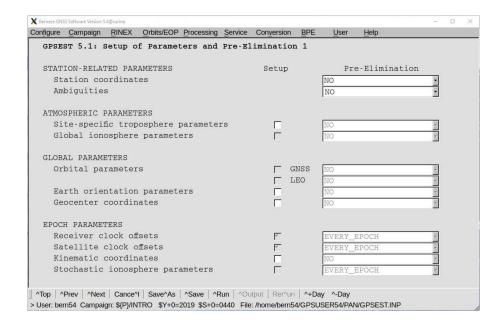
The models are selected to be fully consistent with the simulation:



No other parameters than station coordinates and ambiguities are estimated:



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The results are analogue to the zero difference case previously described:

```
Statistics:
{\tt Total\ number\ of\ authentic\ observations}
                                               118515
Total number of pseudo-observations
                                                   16
Total number of explicit parameters
                                                 2763
Total number of implicit parameters
Total number of observations
                                               118531
Total number of adjusted parameters
                                                2763
                                               115768
Degree of freedom (DOF)
A posteriori RMS of unit weight
                                            0.000000 m
                                                 0.00
Total number of observation files
                                                   16
Total number of stations
                                                   17
```

The ambiguities are set up in a way that the correct resolution for all ambiguities is zero in any case. This is an easy way to verify ambiguity resolution strategies.

## 7.3.4 Final Remarks

There are many opportunities to use this simulation tool. It depends on your needs and the concrete target of the simulation study to define the experiment. As it was just demonstrated the full consistency between the processing and the simulation programs is guaranteed by the *Bernese GNSS Software*.

The big advantage of a simulation is that the correct solution is known a priori. On the other hand, you have to keep in mind that the simulated data can only contain effects included in the simulation model. If a receiver for instance introduces a significant variation of the inter–system bias between GPS and GLONASS data – an effect that is not considered in the simulation model – the influence of such an effect on the results cannot be evaluated by the simulation.

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