

# **Present and Future Impact of GNSS on Spaceborne Scientific Applications – *in particular on Orbit and Gravity Field Determination***

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- **Global Navigation Satellite Systems (GNSS): Current status and impact on science**
- **Impact of *different* GNSS on science**
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# Global Navigation Satellite Systems 2010

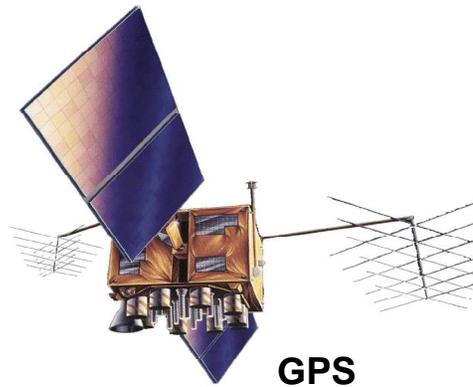
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Global Positioning System (GPS): **USA** , ~ 30 satellites

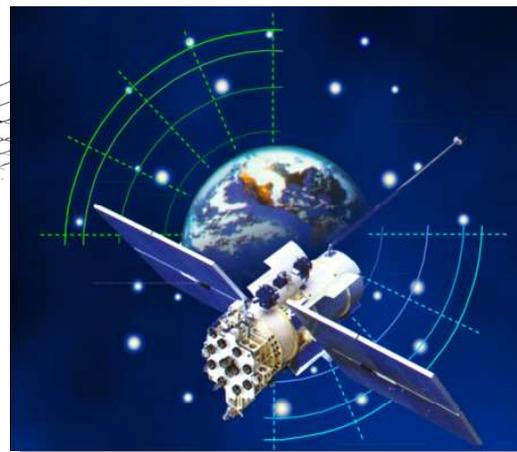
Global Navigation Satellite System  
(GLONASS): **Russia**, ~ 20 satellites

GALILEO: **Europe**, 2 satellites

... Currently, GPS and GLONASS  
are routinely used for science and  
other demanding applications



GPS



GLONASS



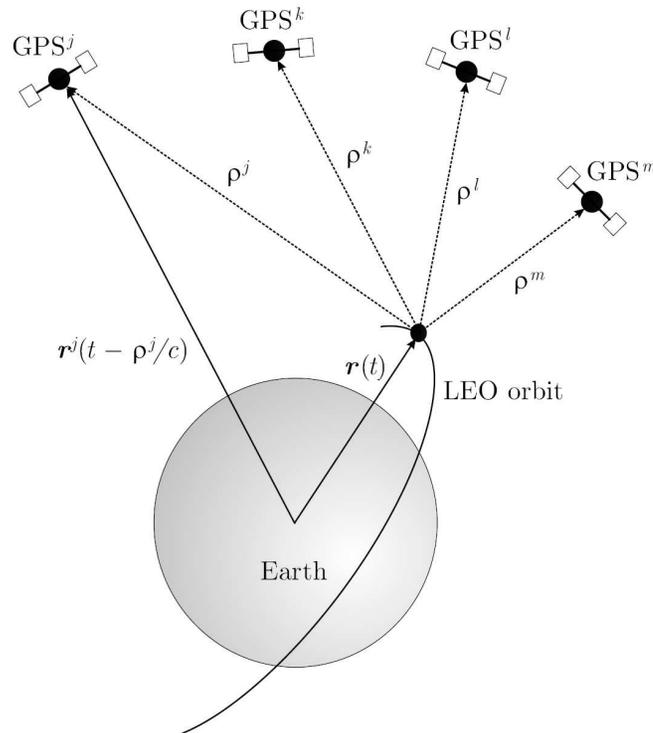
GALILEO

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# Global Navigation Satellite Systems 2010



**GNSS** have been designed for use on the Earth's surface or in the Earth-near space.

Each satellite is equipped with a stable oscillator generating at least two coherent carriers. Code info is modulated on the carrier.

The travelling time of signals (and its changes in time) between the **GPS satellite and the receiver** are the **basic measurements**.

→ With the speed of light  $c$  the distances  $\rho$  (and their time evolution) between satellite and receiver may be reconstructed.

# Global Navigation Satellite Systems 2010

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The so-called pseudo-range is composed of

$$c (t_r - t^s) = \rho + c (\Delta t_r - \Delta t^s) + \Delta \rho_I(\lambda) + \Delta \rho_t$$

- $\rho$  is used to determine the **position** of the receiver, but also the **orbit** of the satellite.
- $c (\Delta t_r - \Delta t^s)$  is used for **clock-synchronisation**.
- $\Delta \rho_I(\lambda)$ , the delay of the signals due to the ionosphere, may be used for **ionosphere modelling**.
- $\Delta \rho_t$ , the signal delay in the neutral atmosphere, may be used in **meteorology** (in particular to determine the rapidly varying water vapor content).

# Global Navigation Satellite Systems 2010

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**The scientific use of GNSS is coordinated by the IGS.**

**Created in 1989** with I.I. Mueller, G. Mader, B.

Melbourne, and **Ruth Neilan** as protagonists.

The IGS became an **official IAG service** in 1994.

The IGS **first** was **a pure GPS Service**, it was renamed as the **International GNSS Service** in 2004.

Today the IGS is an **interdisciplinary service in support of Earth Sciences and Society**, making optimum use of the data from all GNSS.

Since its creation the **IGS Central Bureau** is located at JPL, Pasadena, with **Ruth Neilan** as director.

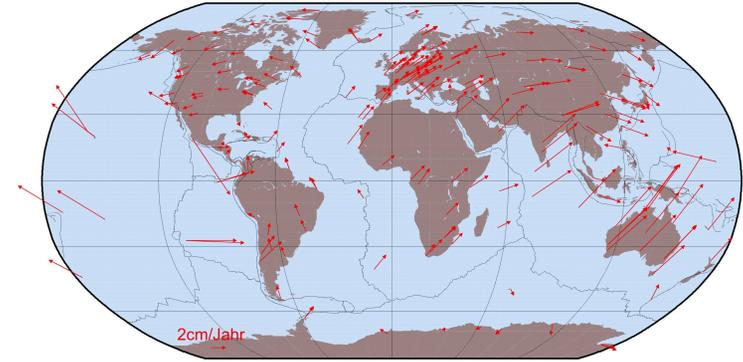
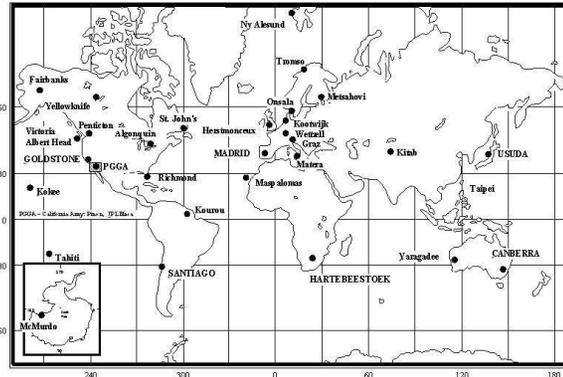
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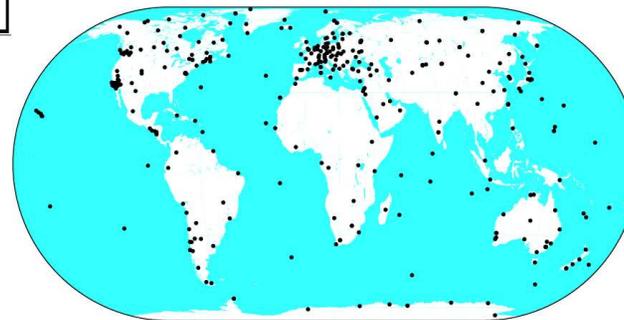
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# Global Navigation Satellite Systems 2010

Station Locations for the IGS Pilot Campaign, 1992



Monitor station motion  
in „real time“



IGS Network in March 2010

In 1992 the IGS was based on about 20 geodetic receivers, 400+ receivers are active and their data openly available today

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# Global Navigation Satellite Systems 2010

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In 1992 the IGS started off as an orbit determination service (few dm accuracy) for about 20 GPS satellites.

**Today**, the IGS provides **ephemerides** (?accuracy/consistency? of 2-4 cm) for about **30 GPS satellites** and for all **GLONASS satellites**, i.e., for all currently active GNSS satellites.

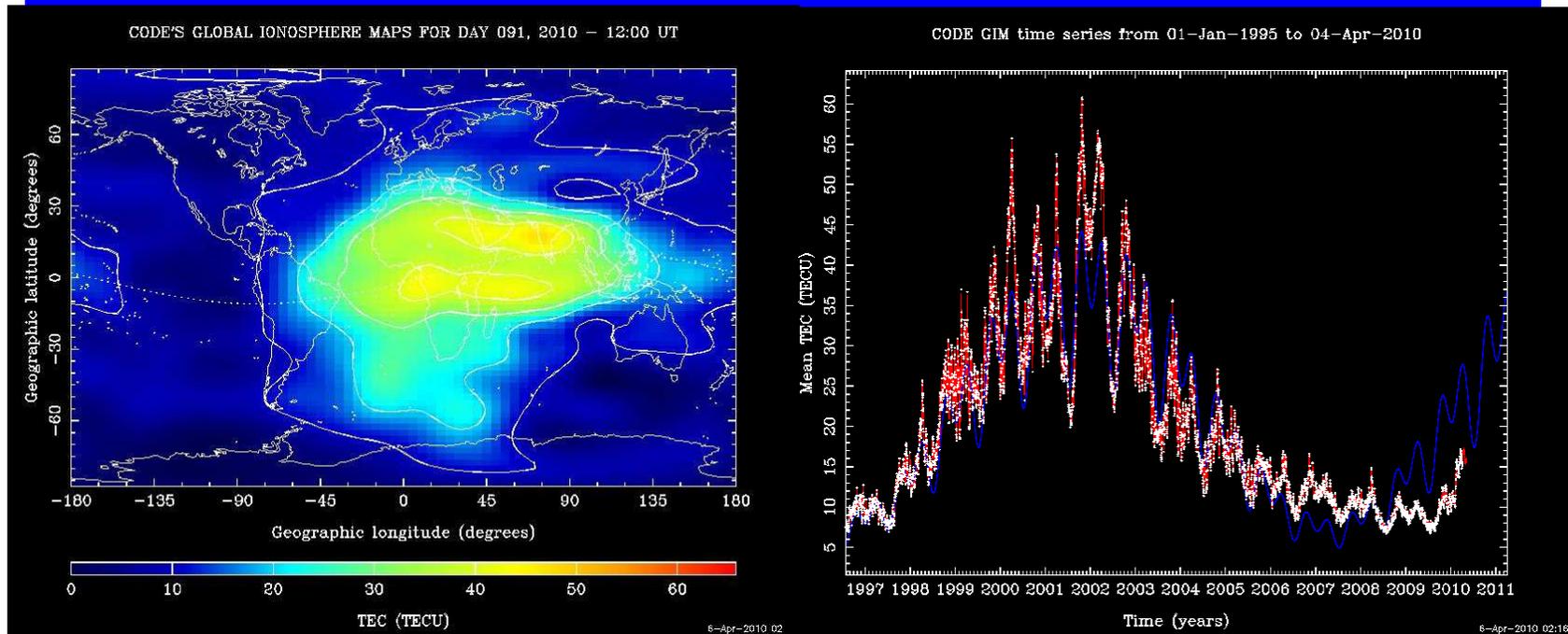
In addition the IGS provides

- archive of all globally relevant GNSS data since 1991
- satellite and receiver clock corrections (sub-ns accuracy)
- polar motion (PM) and length of day (lod) (mm-cm [/day] accuracy)
- coordinates & velocities for 200+ sites (mm-cm / mm/y accuracy)
- atmosphere information (ionosphere and troposphere)

The IGS products are

- **accurate**, **reliable** and **robust**,
- available in a **timely** manner.

# Global Navigation Satellite Systems 2010



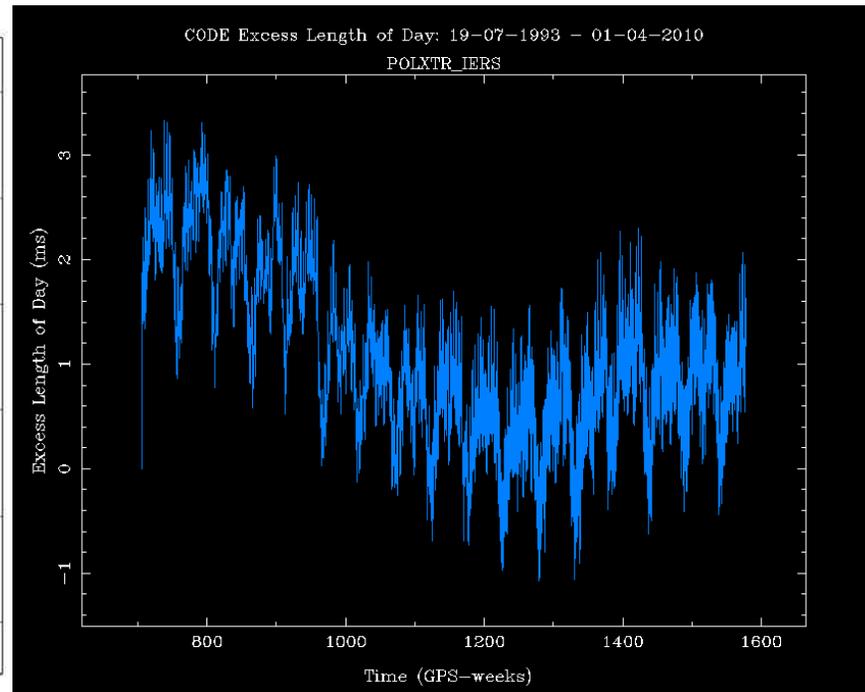
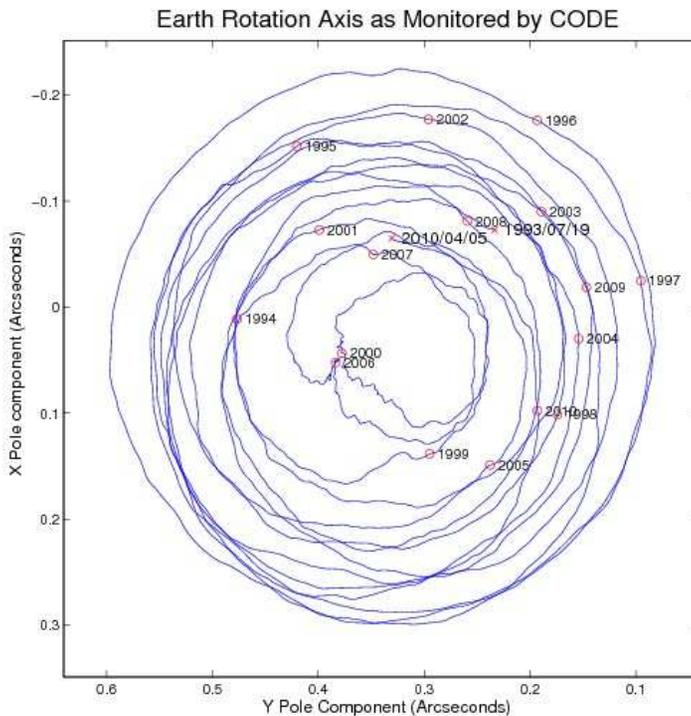
**Left: Ionosphere map (2<sup>h</sup> time resolution) routinely derived by IGS Analysis Centers**

**Right: Development of mean daily TEC since mid 1996 (figures from CODE/IGS Analysis Center)**

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# Global Navigation Satellite Systems 2010



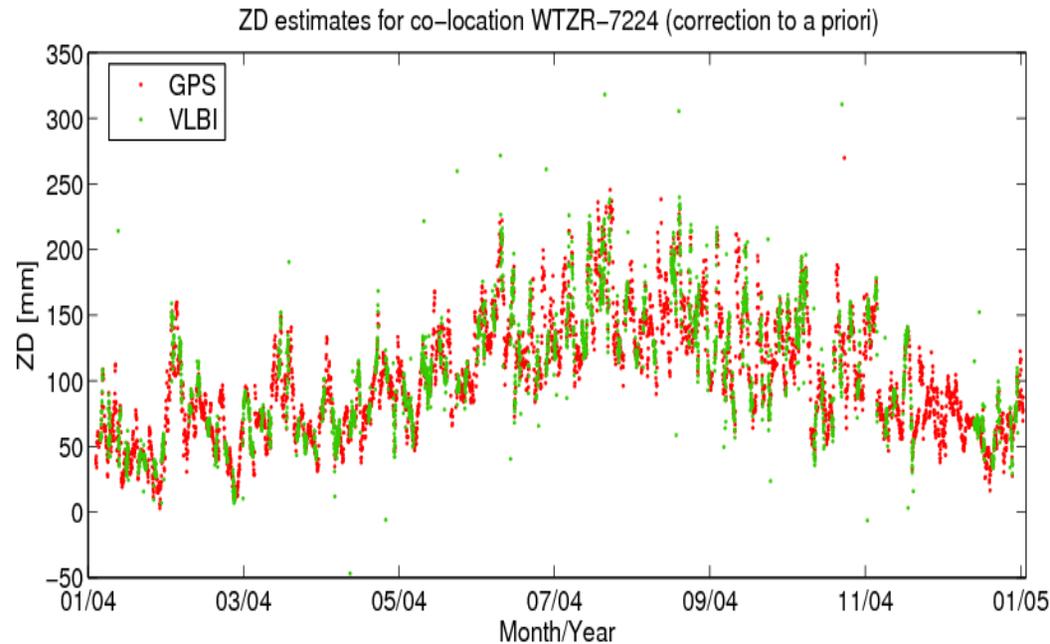
**Left: Polar motion (~motion of rotation axis on Earth's surface) since 1993 (diameter of figure about 14m)**

**Right: Excess Length of day since 1993 (milliseconds) (from CODE Analysis Center)**

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# Global Navigation Satellite Systems 2010



**Tropospheric zenith delays for 2004 as estimated by GPS and VLBI for the collocated GPS and VLBI sites in Wettzell. The variations are mainly due to the atmospheric water vapor content.**

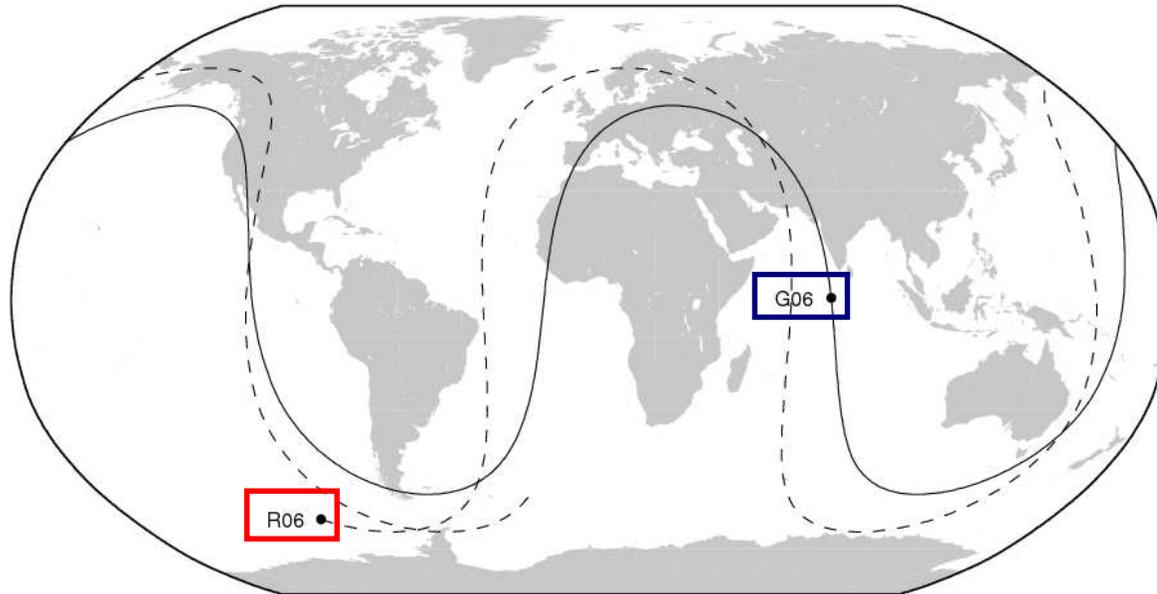
From: D. Thaller, M. Rothacher, and M. Krügel (2009). Combining one year of homogeneously processed GPS, VLBI and SLR data. In H. Drewes (editor), Proceedings of the IAG Symposium on Geodetic Reference Frames GRF2006, number 134 in IAG Symposia. Springer-Verlag.

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# Impact of Different GNSS Constellations

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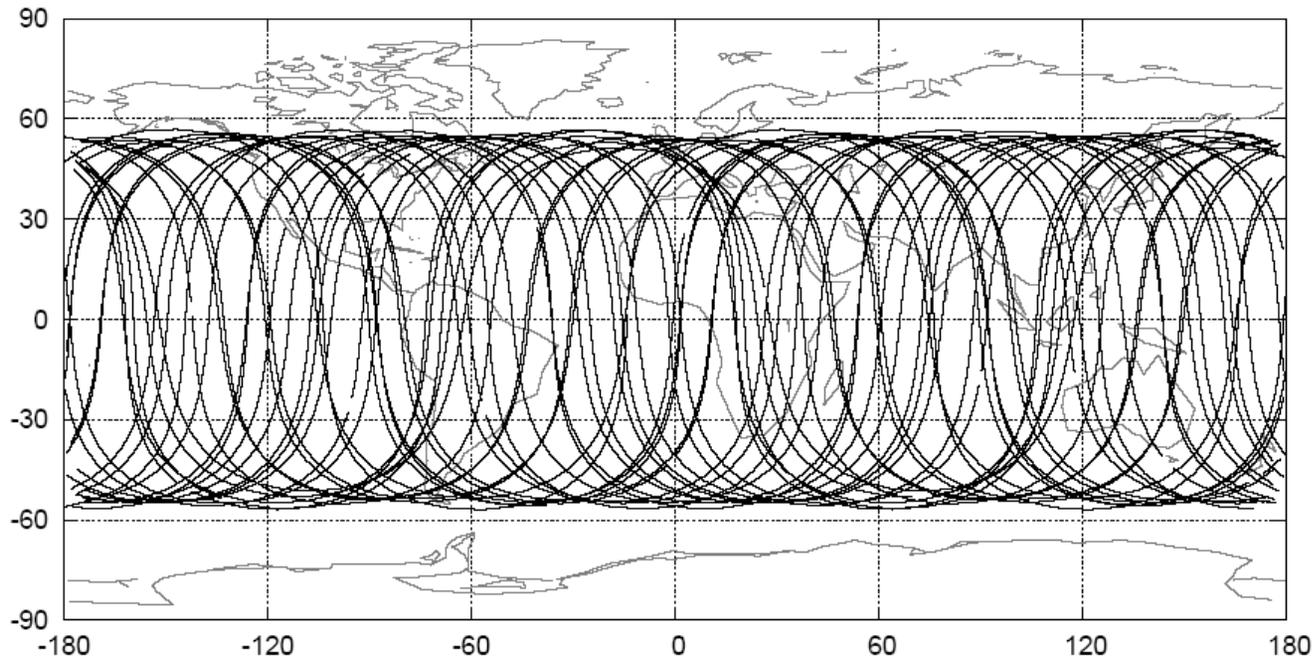
July 7, 2006: **sub-satellite tracks** of:  
**GPS PRN 06**, with daily repeat orbit and  
**GLONASS R06**, orbit repeating after 8 days.

The GNSS constellations differ considerably (inclinations, daily vs. 8-day repeat orbits for GPS and GLONASS, respectively)

**Different constellations improve the geometry, help to understand systematic errors**

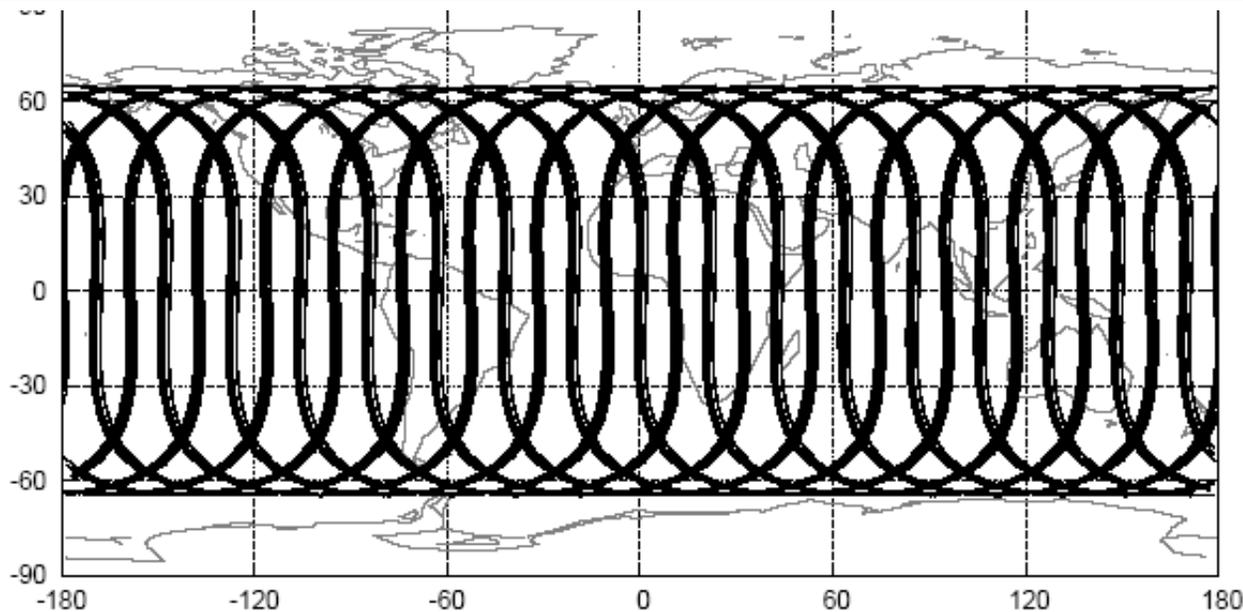
# Impact of Different GNSS Constellations

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**Mean observation geometry of a particular satellite, as viewed from a site at a particular latitude  $\phi$ , is longitude-dependent.**

# Impact of Different GNSS Constellations



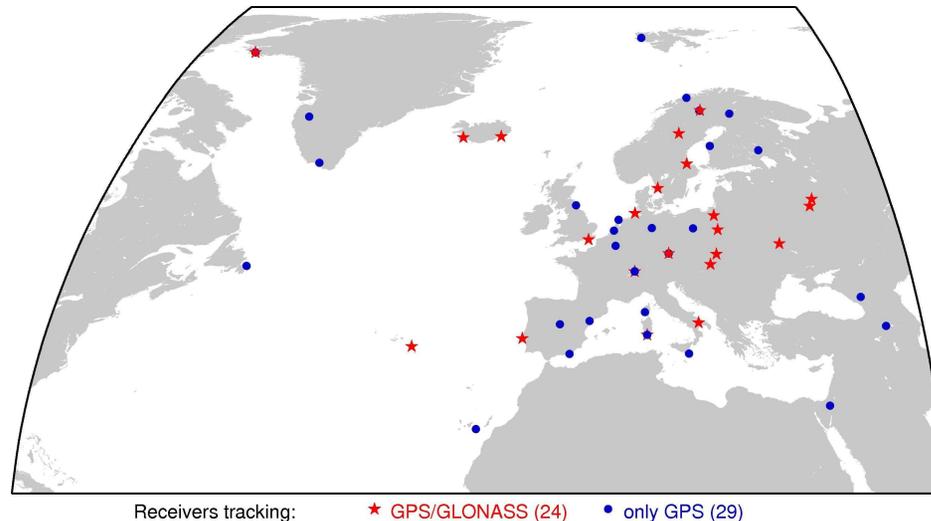
Only two (of three) orbital planes filled with satellites in 2006.

Figure 3.5: Ground-tracks of 12 GLONASS satellites over 17 orbital revolutions, for October 28 - November 4, 2006

**Mean observation geometry of a particular satellite (in the average over 8 (or more) days), as viewed from a site at a particular latitude  $\phi$ , is (almost) longitude-independent.**

# Impact of Different GNSS Constellations

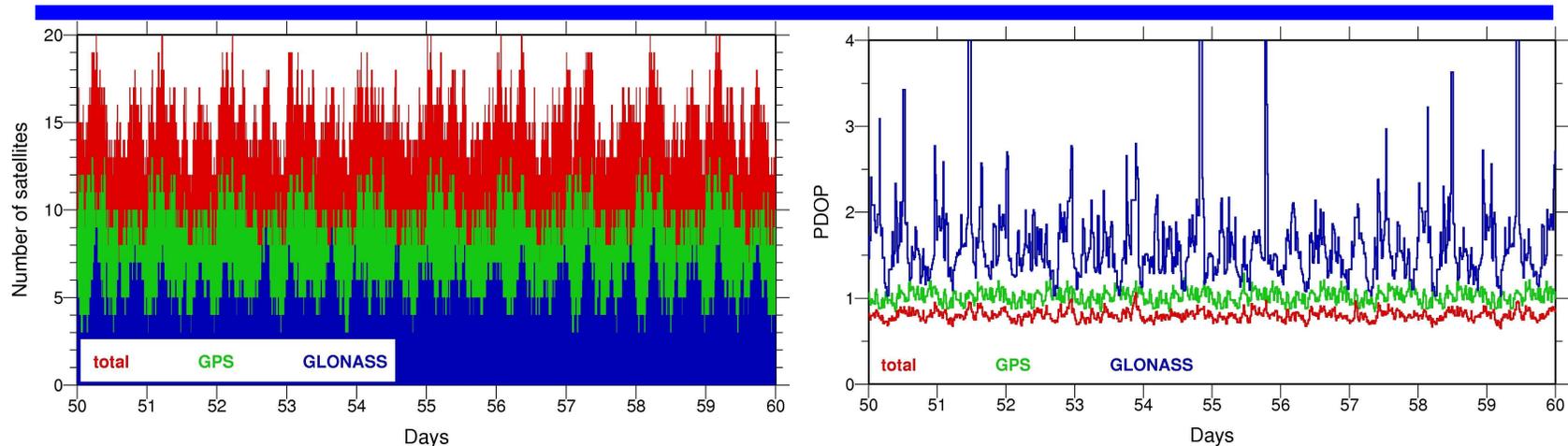
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**A regional network consisting of GPS and combined GPS/ GLONASS receivers was analyzed in (short) batches of 0.5, 1, 2, 3, etc min. Otherwise: network analysis with settings of CODE Analysis Center of the IGS. Details see**

**Dach, R., E. Brockmann, S. Schaer, G. Beutler, M. Meindl, L. Prange, H. Bock, A. Jäggi, L. Ostini; (2009) *GNSS processing at CODE: Status report*. JoG, Vol. 83(3-4), pp. 353-366.**

# Impact of Different GNSS Constellations



**Left: Number of satellites (left) and Right: PDOP (right) for Zimmerwald.**

**These indicators are promising. The mentioned paper revealed, however, that the improvement of station coordinates and Earth rotation parameters is moderate to not existing.**

**Clear improvements were observed when generating solutions GPS and combined solutions using only short data spans (of up to 30 minutes). Improvements of > 10% resulted.**

**→ The heaviest impact of multi-GNSS will occur in kinematic positioning.**

# Expectations concerning GALILEO

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The **scientific community** will not switch from one GNSS to another, but combine the measurements from all systems (the IGS is already doing that with GPS and GLONASS).

It is assumed that at least the same information as in the case of GPS (code and phase on at least two carriers) will be *openly* available for all GNSS.

The obvious advantages of combining GNSS are:

- When using  $n$  different GNSS with the same number of satellites the common parameters should at least **improve by a factor of  $n^{1/2}$**  ... (not yet realized in practice)
- Inconsistencies in the **reference frames** „cannot“ occur.
- System-specific **systematic errors** may be detected more easily (and hopefully removed).
- **Better coverage for (spaceborne) kinematic positioning and atmosphere sounding**

# Expectations concerning GALILEO

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**Kinematic applications** of GNSS are delicate because of

- the huge number of parameters (four unknowns, namely position  $x$  and receiver clock correction  $\tau$  per observation epoch in addition to general parameters like ambiguities, troposphere, etc.).

**Spaceborne kinematic applications** are particularly delicate because of

- the rapid changes of the GNSS observation scenario: A particular GNSS satellite is above horizon for 4-5 hours for Earth-fixed receivers, for 10-20 minutes for spaceborne receivers.

# Expectations concerning GALILEO

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**GALILEO has the potential to revolutionize the scientific and other high accuracy applications of GNSS because of its passive hydrogen space masers.**

**The well-known explicit or implicit double-differencing of observations may be left in favor of models for the GALILEO space clocks as functions of time. The consequences are**

- **Substantially reduced number of parameters → increased degree of freedom.**
- **better separation of clock errors and model errors attributed, e.g., to orbits.**
- **Need to better model the orbits, time-dependent cable delays, ...**
- **Correlation of observations in time → use of weight matrices covering the entire observation time span.**

# Expectations concerning GALILEO

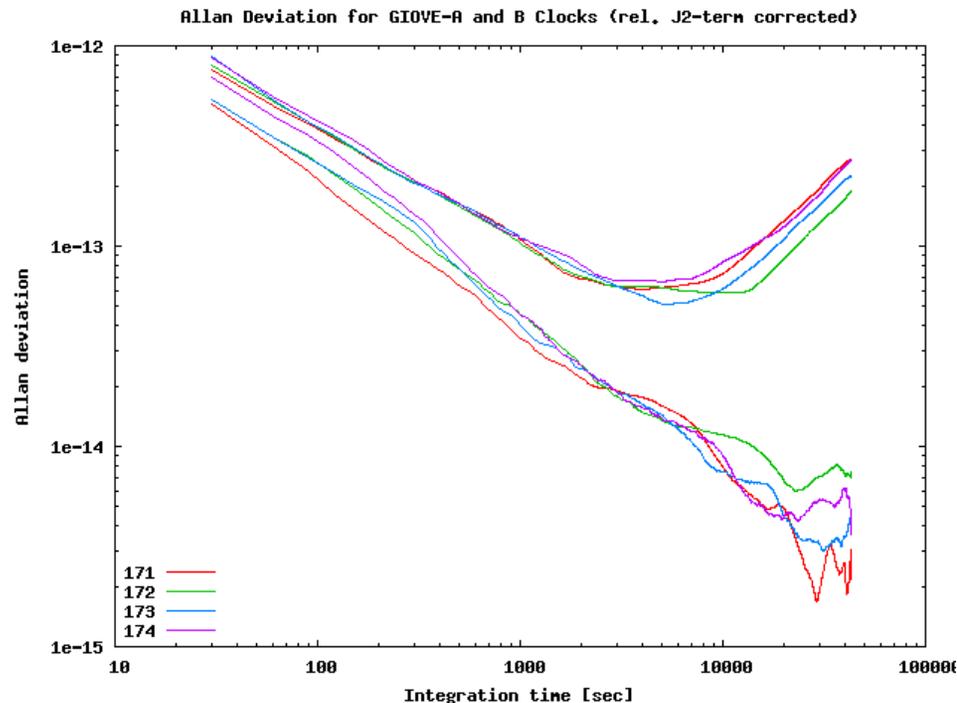
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In 1984 G. Beutler et al. wrote:

*„In the authors' opinion, the best way of modelling the clocks is the following: define a statistical model of clock performances using available information on clock offset, drift, jitter. This leads to a simple stochastic differential equation ... For the clock synchronization error as a function of time ...“*

From: G. Beutler, D.A. Davidson, R.B. Langley, R. Santerre, P. Vanicek, D.E. Wells (1984) *„Some theoretical and practical aspects of geodetic positioning using carrier phase difference observations of GPS satellites“*, Technical Report, Dept. Of Surveying Engineering, University of New Brunswick

# Expectations concerning GALILEO



**Potential impact:**

**Clock prediction** for real-time applications: Orbit errors might become the limiting factor for prediction.

**Double-difference approach** may have to be left.

**But:** The algorithms have to be kept manageable for the user! → Challenge

From Hugentobler et al. (2010) „Evaluation of GIOVE Satellite Clocks using the CONGO Network “. Allan deviations derived from an orbit & clock analysis using the data of the CONGO network of GNSS receivers (deployed by DLR and BKG, Germany → **Paper in Session 3, GNSS Timing I**).

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# The Galileo Science Advisory Committee

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The Galileo Science Advisory Committee (GSAC) was set up to:

- Recommend improvements to Galileo and EGNOS for scientific applications.
- **Maintain the Galileo Science Opportunity Document (GSOD), highlighting scientific priorities.**
- Support the preparation of announcements of opportunity (AO) for scientific studies – **first announcement imminent!**
- Advise on the use of Galileo and EGNOS data for scientific applications.
- Consider and review ESA-furnished documents related to the scientific use of GNSS signals.

# The Galileo Science Advisory Committee

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The members of the GSAC were nominated by the ESA member countries.

The GSAC was created towards the end of 2008 and so far met four times.

The fourth meeting took place 24 March 2010 in ESA-HQ, where the first version of the GSOD was (almost) finalized.

The purpose of the GSOD is to provide an overview of scientific endeavours benefiting from the Galileo system and to establish a common background for the scientific exploitation of the Galileo system and possible future versions.

The GSOD is maintained by the GSAC to advise the ESA directorate and to stimulate research related to Galileo in the wider scientific community.

# The Galileo Science Advisory Committee

2nd International Colloquium - Scientific and Fundamental Aspects of the Galileo Pro... Page 1 of 1

The screenshot shows the website for the 2nd International Colloquium on Scientific and Fundamental Aspects of the Galileo Program. At the top, there is a blue header with the ESA logo and the text "Congrex ESA's professional conference organiser European Space Agency". Below the header is a navigation menu with items like "Main page", "General Information", "Objectives", "Topics", "Abstract Submission", "Important Dates", and "Contact". The main content area features a "Call for Papers" section with the following text: "2nd International Colloquium - Scientific and Fundamental Aspects of the Galileo Program", "COSPAR Colloquium", "14 - 16 October 2009", and "University of Padova Padua, Italy". Below this text is an image of a funnel with stars, representing the Galileo program. At the bottom, there is a grid of logos for the organizing institutions: ESA, COSPAR, Bureau des Longitudes, ASI (Agenzia Spaziale Italiana), IAG, and URSI. The footer contains the text "Copyright 2000 - 2008 © European Space Agency. All rights reserved."

The second International Colloquium on *Scientific and Fundamental Aspects of the Galileo Program* gave detailed insight into the science aspects considered by the GSAC.

More information see:  
<http://www.congrex.nl/09c10/>

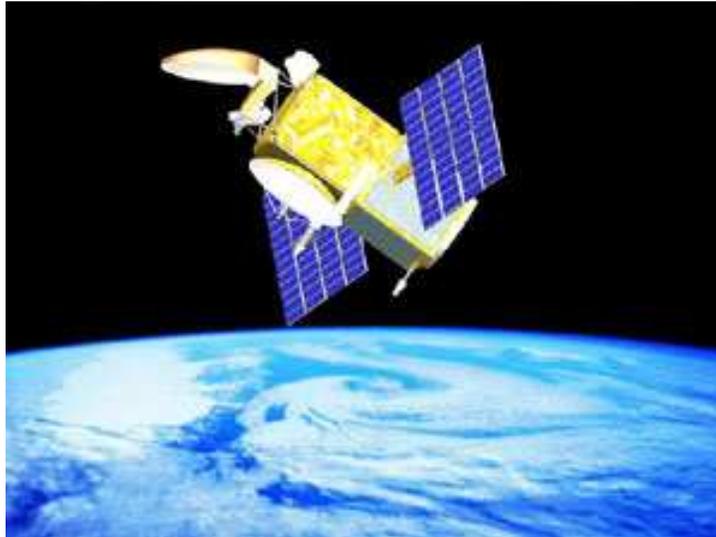
A third symposium is expected for 2011.

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# Spaceborne Applications of GNSS

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Jason-2,

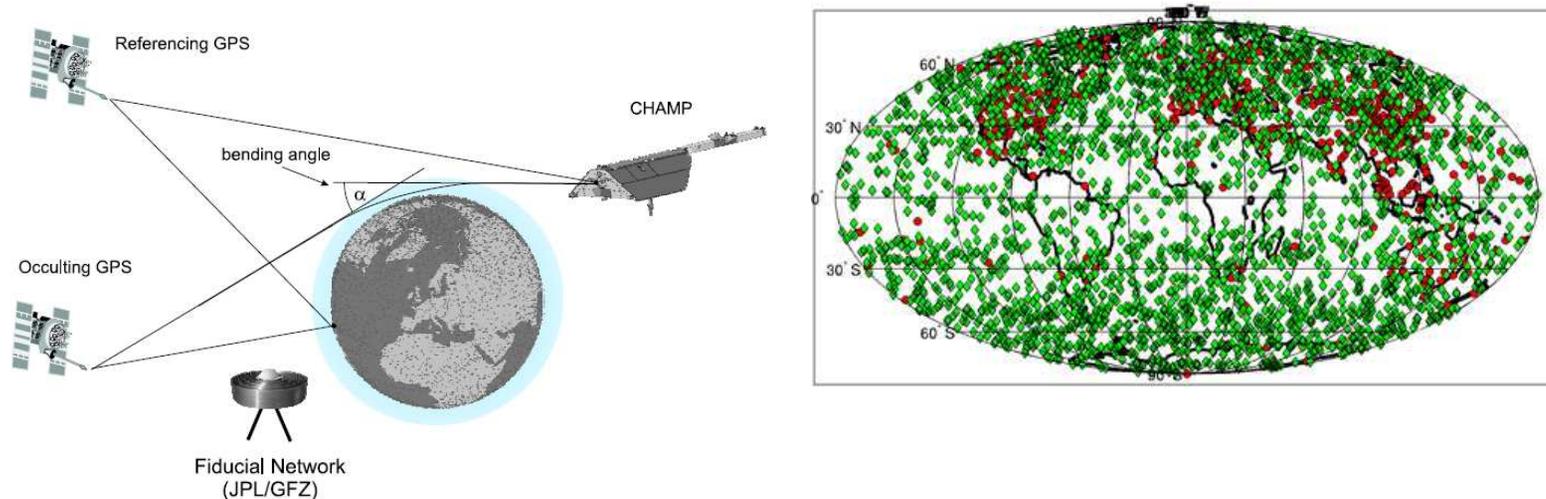


TerraSAR-X

Most Low Earth Orbiters (below 2000km) are today equipped with GPS receivers. **Most receivers are dedicated to Precise Orbit Determination (POD).**

Examples for LEOs equipped with GPS: Altimeter satellites Topex-Poseidon, Jason-1, Jason-2, gravity field satellites CHAMP, GRACE-A, GRACE-B, GOCE, SAR mission (TerraSAR-X)

# Spaceborne Applications of GNSS



**Atmosphere sounding using the GPS occultation technique exploiting the rising/setting of individual GPS satellite (proof of concept GPS/MET, CHAMP, GRACE-A, Formosat-3/COSMIC mission with a constellation of six micro-satellites carrying GPS radio occultation (RO) receivers gaining in total about 2500 profiles/day,**

**Left: from G. Beyerle, J. Wickert, R. Galas, K. Hocke, R. König, C. Marquardt, A. G. Pavelyev, Ch. Reigber and T. Schmidt (2001) „GPS Occultation Measurements with GPS/MET and CHAMP“,**

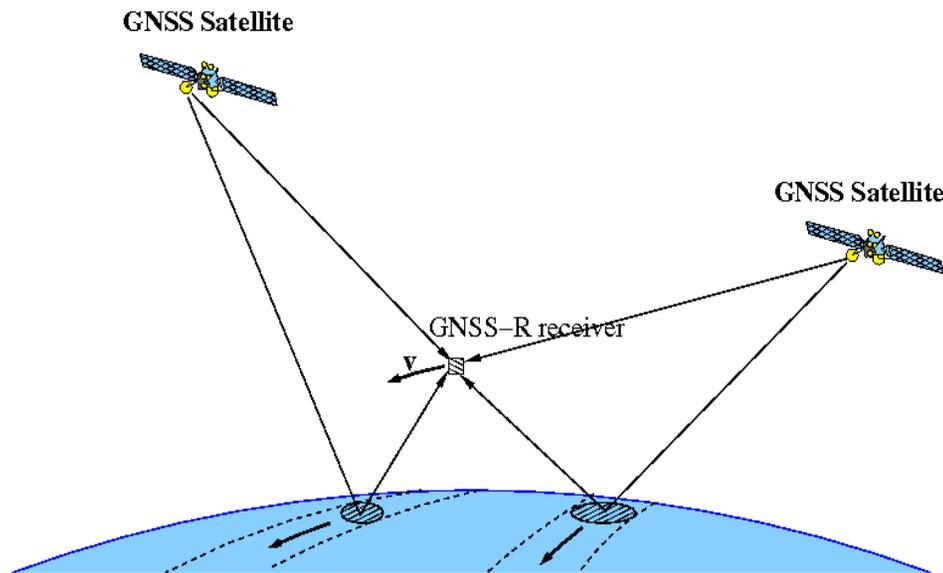
**Right: from <http://www.ucar.edu/news/releases/2006/>).**

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# Spaceborne Applications of GNSS

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From „Galileo Science Opportunity document“ prepared by the Galileo Science Advisory Committee (GSAC)

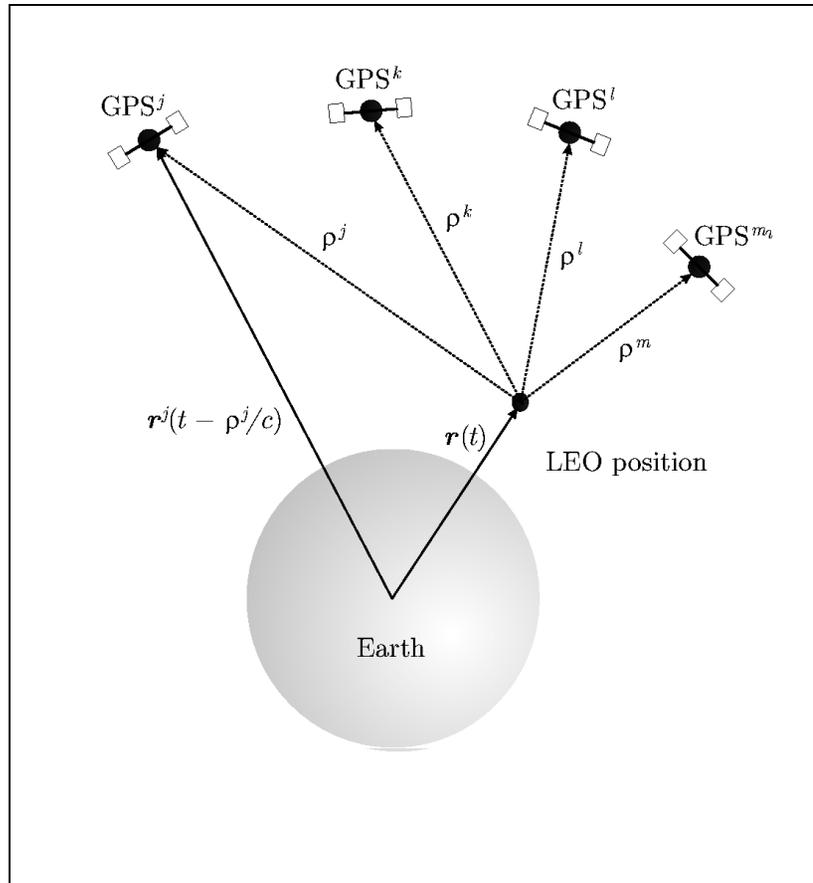
**GNSS reflectometry: A vehicle flying over the Earth's surface observes the GNSS signals emitted by the GNSS constellations, both direct and reflected, using a GNSS-R receiver. The correlation function of the reflected and direct signals (or suitable replicas) provides altimetric and scatterometric information. Several mission concepts are considered, no mission is flying.**

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# GNSS for Orbit Determination



GNSS Observation equations:

$$P^k = \rho^k + c (\Delta t - \Delta t^k) + \Delta \rho^{lk}$$

$$L^k = \rho^k + c (\Delta t - \Delta t^k) - \Delta \rho^{lk} + \lambda B^k$$

Code- und Phase-Measurements, respectively (on two carriers), contain information about the distances between GNSS satellites and LEO .

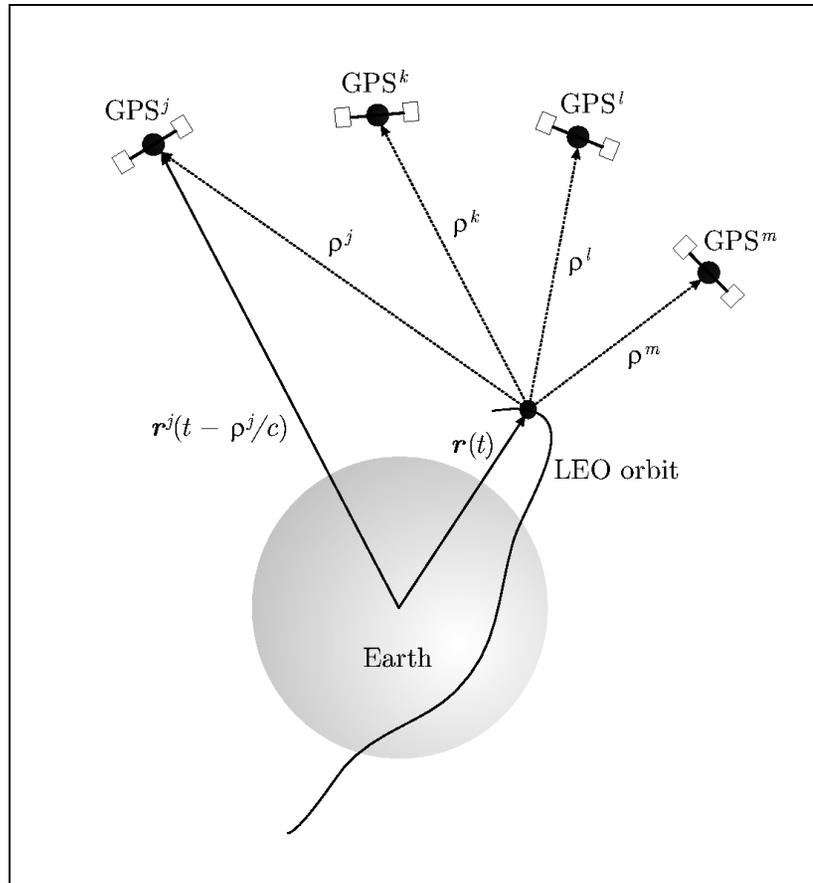
$$\rho^k = |r(t) - r^k(t - \rho^k/c)|$$

$r(t)$  : LEO position (**unknown**)

$r^k(t)$  : GNSS position (**known**)

Different options to parametrize the orbit  $r(t)$ .

# GNSS for Orbit Determination



## Kinematic orbit $r(t_k)$ :

LEO Positions are estimated for each epoch without making use of the equations of motion.

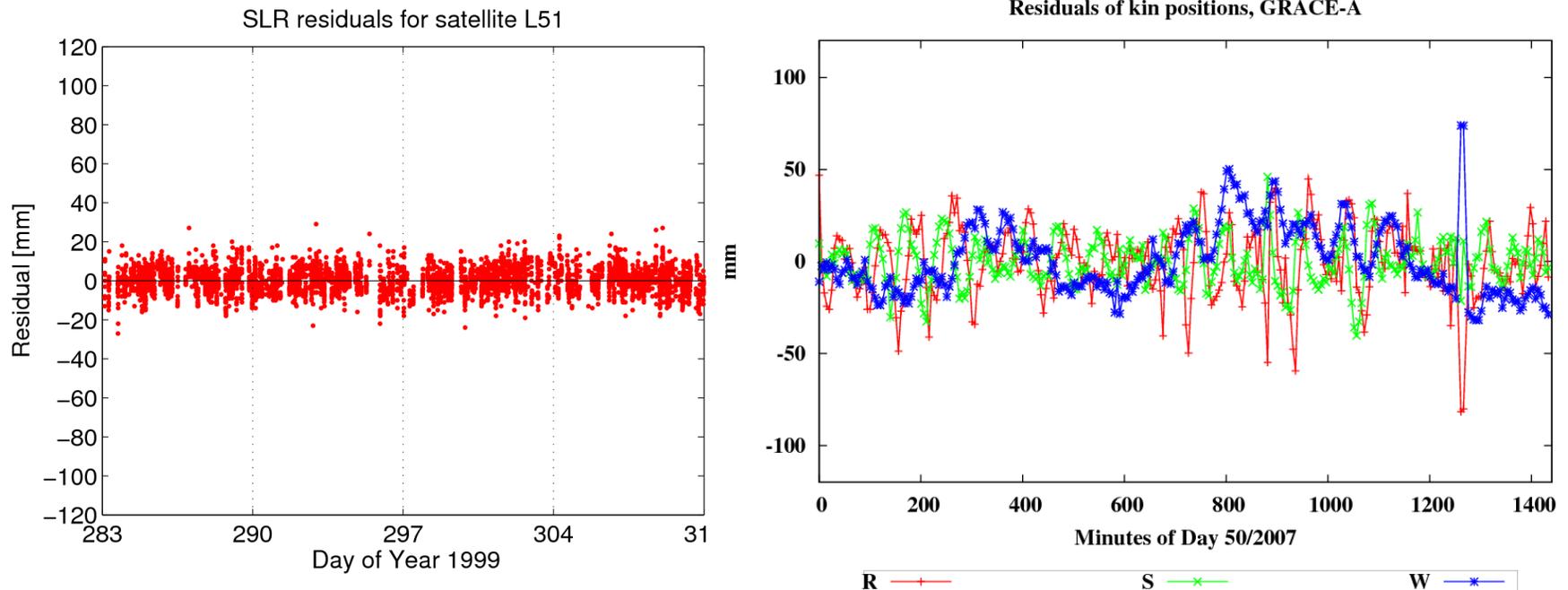
## Dynamic Orbit $r(t)$ :

Equations of motion (with known force model) determine orbit.

## (Pseudo-)stochastic Orbit $r(t)$ :

Piecewise constant accelerations or velocity changes at known epochs give the orbit stochastic properties (generalized short arc).

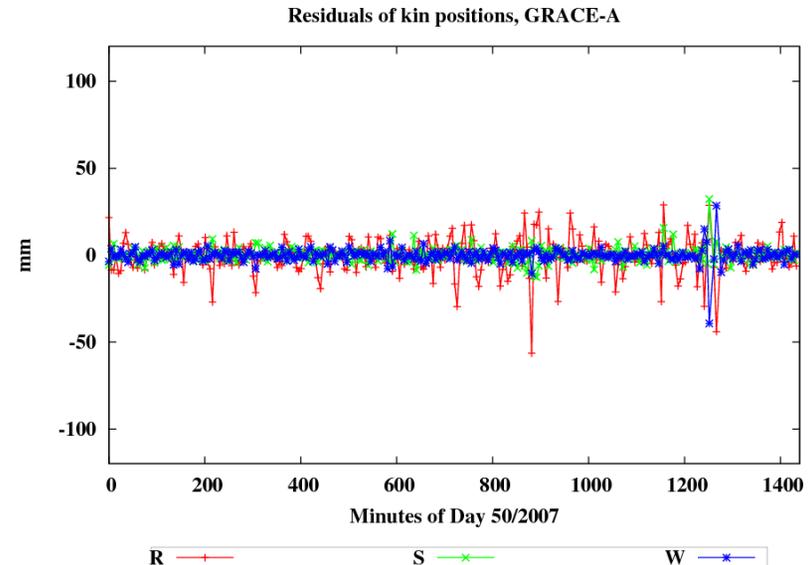
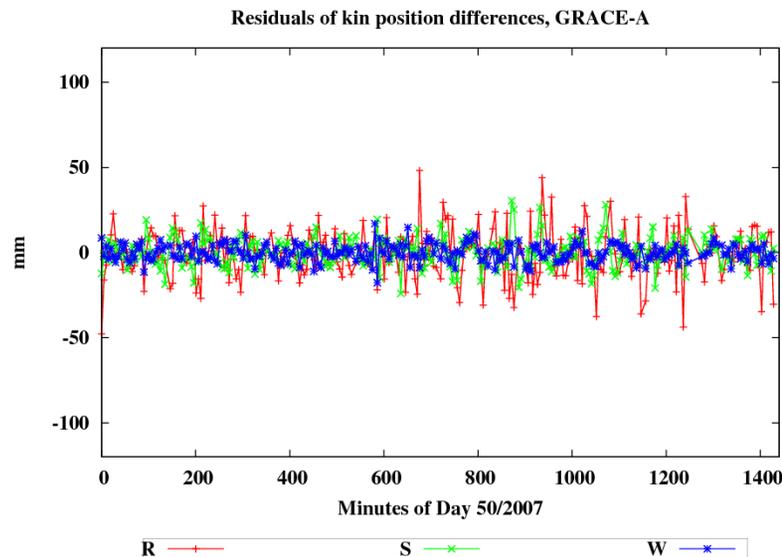
# GNSS for Orbit Determination



**Left: SLR residuals of four contiguous 7-day LAGEOS-orbits (from ILRS Benchmark test, < 20 parameters/arc, CODE)**

**Right: Residuals of fitting 30s kinematic positions with contiguous arcs of 90 min  
→ shorter arcs or use position differences between subsequent positions  
(GRACE gravity field up to  $n=m=150$ ).**

# GNSS for Orbit Determination

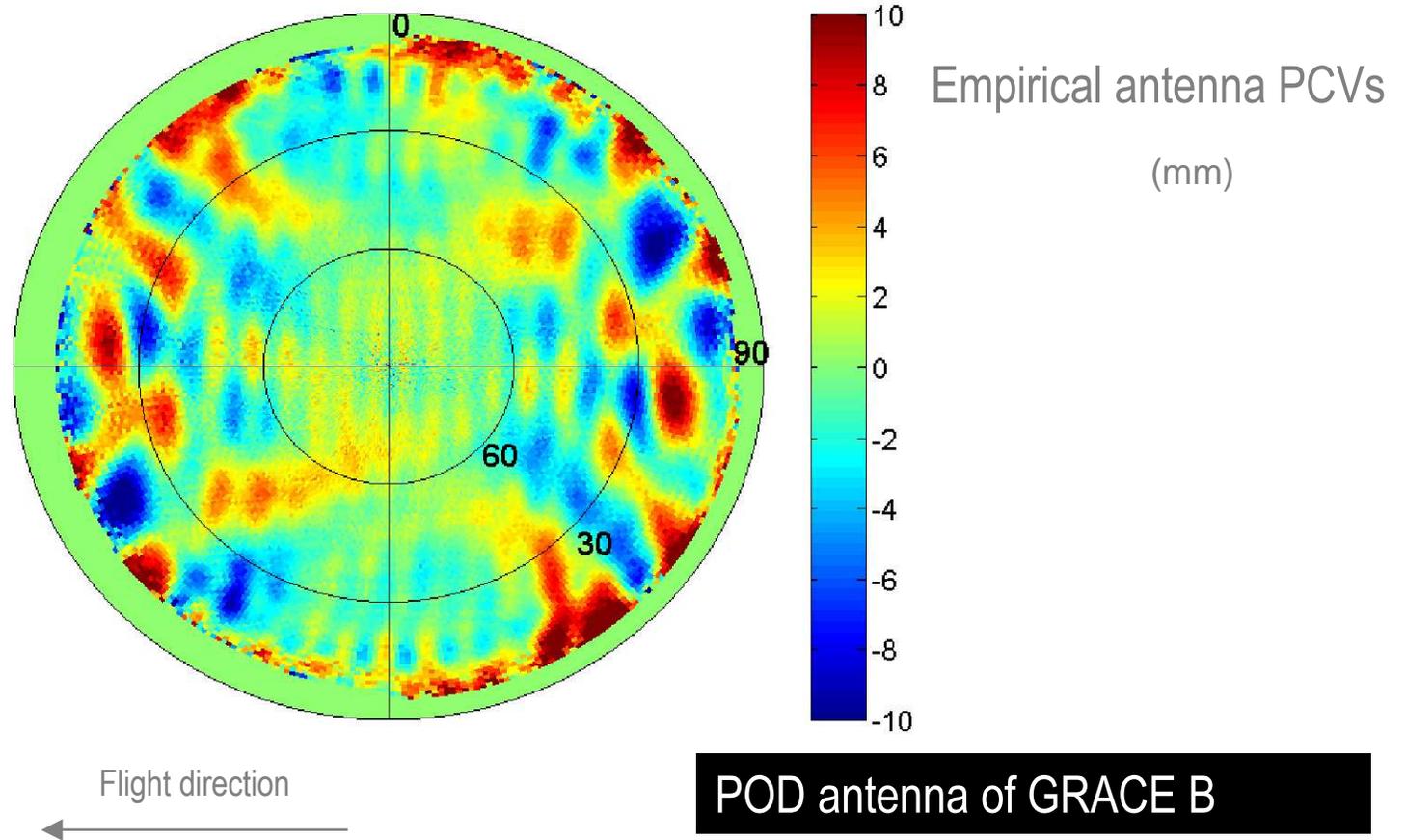


Left: Residuals of **position differences** using 90 min contiguous arcs

Right: Residuals of **positions** using 15 min contiguous short-arcs

Accuracy of kinematic positions (validated with SLR): 1-3 cm rms.

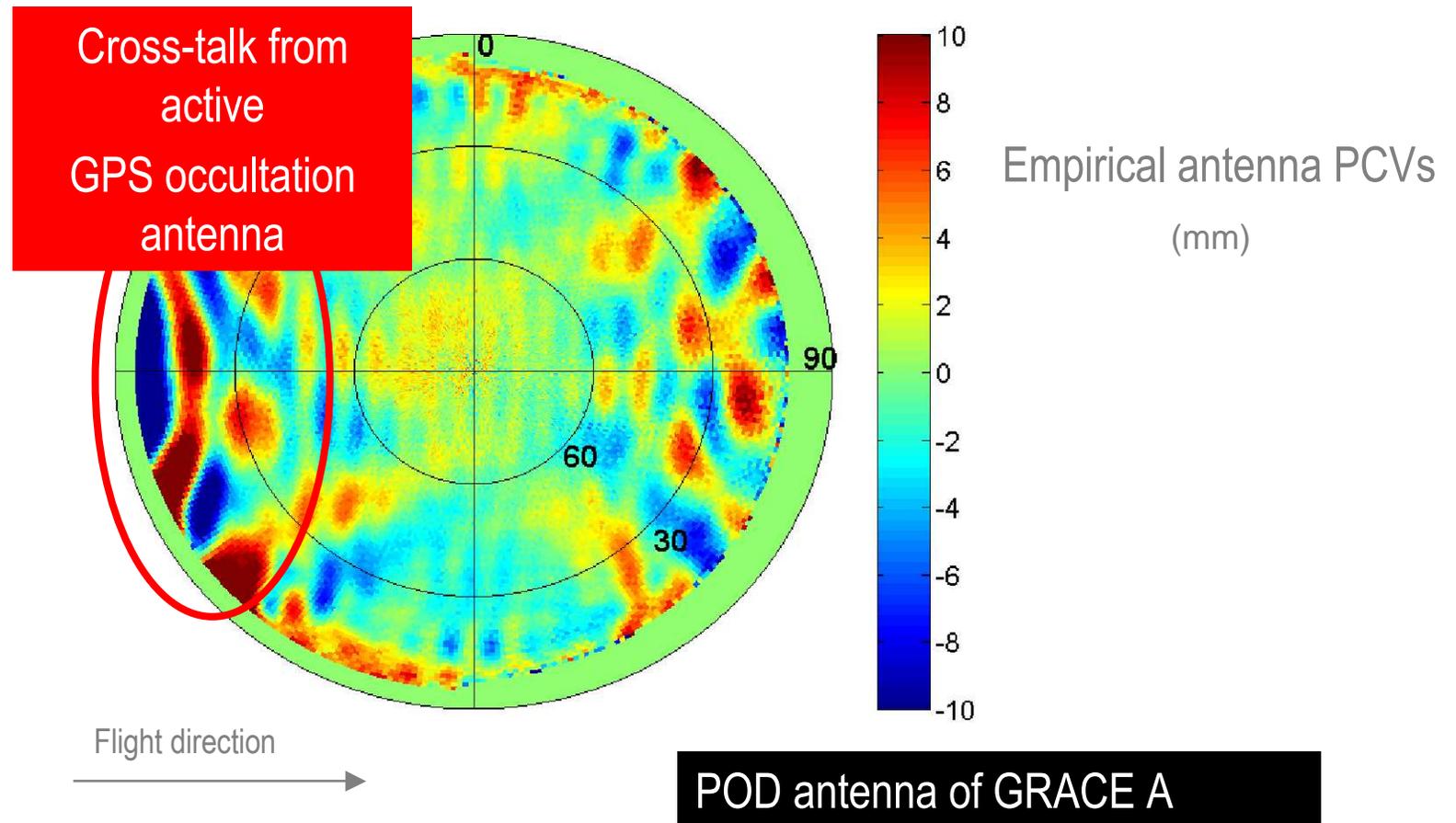
# GNSS for Orbit Determination



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# GNSS for Orbit Determination

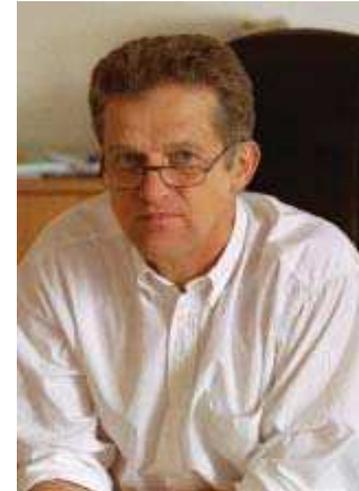
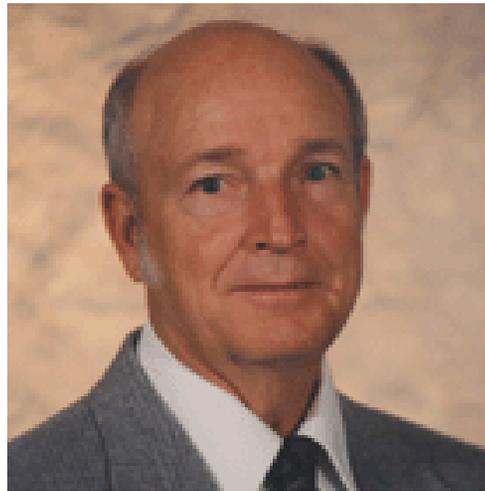


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# Gravity Field Missions 2000-2010

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From the beginning of space age to the end of the 20th century the information about the Earth's gravity field emerged from the SLR (+astrometry and altimetry).

Lacking temporal and spacial coverage: → Need for dedicated LEO missions.

The protagonists of the missions **CHAMP**: Christoph Reigber, **GRACE**: Byron Tapley & Ch. Reigber, **GOCE**: Reinhard Rummel (left to right).

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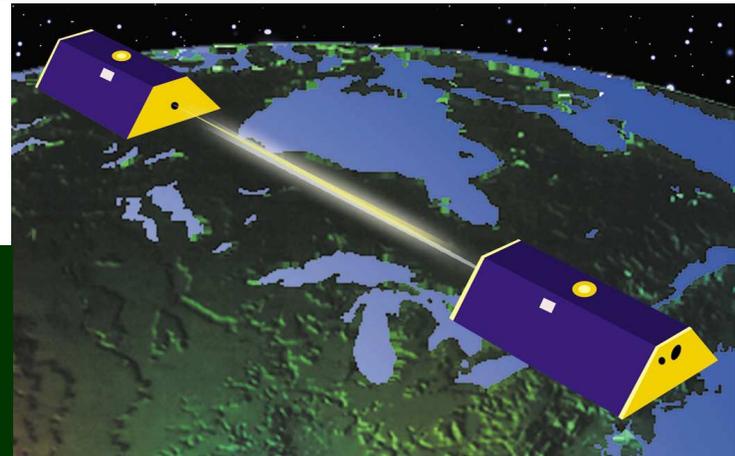
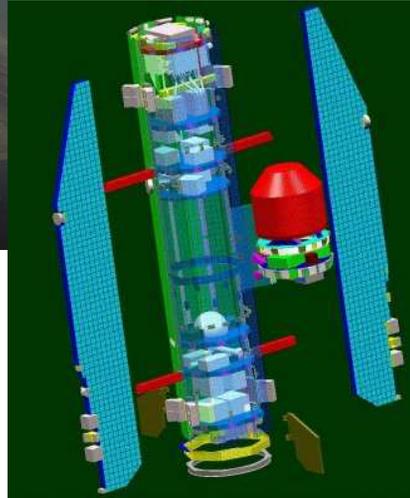
# Gravity Field Missions 2000-2010



CHAMP



GOCE



GRACE A and B

The **golden age of Gravity field determination** was initiated with the start of **CHAMP** in July 2000. **GRACE** was launched in 2002 in particular to study the time varying gravity field. ESA's **GOCE** mission was launched on March 17, 2009 with a European GPS receiver and a three-dimensional gradiometer (three pairs of 3-d accelerometers).

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# Gravity Field Missions 2000-2010

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$$V(r, \lambda, \phi) = \frac{GM}{r} \sum_{i=0}^{\infty} \left(\frac{a}{r}\right)^i \sum_{k=0}^i P_i^k(\sin \phi) \{ C_{ik} \cos k\lambda + S_{ik} \sin k\lambda \}$$

$$N(\phi, \lambda) = a \sum_{i=2}^{\infty} \sum_{k=0}^i P_i^k(\sin \phi) \{ C_{ik} \cos k\lambda + S_{ik} \sin k\lambda \}$$

The Earth's gravity field is usually expressed in spherical harmonics (top) with  $r$ =Absolute value of geoc. Radius vector,  $\phi$ =latitude,  $\lambda$ =longitude.

For  $r=a$ =Earth's equatorial radius) the **geoid height  $N$**  (bottom) above the selected reference ellipsoid) may be easily established.

# Gravity Field Missions 2000-2010

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The strength of the signals of degree  $i$  are expressed by the so-called **degree amplitudes**  $\sigma_i$ , where:

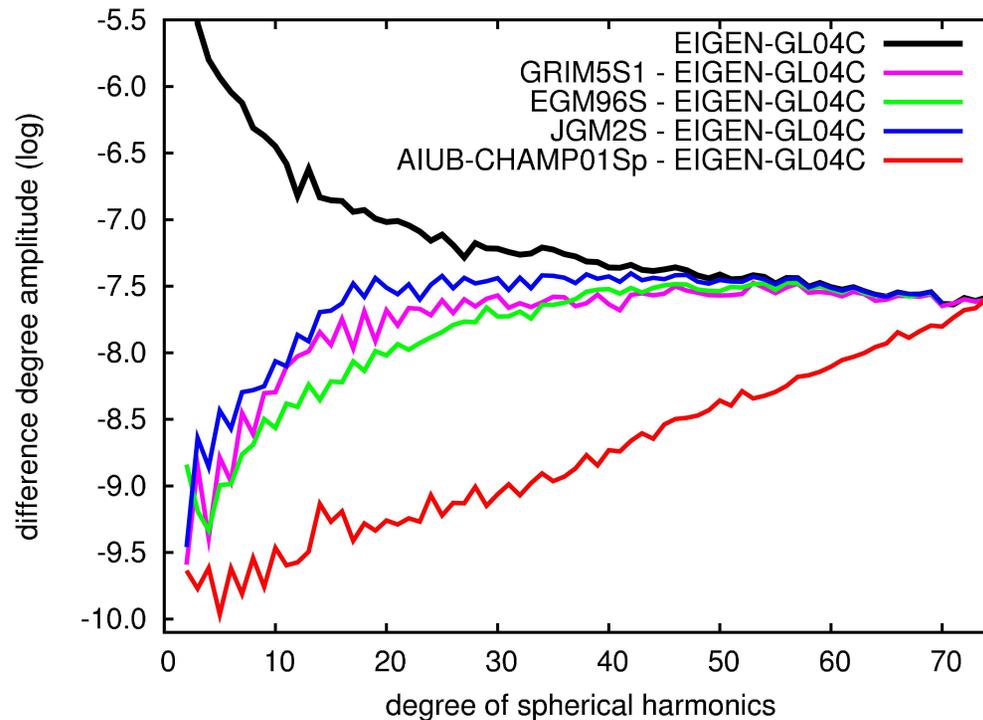
$$\sigma_i^2 = \sum_{k=0}^i \{C_{ik}^2 + S_{ik}^2\}$$

The degree-specific differences between two gravity fields is conveniently expressed by the **difference degree amplitudes**  $\Delta\sigma_{ik}$ :

$$\Delta\sigma_i^2 = \sum_{k=0}^i \left\{ (C_{ik,1} - C_{ik,2})^2 + (S_{ik,1} - S_{ik,2})^2 \right\} = \sum_{k=0}^i \{ \Delta C_{ik}^2 + \Delta S_{ik}^2 \}$$

Usually, the difference degree amplitudes are represented in a logarithmic scale. The  $\Delta\sigma_{ik}$  may either be analyzed directly or transformed into geoid height differences.

# Gravity Field Missions 2000-2010



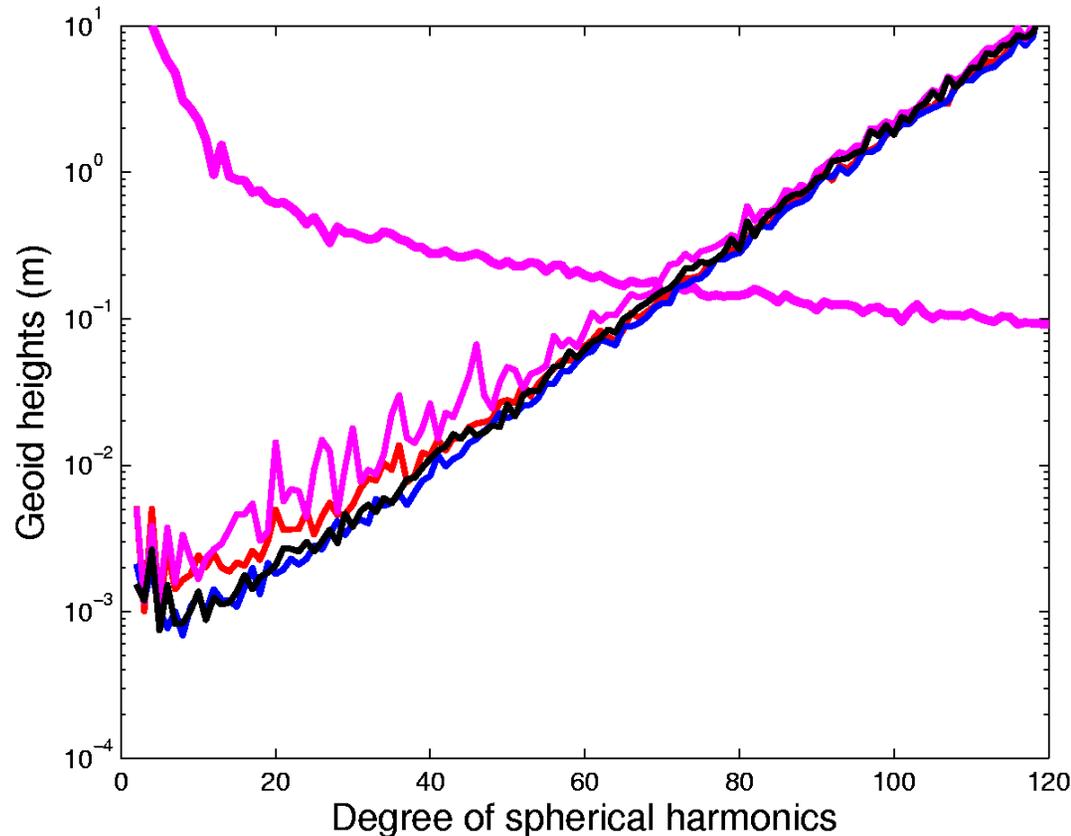
Difference degree amplitudes w.r.t. a GRACE multi-annual solution of three of the best 20<sup>th</sup> century gravity field solutions (mainly based on SLR and terrestrial measurements) and a CHAMP-only solution using „only“ one year of GPS-data: A **quantum jump in gravity field determination** is obvious!

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# GNSS Contribution to the Gravity Fields

PCVs are very important for GPS-only solutions



Reference field:

ITG-GRACE03S

Differences:

GRCB (w/o PCV)

GRCB (with PCV)

GRCA (w/o PCV)

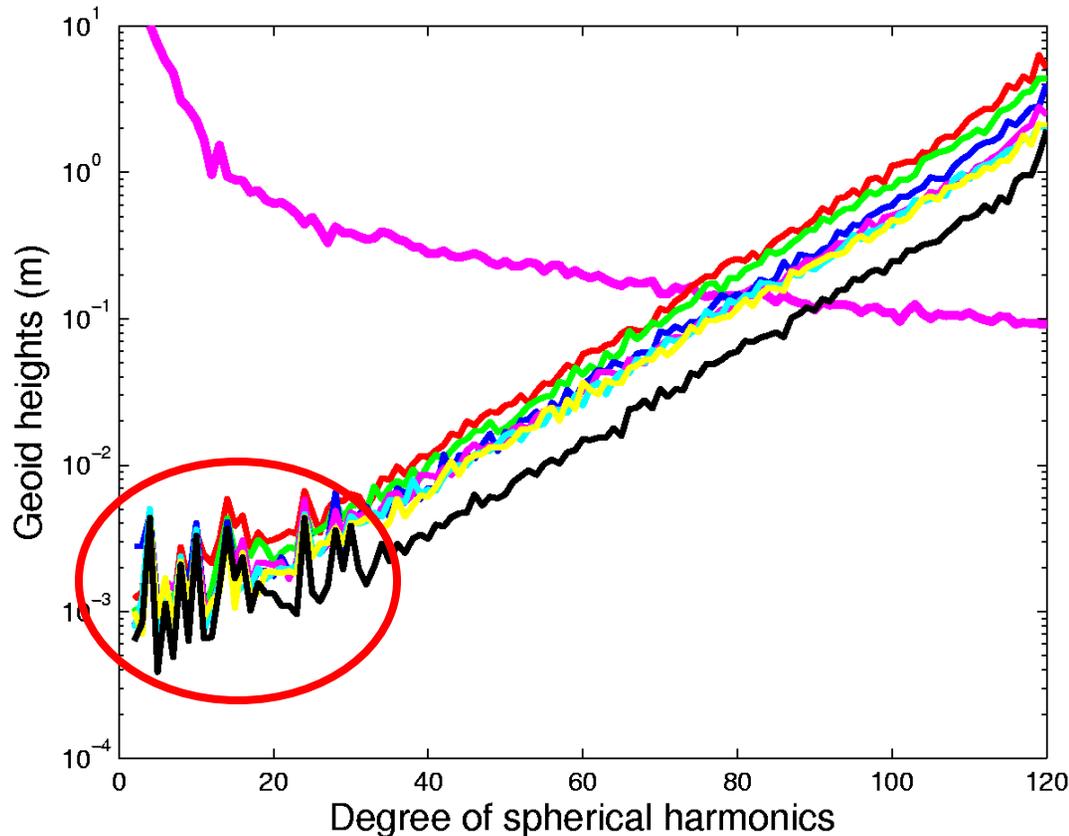
GRCA (with PCV)

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# GNSS Contribution to the Gravity Fields

CHAMP solutions with “normal” modeling



Reference field:

ITG-GRACE03S

Differences:

Year 2002

Year 2003

Year 2004

Year 2005

Year 2006

Year 2007

Combination:

6-year solution

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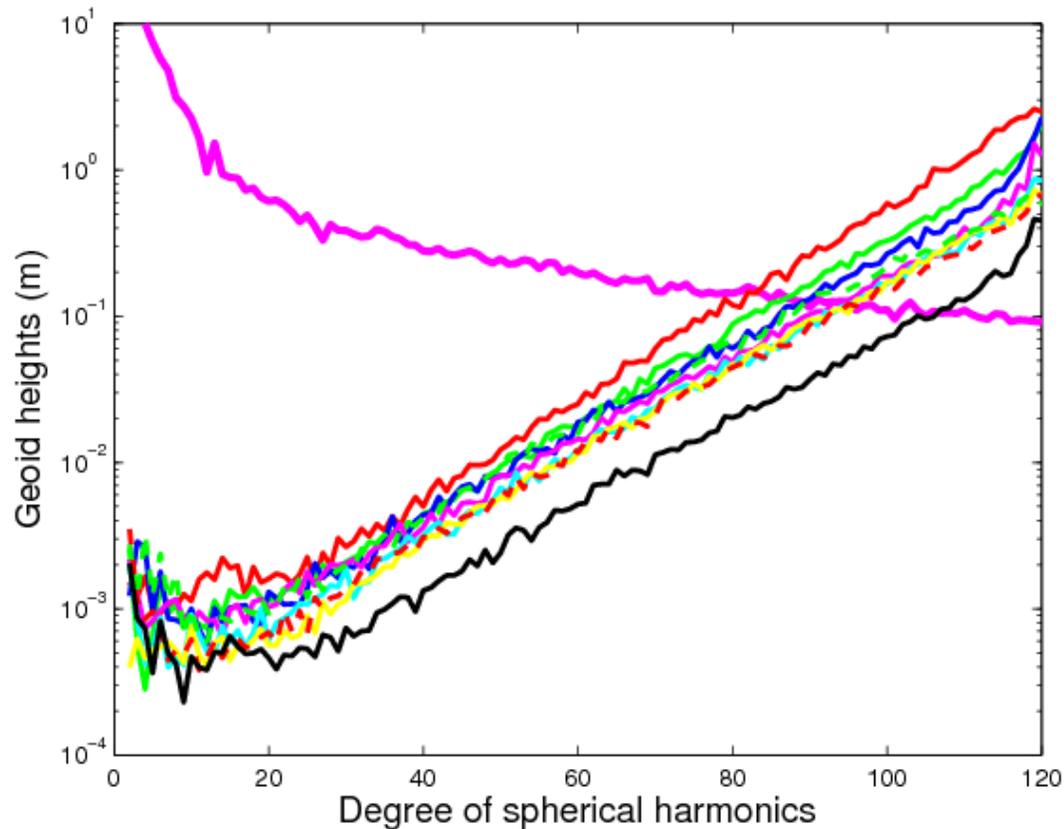
*AIUB*

12-Apr-10

Low degree coefficients disagree with the reference field

# GNSS Contribution to the Gravity Fields

CHAMP solutions with “advanced” modeling



Reference field:

ITG-GRACE03S

Differences:

Year 2002

Year 2003

Year 2004

Year 2005

Year 2006

Year 2007

Year 2008

Year 2009

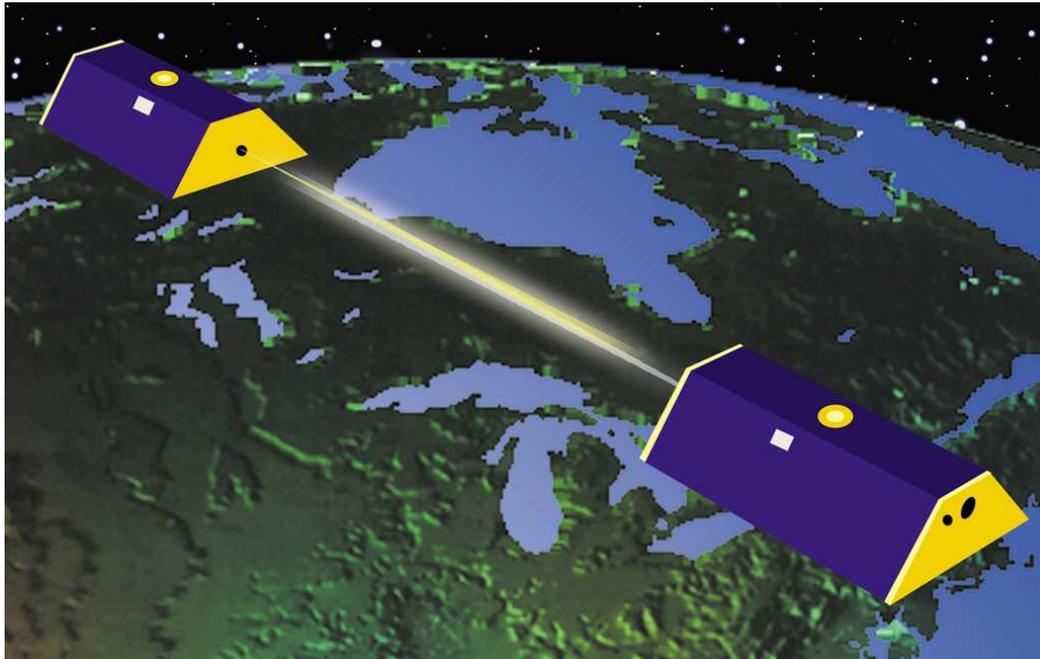
Combination:

8-year solution

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Excellent solution (PCVs, COVs, 10s sampling, ...)

# Gravity Fields from GNSS and K-Band

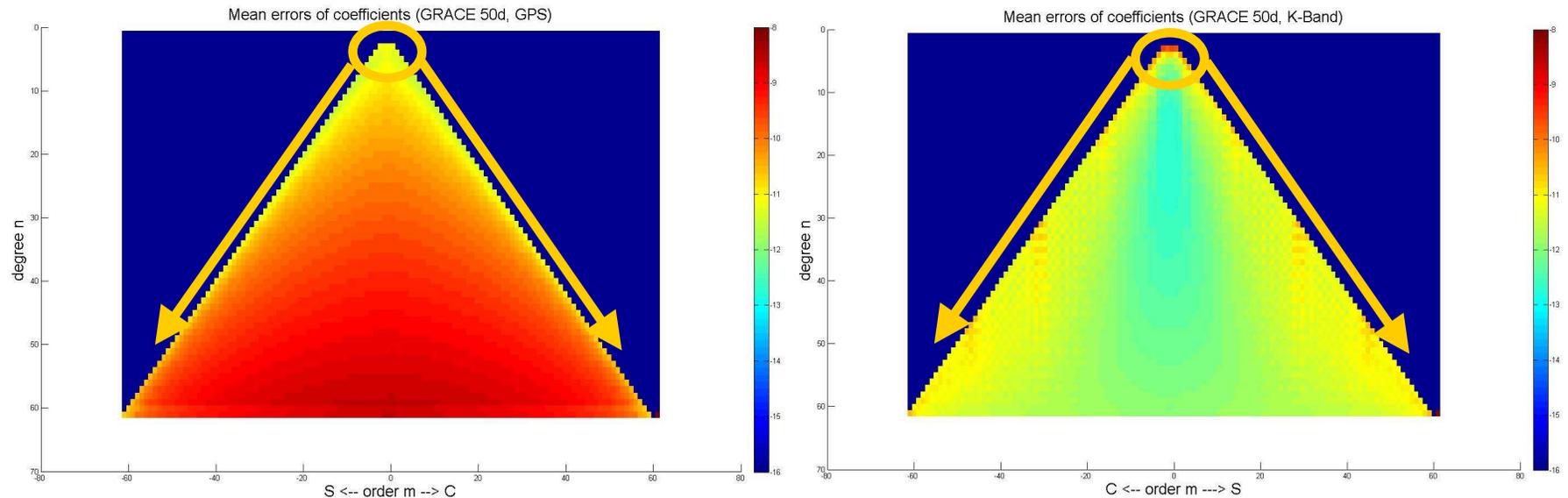


GRACE-A and -B are in close to circular and polar orbits about 450km above the Earth's surface.

The gravity field may be established from GPS alone, K-Band alone (?) and from a combination of the two.

The temporal evolution of the inter-satellite distance is measured with an accuracy of better than  $10^{-6}$ m/s using a microwave link (in the K-Band) between the two satellites *in addition to* the absolute positions established by GPS → „Accuracy-wise“ K-Band dominates GPS but K-Band alone cannot determine the orbits.

# Gravity Fields from GNSS and K-Band



**K-Band superior for all terms, except for low degree terms and sectorials!**

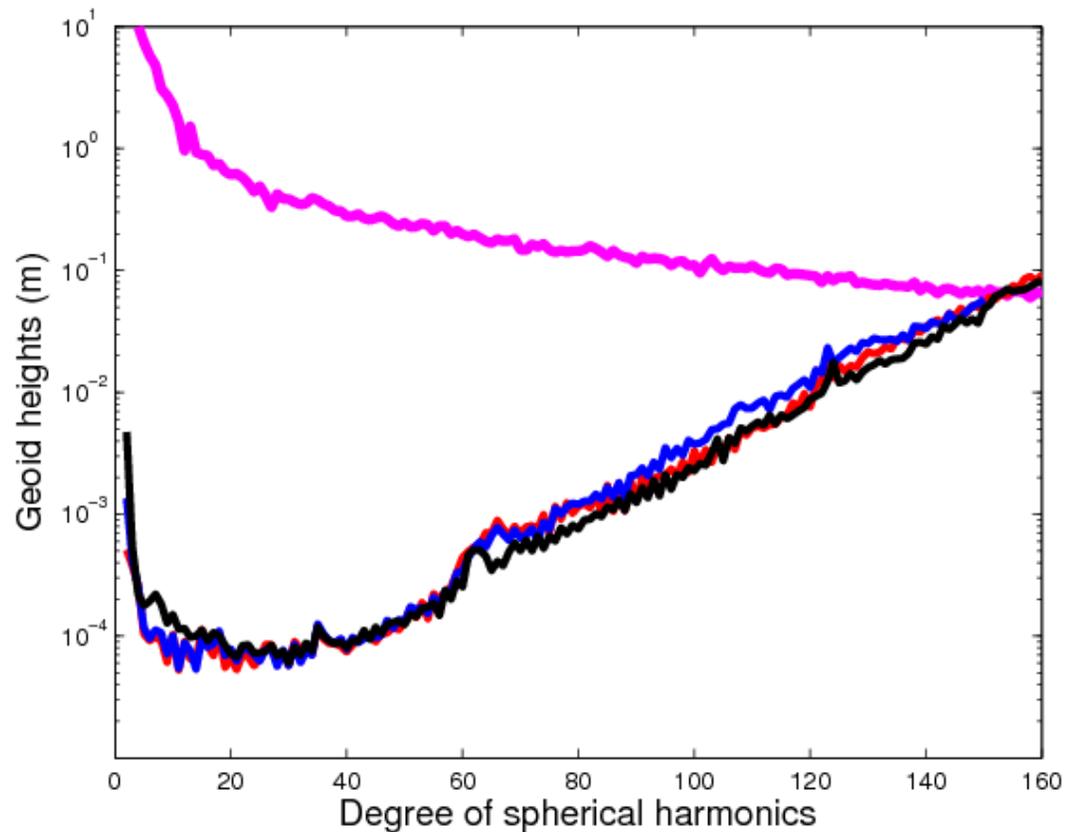
**→ Errors in GPS analysis are „well hidden“ in combined GRACE solutions**

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# Gravity Fields from GNSS and K-Band

Consistency of multi-annual GRACE solutions (GPS+K-Band)



Reference field:

ITG-GRACE03S

Differences:

GGM03S (CSR)

EIGEN5S (GFZ)

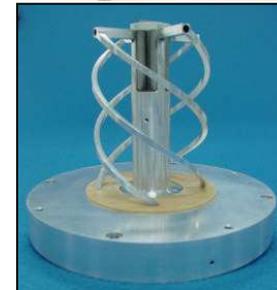
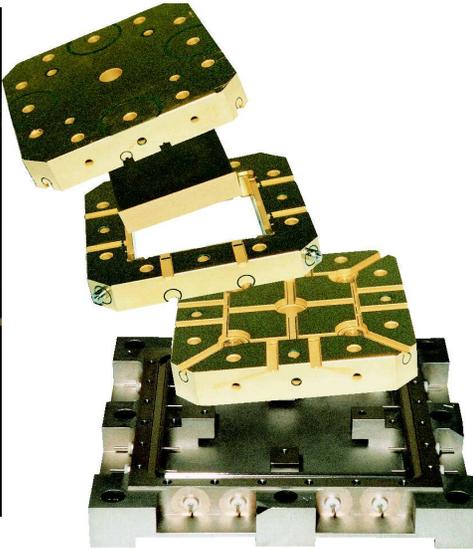
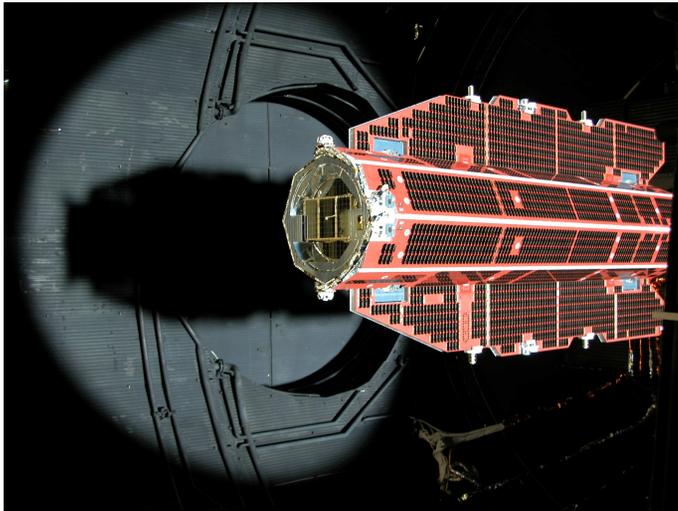
Differences:

AIUB-03Sp

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# Gravity Fields from GOCE

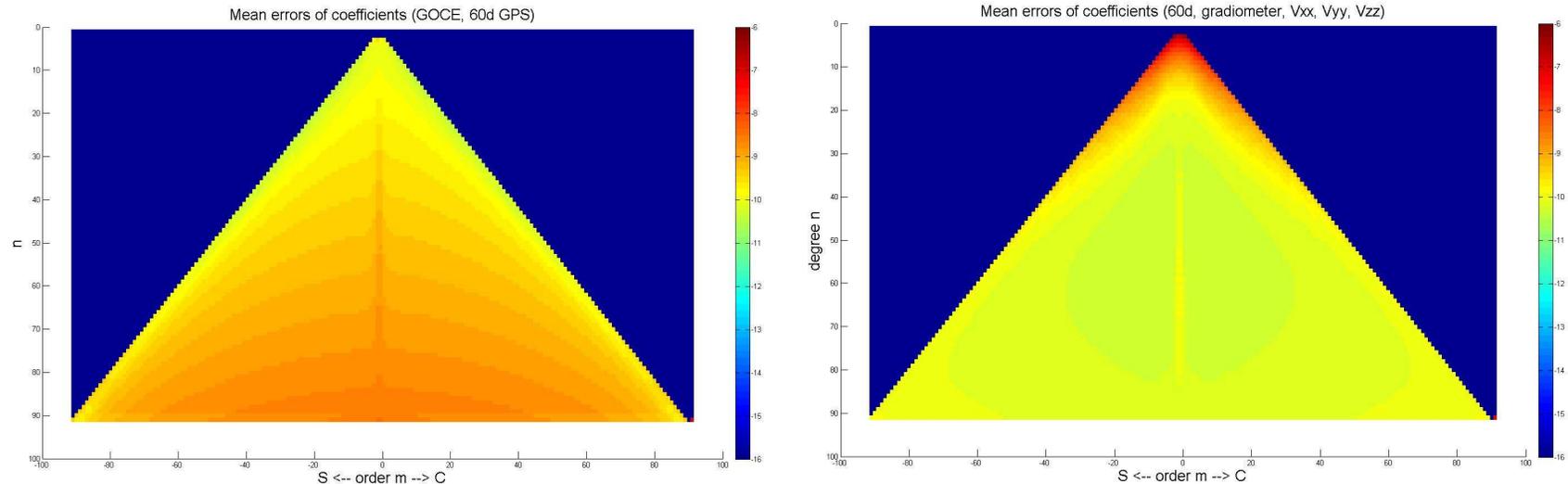


The GOCE satellite (left) carries an advanced European GPS receiver (right, bottom), a gradiometer consisting of three pairs of three-dimensional accelerometers (center). In addition it continuously corrects for atmospheric drag using ion thrusters (right, top).

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# Gravity Fields from GOCE



**Approximate error behavior (from simulation of a 60<sup>d</sup> measurement period) GPS-only (left) and gradiometer-only (right). GPS will influence combined solution up to higher degrees. Gradiometer dominates from  $n=20-30$  onwards.**

**Combination of GPS and gradiometer will be a challenge.**

# Summary I

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**GNSS** has since 20 years a heavy impact on the determination of:

- the terrestrial reference frame
- polar motion and length of day
- ionosphere mapping
- local/regional troposphere monitoring

**Spaceborne GNSS** has a heavy impact on science since 10 years

- Orbit determination of LEOs and LEO constellations
- Atmosphere sounding using the GNSS occultation technique
- gravity field determination

# Summary II

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**Spaceborne GNSS** has revolutionized gravity field determination.

- **GPS** plays an important role in all currently active missions (CHAMP, GRACE, GOCE)
- The **GNSS** contribution, together with the **SLR** contribution, is critical to determine the low degree terms of the gravity field.
- **Multi-GNSS (GPS+GLONASS+GALILEO)** will significantly advance orbit and gravity field determination with **GNSS**.
- **GNSS** might be even more successful if constellations of **LEOS** with significantly different inclinations would be used.

# Summary III

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**Spaceborne GNSS** has not yet exploited its full potential: GNSS reflectometry should, e.g., be added to the list of spaceborne applications.

**Galileo**, with its rich ensemble of frequencies and high-quality signals,

- will significantly contribute to all aspects of science enabled by terrestrial and spaceborne GNSS,
- has the potential to invoke a revolution of GNSS methodology and will, if fully exploiting the options offered by its space clocks, trigger important new modeling activities (e.g., for orbits).

The **GSAC** will do its best to stimulate the necessary developments.