

Influence of Low Callisto Orbit design on gravity field recovery



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| Orbit characteristics | Max. degree |
|-----------------------|-------------|
| 88° 200x200km | 70 |
| 88° 400x400km | 45 |
| 112° 400x400km (SSO) | 18 |
| 88° 400x1400km | 19 |

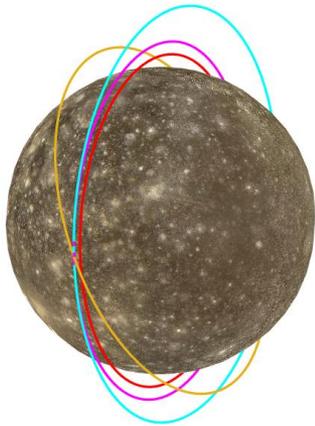
Maximum spherical harmonic degree of gravity field recoverable for low Callisto orbits, with different inclinations and altitudes for a duration of 90 days

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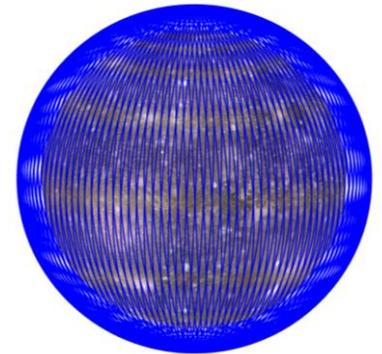


Introduction and Background

- Gan De is a Chinese exploration mission under study, that would fly to Jupiter in the 2030's [3]. An orbiter would be injected into a Low Callisto Orbit to perform an extensive characterization of its surface and interior, investigate its degree of differentiation and search for the possible existence of an internal ocean.
- After an extended tour of the Jupiter system, a first polar **elliptic orbit** is foreseen for capture around Callisto. Then two polar circular orbits could be used for science investigation. A **first one** for at least 6 months, and a **second one** with lower altitude, with the possibility of regular manoeuvres to counteract orbit decay.
- Here, more specific orbits are also investigated due to their relevance for mission design:



- **Sun synchronous orbits (SSO)**: constant angle between Sun and orbital plane, but with an important polar gap and highly dependent on the gravity field knowledge at low altitude.
- **Repetitive Ground Track Orbits (RGTO)**: defined by an integer triplet (N,P,Q) [4], fixed phase grid defined for $N \cdot P + Q$ orbit revolutions during P Callisto days [5].



- Orbit propagations in a full force model, as well as the whole gravity field recovery process were done using a development version of the Bernese GNSS Software [6].

Set of orbits and simulation setup

| | Altitude | Inc. | RGTO | SSO |
|---|------------|------|-----------|-----|
| ■ | 200x200km | 88° | No | No |
| ■ | 200x200km | 88° | (146,1,0) | No |
| ■ | 197x197km | 88° | (146,5,1) | No |
| ■ | 395x395km | 88° | (131,1,0) | No |
| ■ | 401x401km | 112° | No | Yes |
| ■ | 400x1400km | 88° | No | No |

Set of 6 orbits under study. All have a $45^\circ \beta_{Earth}$ angle (between orbital plane and Earth)

**Simulation flow chart
(for each orbit)**

Initial condition

90/200 days propagation
from 01-May-2031

Daily initial
conditions

2-way Doppler
X-band obs. *

$\sigma_p = 50$ m
 $\sigma_v = 1$ mm/s

$\sigma_{obs} = 0.1$ mm/s
at $\tau = 60$ s

- 3rd body perturbations:
Sun, planets, Galilean moons
- Jupiter gravity field: J_2 to J_6
- Tides from Jupiter: $k_2 = 0.0$

Generalized orbit determination
(Celestial Mechanics Approach [2])

Reference Callisto gravity field:
- d/o 2: Anderson et al (1998) [1]
- d/o 3 to 50/90: scaled Moon's field

Comparison: $\Delta \bar{C}_{nm}, \Delta \bar{S}_{nm}, \Delta g_{\theta, \phi}$

(***)

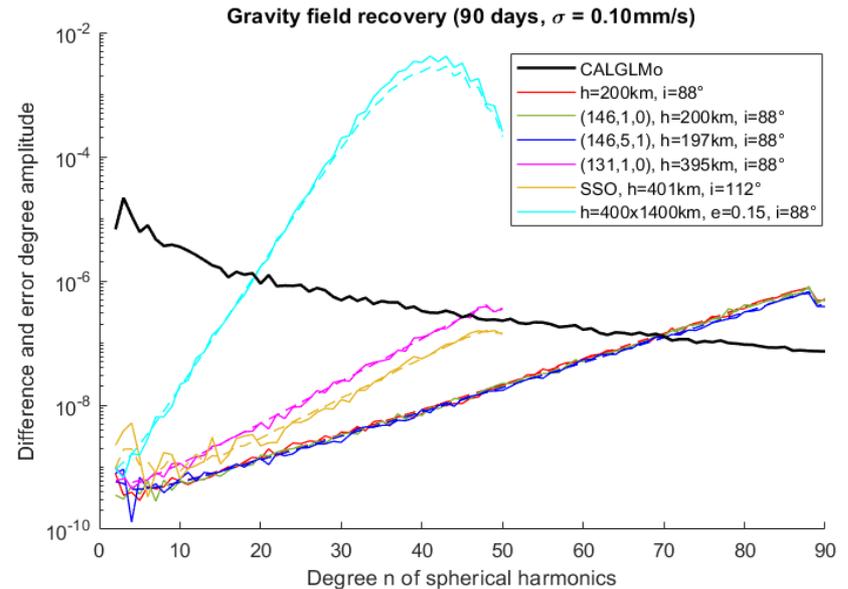
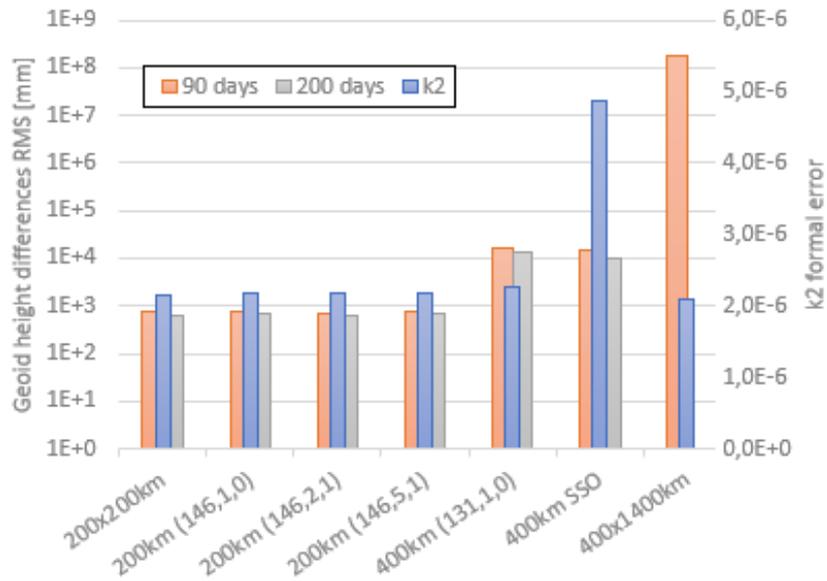
k_2 and gravity field solution **

Stacked normal equation
(90/200 days)

Daily normal equations

* : Generated with a full coverage of 3 Deep Space Network stations
 ** : Coefficients are estimated freely in only one iteration
 ***: Tests have been made with degraded a priori gravity field, requiring then several iterations

Gravity field recovery



Weighted RMS of geoid height differences $\Delta g_{\theta,\phi}$ for 90/200 days $\left(\sqrt{\frac{\sum_{\theta,\phi} \cos(\theta) \Delta g_{\theta,\phi}^2}{\text{gridsize}}} \right)$ and k_2 Love number formal error for 90 days mission computed using an a priori d/o 50 field.

Difference (solid) and error (dashed) degree amplitudes $(M_n = \sqrt{\frac{\sum_{m=2}^n (\Delta \bar{C}_{nm}^2 + \Delta \bar{S}_{nm}^2)}{2n+1}})$. For 200km orbits, the gravity field was estimated up to d/o 90

- 22° polar gap is omitted for the Sun Synchronous Orbit
- With face-on orbit, the gravity field recovery is worse. As an example, the (146,1,0) orbit leads to a larger weighted RMS of geoid height difference for $\beta_{Earth}=90^\circ$ (153cm) than for $\beta_{Earth}=45^\circ$ (88cm).
- Using a d/o 40 truncated gravity field with the 200km (146,1,0) orbit, 4 iterations on the gravity field solution are needed to reach the solution computed with a full d/o 50 a priori gravity field.

Conclusions

- A highly eccentric orbit over a time span of 90 days can already improve the knowledge of Callisto's gravity field (up to d/o 19 for a 400x1400km orbit). However, as the eccentricity increases significantly with time, such an orbit is not stable for more than 3 months.
- Sun synchronous orbits suffer from a large polar gap, the recovery of zonal coefficient is then largely impacted, just as Love number k_2 recoverability.
- For all non-Sun synchronous orbits, β_{Sun} does not vary much (max. 1.2°/month). A SSO for maximum illumination might then not be compulsory.
- Low altitude polar orbits are the best suited for gravity field recovery. At 400km altitude, one can expect to recover the gravity field up to d/o 45 after 90 days.
- Lower orbits are even more beneficial, but will require manoeuvres to increase the orbit lifetime. Repetitive Ground Track Orbits are well suited to efficiently plan station keeping manoeuvres.
- For 200km polar orbits a sensitivity up to d/o 70 was found after 90 days. In the case of Callisto, the effect of low density ground tracks (for RGTO) is negligible.

Acknowledgements & References

Acknowledgments

This study has been funded with the support of the Swiss National Foundation (SNF).

References

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